

Exploring the reliability of chemical traits for RGB stars using accurate asteroseismic ages

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Spectral fidelity

Florence, 4-8 September





THE POWER OF STELLAR SPECTROSCOPY



Spectroscopic surveys (APOGEE, GALAH, Gaia+ESO) + space missions (Gaia) + simulations

Many applications

- Formation and evolution of the Galaxy
- Characterize stellar populations
- Understand star formation history
- Disentangle galactic components
- Age calibration



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Chemical abundance ratios

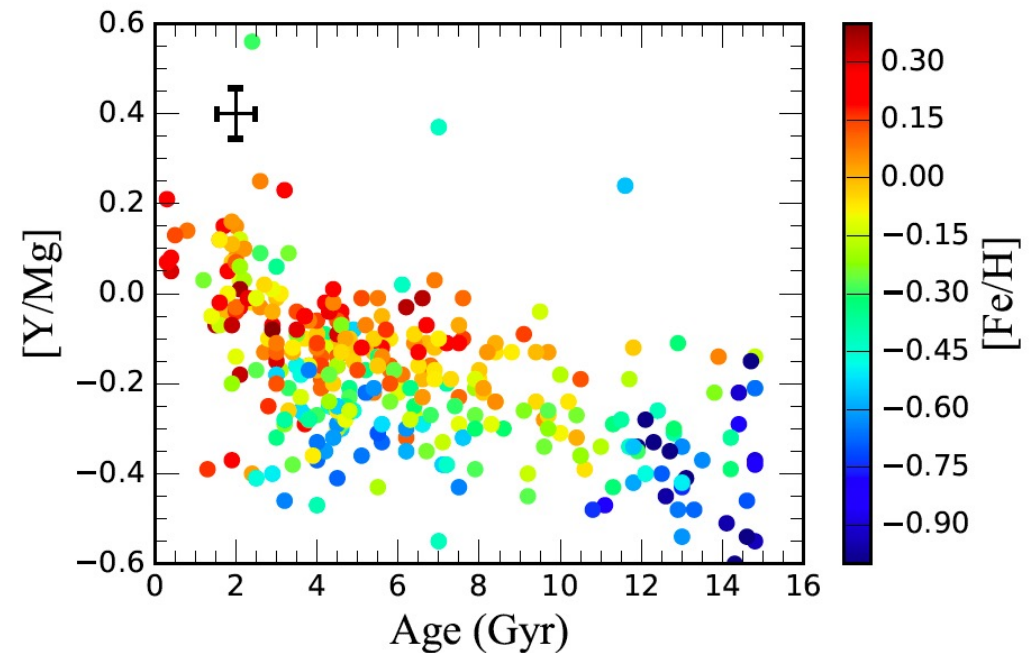
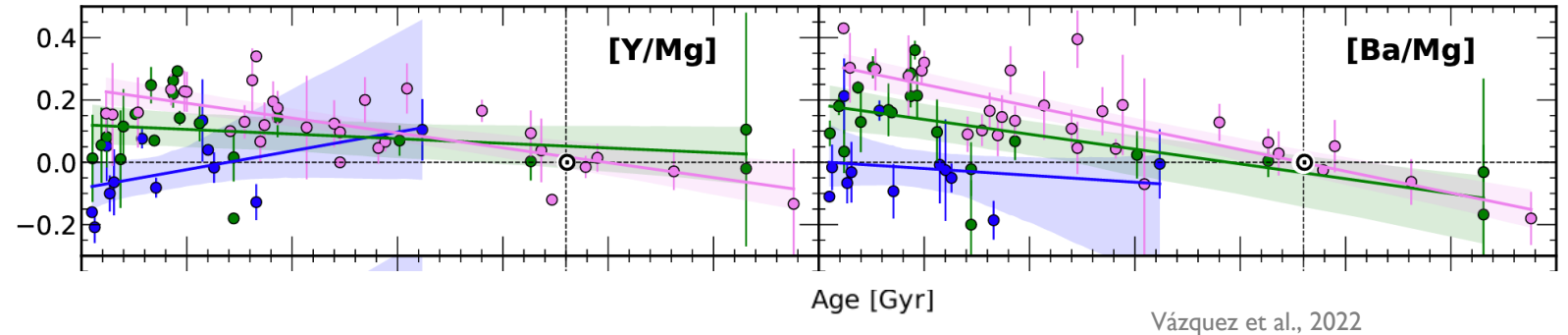
🕒 CHEMICAL CLOCKS: HOW MUCH WE CAN TRUST THEM?

Chemical abundance ratios
sensitive to ages
(Nissen+2015/2020, da Silva 2012,
Tucci Maia 2016...)

Their reliability and homogeneity
have been investigated

Dependence on:

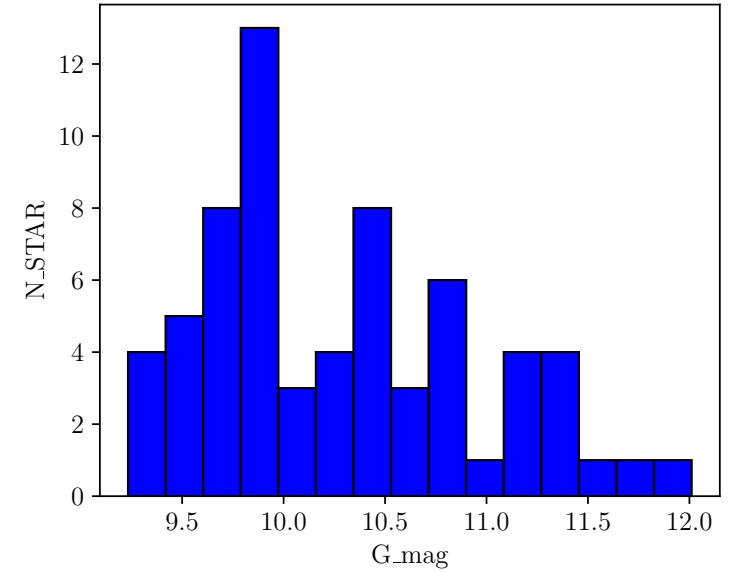
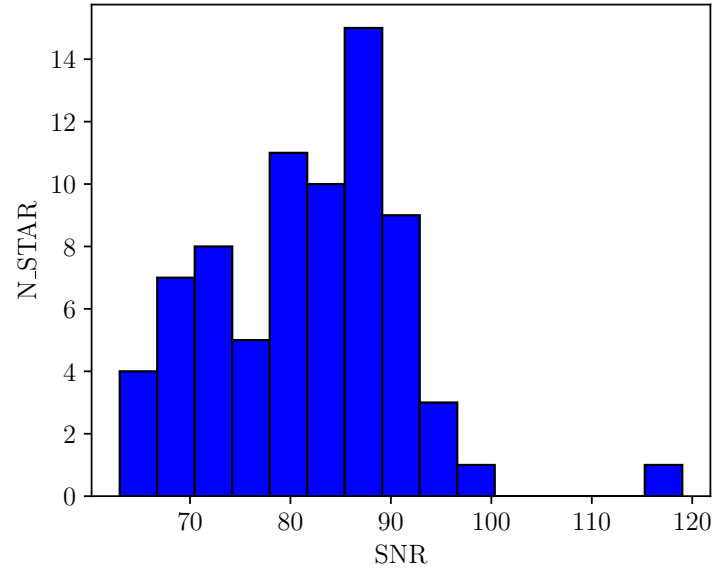
- Metallicity
- Environment
- Most informative traits?
- Which are their dependencies (metallicity, evolutionary stage, position...)?





THIS SAMPLE

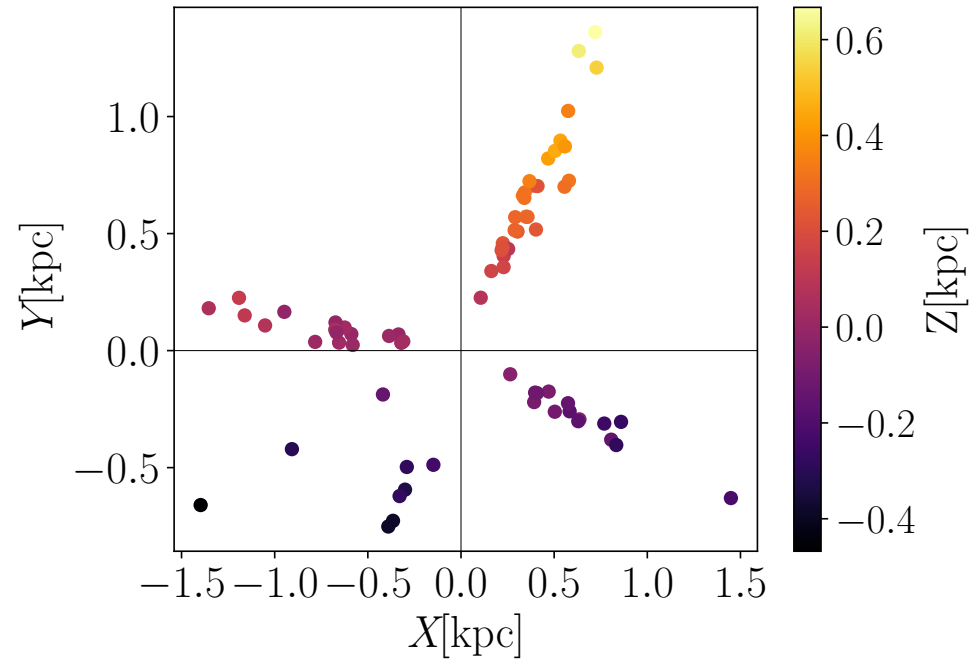
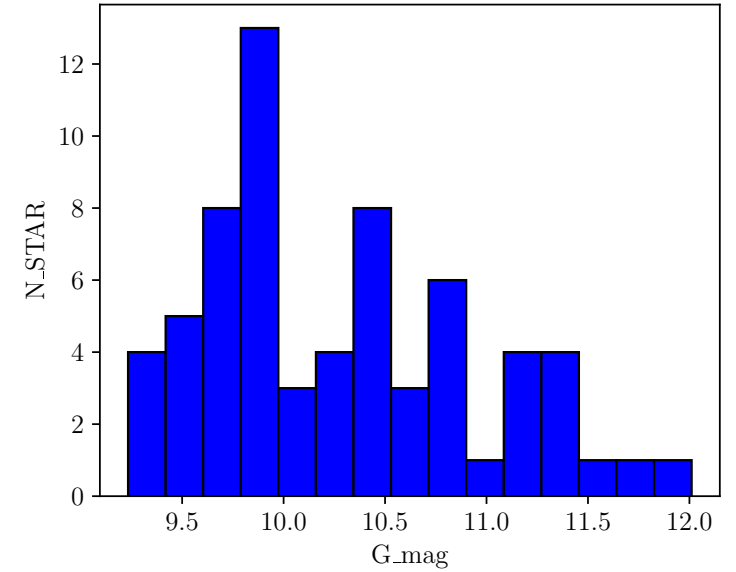
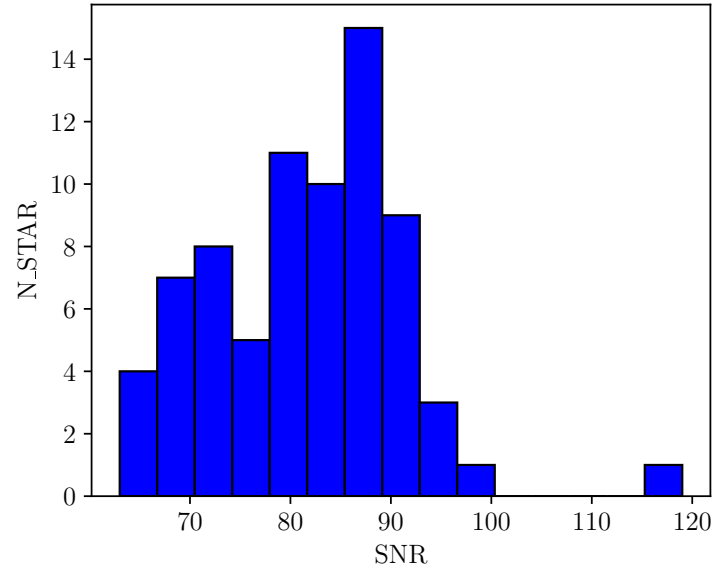
73 UVES spectra, $R \sim 110\,000$



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73 UVES spectra, $R \sim 110\,000$

Disk bright field giants



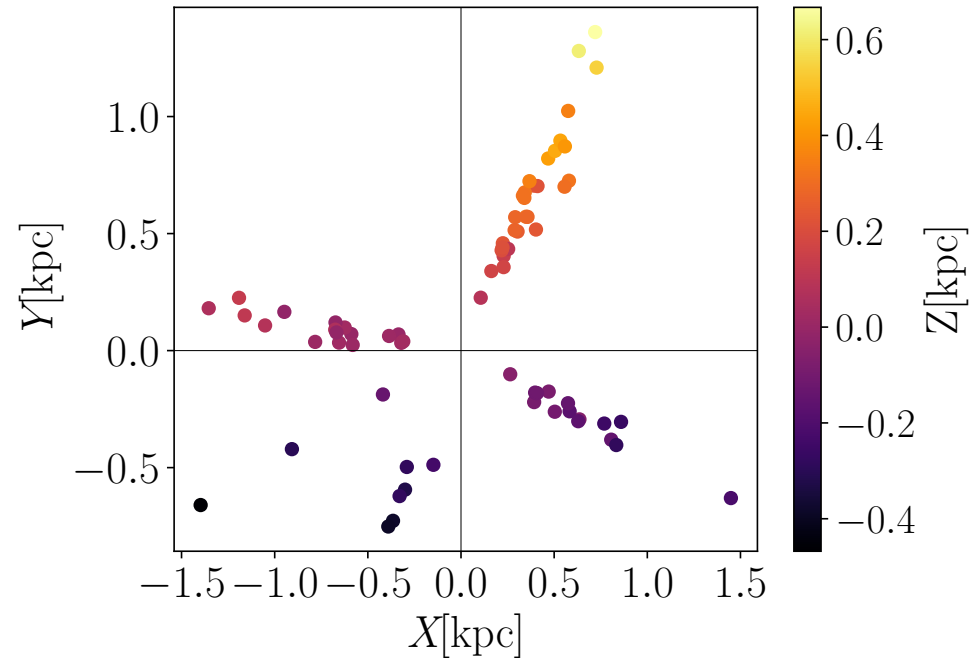
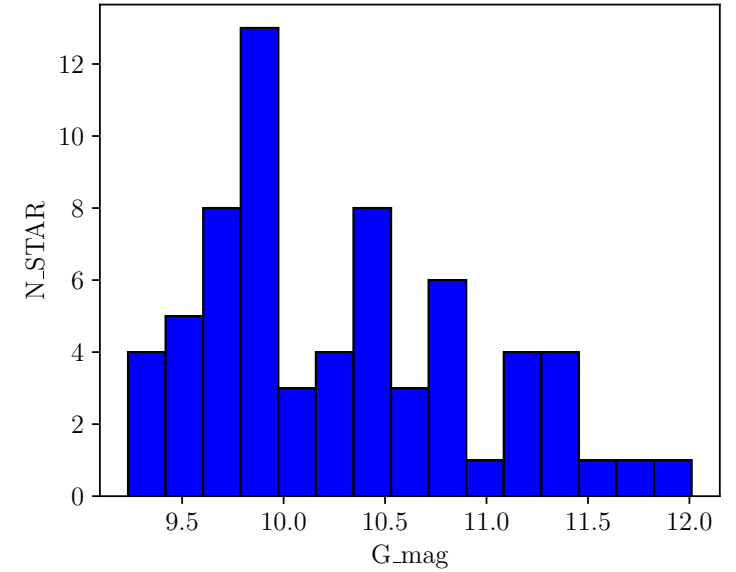
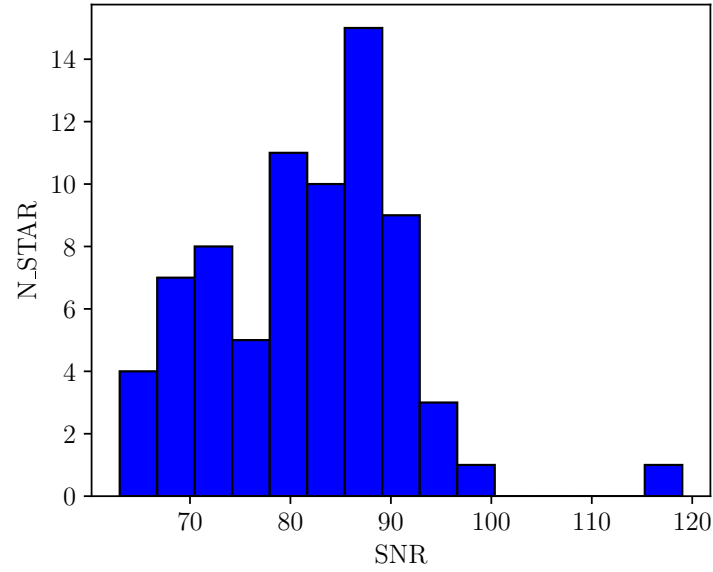
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Disk bright field giants

APOGEE fields

Wide metallicity range



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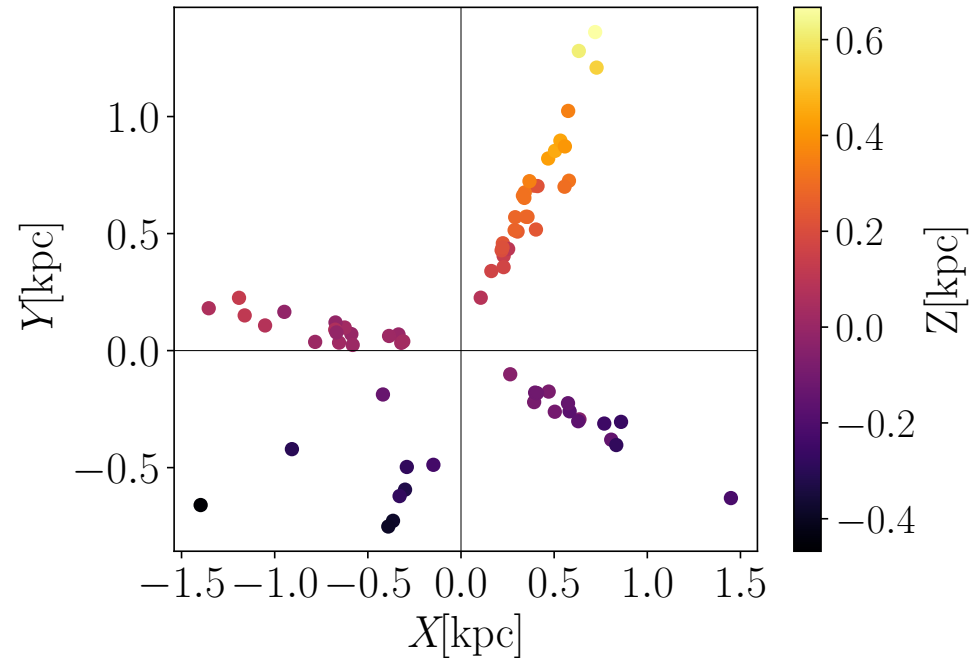
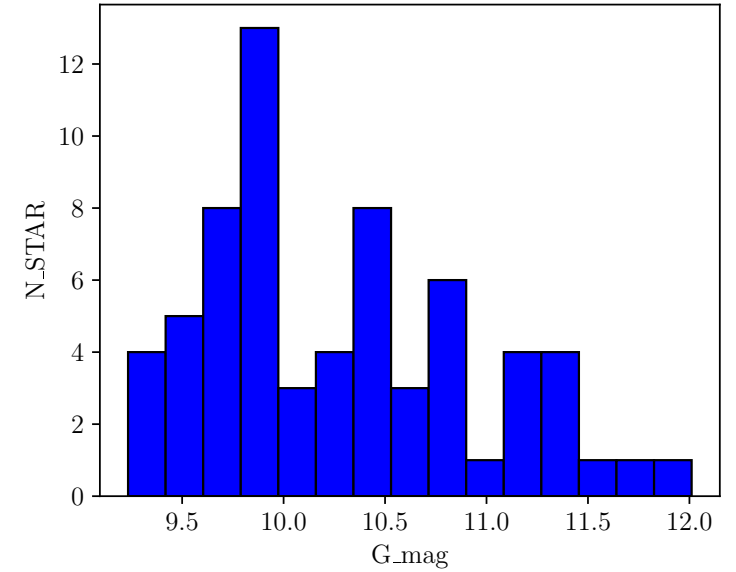
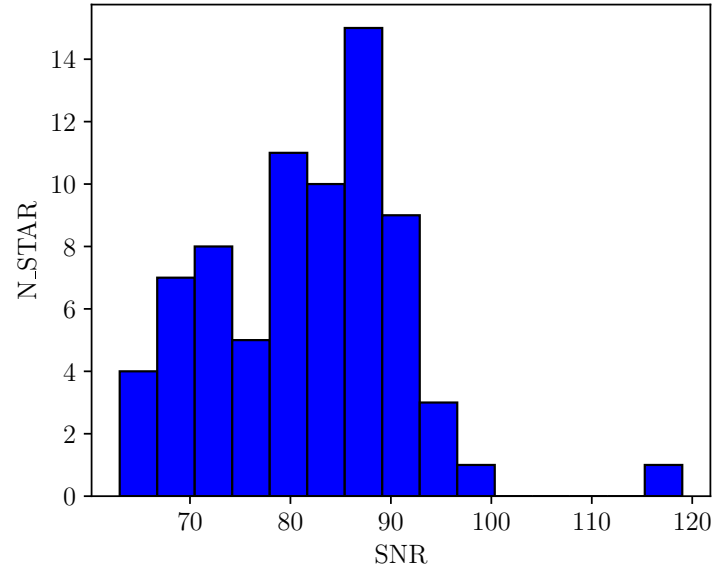
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APOGEE fields

Wide metallicity range

K2 fields

Asteroseismology



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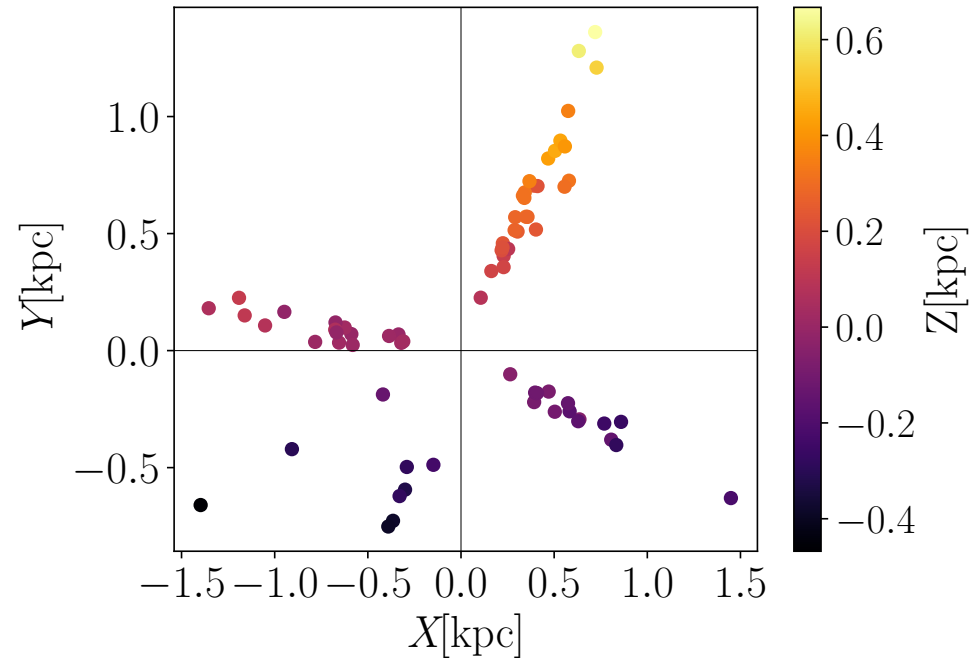
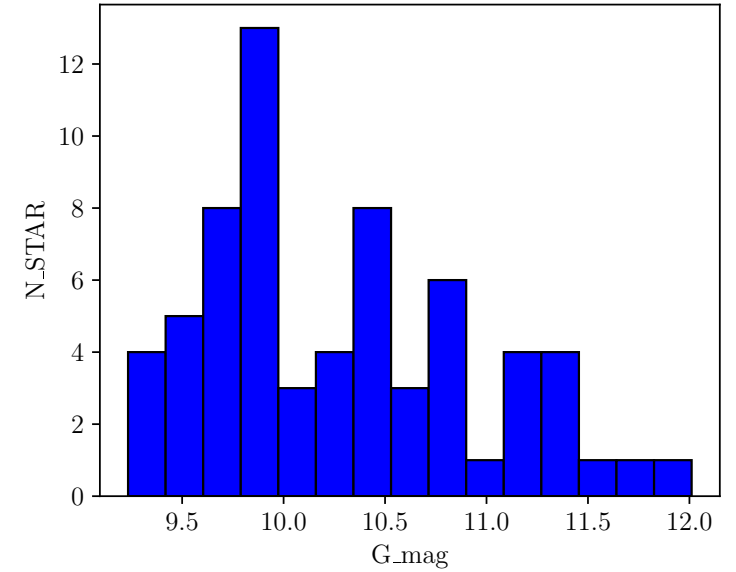
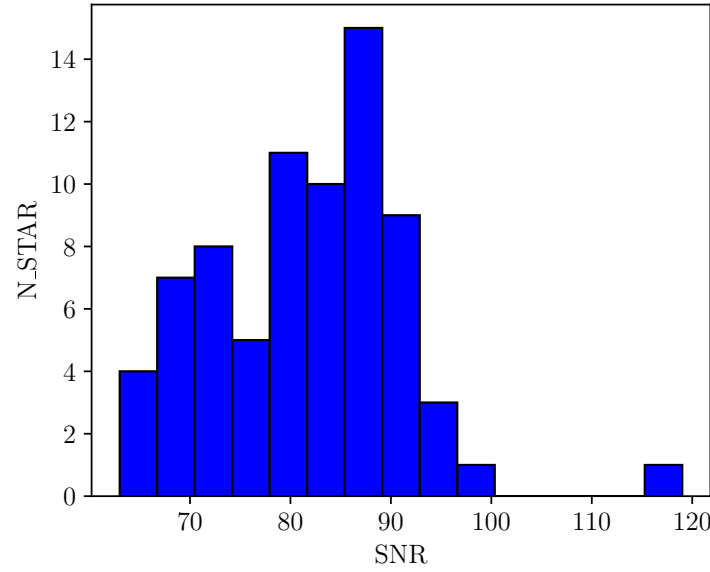
Wide metallicity range

K2 fields

Asteroseismology

High precision abundances

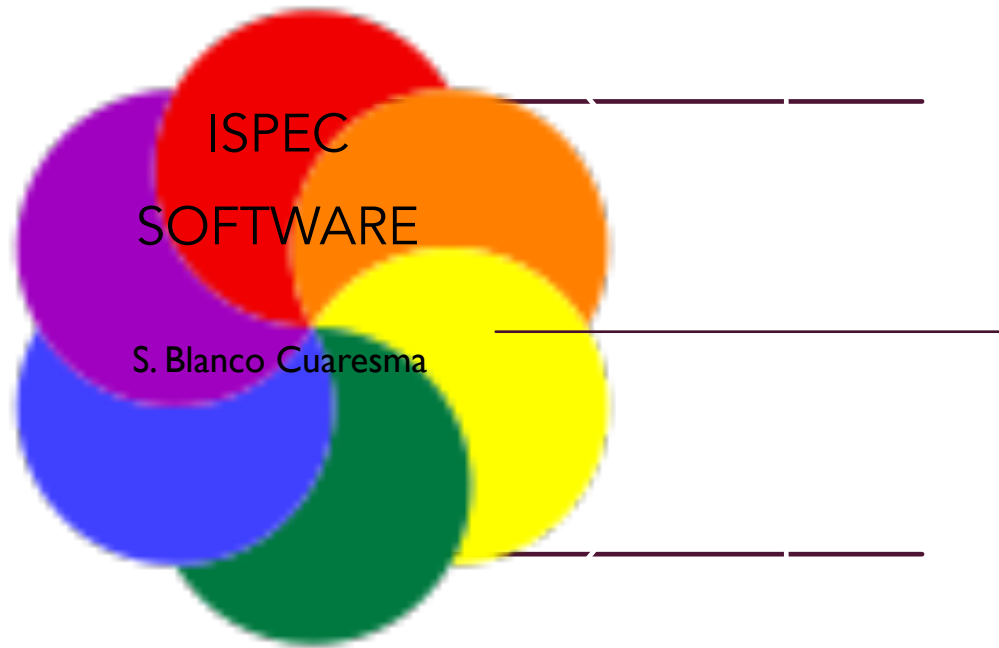
Accurate ages





SPECTRAL ANALYSIS

ESO pipeline ([VLT pipeline](#)) for data reduction



PRE-PROCESSING: SKY-SUBTRACTION
NORMALIZATION

ATMOSPHERIC PARAMETERS

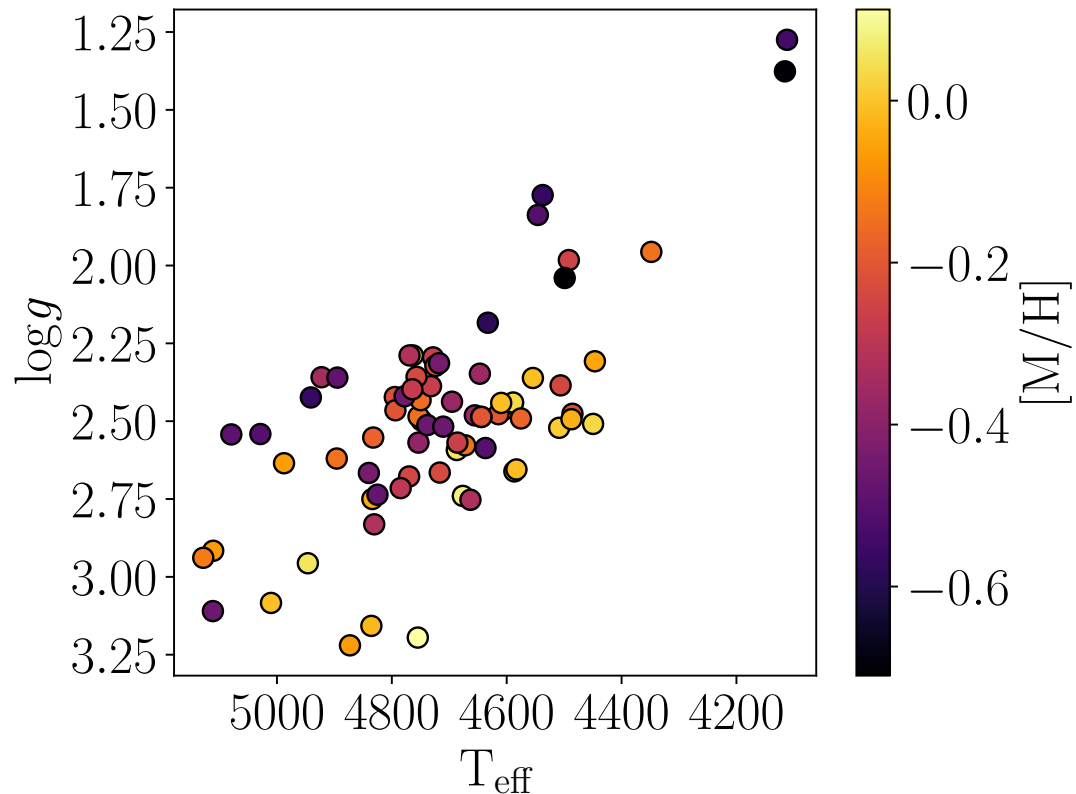
CHEMICAL
ABUNDANCES

} STELLAR
SYNTHESIS



ATMOSPHERIC PARAMETERS

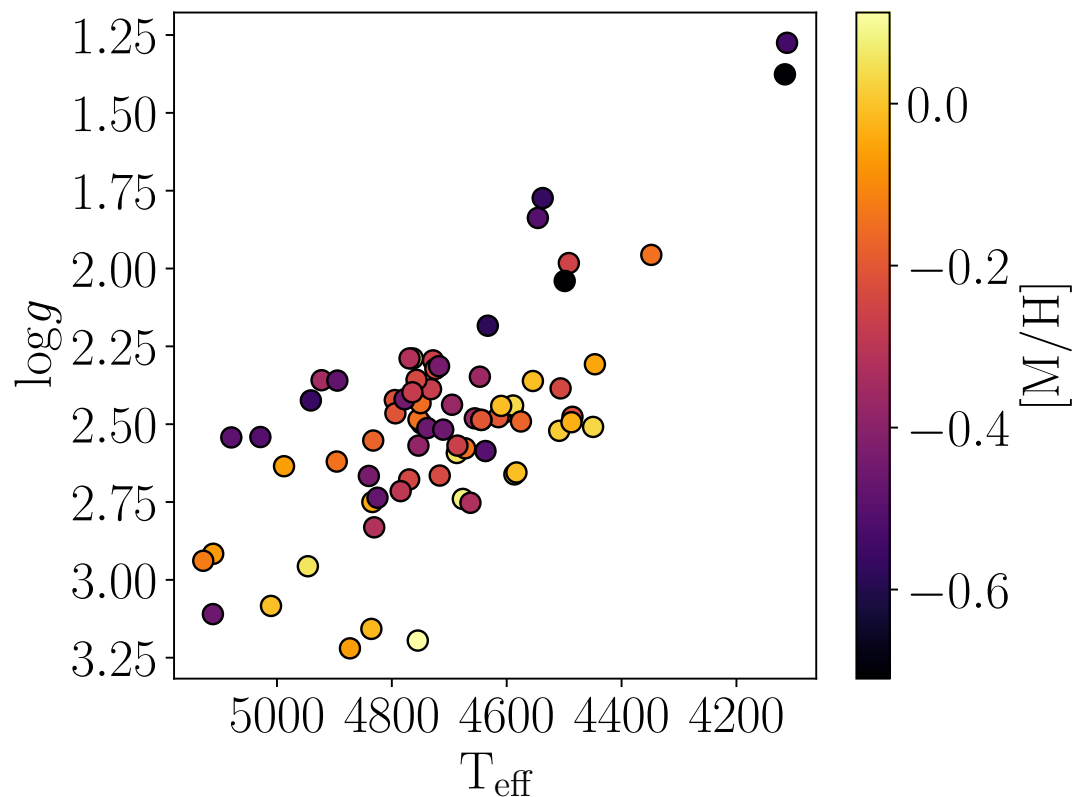
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- Radiative code: TURBOSPECTRUM
- MARCS atmospheric model
- *Gaia*-ESO line-list (v6)





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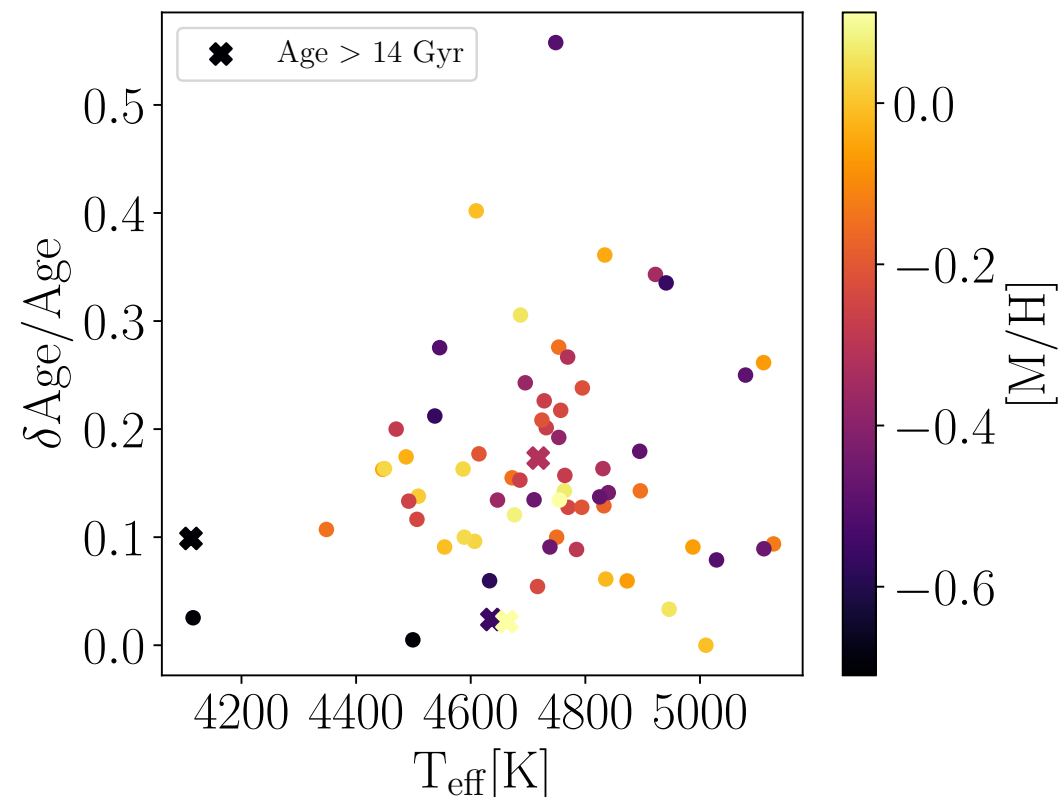
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AGES

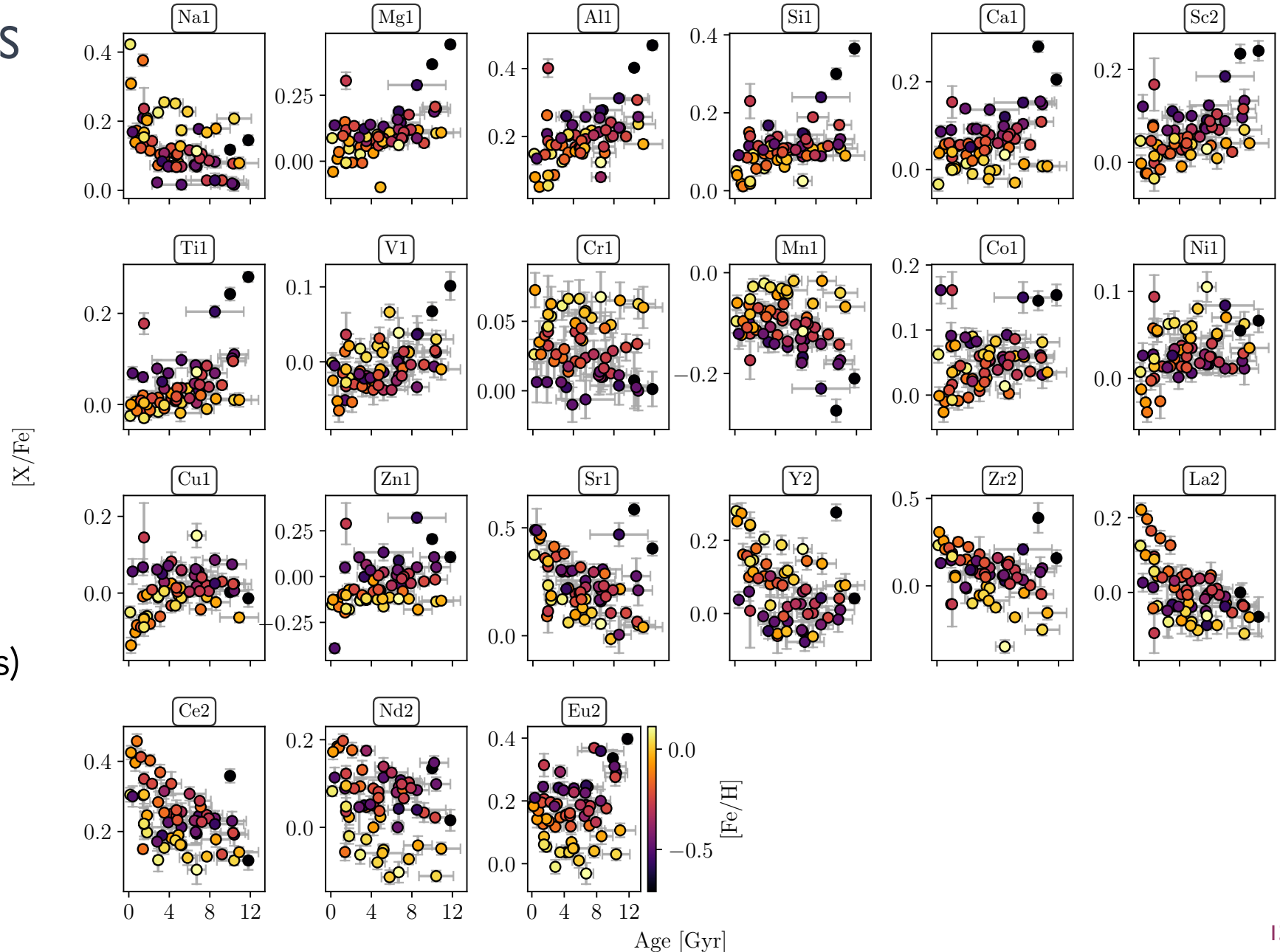
BASTA - BAYesian STellar Algorithm (Silva-Aguirre et al., 2015)
Grids of stellar models (BaSTI isochrones)

- asteroseismic information (ν_{max} , $\Delta\nu$) + stellar parameters
- Diffusion + overshooting



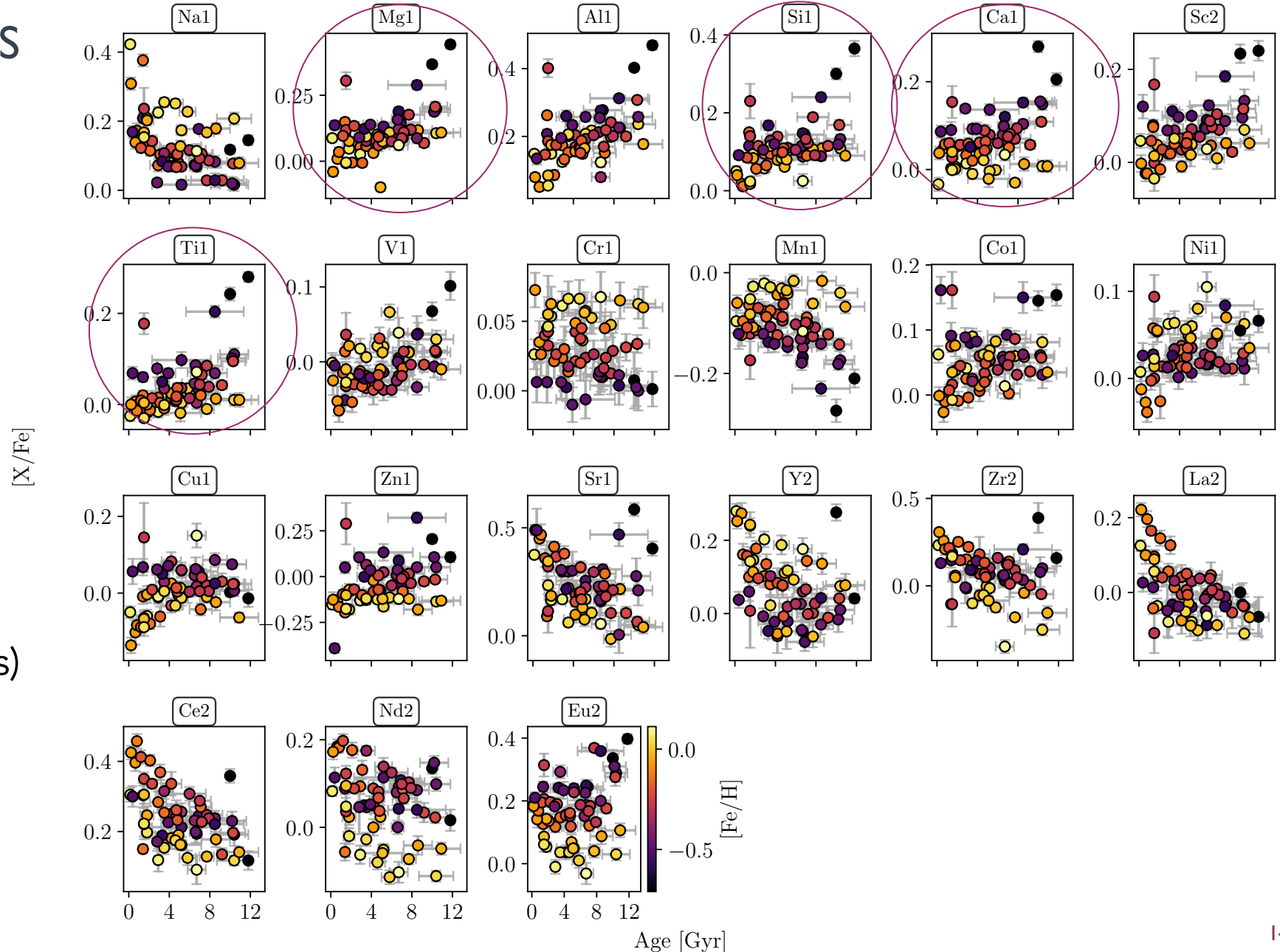
CHEMICAL-AGE TRENDS

- Stellar synthesis
 - Differential abundances with respect to the Sun
 - GES line
- α -elements
 - odd-z
 - iron-peak
 - n-capture (s and r process)



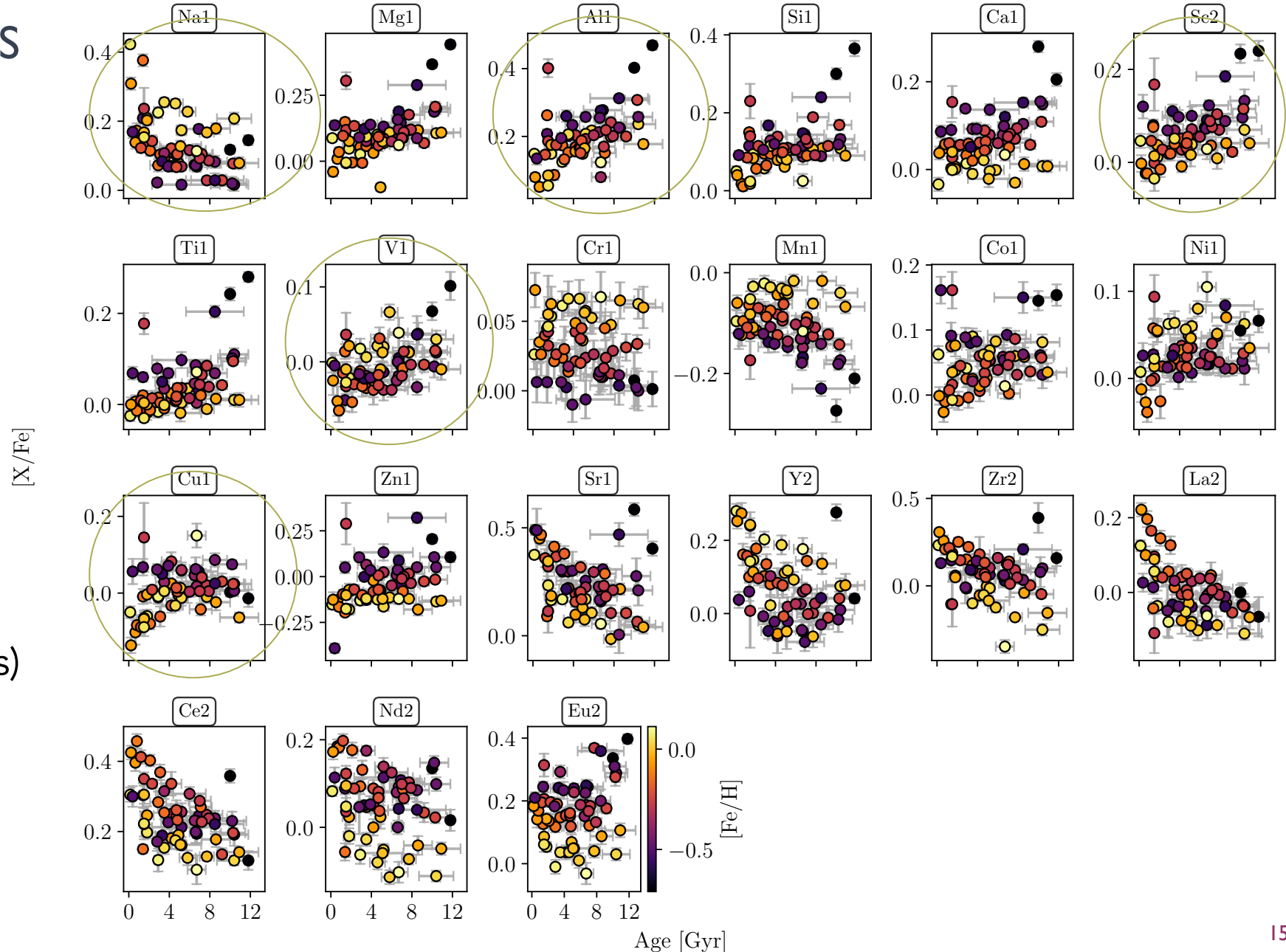
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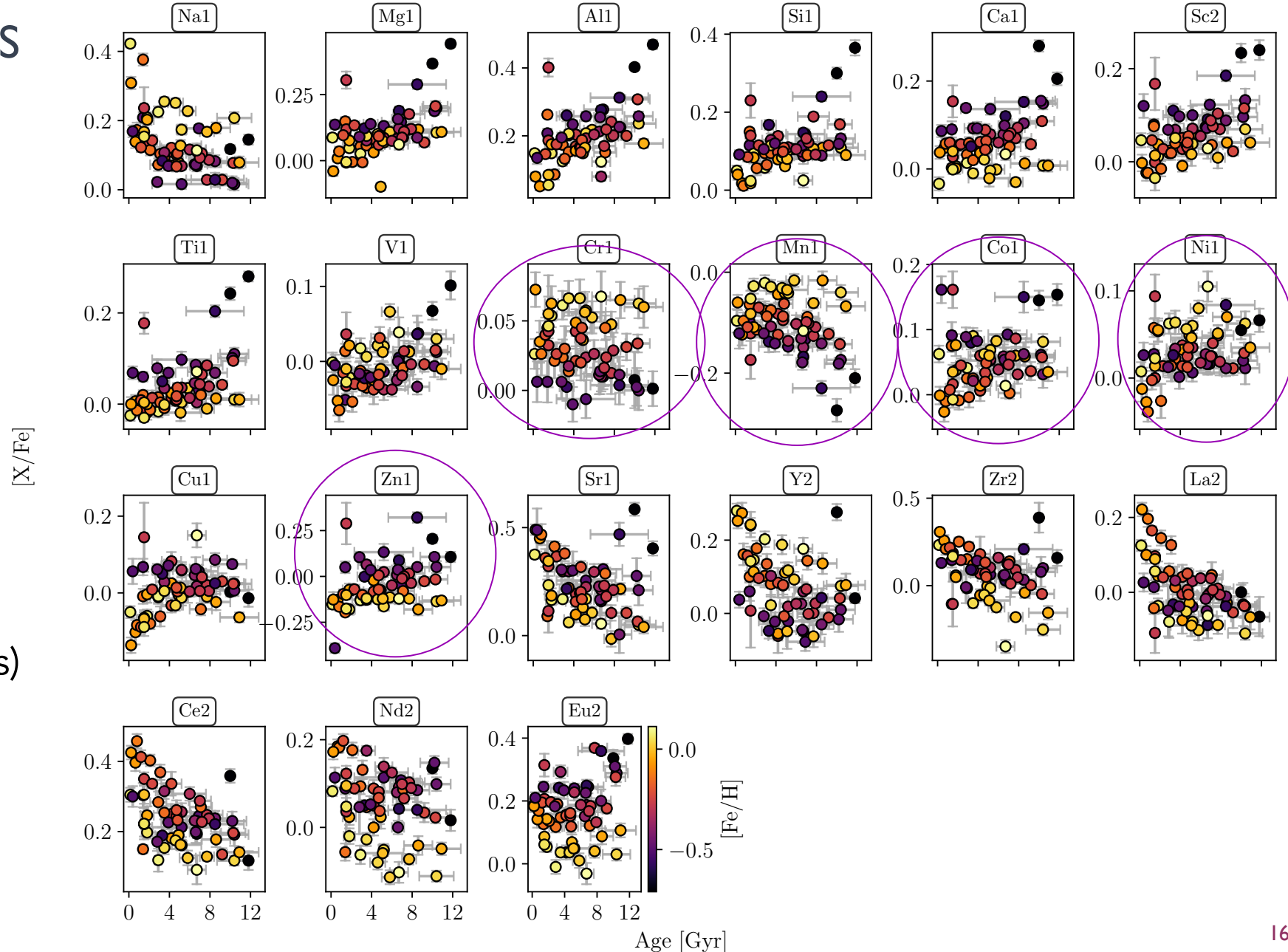
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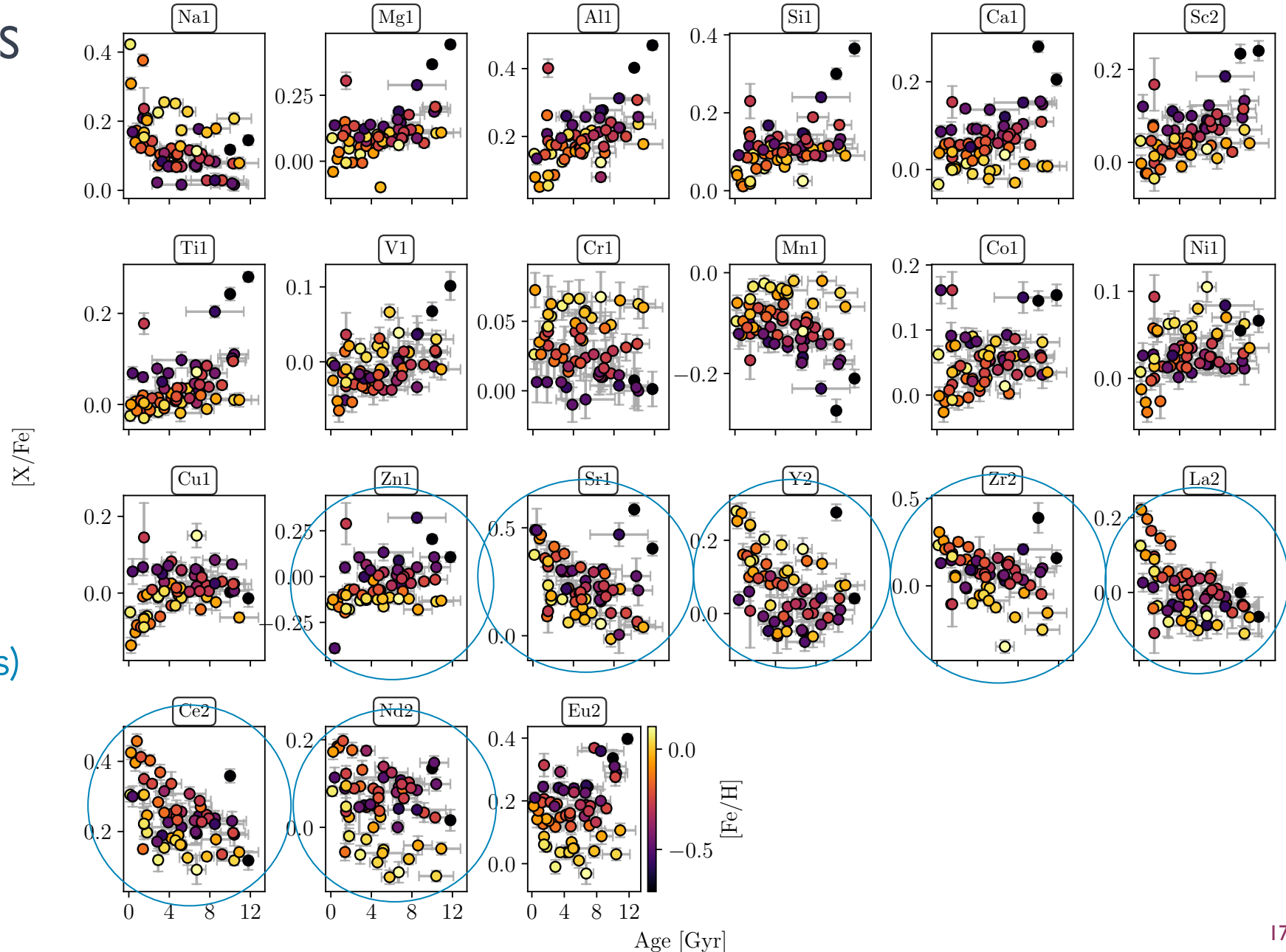
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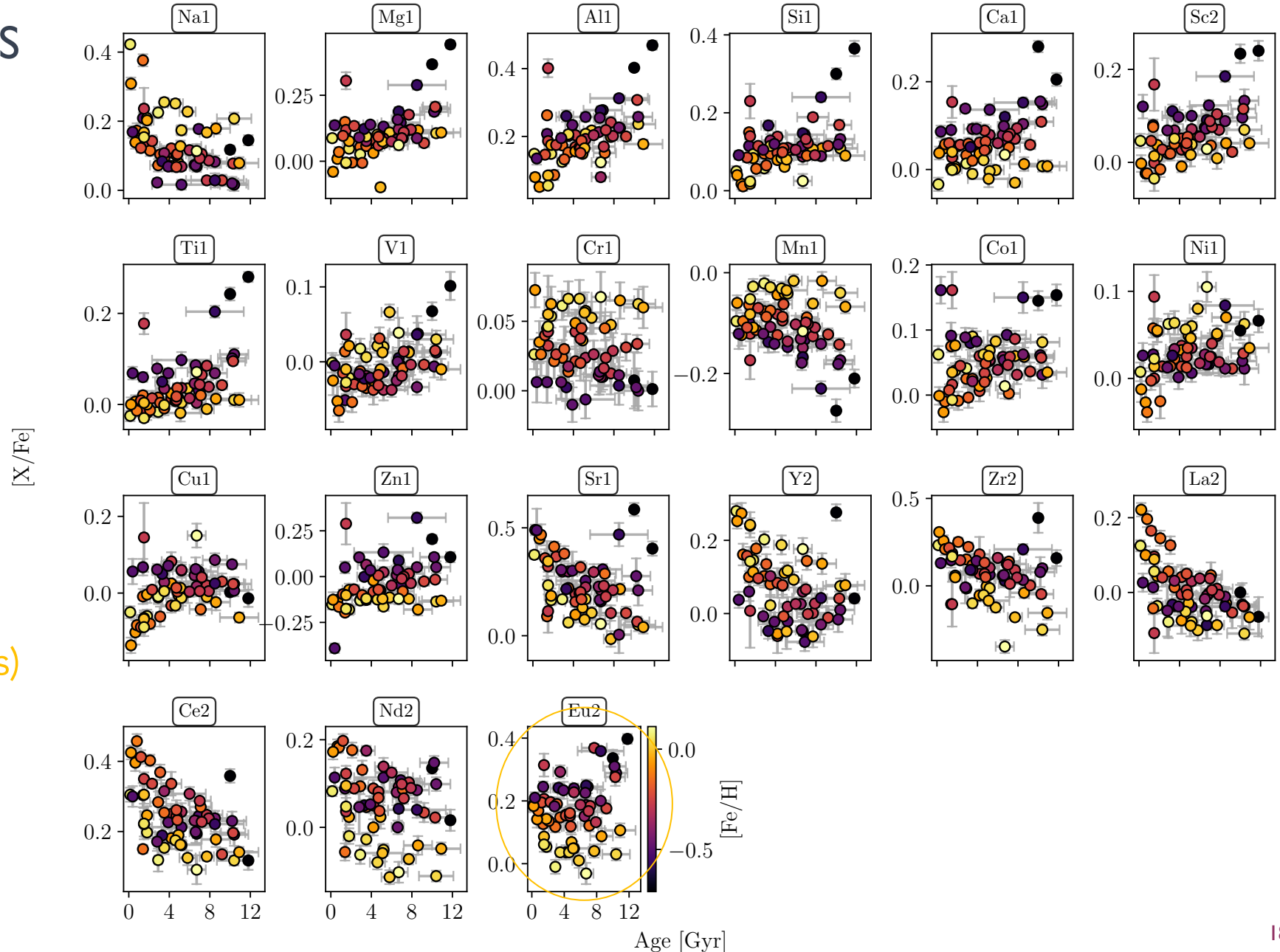
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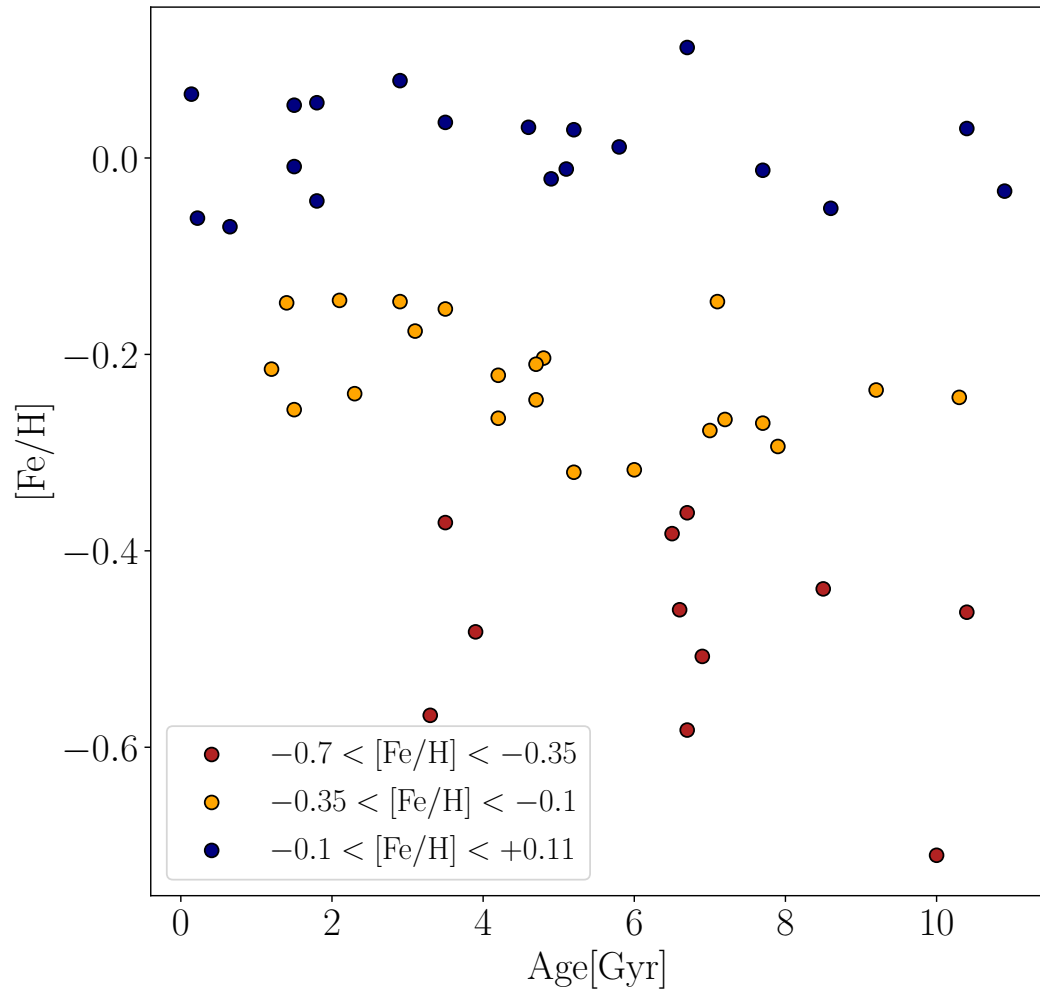
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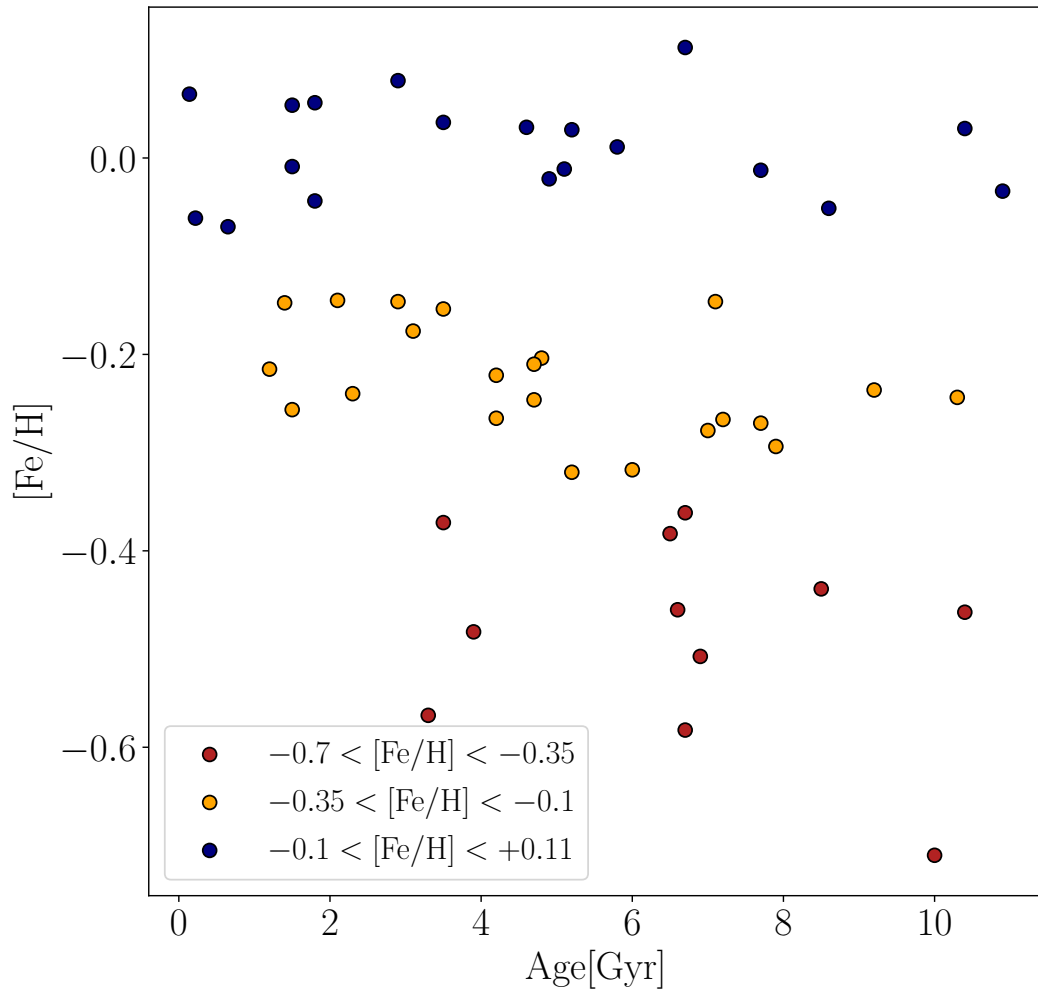
METALLICITY DEPENDENCE

3 metallicity bins

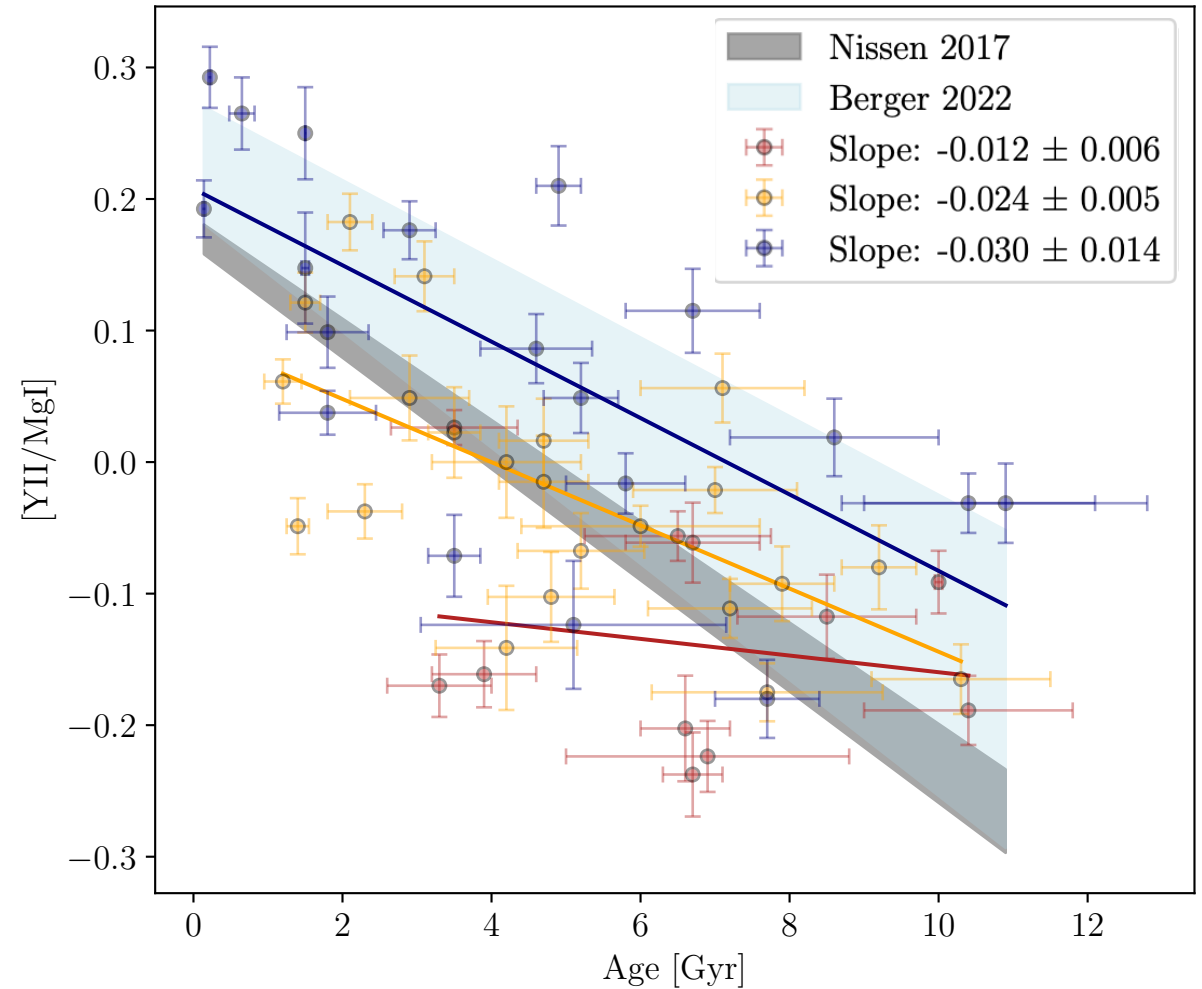


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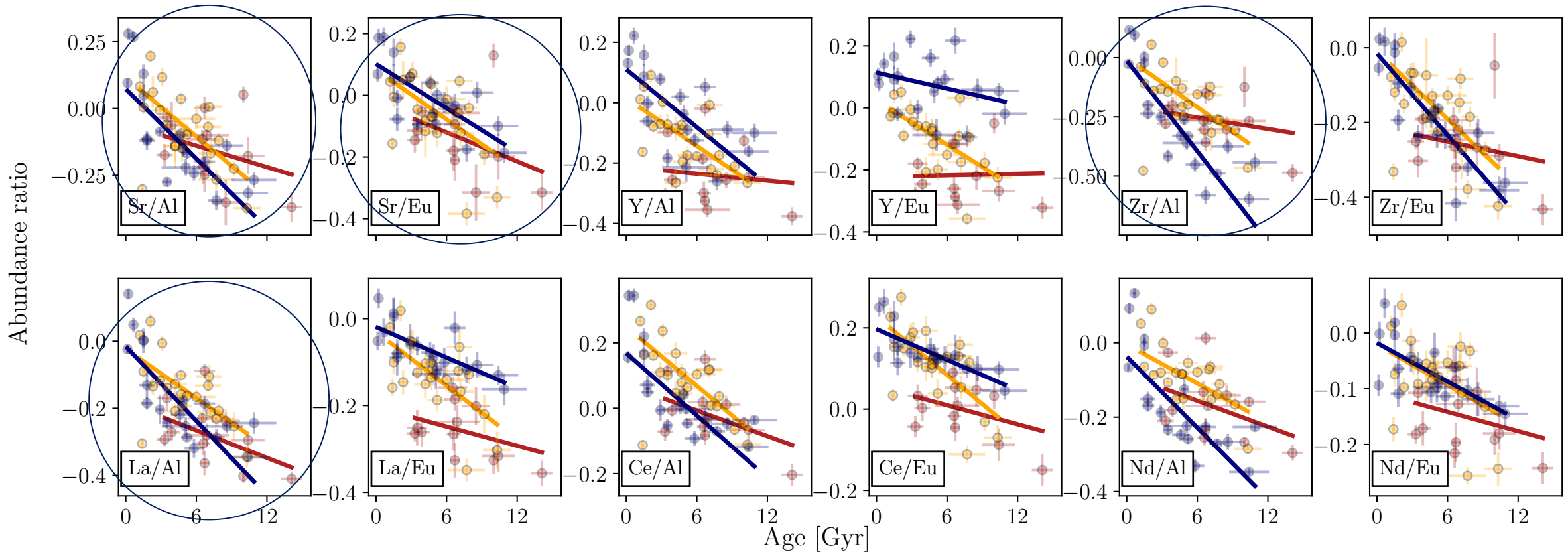


[s-processed/alpha] element



CHEMICAL TRAITS

[n-capture/Al] & [s/r] processed elements

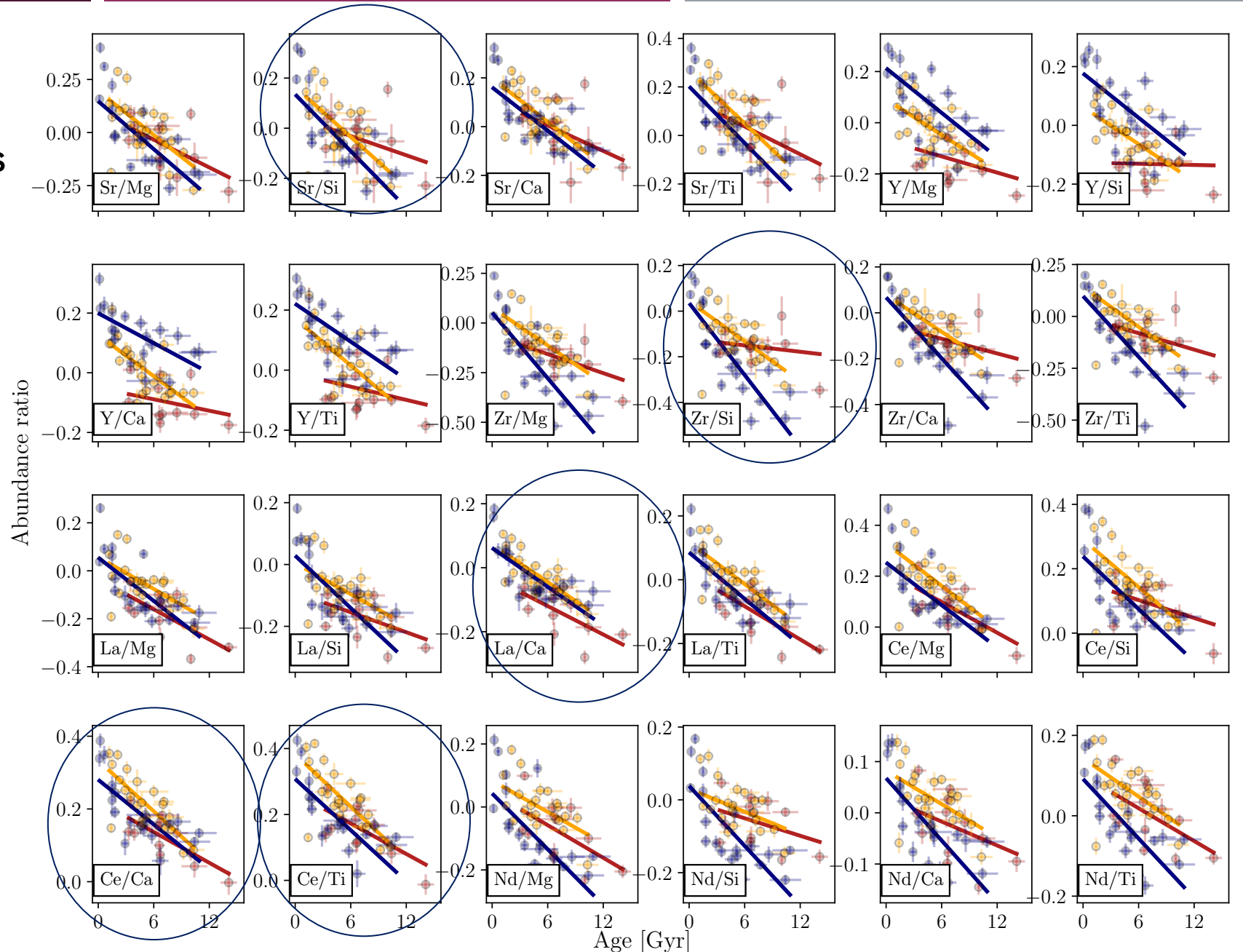


Loss of correlation for lower metallicities

CHEMICAL TRAITS

[n-capture/alpha] elements

- Trends \Leftrightarrow Different production timescales
- Influence of $[\text{Fe}/\text{H}]$
- Strong relations for Zr and Sr
- Different production paths

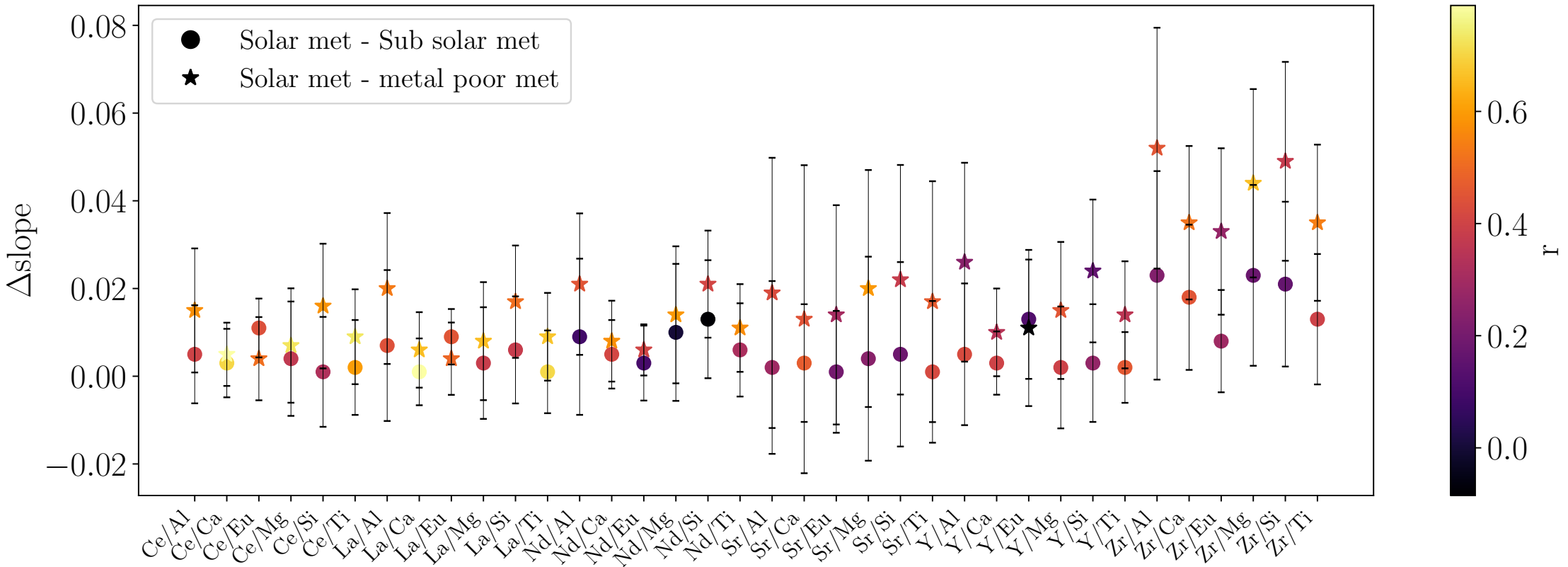


INFORMATIVE AND UNIVERSAL TRAITS

- Most informative traits (higher correlation)
- More universal (smaller Δ slopes)



Solar met: $-0.1 < [Fe/H] < +0.11$
 Sub solar: $-0.35 < [Fe/H] < -0.1$
 Metal poor: $-0.7 < [Fe/H] < -0.35$





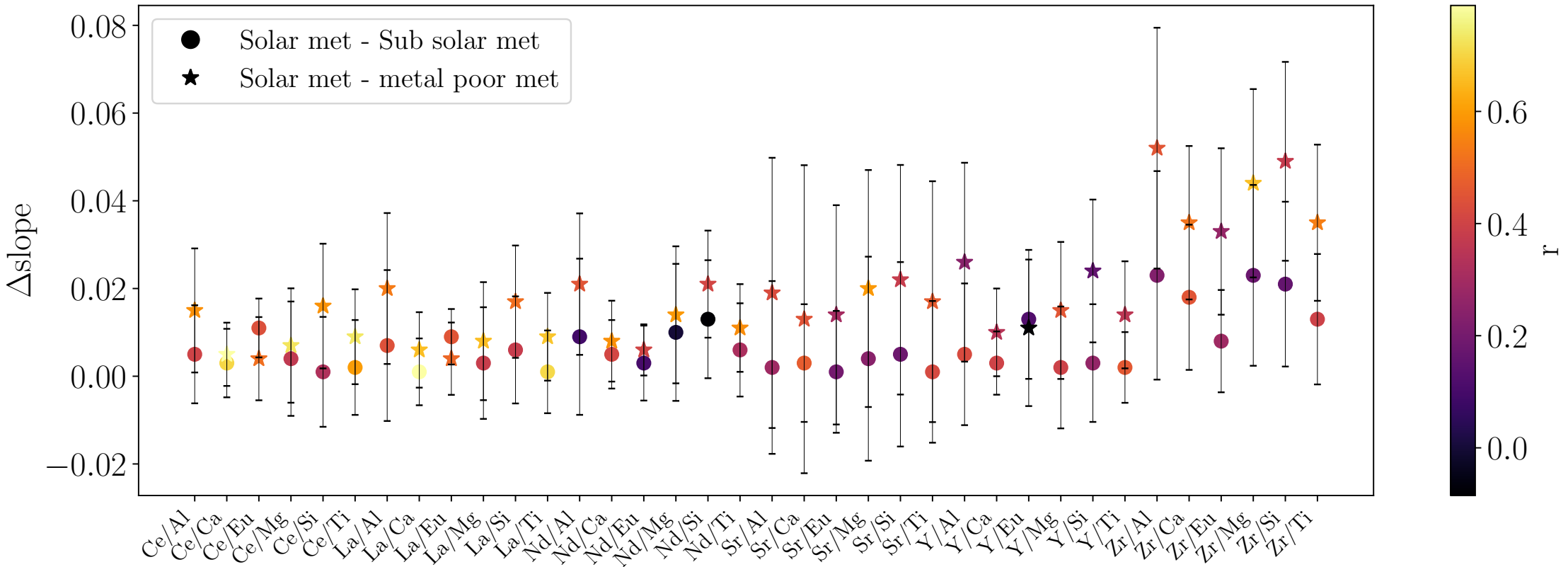
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THANK YOU FOR YOUR
ATTENTION!

