## Probing the origin of VHE emission from M 87 with H.E.S.S., MAGIC, VERITAS, VLBA, & Chandra

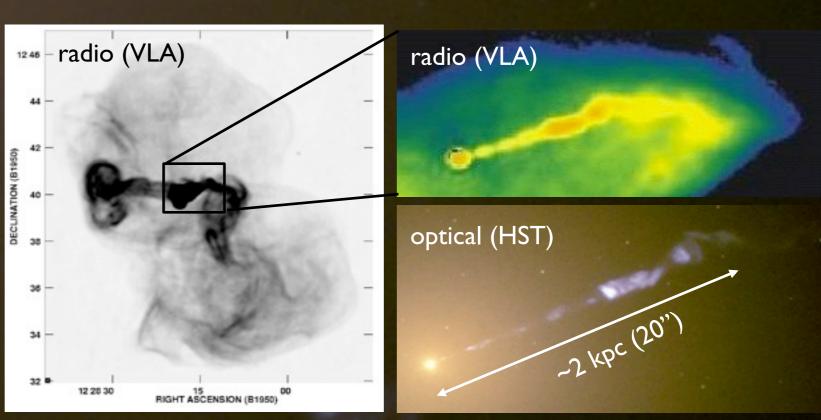
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for the H.E.S.S., MAGIC, and VERITAS Collaborations, the VLBA 43 GHz M 87 Monitoring Team, Chandra M 87 Monitoring Team

Cospar 2010 - July 18-24, 2010 - Bremen, Germany

### Radio Galaxy M 87

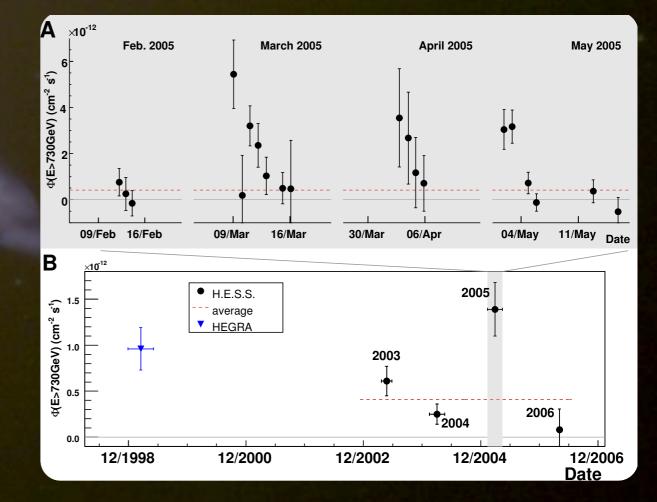
- Distance:
  ~16 Mpc
- Central BH:  $M_{BH} \sim 6 \times 10^9 M_{\odot}$
- Jet angle: ~30°
- Giant outer radio lobes
  0.2 x 0.2 deg
- Highly structured inner jet (knots) resolved in radio, optical, X-ray
- Apparent superluminal motion in jet
  relativistic particle population





### M 87 at very high energies

- First evidence for VHE emission 1998/99 (HEGRA)
- Detection by H.E.S.S. (2003-2006) Signal confirmed by MAGIC & VERITAS
- VHE flare and fast variability (2005)
  - Discovery of vast variability on time-scales of days strongly constrains emission region R < 5×10<sup>15</sup> cm ~ 5×δ×R<sub>S</sub>
- Detected at HE by Fermi (2008/9)



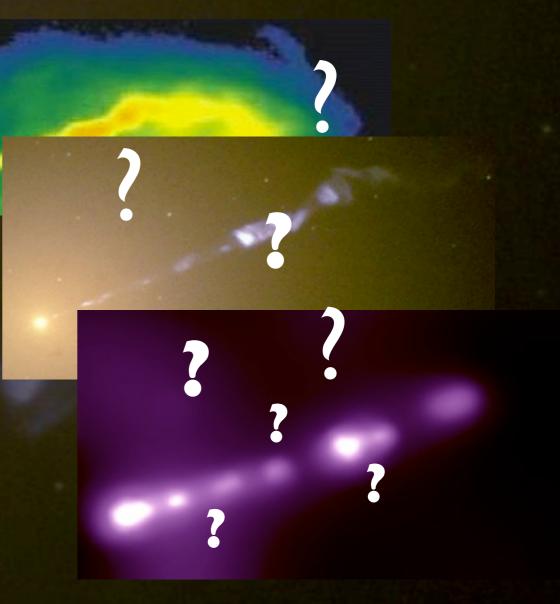
### Origin of the VHE emission?

#### • Possible mechanism/sites for VHE emission

- Dark matter annihilation #
- ▶ CR in M 87 🖊
- Intra-cluster gas #
- Large scale jets #
- Inner/Outer-jet (Hadronic/Leptonic)
- Knots? HSTI?
- Vicinity of black hole
- **)** ...

#### Further observations to ..

- .. confirm short-term VHE variability
- .. pin down location of VHE emission



#### 2008 Campaign - Instruments



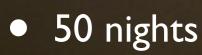
#### H.E.S.S., MAGIC, & VERITAS collected >120 h of data from Jan to Jul 2008

Acciari et al. 2009, Science Express, Jul 2

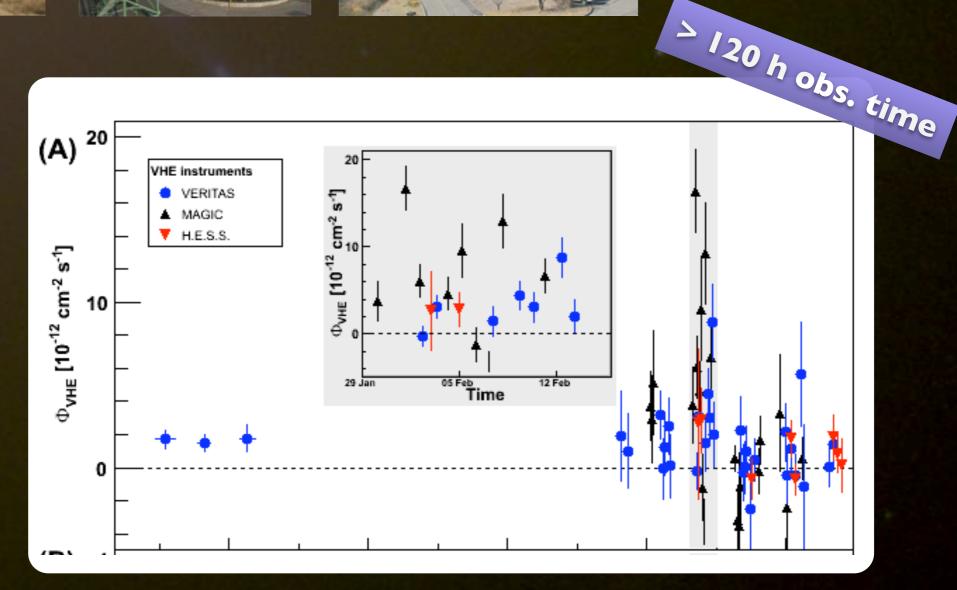
### 2008 Campaign - VHE Results







- Feb 08: increased VHE activity
- MAGIC and VERITAS detected rapid flares with t<sub>var</sub> ≃ I day

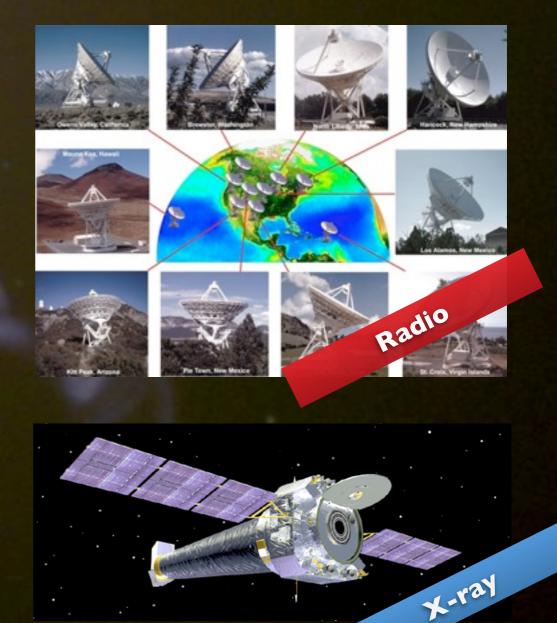


VERITAS

Albert et al. (MAGIC Collaboration) 2008, Beilicke et al. (HESS/MAGIC/VERITAS) 2008

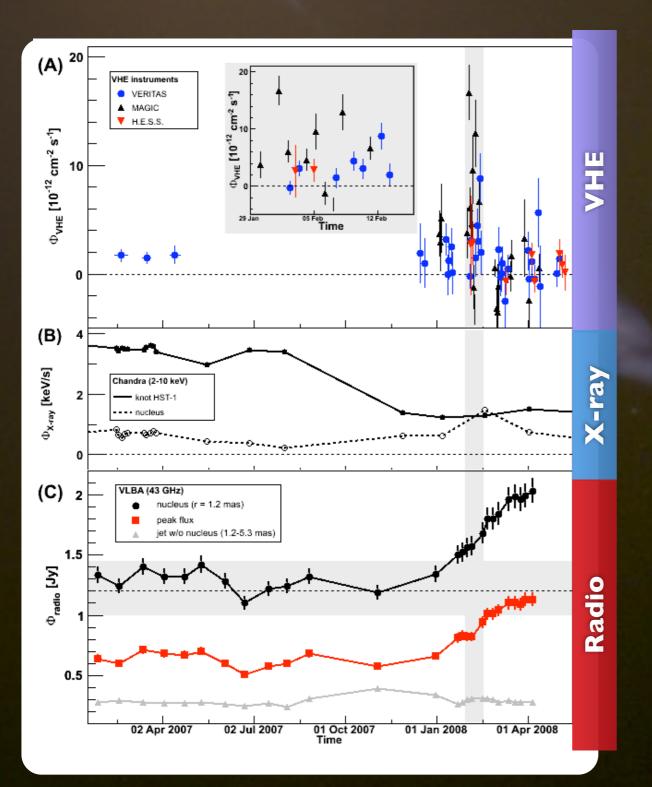
### 2008 Campaign - Instruments

- VLBA observations at 43 GHz C.Walker et al. 2009
  - Resolution: 0.21x0.43 mas (30x60 R<sub>s</sub>)
  - Scale: <pc im direct vicinity of the BH</p>
  - Movie project: 13 obs. in 2007, every three weeks; 14 obs. in 2008, every 5 days
- Chandra observations in X-rays since 2002
   D. Harris et al. 2007/8/9
  - Scale: kpc 🗰 inner jet
  - Every 6 weeks



Acciari et al. 2009, Science Express, Jul 2

### 2008 Campaign - MWL Results

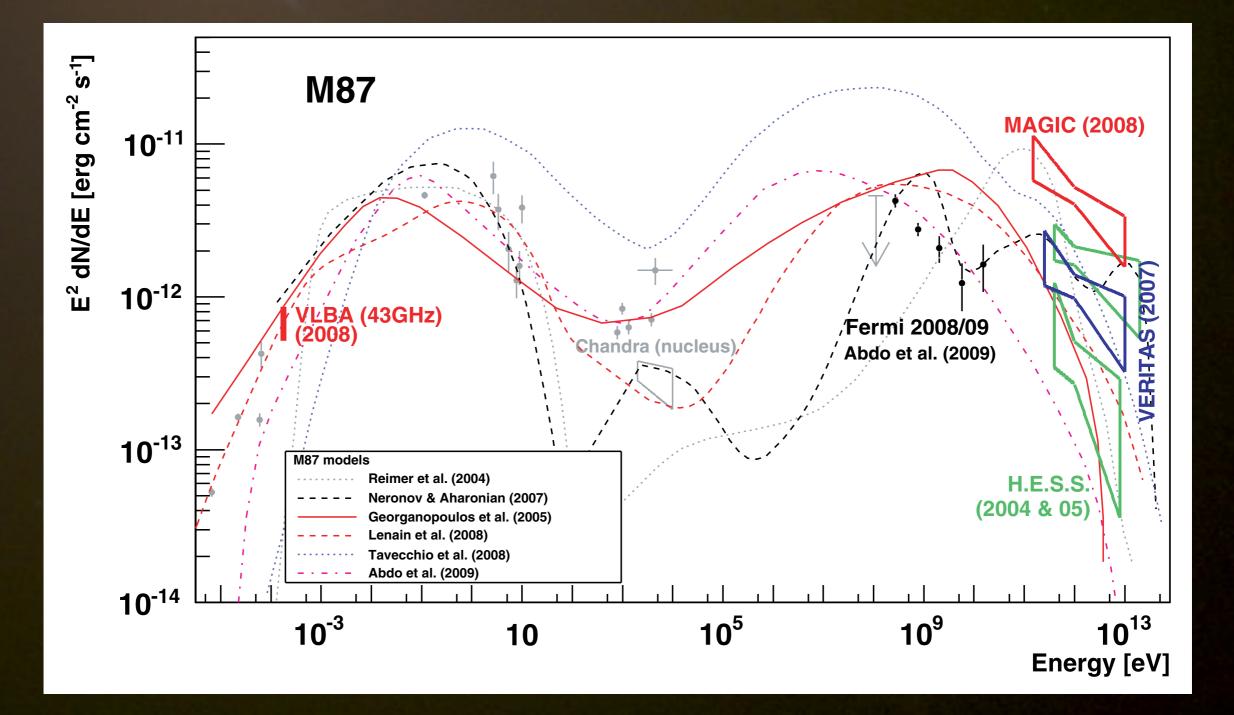


- **VHE** activity in Feb '08
- X-ray: core flux increases
- Radio: inner core flux increases coinciding with the VHE flare

Strong indication for VHE emission from the direct vicinity of the SMBH

Acciari et al. 2009, Science Express, Jul 2

#### M 87 - SED



Beilicke 2010, IJMPD, 19, 06

### 2010 Campaign

- Joint VHE monitoring campaign by MAGIC & VERITAS with tied in ToO observations triggered by a **VHE high state**
- Ist trigger by MAGIC in Feb. (ATel #2431)
  - Higher flux state but no strong flare
  - EVN follow up observations (see next talk by Gabriele Giovanni)
- 2nd trigger by MAGIC & VERITAS in April (ATel #2542)
  - H.E.S.S. started observation Apr. 9 and could confirm high state
  - Full MWL campaign triggered
    - Chandra (D. Harris), Swift (B. Giebels), RXTE, Fermi, VLBA (C. Walker et al.), EVN, optical polarimetry (B. del Almeida), ...



to observe M87 and multi-frequency observations are strongly encouraged.

are encouraged.

#### Summary & Conclusion

#### • Radio galaxy M 87

- Laboratory to probe the origin of VHE emission from SMBH and relativistic jets
- 2008: H.E.S.S., MAGIC, VERITAS + VLBA, Chandra
  - VHE flare detected, coinciding with rise in radio core and enhanced X-ray core flux
  - Indication for origin of VHE emission very close to the SMBH (<100Rs)</p>
- 2010: MAGIC, VERITAS, + ToO (H.E.S.S., VLBA, Chandra, ...)
  - VHE high state (Feb) and flare (April) detected (ATel #2431,#2542)
  - MWL campaign triggered (Radio, X-ray, optical polarimetry, ...)
  - Analysis ongoing

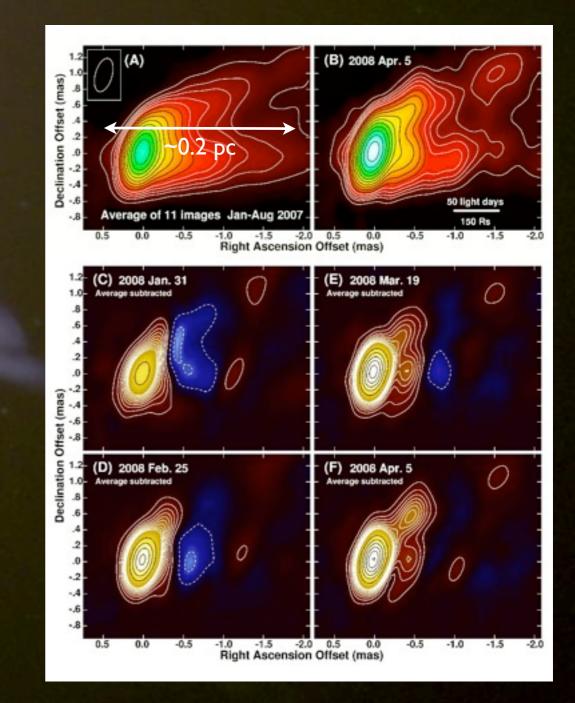
# Backup

#### VLBA radio observations



#### • VLBA observations at 43 GHz

- Movie project: 13 shots in 2007, every three weeks; 14 shots in 2008, every 5 days
- Resolution: 0.21x0.43 mas
  ~30x60 R<sub>s</sub> !!!



Acciari et al. 2009, Science Express, Jul 2

C.Walker et al. 09