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## A comparative study of dark matter flow & hydrodynamic turbulence and its applications

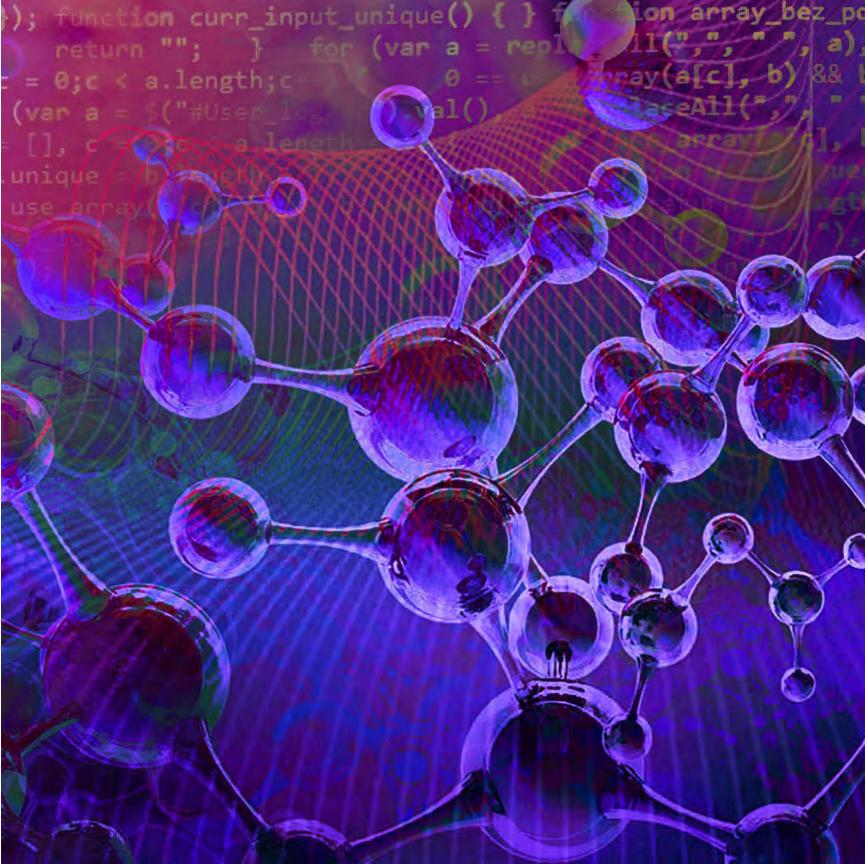
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#### Zhijie (Jay) Xu

Multiscale Modeling Team Computational Mathematics Group Physical & Computational Science Directorate <u>Zhijie.xu@pnnl.gov; zhijiexu@hotmail.com</u>



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Dark matter, if exists, accounts for five times as much as ordinary baryonic matter. Therefore, dark matter flow might possess the widest presence in our universe. The other form of flow, hydrodynamic turbulence in air and water, is without doubt the most familiar flow in our daily life. During the pandemic, we have found time to think about and put together a systematic comparison for the connections and differences between two types of flow, both of which are typical non-equilibrium systems.

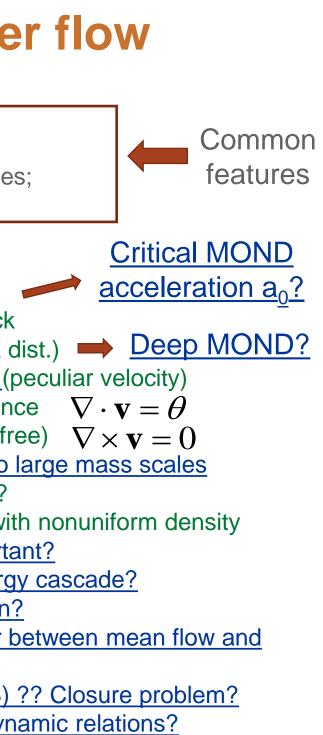
The goal of this presentation is to leverage this comparison for a better understanding of the nature of dark matter and its flow behavior on all scales. Science should be open. All comments are welcome.

Thank you!



#### Pacific Northwest Hydrodynamic turbulence vs. dark matter flow

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Key attributes of hydrodynamic turbulence:	Key attributes of dark matter flow:	
<ul> <li>Disorganized, chaotic, random;</li> </ul>	<ul> <li>Disorganized, chaotic, random;</li> </ul>	
<ul> <li>Nonrepeatability (sensitivity to initial conditions);</li> </ul>	<ul> <li>Nonrepeatability;</li> </ul>	
<ul> <li>Multiscale in length and time scales;</li> </ul>	Multiscale in mass/length/time scales	
<ul> <li>Intermittency in space and time;</li> </ul>	<ul> <li>Intermittency in space and time;</li> </ul>	
<ul> <li>Dissipative and collisional</li> </ul>	<ul> <li>Dissipationless and collisionless</li> </ul>	
<ul> <li>Short-range interaction</li> </ul>	<ul> <li>Long-range gravity</li> </ul>	
Velocity fluctuation	Velocity & acceleration fluctuation	
<ul> <li>Vortex as fundamental building block</li> </ul>	<ul> <li>Halos as fundamental building block</li> </ul>	
<ul> <li>Maximum entropy distribution (Gaussian)</li> </ul>	Maximum entropy distribution?? (X d	
• Incompressible on all scales $\nabla \cdot \mathbf{v} = 0$	Flow behavior is scale-dependent (pressure of the second secon	
<ul> <li>Divergence-free</li> </ul>	Small scale: constant divergend	
<ul> <li>Constant density</li> </ul>	Large scale: irrotational (curl-free	
Energy cascade from large to small length scales	Mass/energy cascade from small to	
Vortex stretching responsible for energy cascade	Role of halos for energy cascade??	
<ul> <li>Volume conserving</li> </ul>	<ul> <li>Halos are growing, rotating, wit</li> </ul>	
<ul> <li>Shape changing</li> </ul>	Is halo shape changing importa	
<ul> <li>Uniform density</li> </ul>	Mass cascade facilitates energy	
Reynolds decomposition	Velocity/acceleration decomposition?	
Reynolds stress for energy transfer between mean	What facilitates the energy transfer b	
flow and random motion (turbulence)	random motion??	
Closure problem, eddy viscosity, etc	Self-closed model (analogue of NS)	
Statistical theory: correlation/structure functions	Statistical theory: Kinematic and dyn	
scaling laws in inertial range	Scaling laws?	
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# Data repository and relevant publications

#### Structural (halo-based) approach:

- 0. Data <u>https://dx.doi.org/10.5281/zenodo.6541230</u>
- 1. Inverse mass cascade in dark matter flow and effects on halo mass functions <u>https://doi.org/10.48550/arXiv.2109.09985</u>
- 2. Inverse mass cascade in dark matter flow and effects on halo deformation, energy, size, and density profiles <u>https://doi.org/10.48550/arXiv.2109.12244</u>
- 3. Inverse energy cascade in self-gravitating collisionless dark matter flow and effects of halo shape <u>https://doi.org/10.48550/arXiv.2110.13885</u>
- 4. The mean flow, velocity dispersion, energy transfer and evolution of rotating and growing dark matter halos <u>https://doi.org/10.48550/arXiv.2201.12665</u>
- 5. Two-body collapse model for gravitational collapse of dark matter and generalized stable clustering hypothesis for pairwise velocity <u>https://doi.org/10.48550/arXiv.2110.05784</u>
- 6. Evolution of energy, momentum, and spin parameter in dark matter flow and integral constants of motion <u>https://doi.org/10.48550/arXiv.2202.04054</u>
- 7. The maximum entropy distributions of velocity, speed, and energy from statistical mechanics of dark matter flow https://doi.org/10.48550/arXiv.2110.03126
- 8. Halo mass functions from maximum entropy distributions in collisionless dark matter flow <u>https://doi.org/10.48550/arXiv.2110.09676</u>

#### Statistics (correlation-based) approach: 0. Data https://dx.doi.org/10.5281/zenodo.6569898

	0.	Data https://dx.doi.org/10
	1.	The statistical theory of da and potential fields <u>https://doi.org/10.48550/ar</u>
	2.	The statistical theory of da kinematic and dynamic relacorrelations <u>https://doi.org/</u>
	3.	The scale and redshift vari distributions in dark matter pairwise velocity <u>https://do</u>
	4.	Dark matter particle mass and energy cascade in dar https://doi.org/10.48550/ar
	5.	The origin of MOND acceleration fluctuation and flow <u>https://doi.org/10.4855</u>
	6.	The baryonic-to-halo mass cascade in dark matter flow https://doi.org/10.48550/ar

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#### Xiv.2202.00910

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iation of density and velocity flow and two-thirds law for <u>i.org/10.48550/arXiv.2202.06515</u>

and properties from two-thirds law rk matter flow

Xiv.2202.07240

eration and deep-MOND from d energy cascade in dark matter 50/arXiv.2203.05606

s relation from mass and energy w

Xiv.2203.06899





- Some fundamentals of dark matter research
- Basic concepts in hydrodynamic turbulence
- Dark matter flow (SG-CFD) vs. hydrodynamic turbulence
- Theory of dark matter flow
  - Structural (halo-based) approach
  - Statistical (correlation-based) approach
- Applications of dark matter flow
  - Predicting dark matter particle properties
  - Understanding the origin of MOND
  - The baryonic-halo mass ratio and total baron faction







## PROFILE: Zhijie (Jay) Xu

- Computational Scientist
- Team lead

## **EXPERIENCE**:

Idaho National Laboratory (INL) Pacific Northwest National Laboratory (PNNL)

# Grand canyon Yellowstone

## **EDUCATION:**

- Zhejiang University

### CONTACT:

- zhijiexu@hotmail.com
- Zhijie.xu@pnnl.gov

## **INTERESTS**:

- Fluid dynamics
- Cosmological flow
- **Multiscale Modeling**

HOBBIES:

- Travel
- Hiking
- **Biking**

**Civil Engineering**  National University of Singapore Structural Engineering Rensselaer Polytechnic Institute Mechanical Engineering