AMoRE

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Yoomin Oh on behalf the AMoRE Collaboration



$0\nu\beta\beta$ search using Mo-100

•Neutrinoless double beta decay:

- -Direct measure of Majorana nature of neutrino.
- -Lepton number violation process.
- -Effective neutrino mass.
- •Molybdenum-100:
 - $-Q_{\beta\beta} = 3034 \text{ keV}.$
 - -Natural abundance = 9.74%.
 - -Scintillation crystals with ¹⁰⁰Mo enrichment > 95% XMo_aO_b (XMO):
 - · X=Ca, Li₂, Na₂, Zn, Sr, Pb, ...
 - · Detection of light/heat signal \rightarrow rejection of surface- α background.
- •Key parameters for the experimental sensitivity:
 - -Signal ~ efficiency × [isotope mass × time] exposure.
 - -Background ~ radioactivity level at around $Q_{\beta\beta}$ and energy resolution.

Detector module

- Cylindrical CMO and LMO crystals, sizes vary $\Phi \ge 4$ cm / H $\lesssim 5$ cm.
- CMO: ⁴⁸Ca depleted, $Q_{\beta\beta}$ (⁴⁸Ca) = 4271 keV.
- Metallic magnetic calorimeter (MMC) + SQUID:
 - Fast signal timing: a few millisecond rise-time for phonon signals at mK.
 - Low random coincidence background.
 - Energy resolution ~ 10 keV FWHM at 2.6 MeV.



Cryostat / shielding



+ More enhanced shielding for AMoRE-I

- Cryogen-free dilution refrigerator.
- For AMoRE-pilot and AMoRE-I.
- Now operating at 10 mK with 1.2 μ W cooling power.
- Pb (γ), boron, and polyethylene (n).
- Plastic scintillator muon counter.
- Yangyang Underground Laboratory (Y2L) at 700 m depth.



Signal processing and analysis



- timings (rise/fall): pulse shape discrimination (PSD).
- Reconstruction for improving energy resolution and β/a discrimination power (DP):
 - Butterworth bandpass filter mainly for noise suppression:
 - \cdot pulse amplitude: pulse height or a least square fit to the template signal.
 - Stabilization heater signal every 10 seconds for gain drift corrections.

AMoRE-pilot final result



- Understanding of the background components and reduction of them.
- Background level of ~0.5 counts/keV/kg/yr at 2.8-3.2 MeV.
 - neutron-induced γ , crystals' internal contamination, rock/air-radon γ .
 - Internal background arXiv:2107.07704
- $T_{1/2}^{0\nu}$ > 3.2 × 10²³ years at 90% CL.



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AMoRE-pilot → AMoRE-I

- •6 CMO (1.89 kg) → 13 CMO (4.58 kg) + 5 LMO (1.61 kg)
 - Total crystal mass = 6.19 kg, ¹⁰⁰Mo mass = 3.0 kg
- Stabilization heater for all crystals.
- MMC sensor: Au:Er → Ag:Er.
- Using same cryostat + two stage temperature control: $\langle \Delta T \rangle$ < 1 $\mu K.$
- Shielding enhancements:
- -Outer Pb: 15 \rightarrow 20 cm; neutron shields: boric acid silicon + more PE / B-PE.
- More muon counter coverage.
- More supply of Rn-free air.





AMoRE-I data taking



- DAQ duty factor ~ 90%, good physics data ~ 75%.
- Data taking until end of 2022 (at least).
- 3.44 kg·yr crystal (1.67 kg·yr ¹⁰⁰Mo) exposure is presented here (selected data in blue dotted boxes).

Energy calibration



Particle IDentifications, CMO and LMO



- CMO shows better discrimination power light yield: CMO > LMO.
- LMO has much less *a* contamination.

Background spectrum



- All crystal excluding 1 LMO for very poor β/a discrimination power:
- 13 CMO + 4 LMO: exposure = 3.44 kg_{XMO} ·yr = 1.67 kg_{ISO} ·yr.
- Anti-coincidence cuts reject events:
- coincident at multiple crystals within 2 ms ($\varepsilon \sim 99\%$),
- -within 10 ms after a muon counter event ($\varepsilon \sim 99.7\%$),
- -within 20 minutes after a ²¹²Bi *a*-decay event candidate ($\varepsilon \sim 98\%$).

Preliminary 0vββ limit from AMoRE-I



- ROI to contain most (> 99%) of the $0\nu\beta\beta$ signal peak, $\varepsilon_{\text{containment}} \sim 81\%$.
- Background = 0.039 ± 0.006 counts/keV/kg/year, from ROI side-band.
- Combining the result of counting analysis at ROI, with a flat background constraint from the side-band events for each crystal.
- $T_{1/2}^{0\nu}$ > 1.2×10²⁴ years at 90% CL.

AMoRE-II in preparation

AMoRE-II Detector module







90 modules (~27 kg LMO) for the first stage







AMoRE-II in Yemilab



Limits and sensitivities



- Final results of AMoRE-I with doubled data and further improved analysis.
- AMoRE-II for $T_{1/2}^{0\nu}$ > ~5×10²⁶ years by 100 kg of Mo-100 × 5 years running.
- Reduction of background level down below 10⁻⁴ ckky.

Summary

- AMoRE searches for 0vbb using Mo-100 based scintillation crystals at the low temperature detector system.
- Preliminary result of AMoRE-I at its mid-point:
 - -Massxtime exposure: 3.44 (1.67) kg·yr XMO (100 Mo).
 - -Background level ~ 0.04 counts/keV/kg/year at 2860-3200 keV.
- $-T_{1/2}^{0\nu} > 1.2 \times 10^{24}$ years.
- AMoRE-I data taking will continue at least until end of 2022.
- AMoRE-II starts its data taking soon to head for $T_{1/2}^{0\nu} > 5 \times 10^{26}$ years.

- [DT04-774] Hanbum Kim, "AMoRE-I data analysis"
- [DT04-398] Hyejin Lee, "Detector sensors and modules for AMoRE-II"
- [DT04-377] Wootae Kim, "Detector R&D for AMoRE-II"
- [DT04-385] Jeewon Seo, "Radioassay and simulation for AMoRE-II"
- [DT04-560] Jaison Lee, "AMoRE-II construction"

Posters

(Session 1-b)

4F. Dirac

Good afternoon, good evening and good night.

back up slides >>

CMO internal background



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Background budget for AMoRE-II



Two stage temperature control

