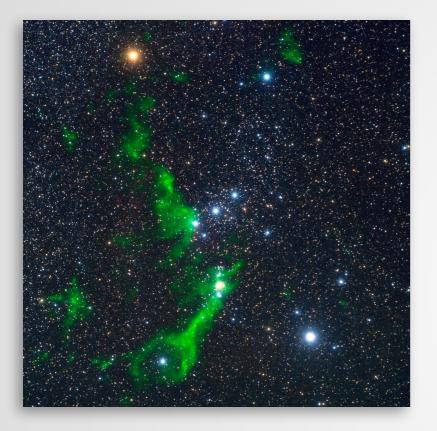
VISTA Star Formation Atlas







Stefan Meingast

João Alves (PI), Herve Bouy

Joana Ascenso, Amelia Bayo, Emmanuel Bertin, Anthony Brown, Jos de Bruijne, Jan Forbrich, Josefa Großschedl, Alvaro Hacar, Birgit Hasenberger, Jouni Kainulainen, Jens Kauffmann, Rainer Köhler, Karolina Kubiak, Kieran Leschinski, Marco Lombardi, Diego Mardones, Andre Moitinho, Karla Peña Ramirez, Monika Petr-Gotzens, Timo Prusti, Luis Sarro, Thomas Robitaille, Ronny Ramlau, Rainer Schödel, Paula Teixeira, Eleonora Zari, Werner Zeilinger

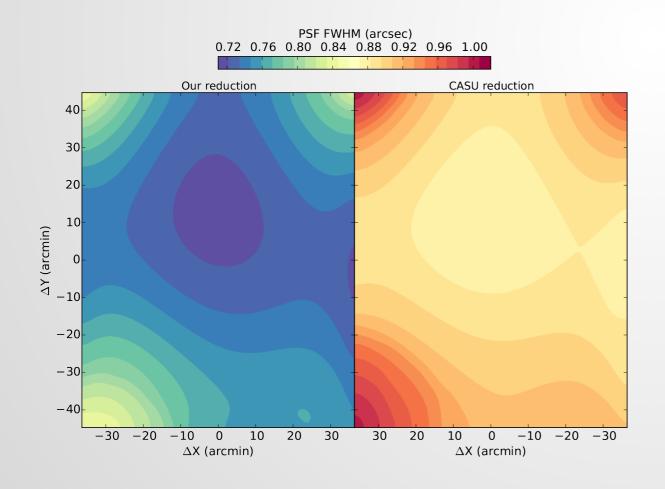


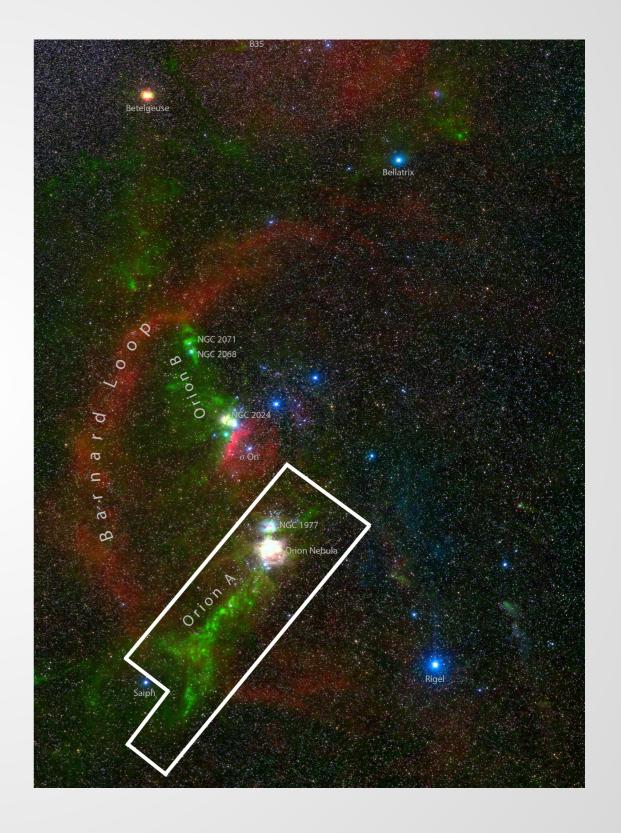
VISIONS - Stefan Meingast

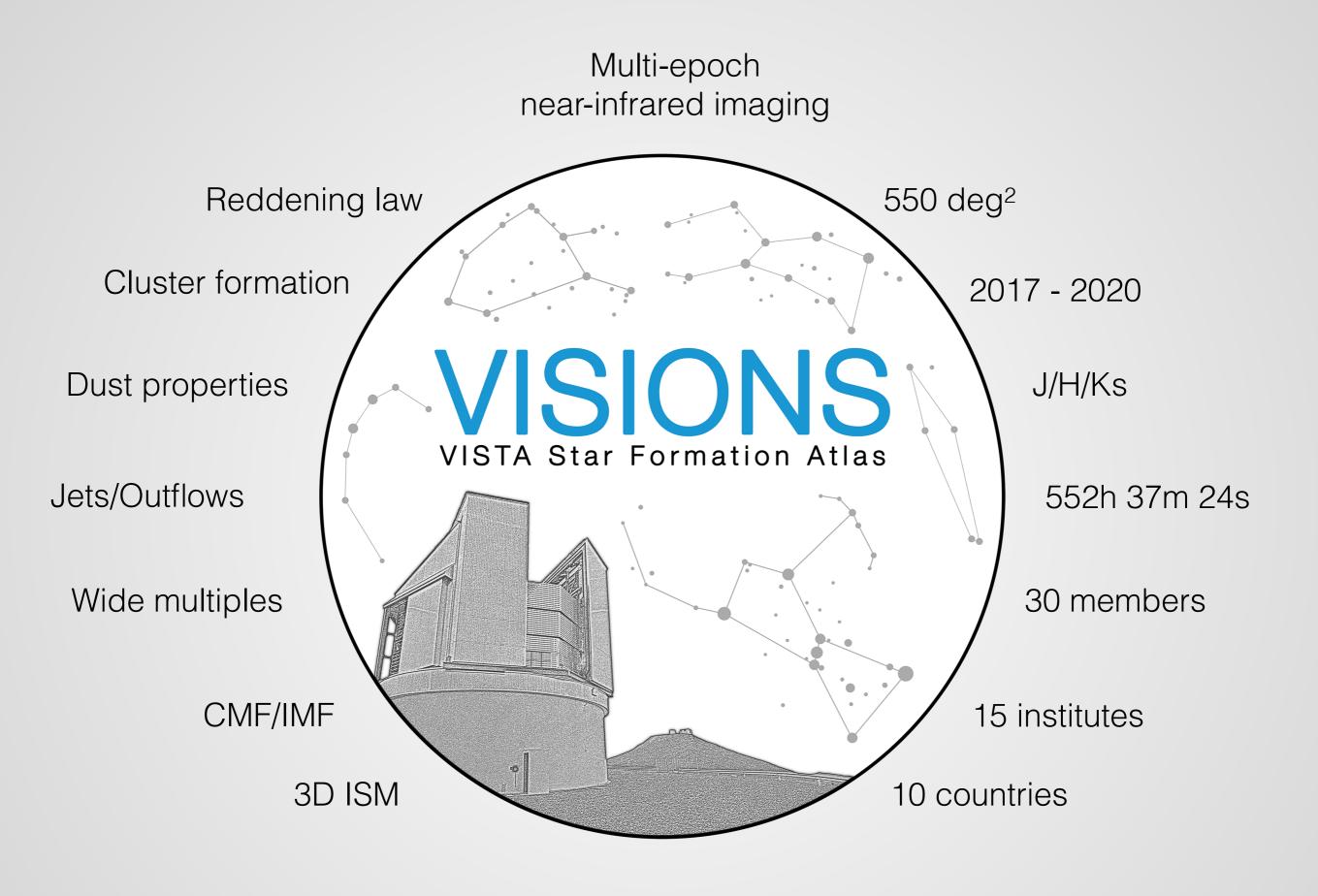
Star Formation from Cores to Clusters - Santiago 9.3.2017

VISIONS - How it started

- VISION (Meingast et al. 2016) 33h VIRCAM@VISTA J/H/Ks survey of Orion A (18.2 deg²)
- Data reduction improves pipeline 20% better image quality
- Completeness @19-20 mag (3-4 mag deeper than 2MASS)







VISIONS in a nutshell

o Wide survey

6 epochs proper motions complementary to VHS 7 - 10 yrs baseline

o *Deep* survey

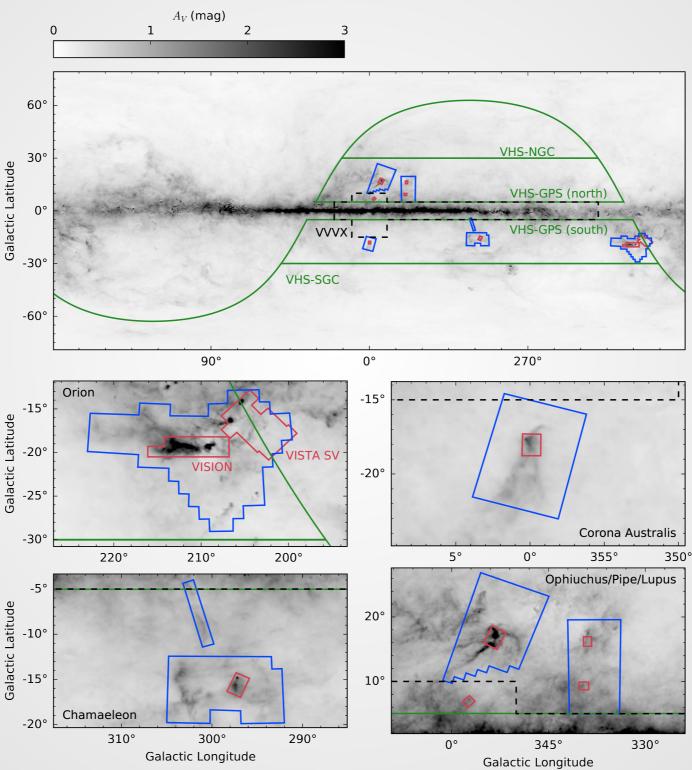
VISION-like Single epoch high S/N

• Control survey high S/N extinction-free

• P99

Ophiuchus, Lupus, Corona, Chamaeleon

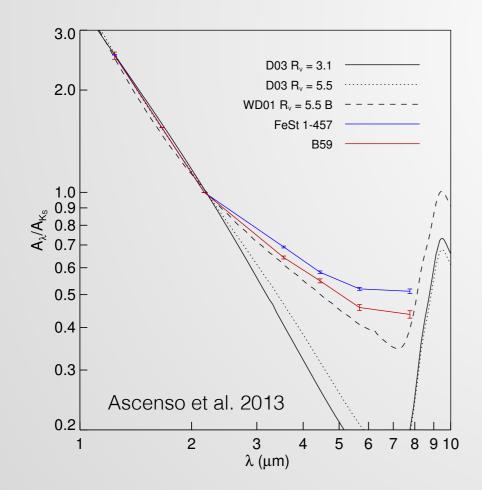
• DR1 end of 2018 Final DR 2020-2021

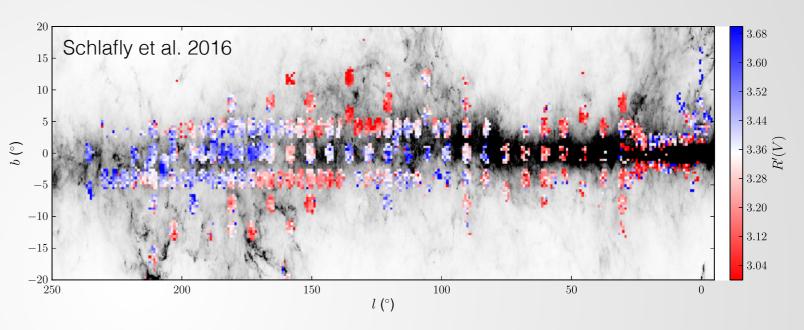


- Orion
 171 deg², 152h
- Chamaeleon 84 deg², 90h
- Corona Australis
 38 deg², 42h
- Ophiuchus 140 deg², 150h
- Lupus
 105 deg², 111h
- Pipe
 1.5 deg², 6h

Molecular cloud properties

- Wide-area R(V) variations Schlafly+ 2016, Nataf+ 2016
- Star-forming regions excluded Environmental dependency





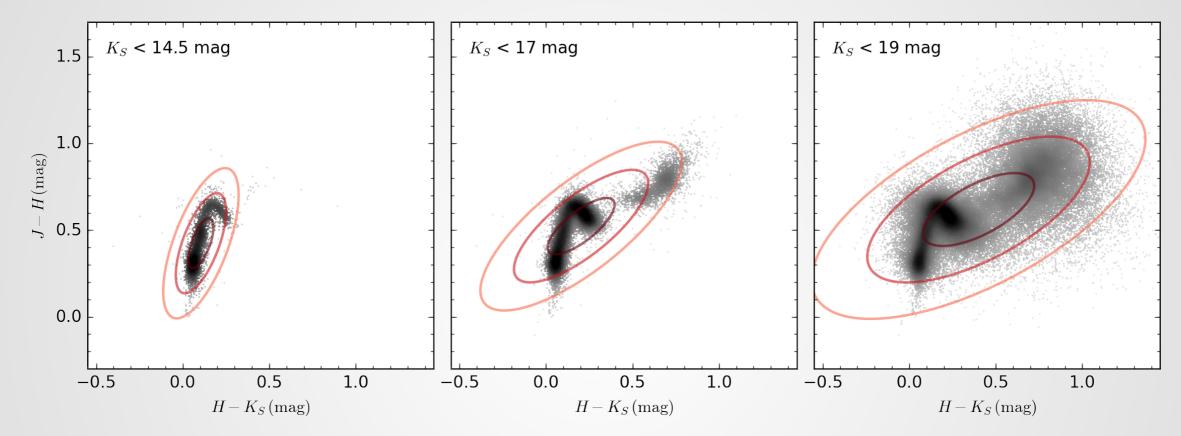
• Universal NIR/MIR extinction law? Star-forming vs. starless cores

• Origin of variations?

Grain growth, influence of feedback

• Synergy with Spitzer Comparing NIR and MIR

Estimating Extinction using Unsupervised Machine Learning (Meingast et al. 2017)



Dust thermal emission

- + large dynamic range
- depends on dust properties

Molecular line emission

- + high resolution
- optically thick emission

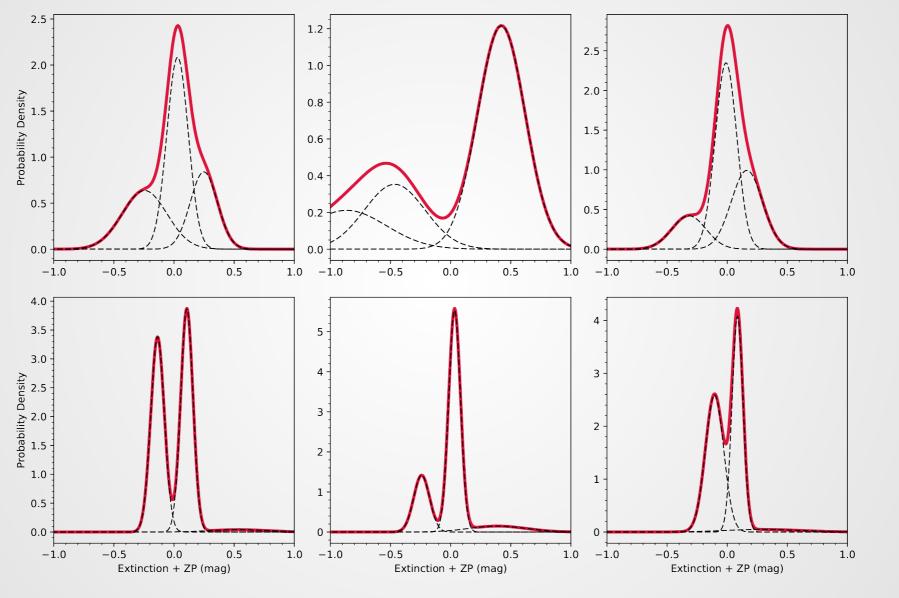
Dust extinction

- + independent of dust properties (NIR)
- number of background sources

$$E(m_1 - m_2) = (m_1 - m_2) - (m_1 - m_2)_0$$

= $(m_1 - m_{1,0}) - (m_2 - m_{2,0}) = A_{m_1} - A_{m_2}$

Estimating Extinction using Unsupervised Machine Learning (Meingast et al. 2017)



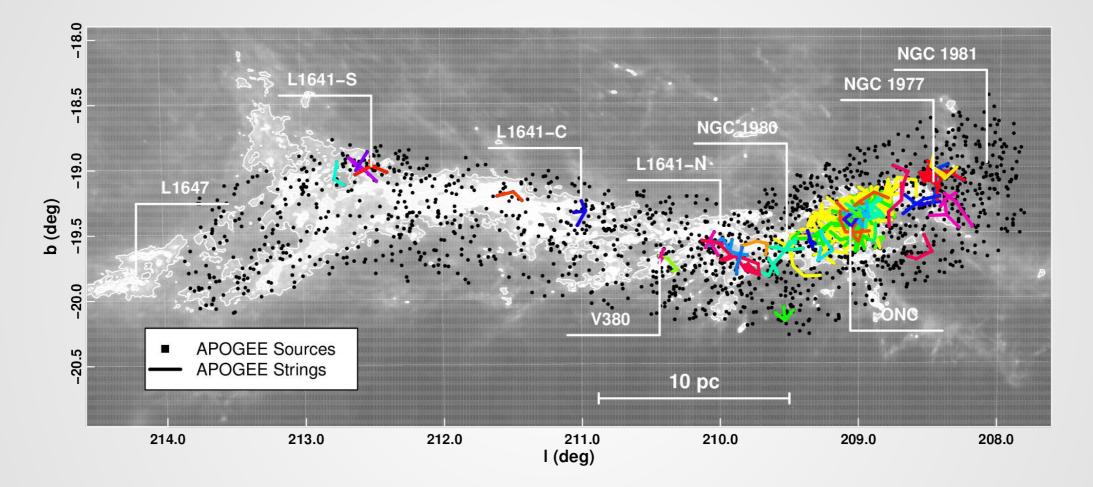
• Unsupervised Machine Learning

Statistically derived probability densities Arbitrary feature combinations • Open source Python code Fast and simple front-end

visions.univie.ac.at

ESO photo release, http://www.eso.org/public/news/eso1701/

Evolution of star-forming regions



Synergy with Apogee and Gaia
 <10% of Protostars in Gaia
 3D space motion

0.1-0.3 km/s transverse velocity

• Substructure at small scales Apogee strings RV correlation of YSOs and CO (Hacar et al. 2016)

VISIONS - Selected science cases YSO imaging



Wide multiple systems

YSO Morphology

Jets and outflows