



# Fe, Zn and Mg stable isotope records of early differentiation and core formation of the ureilite parent body Stepan M. Chernonozhkin, Steven Goderis, Genevieve Hublet, Stefan Weyer,

Ingo Horn, Lidia Pittarello, Philippe Claeys, Vinciane Debaille and Frank Vanhaecke











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### **WHY UREILITES?**





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Ureilites are achondrite meteorites
Ultramafic olivine-pyroxen rocks
Interstitial carbon, Fe-Ni metal, sulfides
Olivine in contact with C is zoned
Metal blebs in olivine rims



## **WHY UREILITES?**



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- ► The Fo# is ~74 to 97 among ureilites
- Oxygen isotopes fall close to the CCAM line
- Oxygen isotopes correlate with Fo#
- Extraction of PGE with a S-bearing metallic melt
- Lithophile elements suggest a ~25% removal of silicate partial melt
- Depleted in volatile elements

Goodrich 1992, Rankenburg et al. 2008, Barrat et al. 2016, Ray et al. 2020



- Accretion of inhomogeneous PB from MgO-rich and FeO-rich chondrules
- Ureilites formed as ultramafic mantle restites
- The parent body did not reach the magma ocean stage
- Episode of reduction
- Episode of loss of volatiles (during the catastrophic break up or in a diffusion limited regime?)



● bulk URE, Brugier&Moynier ● Zn in bulk URE × CI, Brugier&Luck



Brugier et al. 2019, Moynier et al. 2010, Luck et al. 2005



## **SEPARATION OF METAL AND SILICATE MINERALS**

- Crushing to 150 250 μm
- Handmagnet
- Franz isodynamic separation
- Handpicking





"Silicate fractions" = cores of silicate grains

"Metal fractions" = mixture of interstitial metal and sulfide



### **Zn ISOTOPES IN UREILITES AND THEIR COMPONENTS**



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# **EVAPORATION OR METAL FRACTIONATION?**



× CI, Fischer-Godde, Moynier .....Batch melting

- The metal fractions of ureilites fall onto the batch melting curve in the δ<sup>66/64</sup>Zn – Ir/Pd space
- GRA95205 is the only "outlier", but it has likely underwent localized evaporation event
- CI chondrites can not be the primitive material for the ureilite PB



Brugier *et al.* 2019, Moynier *et al.* 2010, Chabot *et al.* 2003, Warren and Huber 2006, Rankenburg *et al.* 2008, Fischer-Gödde *et al.* 2010, Dhaliwal *et al.* 2017



## **Fe ISOTOPIC SYSTEMATICS**



#### The early core formation resulted in extraction of heavier Fe isotopes and lighter Zn isotopes with the S-rich metal melts



Craddock *et al.* 2011, Chabot *et al.* 2003, Warren and Huber 2006, Rankenburg *et al.* 2008, Fischer-Gödde *et al.* 2010, Dhaliwal *et al.* 2017

## Fe ISOTOPIC DISEQUILIBRIUM



◆ ureilites ● URE average

Likely metal silicate assemblage did not equilibrate after the extraction of metal melts





## **Mg ISOTOPIC SYSTEMATICS**



-0.4 -0.3 -0.2 -0.1 0 0.1 0.2 0.3  $\delta^{26/24} Mg_{DSM3}$ 

• bulk ureilites A URE silicate • URE olivine

van Kooten et al., 2017

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Mg isotopic signatures are heavier than Cl (?)
Fractionation during magmatic differentiation
Partial melting of silicate in the presence of magnesite
Kinetic isotopic fractionation during silicate partial melting
Preferential evaporation of lighter isotopes during heating



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# Fe, Mg ISOTOPIC OLIVINE PROFILE (GRA95205)



#### Zoned olivine which experienced significant reduction by carbon:

- Heavily reduced olivine grains are not in equilibrium with the silicate assemblage
- Complex superimposed lateral Fe and Mg isotope ratio profiles
- Cannot be explained solely by interdiffusion during cooling

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Diffusive cooling superimposed by metal-olivine re-equilibration (redox) reactions

## CONCLUSIONS

- **Heavier than chondritic Fe, Zn, and Mg** (?) isotopic signatures
- The isotope ratios show no correlation with proxies of nebular processes
- > Zn and Fe isotopic signatures are controlled by **metal melt fractionation** 
  - Heavier Fe and lighter Zn isotopes are extracted into the core
- > This process left the remaining **metal reservoir in disequilibrium with silicate**
- Metal differentiation is overprinted by effects of evaporation (GRA95205)
- **PB scale evaporation** fractionating Fe, Mg and Zn also cannot be dismissed
- The olivine grains, which experienced reduction, are significantly affected by kinetic isotopic fractionation

