## A Systematic Search for

## Multiplanet Systems in TESS Full-Frame Images



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New Planet Candidates!
Multiplanet systems help us understand planetary system formation \& evolution. Currently, no search has been conducted on the millions of existing Quick-Look Pipeline (QLP) Full Frame Image (FFI) light curves. ${ }^{2}$ We are conducting the first systematic search for multiplanet systems in FFI light curves.
250k light curves processed so far -> a few new candidates from past sectors!

- Code added to QLP in S33.

Outline

1. Generate QLP LCs from TICA-calibrated FFIs (~14 M)
2. List stars with Threshold Crossing Events ( $\sim 2.5 \mathrm{M}$ )

- Sectors 1-36
- Tmag < 12.5

Signal-to-Pink Noise Ratio > 7
3. Search for multiplanet signals with iterative Box Least Squares (BLS) algorithm
4. Filter results with neural-network model Astronet ${ }^{3}$
5. Review remaining candidates

References

1. BLS (Kovacs et al. 2002)
2. QLP (Huang et al. 2020)
3. Astronet (Moldovan et al. in prep)

TIC ID 354944123 (Tey et al. in prep)

- M dwarf 2-planet system near 2:1 resonance


Updated Multiplanet TOI Demographics



TOI 696.03 (Silverstein et al. 2021, submitted)
Third candidate found near 4:3 resonance with 696.02


Hours from mid-transit

## Far left: Period \& radius ratio

 comparisons between Kepler and TESS multiplanet systems.Left: Scatter plots of period \& radius ratio vs planet radii zoomed in around 2:1 and 5:4 period resonances

Right: Mosaic of all multiplanet TOI systems. Dot sizes are scaled by the planet radius. Dot colors reflect the host star's $\mathrm{T}_{\text {eff }}$ Arrow mark new candidates.

