

### Stellar outflows with the NOT





Physics of Extreme Massive Stars

Marie-Curie-RISE project funded by the European Union

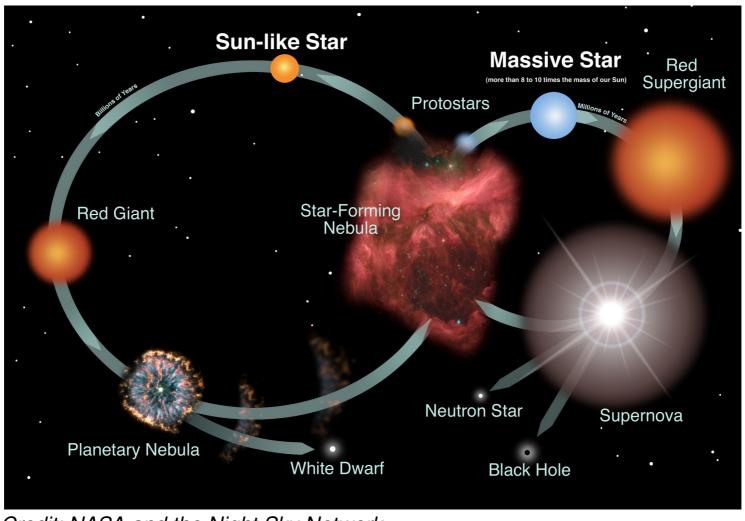


Tiina Liimets (ASU, Czech)

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B. Balick (UW, USA)
M. Kraus (ASU, Czech)
D. Jones (IAC, Spain),
A. Djupvik (NOT)

# It all starts with a nebula and ends with a nebula!

Most of stars are binaries!

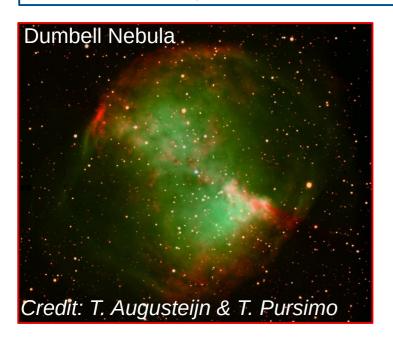


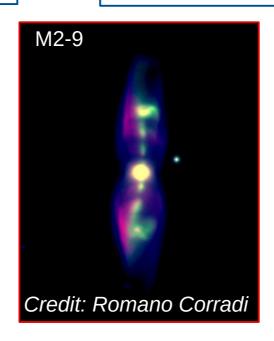
Credit: NASA and the Night Sky Network

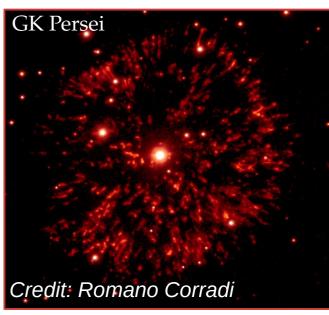
### WHY study stellar outflows?

### **Astronomy in real time!**

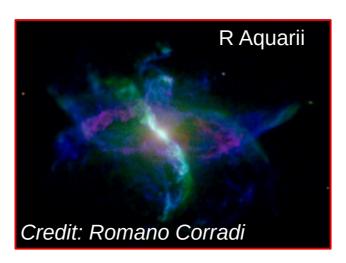
Proper motions! + spectra = 3D

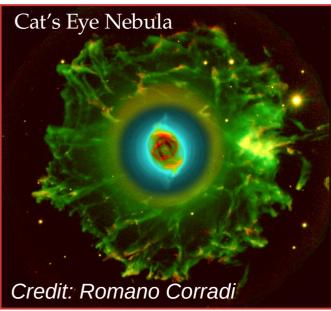






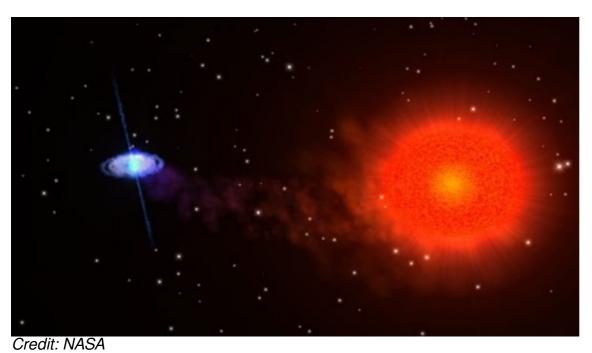






Most of the images from the NOT archive

### Symbiotic nebulae



- ★ Interacting wide binary system: white dwarf + red giant
- ★ Mass transfer mechanisms: stellar winds, wind-Roche-lobe overflow.
- ★ Nebulosities: stellar winds, jets, nova-like explosions,

★ Symbiosis: absorption lines (red giant)

nebular emission lines (hot component)

### Symbiotic nebulae: Legacy of the NOT

~ 30 yrs of NOT

3 generations of astronomers.

"A narrowband imaging survey of symbiotic stars"

R.L.M. Corradi, E. Brandi, O. E. Ferrer, and **H. E. Schwarz** A&A, 1999, 343, 841







Survey of symbiotic nebulae: <u>NOT was the main telescope to observe the Northern targets.</u> Before: only 3 known symbiotic nebulae (2 discovered by Hugo and Romano)

During 1<sup>st</sup> observing night at the NOT, 3 more nebulae were discovered!

Result: enlarged ~4 times the known extended nebulae around Symbiotic stars.

# Symbiotic binary R Aquarii

A&A 612, A118 (2018) https://doi.org/10.1051/0004-6361/201732073 © ESO 2018



#### New insights into the outflows from R Aquarii

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#### PROCEEDINGS OF SCIENCE

### **Nebulosities of the Symbiotic Binary R Aquarii - A Short Review**

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#### Romano M. L. Corradi

GRANTECAN, Cuesta de San José s/n, E-38712, Breña Baja, La Palma, Spain

NOT: [OI] [OII] [OIII]

Distance: 180 pc Liimets+ 2018 R Aquarii
Cool Mira + hot WD

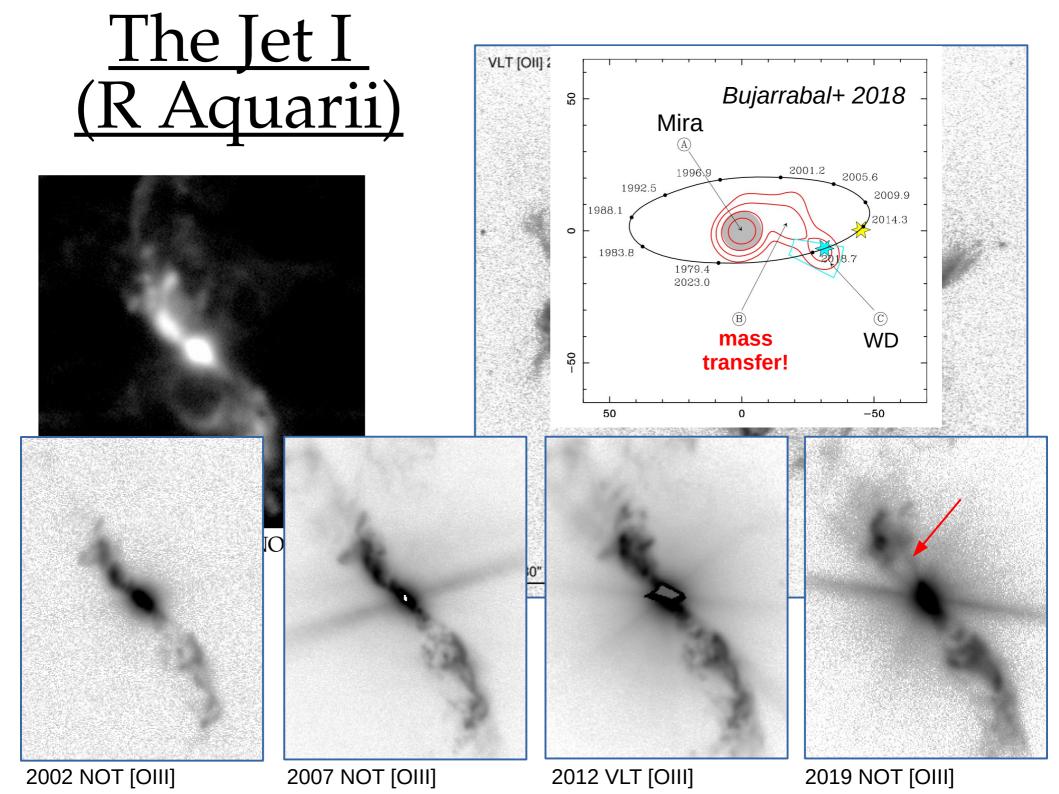
Jet

Hourglass nebula ~670 yrs Liimets+ 2018 ~300 to 0 yrs Liimets+ 2018 Schmid+ 2018

MAIN MOTIVATION: Jet

- Continuous/clumpy?
- Alternates between poles?

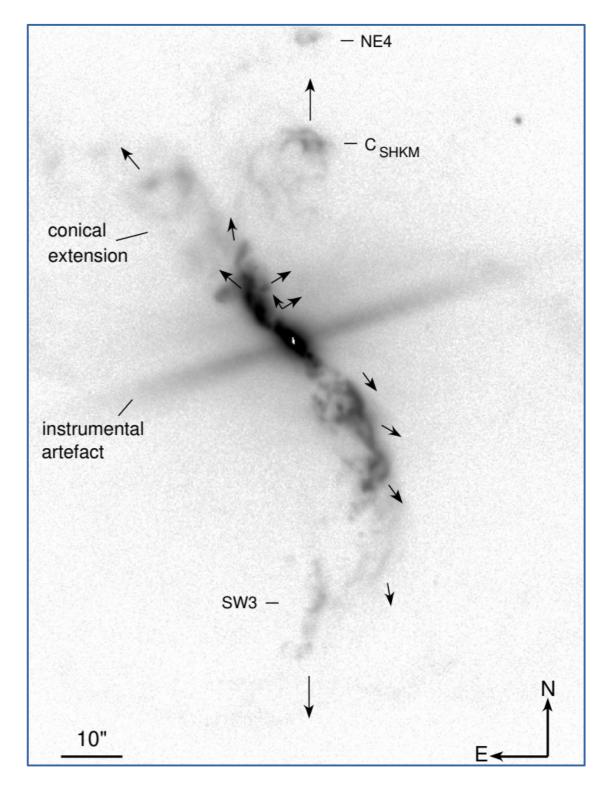
3 decades of IMAGING (mostly NOT!) Hα+[NII], [OIII]5007Å, [OII]3727Å, [OI]6300Å



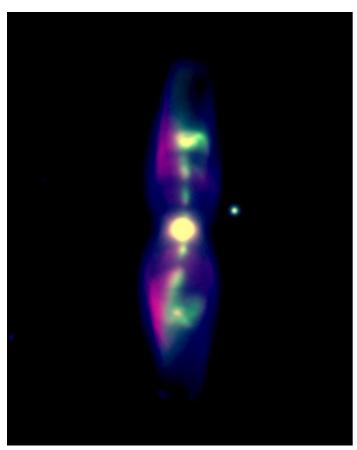
## The Jet II (R Aquarii)

- **★**Overall ballistic expansion
- ★Fast moving lateral features. Up to 900 km/s
- ★The evolution of the jet is complex:
  - physical matter moving,
  - changing ionisation,
  - changing illumination,
  - shocks,
  - precession, ...

Liimets et al. 2018, A&A, 612, 118 Liimets et al. 2021, Pos-SISSA, 41



# M 2-9 (Butterfly Nebula) Symbiotic/planetary nebula?



Credit: R. Corradi (NOT: Ha Ha [OIII]) 2006

~4 decades of IMAGING (large part with the NOT!) Hα+[NII], [OIII]5007Å

Astron. Astrophys. 319, 267–273 (1997)

ASTRONOMY AND ASTROPHYSICS

#### M 2-9: moving dust in a fast bipolar outflow\*

Hugo E. Schwarz<sup>1,2</sup>, Colin Aspin<sup>1,3</sup>, Romano L.M. Corradi<sup>4</sup>, and Bo Reipurth<sup>2</sup>

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- <sup>3</sup> Joint Astronomy Centre, 660 A'ohoku Pl., Hilo, Hawaii 96720, USA
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Received 29 April 1996 / Accepted 30 July 1996

THE ASTRONOMICAL JOURNAL, 119:1339–1344, 2000 March
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#### THE EVOLVING MORPHOLOGY OF THE BIPOLAR NEBULA M2-9

SEAN DOYLE<sup>1</sup> AND BRUCE BALICK<sup>2</sup>
Department of Astronomy, University of Washington, Box 351580, Seattle, WA 98195

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Instituto de Astrofísica de Canarias, Calle Vía Láctea, E-38200 La Laguna, Tenerife, Spain

AND

H. E. SCHWARZ<sup>4</sup>

Nordic Optical Telescope, Apdo. 321, E-38780 Santa Cruz de la Palma, Spain Received 1999 October 6; accepted 1999 December 3

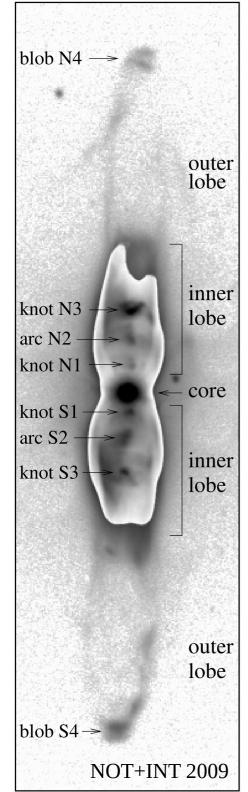
A&A 529, A43 (2011) DOI: 10.1051/0004-6361/201016361 © ESO 2011 Astronomy Astrophysics

#### The evolution of M 2-9 from 2000 to 2010\*,\*\*

R. L. M. Corradi<sup>1,2</sup>, B. Balick<sup>3</sup>, and M. Santander-García<sup>1,2,4</sup>

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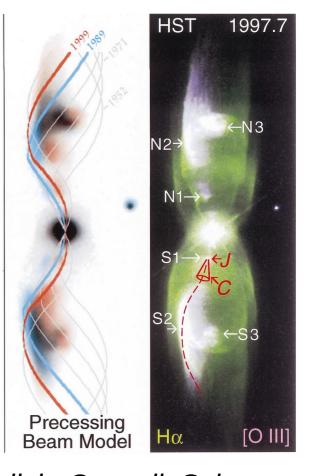
Received 20 December 2010 / Accepted 4 February 2011



## M 2-9: main results

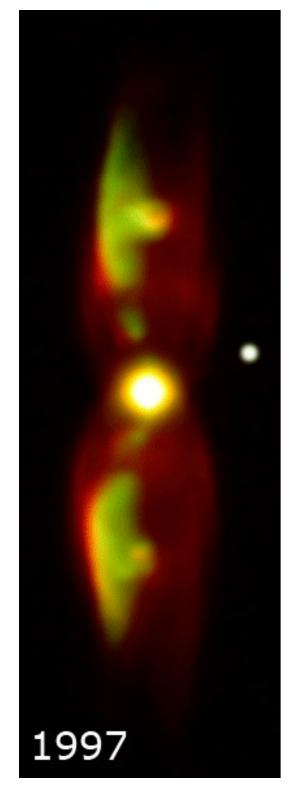
### **★Nested system of lobes:**

- Neutral shell >
   outer lobes + dusty blobs
   (distance 1.3 +- 0.2 kpc
   inclination i=18 deg)
- Inner lobes + knots/arcs(ionized gas)
- **★**Corkscrew geometry of the inner rotating features.

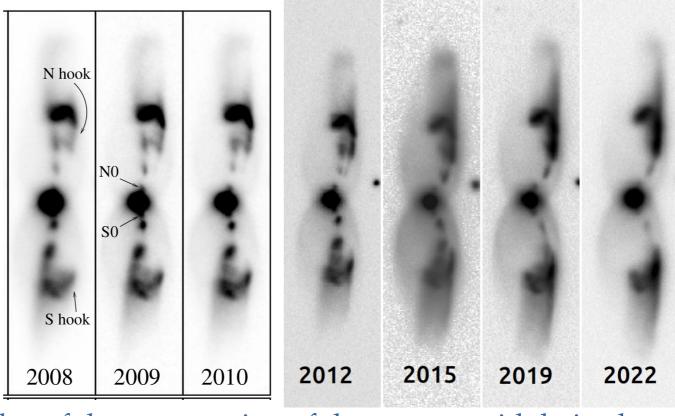


Doyle, Balick, Corradi, Schwarz AJ, 2000, 119, 1339

Corradi, Balick, and Sandander-Garcia A&A, 2011, 529, A43



# The lighthouse effect



- ★Delay of the propagation of the pattern with latitude > spray of high velocity particles (11 000 16 000 km/s)
- **★**No other ionizing mechanism needed!
- **★**Mirror symmetry of the rotating pattern!
- ★The pattern rotates P~90 years.

ALFOSC: Ha Ha [OIII]

Corradi, Balick, and Sandander-Garcia A&A, 2011, 529, A43

### Extended shells around B[e] stars.



Galaxies 2022, 10, 41



Article

### Follow-Up of Extended Shells around B[e] Stars

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Ha survey + individual targets

NOT + ALFOSC (NOTCAM)

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https://doi.org/10.3847/1538-3881/ac1355



### Resolving the Circumstellar Environment of the Galactic B[e] Supergiant Star MWC 137. II. Nebular Kinematics and Stellar Variability\*

Michaela Kraus<sup>1</sup>, Tiina Liimets<sup>1,2</sup>, Alexei Moiseev<sup>3</sup>, Julieta P. Sánchez Arias<sup>1</sup>, Dieter H. Nickeler<sup>1</sup>, Lydia S. Cidale<sup>4,5</sup>, and David Jones<sup>6,7</sup>, Astronomical Institute, Czech Academy of Sciences, 251 65 Ondřejov, Czech Republic; michaela.kraus@asu.cas.cz

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https://doi.org/10.3847/1538-3881/aa8df6



### Resolving the Circumstellar Environment of the Galactic B[e] Supergiant Star MWC 137 from Large to Small Scales\*

Michaela Kraus<sup>1,2</sup>, Tiina Liimets<sup>2,3</sup>, Cristina E. Cappa<sup>4,5</sup>, Lydia S. Cidale<sup>5,6,7</sup>, Dieter H. Nickeler<sup>1</sup>, Nicolas U. Duronea<sup>4</sup>, Maria L. Arias<sup>5,6</sup>, Diah S. Gunawan<sup>7</sup>, Mary E. Oksala<sup>8,9</sup>, Marcelo Borges Fernandes<sup>10</sup>, Grigoris Maravelias<sup>7</sup>, Michel Curé<sup>7</sup>, and Miguel Santander-García<sup>11</sup>

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# Protostellar jets in Serpens

A&A 587, A75 (2016) DOI: 10.1051/0004-6361/201526532 © ESO 2016 Astronomy Astrophysics

#### Amanda Djupvik (NOT)

#### Proper motions of embedded protostellar jets in Serpens\*

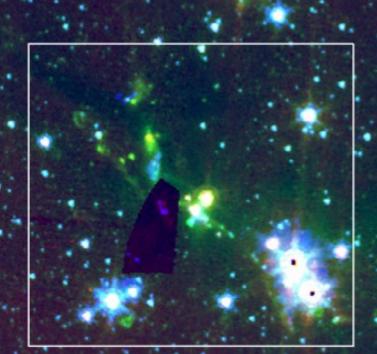
A. A. Djupvik<sup>1</sup>, T. Liimets<sup>2,3</sup>, H. Zinnecker<sup>4,5</sup>, A. Barzdis<sup>6</sup>, E. A. Rastorgueva-Foi<sup>7</sup>, and L. R. Petersen<sup>8</sup>

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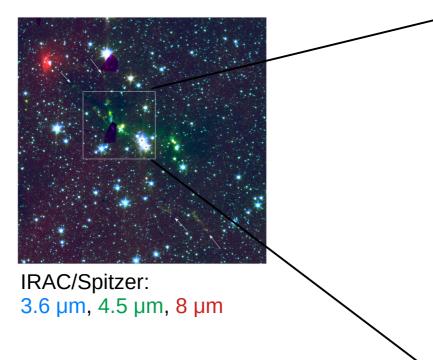
Received 14 May 2015 / Accepted 18 December 2015

Two decades of NOTCAM H<sub>2</sub> imaging + *K*-band long-slit spectra (2011, 2014)

Image: IRAC/Spitzer: 3.6 μm, 4.5 μm, 8 μm



# Serpens/G3-G6 region

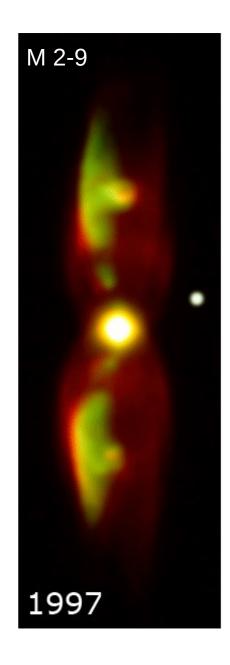


Class I binary Ser-emb-17

Class I Ser-emb-17

NOTCam: H<sub>2</sub> 2.12 µm line of shocked molecular hydrogen

- ★ 3 partly overlapping deeply embed flows.
- ★ Mass flow rate is two order of magnitude higher in Class 0 compared to Class I jets, while their bolometric luminosities are equal.
- ★ Episodic behaviour: appearing, disappearing, tima-variable velocities.



### To take home

**★** Scientific legacy of the NOT.

- **★** Johannes Andersen Student Programme.
- \* NOT family.



### The end