

A comparative study of dark matter flow & hydrodynamic turbulence and its applications

May 2022

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PNNL is operated by Battelle for the U.S. Department of Energy





Dark matter, if exists, accounts for five times as much as ordinary baryonic matter. Therefore, dark matter flow might possess the widest presence in our universe. The other form of flow, hydrodynamic turbulence in air and water, is without doubt the most familiar flow in our daily life. During the pandemic, we have found time to think about and put together a systematic comparison for the connections and differences between two types of flow, both of which are typical non-equilibrium systems.

The goal of this presentation is to leverage this comparison for a better understanding of the nature of dark matter and its flow behavior on all scales. Science should be open. All comments are welcome.

Thank you!



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Data repository and relevant publications Northwest

Structural (halo-based) approach:

- Data https://dx.doi.org/10.5281/zenodo.6541230 0.
- Inverse mass cascade in dark matter flow and effects on halo mass 1. functions https://doi.org/10.48550/arXiv.2109.09985
- 2. Inverse mass cascade in dark matter flow and effects on halo deformation. energy, size, and density profiles https://doi.org/10.48550/arXiv.2109.12244
- Inverse energy cascade in self-gravitating collisionless dark matter flow and 3. effects of halo shape https://doi.org/10.48550/arXiv.2110.13885
- The mean flow, velocity dispersion, energy transfer and evolution of rotating 4. and growing dark matter halos https://doi.org/10.48550/arXiv.2201.12665
- Two-body collapse model for gravitational collapse of dark matter and 5. generalized stable clustering hypothesis for pairwise velocity https://doi.org/10.48550/arXiv.2110.05784
- Evolution of energy, momentum, and spin parameter in dark matter flow and 6. integral constants of motion https://doi.org/10.48550/arXiv.2202.04054
- The maximum entropy distributions of velocity, speed, and energy from statistical mechanics of dark matter flow https://doi.org/10.48550/arXiv.2110.03126
- Halo mass functions from maximum entropy distributions in collisionless 8. dark matter flow https://doi.org/10.48550/arXiv.2110.09676

Statistics (correlation-based) approach: .5281/zenodo.6569898

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		6.	The baryonic-to-halo mass cascade in dark matter flow https://doi.org/10.48550/ar

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Xiv.2202.00910

rk matter flow and high order ations for velocity and density 10.48550/arXiv.2202.02991

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and properties from two-thirds law rk matter flow

Xiv.2202.07240

eration and deep-MOND from d energy cascade in dark matter 50/arXiv.2203.05606

relation from mass and energy Xiv.2203.06899



Applications of dark matter flow



The baryonic-to-halo mass relation from mass and energy cascade in dark matter flow

Xu Z., 2022, arXiv:2203.06899v1 [astro-ph.GA] https://doi.org/10.48550/arXiv.2203.06899

Pacific Northwest Introduction

- The existence of dark matter (DM) is supported by numerous astronomical observations:
 - Flat rotation curves of spiral galaxies
 - Motion of galaxies in galaxy clusters
 - Gravitational lensing, Bullet clusters, CMB
- Though the nature of dark matter is still unclear, dark matter is believed to be cold (non-relativistic), collisionless, dissipationless, non-baryonic, barely interacting with baryons except through gravity, and sufficiently smooth with a fluid-like behavior.
- Total galaxy baryonic mass = stellar mass + cold gas.
- Stellar-to-halo mass relation (SHMR)
 - halo abundance matching approach
- Baryonic-to-halo mass relation (BHMR)

- Baryonic Tully and Fisher relation (BTFR): $v_{f_{1}}^{4} = Gm_{b}a_{0}$ \leftarrow observed baryonic mass
- Halo mass m_b can be related to the halo virial radius r_h through constant density ratio Δ_c

$$\underbrace{m_h}_{h} = \frac{4}{3} \pi (r_h)^3 \Delta_c \overline{\rho}_0 ($$

- The BHMR (m_h and m_h) can be obtained only if the relation between v_f and r_h is known.
- The BHMR from the mass and energy cascade of dark matter flow?
- What is the average mass fraction of baryons in all halos?
- What is the fraction of total baryons residing in all galaxies?

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Energy cascade in hydrodynamic turbulence

There exist an inertial range with a scaleindependent rate of energy cascade (ε does not depend on eddy size *l*) for eddy size $\eta < l < L$. η is a dissipative scale determined by viscosity v and ε .

In this range, inertial force is dominant over viscous force. For eddies with a characteristic velocity u and size I , the lifetime (turnaround time) \gtrsim $\log E()$ of eddy is I/u. The rate ε can be computed as the kinetic energy passed per eddy lifetime.



acceleration turnaround time

Big whirls have little whirls, That feed on their velocity; And little whirls have lesser whirls, And so on to viscosity.



Mass/Energy cascade in dark matter flow (SG-CFD)

Collisionless nature and long-range interaction.

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- Long-range gravity requires a broad spectrum of halos to be formed to maximize system entropy. No halo structure for short-range forces.
- A continuous cascade of mass/energy from smaller to larger mass scales with a scale-independent rate of mass transfer ε_m and ε_{μ} in a certain range of mass scales (propagation range).
- The mass/energy cascade is an intermediate statistically steady state for non-equilibrium systems to continuously maximize system entropy.
- The maximum entropy distribution of dark matter flow (the X distribution).

Little halos have big halos, That feed on their mass; And big halos have greater halos, And so on to growth.







Pacific Northwest Mass/Energy cascade in dark matter flow (SG-CFD)

- Collisionless, no dissipation range in SG-CFD.
- The smallest length scale of inertial range is not limited by viscosity.
- This enable us to extend the scale-independent ε_{μ} down to the smallest scale, where quantum effects become important
- Dark matter flow exhibits scale-dependent flow behaviors for peculiar velocity, i.e. a constant divergence flow on small scales and an irrotational flow on large scales.
- The constant divergence flow shares the same even order kinematic relations with those of incompressible (divergence free) flow. This hints to similar scaling laws holds for dark matter.



Constant (time and scale independent) rate of Pacific Northwest energy cascade

Power-law time evolution for energy in terms of rate of energy cascade ε_{..}:

 $K_p = -\mathcal{E}_{\mathbf{u}}t$

 $P_{y} = \frac{7}{5} \varepsilon_{\mathbf{u}} t$

Power-law for Peculiar kinetic energy

Power-law for potential energy

$$\varepsilon_{u} = -\frac{K_{p}}{t} = -\frac{3}{2} \frac{u_{0}^{2}}{t_{0}} \approx -4.6 \times 10^{-7} \frac{m^{2}}{s^{3}}$$

Also see detail analysis for inverse kinetic energy cascade.





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Dimensional analysis for critical mass scales Northwest

The smallest mass scale (dark matter particle mass)

At the smallest scale, three fundamental constants:

Gravitational constant

> Rate of energy cascade

Planck constant

$$\varepsilon_u = -4.6 \times 10^{-7} m^2/s^3$$

 $G = 6.67 \times 10^{-11} m^3 / (kg \cdot s^2)$

$$\hbar = 1.05 \times 10^{-34} \, kg \cdot m^2 / s$$

Simple dimensional analysis predicts:

Mass scale:

$$m_X \propto \left(-\varepsilon_u \hbar^5 / G^4\right)^{\frac{1}{9}} \approx 8.7 \times 10^{-16} kg = 0.5 GeV$$

Length scale: $l_X \propto \left(-G\hbar / \varepsilon_u\right)^{\frac{1}{3}}$
Time scale: $t_X \propto \left(G^2 \hbar^2 / \varepsilon_u^5\right)^{\frac{1}{9}}$

The largest mass scale (critical halo mass) Three fundamental constants: Gravitational constant Rate of energy cascade Velocity dispersion or Hubble constant H Simple dimensional analysis predicts: Mass scale: $m_L \propto -u_0^5 / (G \varepsilon_u) \approx 9.14 \times 10^{13} M_{\odot}$ Length scale: $l_L \propto -u_0^3 / \varepsilon_u \approx 3.14 Mpc$ Time scale: $t_L \propto u_0^2 / \varepsilon_u \approx 8.7 \times 10^9 yr$

 $G = 6.67 \times 10^{-11} m^3 / (kg \cdot s^2)$

 $\varepsilon_{\mu} = -4.6 \times 10^{-7} \ m^2/s^3$

 $u_0 \equiv u(a=1) = 354.61 \, km/s$

Pacific Northwest National Laboratory The baryonic-to-halo mass ratio from energy cascade

Baryonic Tully-Fisher Halo mass and halo relation (BTFR): size relation:

$$v_f^4 = Gm_b a_0$$

$$m_h = \frac{4}{3}\pi r_h^3 \Delta_c \overline{\rho}_0 a^{-3}$$

 $\frac{1}{2} \frac{\partial u}{\partial r_{h}} \mathcal{E}_{u} = -\beta_{f} \frac{u^{2}}{r_{h}/v_{f}} a^{q}$ Small halos <m_L: Baryonic mass in equilibrium with DM, i.e. same kinetic energy u^2

$$v_{cir} = \frac{4}{9} \sqrt{\frac{\Delta_c}{2}} \beta_f v_f a^q \propto (m_h)^{1/3} a^{-1/2}$$

$$r_h = \frac{4}{9} \beta_f v_f H^{-1} a^q \propto (m_h)^{1/3} a^1$$

$$= \frac{9}{4\beta_f} \left(\frac{2}{\Delta_c}\right)^{\frac{1}{3}} (Gm_h H)^{1/3} a^{-q} \propto (m_h)^{1/3} a^0$$

Baryonic Tully-Fisher relation (BTFR):

$$v_f^4 = Gm_b a_0 \qquad m_h$$



Halo mass and halo size relation: $=\frac{4}{2}\pi r_h^3 \Delta_c \overline{\rho}_0 a^{-3}$

Large halos $> m_1$: Baryonic mass and DM are two miscible phases sharing same rate of cascade.

Northwest Critical scales and Baryonic-Halo-Mass Ratio

Critical rotation $v_{fc} = ua^{(q-p)/2} \sqrt{\beta_f / \alpha_f}$ speed:

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Critical circular speed:

Critical Tritical halo $r_{hc} = \frac{4}{\Omega} a^{(3q-p)/2} u H^{-1} \beta_f \sqrt{\beta_f / \alpha_f}$ size:

Critical halo $m_{hc} = \frac{16}{81} \left(\frac{\beta_f^3}{\alpha_f}\right)^{3/2} \left(\frac{\Delta_c}{2}\right) \left(\frac{u^5}{G\varepsilon_u}\right) a^{\frac{3}{2}(3q-1)}$ mass: Critical

mass:

Critical baryonic $m_{bc} = \frac{2}{\Delta_c} \left(\frac{\beta_f}{\alpha_f}\right)^2 \left(\frac{u^5}{G\varepsilon_u}\right) a^{2(q-p)}$ mass:

$$v_{cc} = \frac{4}{9} \sqrt{\frac{\Delta_c}{2}} \sqrt{\frac{\beta_f^3}{\alpha_f}} u a^{(3q-p)/2}$$

The baryonic mass in small halos:

$$m_b = (M_{c1})^{-1/3} (m_h)^{4/3} \qquad M_{c1}(a) =$$

The baryonic mass in large halos:

$$m_{b} = (M_{c2})^{5/9} (m_{h})^{4/9} M_{c2}(a) = \left(\frac{2}{3}\right)^{5/9} M_{c2}(a) = \left(\frac{2}{3}\right)^{5/$$

The baryonic-halo-mass ratio in critical halos:

$$A(z) \equiv \frac{m_{bc}}{m_{hc}} = \left(\frac{M_{c2}}{M_{c1}}\right)^{5/24} = \frac{81}{16(\alpha)}$$

$$\frac{Mass}{cale m_{L}}$$

 $A(z=0) \approx 0.076$



 $= \left(\frac{2}{3}\right)^{16} \left(\beta_f a^q\right)^{12} \left(\frac{\Delta_c}{2}\right)^{\prime} \left(\frac{u^5}{G\varepsilon}\right)$ $\frac{2}{3}\right)^{-\frac{16}{5}} \left(\alpha_f a^p\right)^{-\frac{12}{5}} \left(\frac{2}{\Delta_c}\right)^{\frac{15}{5}} \left(\frac{u^5}{G\varepsilon_u}\right)$

 $\frac{1(2/\Delta_{c})^{2}}{(\beta_{c})^{1/2}(\beta_{c})^{5/2}}a^{-(5q+p)/2}$

Pacific Northwest National Laboratory Relevant parameters for baryonic-to-halo mass ratio

T	Table 2. Parameters for	r derivii	ng baryonic	c-to-halo	mass
Δ_c	200	р	7/4	$M_{\rm c1}$	3.0
\mathcal{E}_{u}	$4.6 \times 10^{-7} m^2/s^3$	q	-1/2	M_{c2}	1.2
H_{0}	$1.62 \times 10^{-18} \mathrm{l/s}$	$\alpha_{_f}$	0.5	m_{hc}	1.3
u _o	354.61 <i>km/s</i>	β_{f}	0.16	m_{bc}	1.0
$a_0(z=0)$	$1.2 \times 10^{-10} m/s^2$	т	4	A(z)	0.0
$\eta_{_0}$	0.76	$q_{\scriptscriptstyle 0}$	0.556	m_h^*	$4 \times$

s ratio $0.01 \times 10^{15} a^{-9/4} M_{sum}$ $9 \times 10^{10} a^{-9/20} M_{sun}$ $3 \times 10^{12} a^{-9/8} M_{sun}$ $1 \times 10^{11} a^{-3/4} M_{sun}$ $761a^{3/8}$ $(10^{13} a^{3/2} M_{sun} [27])$

SPARC (Spitzer Photometry & Accurate Rotation Pacific Northwest Curves) data and model



Halos have different rate of energy cascade with an average around ε_{II} (spatial intermittence in dark matter flow?)









Pacific Northwest SPARC data and model

Baryonic mass in small halos:

$$m_b = \left(M_{c1}\right)^{-1/3} \left(m_h\right)^{4/3}$$

Baryonic mass in large halos:

$$m_b = (M_{c2})^{5/9} (m_h)^{4/9}$$

Model incorporate two limits:

$$\frac{m_b}{m_h} = 2^{\frac{1}{m}} A(z) \left[\left(\frac{m_h}{m_{hc}(z)} \right)^{-\frac{m}{3}} + \left(\frac{m_h}{m_{hc}(z)} \right)^{\frac{5m}{9}} \right]^{-\frac{1}{m}}$$

- Dash line: the stellar-to-halo mass ratio obtained from halo abundance matching approach (require to match the stellar mass function)
- The scaling 4/9 law for both SHMR and BHMR



Pacific Northwest NATIONAL LABORATORY Redshift variation of baryonic-to-halo mass ratio

Models for baryonic-to-halo mass ratio:

$$\frac{m_b}{m_h} = 2^{\frac{1}{m}} A(z) \left[\left(\frac{m_h}{m_{hc}(z)} \right)^{-\frac{m}{3}} + \left(\frac{m_h}{m_{hc}(z)} \right)^{\frac{5m}{9}} \right]^{-\frac{1}{m}}$$

m is a parameter to adjust the transition;

- There exist a maximum BHMR ~0.076 at critical halo mass m_{hc}=1.33x10¹² M_{sun}
- The critical halo mass decreases with time
- The maximum BHMR increases with time



Pacific Northwest National Laboratory Redshift evolution of baryonic-halo-mass relation

Overall cosmic baryonic-to-DM mass ratio (including both halos and out-of-halo) is ~18.8% in ΛCDM model:



Use double-λ mass function to compute:

$$f_1 = \int_0^{v_c} f_{D\lambda}(v) (M_{c1})^{-1/3} (v^{3/2} m_h^*)^{1/3} dv$$
 The baryonic-to-
halo mass ratio
in small halos

 $f_2 = \int_{v_c}^{\infty} f_{D\lambda}(v) (M_{c2})^{5/9} (v^{3/2} m_h^*)^{-5/9} dv$ The baryonic-toin large halos



Redshift evolution of BHMR

Pacific Northwest Summary and keywords

Halo mass function	Mass/energy cascade	Tully-Fisher re
Modified Newtonian Dynamics	Stellar-to-halo mass relation	Baryonic-to-hal relation

- Review <u>direct energy cascade</u> from large to small scales in hydrodynamic turbulence
- Reveal inverse mass and energy cascade that is unique for dark matter flow
- Present a fundamental theory for baryonic-to-halo mass ratio based on the mass/energy cascade in dark matter flow (agrees well with SPARC data)
- Predict a maximum baryonic-to-halo mass ratio ~0.076 for halos with a critical mass (agrees with SPARC data) and an average ratio ~0.024 for all halos
- Predict two distinct regimes for small and large halos, respectively, with critical halo mass and size explicitly derived (agrees with observations of stellar-to-halo mass ratio).
- Predict the fraction of total baryons in all galaxies is ~7.6% and that fraction increases with time (agrees very well with astronomical surveys including optical Sloan Digital Sky Survey and HIPASS). Most baryons (~92.4%) are not in galaxies.

elation o mass