

# IVADO X-TRA A Machine Learning Approach to HII Region Kinematics

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## Introduction

### HII Regions

- HII regions are formed from gaseous clumps ionized by young, hot OB stars
- Characterized by strong recombination or collisional emission lines
- Tracers of different feedback mechanisms

### SITELLE/SIGNALS

- Imaging Fourier Transform Spectrograph at Canada-France-Hawai'i Telescope
- 11'x11' Field-of-View
- Spectral Resolution up to 10,000
- Star formation, Ionized Gas, and Nebular Abundances Legacy Survey

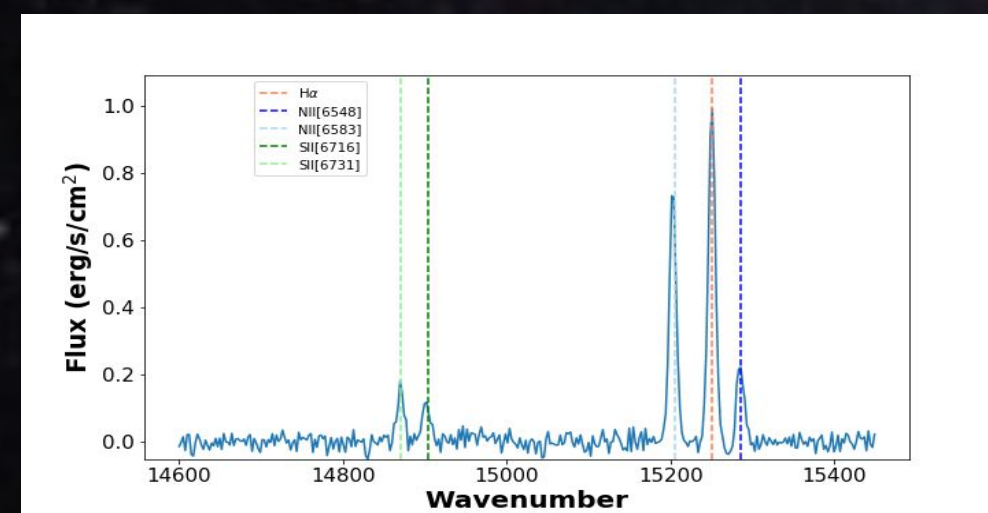
### The Issue

Standard fitting procedures require a priori estimates of kinematic parameters that are often unknown initially. Our solution is to use machine learning to estimate these variables

## Training

SITELLE Filter SN3 (657-685 nm)  
Sincgauss function with added noise  
30,000 Total

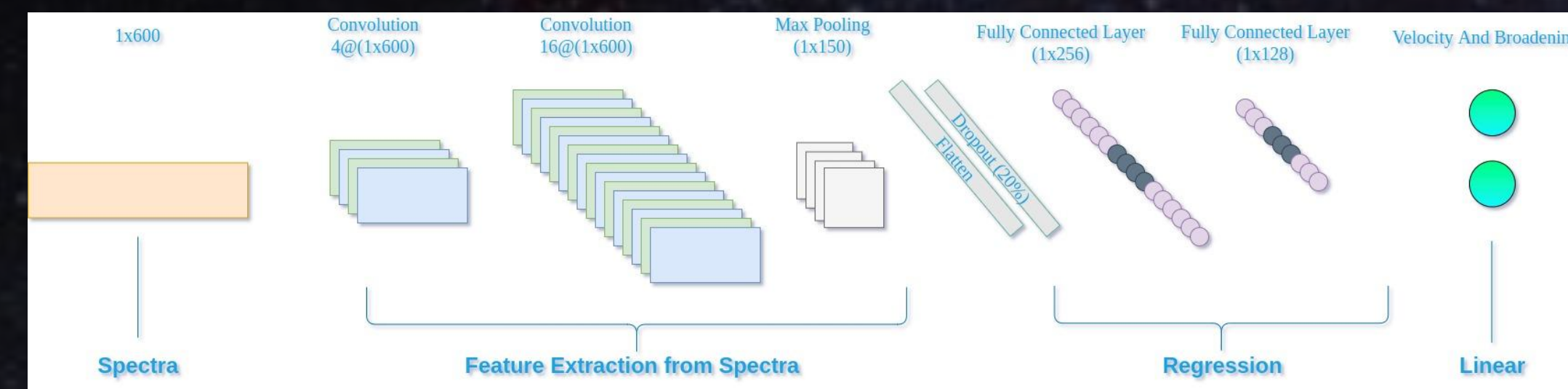
- [NII]λ6548
- [NII]λ6583
- [SII]λ6716
- [SII]λ6731
- Hα6563Å



## References:

- Martin et al. 2016, MNRAS, 463  
Fabbro et al. 2018, MNRAS, 475  
Asari et al. 2016, MNRAS, 460  
Drissen et al. 2019, MNRAS, 485

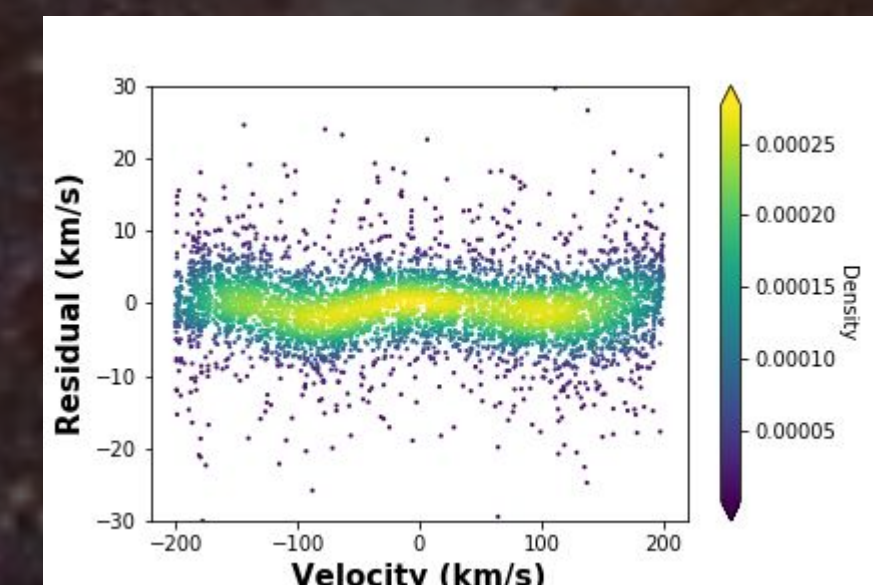
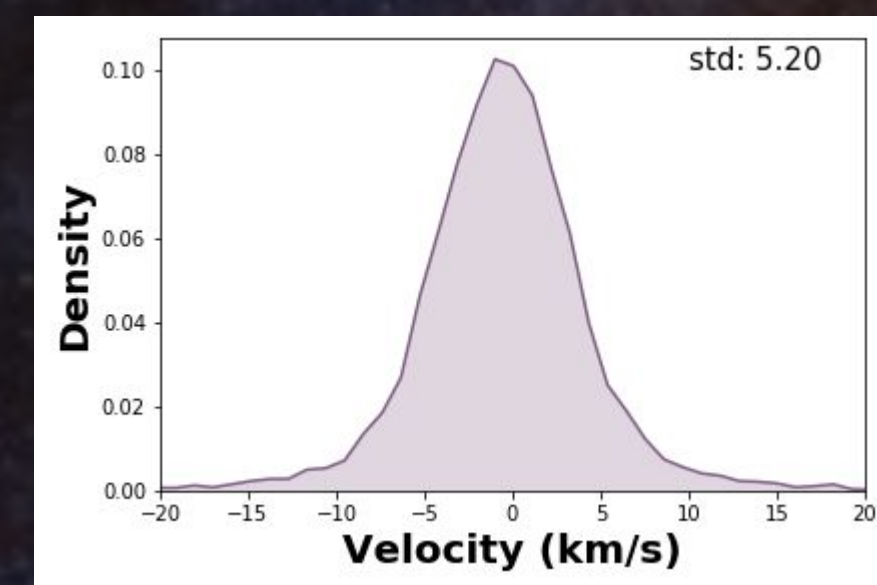
## The Convolutional Neural Network



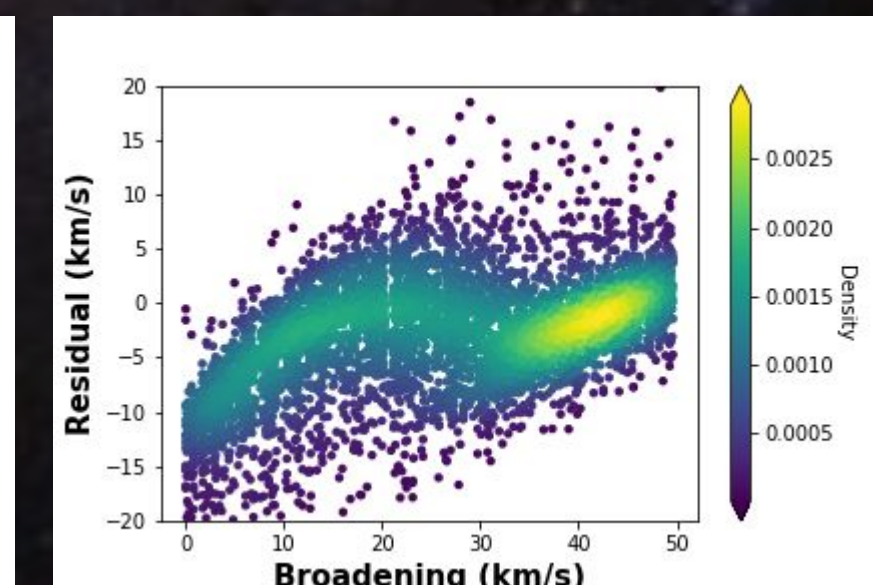
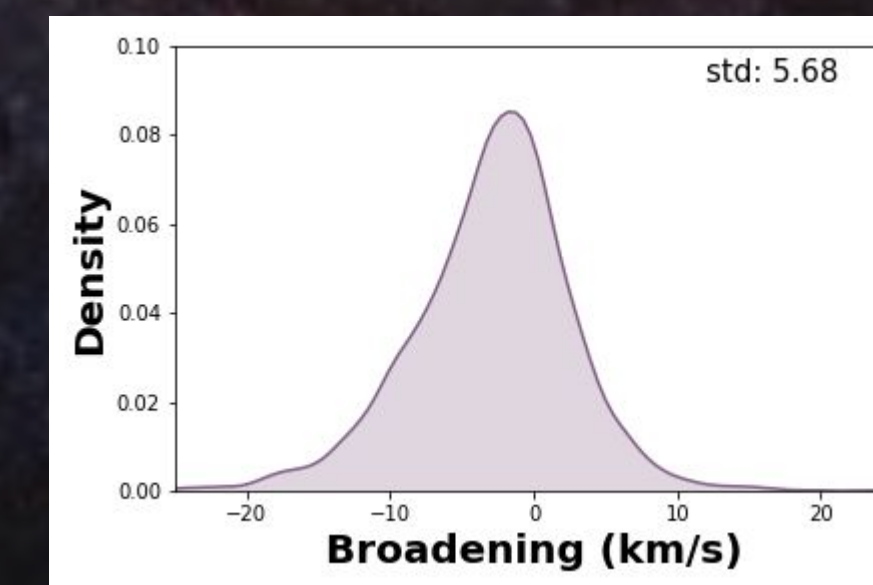
- Input: Spectrum Vector
- Activations: RELU
- Optimizer: ADAM
- Final: Linear
- Output: Velocity and Broadening in km/s

## Velocity & Broadening Results

Divided Synthetic Data into Training Set (70%), Validation Set (20%), and Test Set (10%)

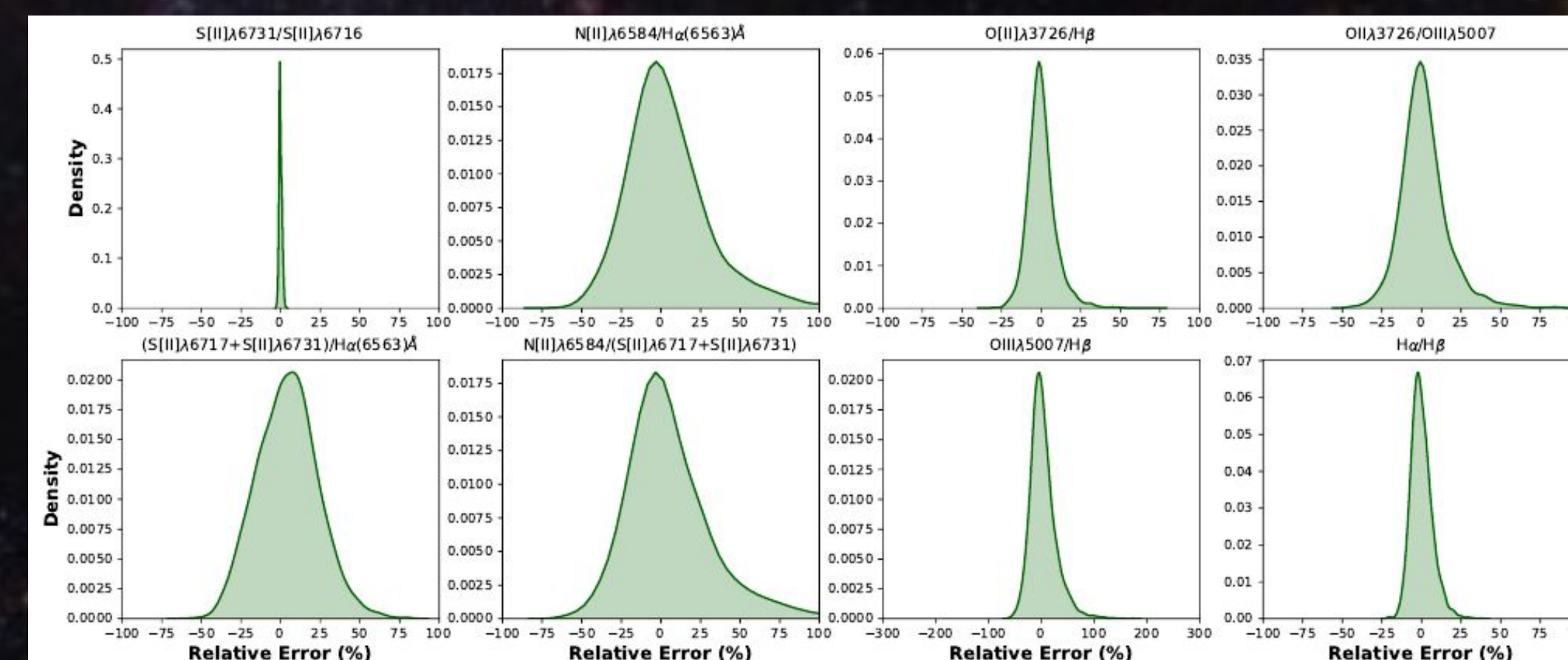


Errors are similar to those calculated using ORCS fitting software



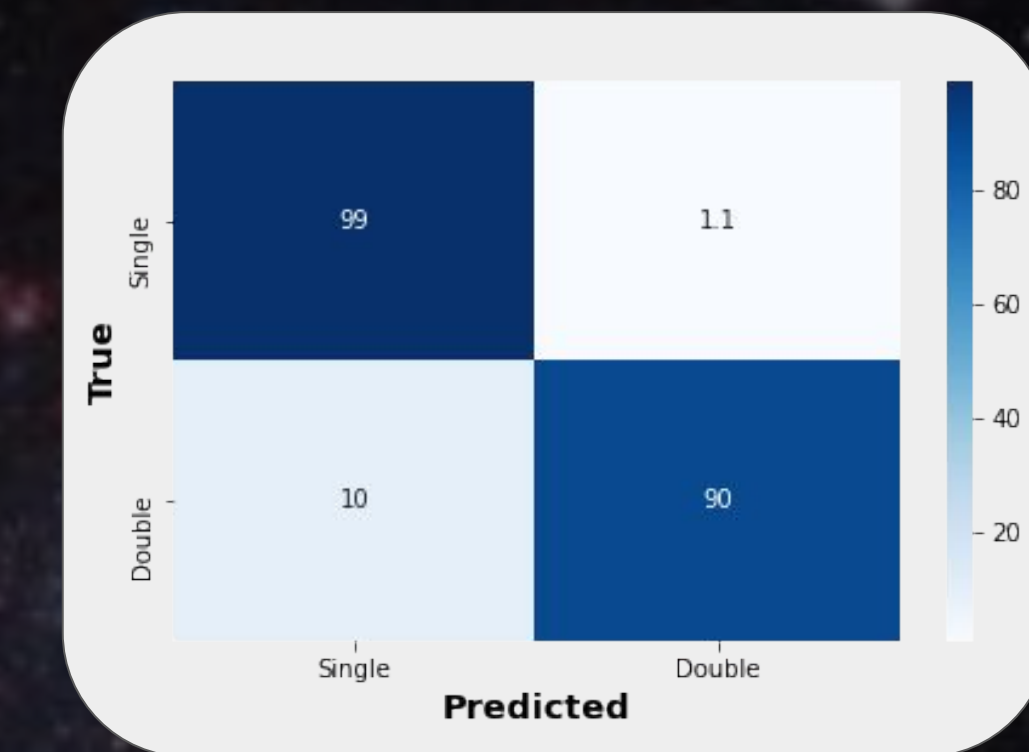
Errors are close to the resolution limit of SITELLE

## Emission Line Ratio Results



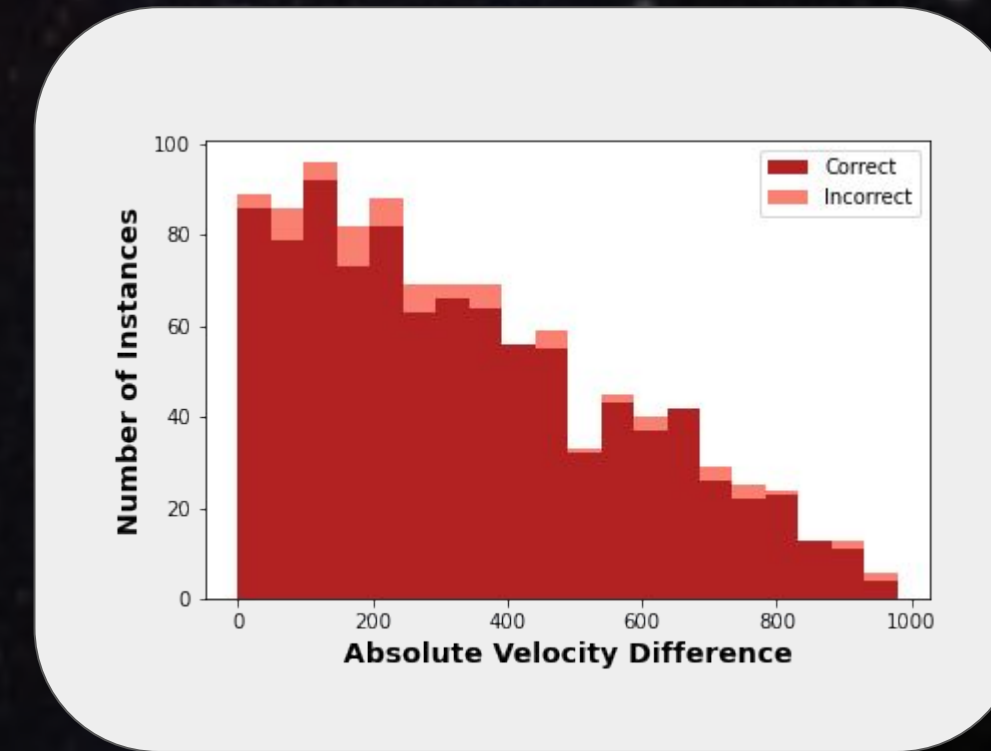
[SII]6716/[SII]6731  
[SII]/[NII]  
[SII]/Hα  
[NII]/Hα  
[OII]/[OII]  
[OII]/Hβ  
[OII]/Hβ  
Hα/Hβ

## Multiple Component Results



Confusion Matrix on the test set demonstrating the network's ability to determine the number of components

Histogram of correct/incorrect classifications as a function as the absolute velocity difference for lines with two components



## Conclusions:

- ★ Machine Learning is a viable method for emission line analysis
- ★ Training set must be developed with care
- ★ Errors are similar to those achieved with traditional methods
- ★ Considerably less computationally expensive than traditional methods

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<https://github.com/sitelle-signals/Pamplemousse>