

# Hunting for young Exoplanets in **Stellar Associations using TESS**

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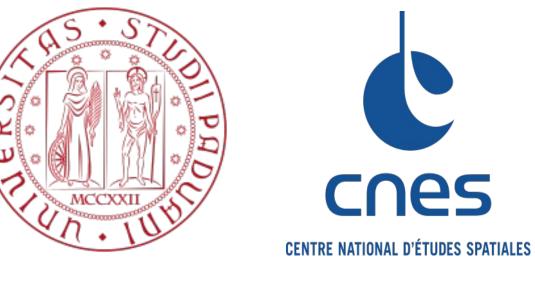
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# **Stellar Associations (SAs)**

- SAs are loose star clusters usually formed by a small number of stars
- Ages of SAs' members can be derived with high accuracy using theoretical

# **Candidate Exoplanets in SAs**

- 11 candidate exoplanets discovered with ages between 10 and 50 Myr.
- All the candidates are giant planets with  $R_P > 7x R_{Earth}$ .





**25 Myr** 

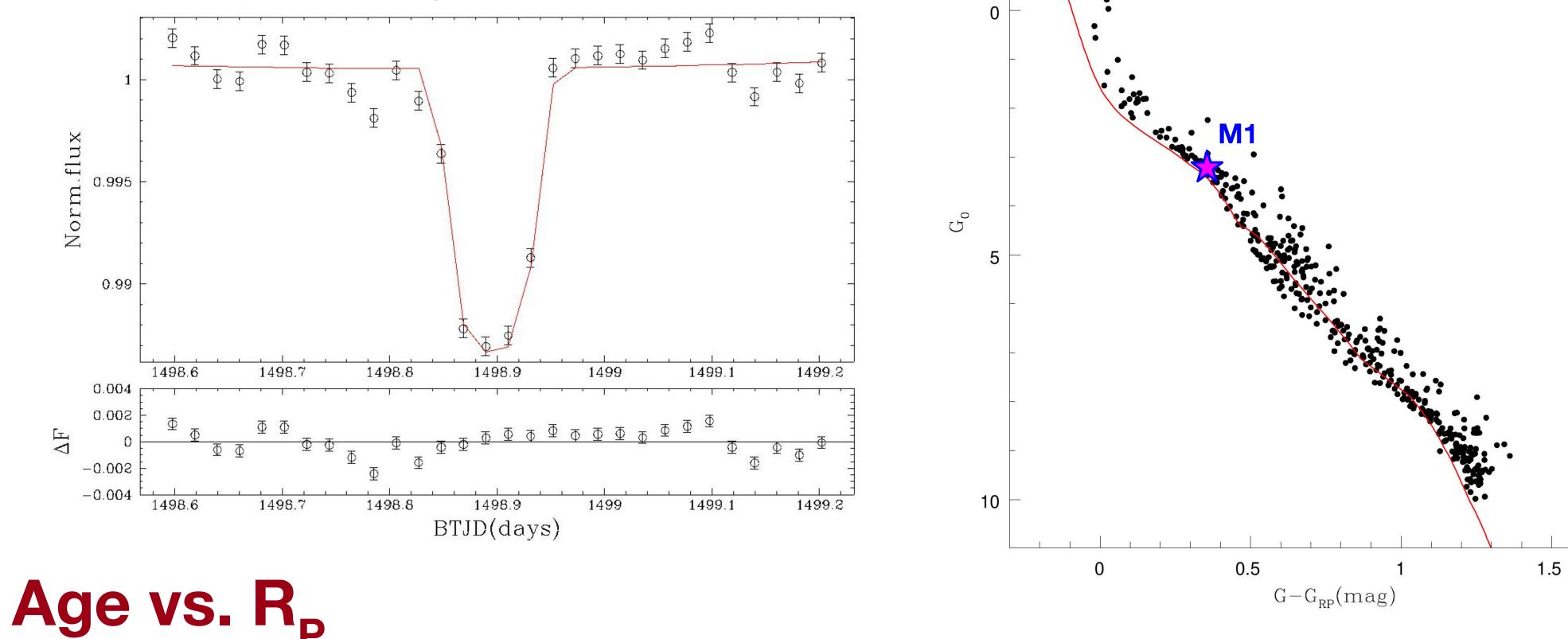
#### models

SAs' members in our sample have ages that span between (1-680) Myr



2 candidates are also TOIs





We analysed the relation between stellar age and  $R_p$  also including data from the literature (see figure), which can allow us to study the evolution of the atmosphere of young, short-period exoplanets.

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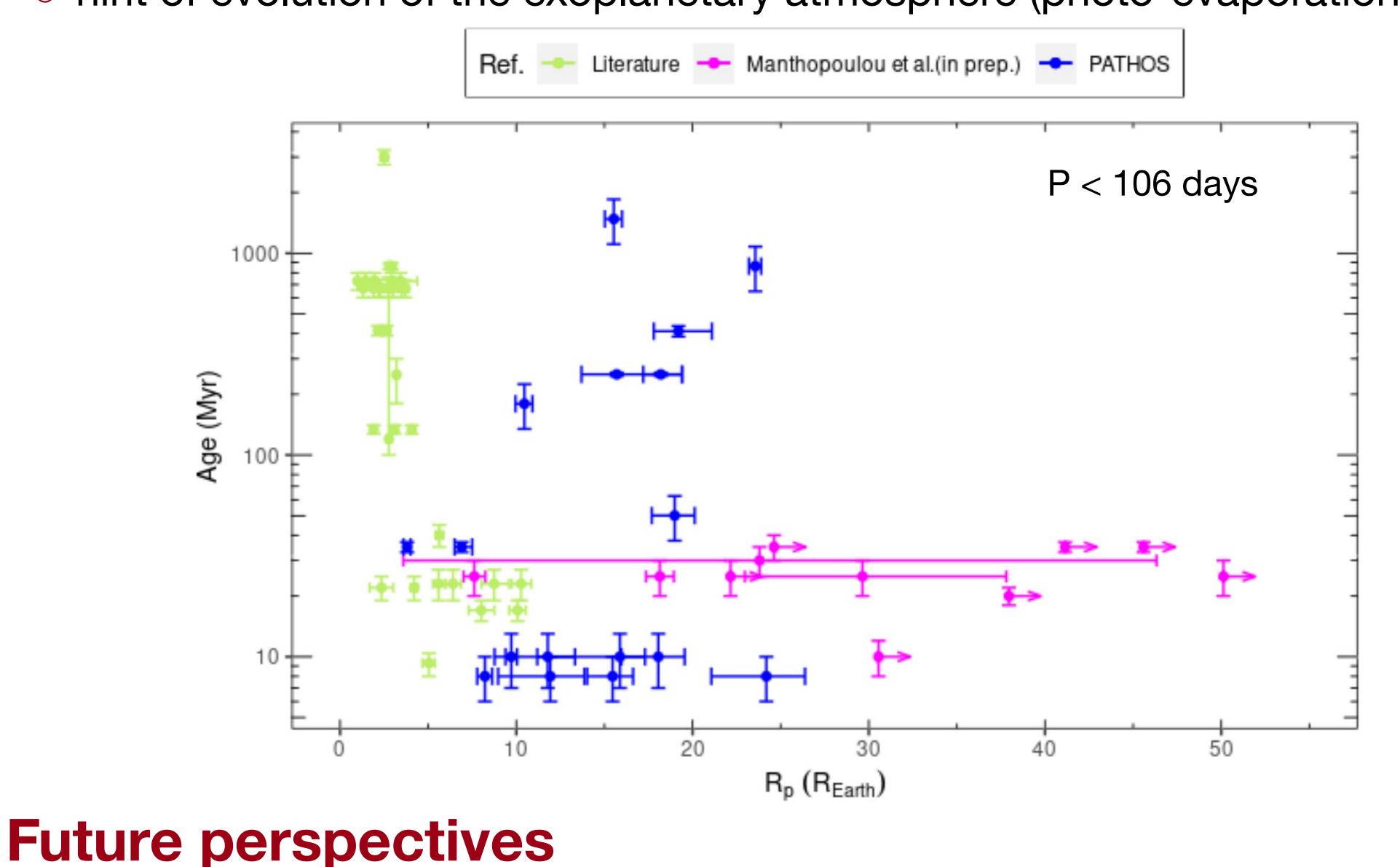
# **Exoplanets in SAs**

- For the vast majority of exoplanets' hosts the age is not well constrained
- The understanding of the formation and evolution mechanisms of exoplanets needs an accurate knowledge of the age of the host star
- Searching for exoplanets in SAs offers the unique opportunity of mapping planetary systems at different stages in their early life

## Finding exoplanets in SAs: a difficult task

• **TESS** has observed many SAs, but the resolution of the instrument sometime makes these objects hard to analyse.

- No peculiar trend for  $R_P > 1 \times R_{Jupiter} \sim 11 \times R_{Earth}$
- For  $R_P < 11 \times R_{Earth}$ :
- older planets (>100-200 Myr) are concentrated at R<sub>P</sub><4 x R<sub>Earth</sub> ~1 x R<sub>Neptune</sub> young planets (<100 Myr) have radii 4x R<sub>Earth</sub> < R<sub>P</sub><11 x R<sub>Earth</sub> hint of evolution of the exoplanetary atmosphere (photo-evaporation)



• We adopted the PATHOS pipeline to extract high-precision light curves from the FFIs for **11380 members** in **41 SAs**.

 We injected simulated transits of planets with radii (0.8-28)R<sub>Farth</sub> in the observed light curves to probe the detection efficiency of our pipeline and used a logistic regression model to classify the detected signals.

• We validated the transiting objects of interest in our sample through a series of vetting tests (e.g. odd/even, secondary transit depth, analysis of the centroid and of different photometric apertures).

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• Estimation of the frequency of exoplanets in SAs (and also stellar clusters) as a function of the age

Analysis of the results with a view to exoplanet formation and evolution

