

COMPARING THE NEW MOG THEORY WITH DARK MATTER APPROACH IN EXTRAGALACTIC CONTEXT

The Galactic rotation curve

Distribution of dark matter in Universe is a long-standing problem of astrophysics and cosmology. Rotation curves of galaxies are considered to be an important source of information on mass distribution within the galaxies. The distribution of mass determines the orbital evolution of individual objects moving in the galaxies. Flat rotation curves of galaxies belong to one of the most significant observations supporting dark matter theory.

Modified Gravitation

A new MOdified Gravitational theory (MOG) based on the observational results of McGaugh et al. [5] was proposed by Klačka et al. [2], [3]. The theory suggests a modified centripetal acceleration relation in the form

$$g_{obs} = \frac{g_{bar}}{1 - \exp{-\sqrt{g_{bar}/g_{\dagger}}}} \,.$$

where $g_{bar} = d\Phi_{bar}/dR$ and $g_{\dagger} = 1.2 \times 10^{-10}$ m s⁻². The circular velocity based on Eq. 1 yields

$$v_c(R) = \sqrt{\frac{R\xi}{1 - \exp\left(-\sqrt{\frac{\xi}{g_{\dagger}}}\right)}}$$

where $\xi = \partial (\Phi_{Total-DMhalo}) / \partial R$ is the potential contribution with the component of the dark matter halo omitted

Galaxy sample

The SPARC catalogue [4] is a database of 175 spiral and irregular galaxies with high-quality rotation curves. We compared the new theoretical results (Eq. 2) and the standard dark matter approach with the rotation curves of 9 SPARC galaxies.

| ID | Туре | $\log(\frac{M_{\rm bar}}{M_{\odot}})$ | $e(M_{\scriptscriptstyle bar})$ | $a_{\scriptscriptstyle bar}$ | $\log(\frac{M_{\rm DM}}{M_{\odot}})$ |
|--|--|---|--|------------------------------|--|
| ESO563-G021 NGC 2841 NGC 3198 NGC 5055 NGC 5985 NGC 5985 NGC 6503 UGC 02953 | Sbc Sbc Sbc Sbc Sbc Sbc Sbc Sbc Sbc Sbc | $ \begin{array}{r} 108(M_{\odot}) \\ 11.40 \\ 11.00 \\ 10.53 \\ 10.96 \\ 11.08 \\ 9.94 \\ 11.18 \\ 10.94 \\ \end{array} $ | ± 0.065 ± 0.045 ± 0.055 ± 0.050 ± 0.120 ± 0.045 ± 0.120 ± 0.120 ± 0.75 | | $12.93 \\ 12.54 \\ 11.67 \\ 11.82 \\ 12.21 \\ 11.28 \\ 12.29 \\ 12.12$ |
| UGC 11455 | Scd | 11.22 | ± 0.010 ± 0.100 | 9.5 | 12.12 12.61 |

E. Puha¹, J. Klačka¹, R. Nagy¹, M. Šturc¹

¹Faculty of Mathematics Physics and Informatics, Comenius University, Bratislava, Slovak republic emil.puha@fmph.uniba.sk

$$\begin{split} \Phi_{\rm halo}(r) &= \frac{-GM_{\rm halo}}{r} - \frac{GM_{\rm halo}}{1.02a_{\rm halo}} \left[\frac{-1.02}{1 + (R/a_{\rm halo})^{1.02}} + \ln\left(\frac{1}{1 + (R/a_{\rm halo})^{1.0$$



Fig. 1: Different results. Left: better fit with the new MOG theory, top: NGC 3198, bottom: NGC 6503. Top right: well-fitted by both approaches, UGC 11455, bottom right: neither of the approaches fit the data, NGC 2841.

Our results imply, that from the selected sample of galaxies (i) 33 % could be fitted with the new MOG theory, (ii) 44 % could be fitted by both approaches and (iii) 23 % was not fitted well with any of the two approaches. All of the selected galaxies are spirals. This couldn't be the cause of the different results. However, the determined masses of the galaxies might have uncertainties due to inclination of the galactic planes, therefore, masses could not be calculated reliably from the M/L ratios. Also, we neglected the galactic components such as the disks, bulge, spiral arms or the stellar halo and we only considered that all the galactic mass is almost spherically symmetrical, therefore we only used a simple Plummer model.



$$e(M_{\text{DM}})$$
 a_{DM}
 ± 0.255 90
 ± 0.135 80
 ± 0.070 90
 ± 0.030 90
 ± 0.080 90
 ± 0.070 90
 ± 0.070 14
 ± 0.190 90
 ± 0.200 30

Remarks

Figure (2) shows the rotation curve of the NGC 3198 galaxy and it's consistency with the theoretical rotation curve of the new MOG theory.



Fig. 2: Rotation curve of the NGC 3198 galaxy.

As we stated, we obtained three different results. However, this might be because of the uncertainties of the determined masses of the galaxies. Our goal was to show, that the new MOG theory proposed by Klačka et al. [2], [3] is able to provide satisfying results when applied to observational data. This goal was achieved, as the results of the new MOG theory are similar or in some cases better than the approach with the use of the dark matter halo. However, further confrontations with the observational data are needed.

Acknowledgements

This work was supported by the Scientific Grant Agency VEGA, Slovak Republic, No. 1/0761/21.

References

- [1] Christine Allen and Alfredo Santillan. "An improved model of the galactic mass distribution for orbit computations". In: *RMAA* 22 (1991), pp. 255–263.
- [2] Jozef Klačka. "Alternative approach to gravity and MOND". In: arXiv preprint arXiv:1904.04074 (2019).
- [3] Jozef Klačka et al. "Milky Way: Rotation curve and new galactic mass models for orbit computations". In: MNRAS, submitted (2022).
- [4] Federico Lelli et al. "SPARC: mass models for 175 disk galaxies with spitzer photometry and accurate rotation curves". In: *ApJ* 152.6 (2016), p. 157.
- [5] Stacy S. McGaugh et al. "Radial acceleration relation in rotationally supported galaxies". In: *PRL* 117.20 (2016), p. 201101.



| • SPARC | | | |
|---------|----------|---------|----|
| | ¥ ¥ | | |
| | | | |
| | •••••••• | | |
| DM halo | | Plummer | |
| | 45 | | 60 |
| -1 | | | |