TFAW K2 SURVEY: AN INSIGHT INTO A FEW HUNDREDS OF NEW EARTH-SIZED PLANETARY CANDIDATES



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OVERVIEW

- We performed a uniform reanalysis of \sim 300K EVEREST 2.0-corrected light curves from K2 campaigns C1-C8 and C12-C18.
- ► We combine TFAW's improved photometric precision and TLS signal detection efficiency (SDE) to search for new transit candidates.
- ► We present a first set of **220** planet candidates validated with VESPA.

TFAW ALGORITHM

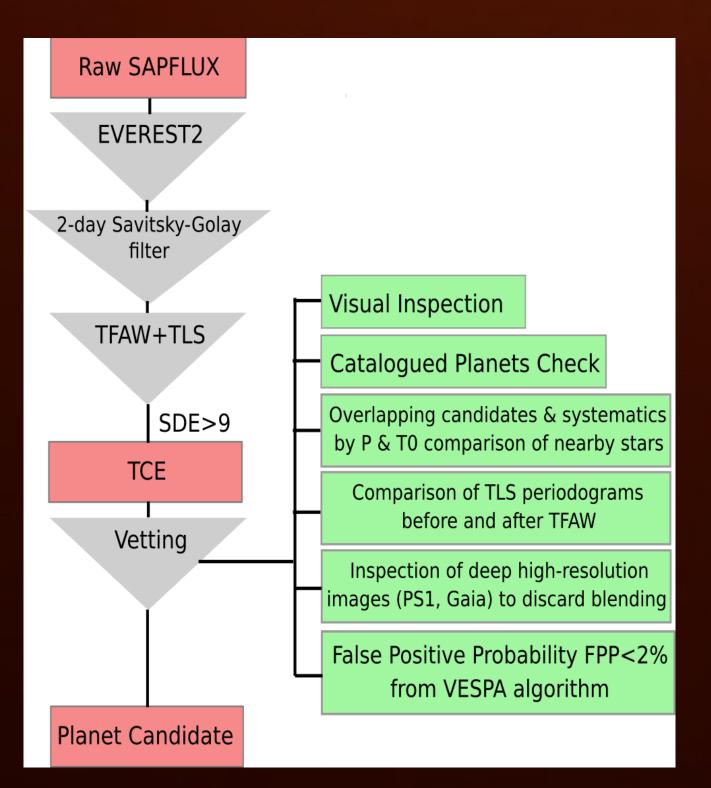
- Wavelet-based modification of the Trend Filtering Algorithm (TFA) [5].
- Divided in two steps:
 - Frequency analysis: Remove systematics, noise and outliers. Perform period search.
 - Iterative signal reconstruction: Once a significant period is found, the final light curve is reconstructed and de-noised.
- TFAW yields improved photometric precision, increased transit detection efficiency and better transit characterization [2].
- ► **K2-327 b & K2-328 b**: TFAW discovers two K2-C1 Earth-sized validated candidates [1].

TRANSIT SEARCH

- Transit Least Squares (TLS) [4] optimized to detect small planets.
- We combine TLS with TFAW's frequency analysis step to enhance the detection of transiting planet candidates.
- Stellar parameters obtained from "K2 star parameters from Gaia & LAMOST" [3] and EPIC catalogs.
- ► We consider **significant periods** those with SDE \geq 9.0 (false-positive rate <10⁻⁴).

VETTING AND VALIDATION

- All light curves that have a significant detection undergo TFAW's iterative signal reconstruction and denoising. We visually inspect them and keep the ones which show transit-like features.
- The full vetting procedure consists on:



- Transit 0.5 days away from the beginning/end of any light curve gaps.
- Rule out eclipsing binary nature by checking odd/even transit depths.
- Catalog comparison to discard already found candidates.
- Check EVEREST 2.0 [6] aperture image for contaminating sources.
- Rule out that no other light curve(s) in the same CCD module present transit-like features with similar periods and T_0 .
- Comparison of TLS periodograms before/after TFAW.
- Inspection of high resolution images from Pan-STARRS1 and Gaia DR2 to discard close stellar companions.
- ▶ VESPA [7] computes the probabilities of the transiting signal to be caused by six different astrophysical false positive scenarios with transiting planets.
- Only targets with FPP lower than 2% are considered as valid planetary candidates.

REFERENCES

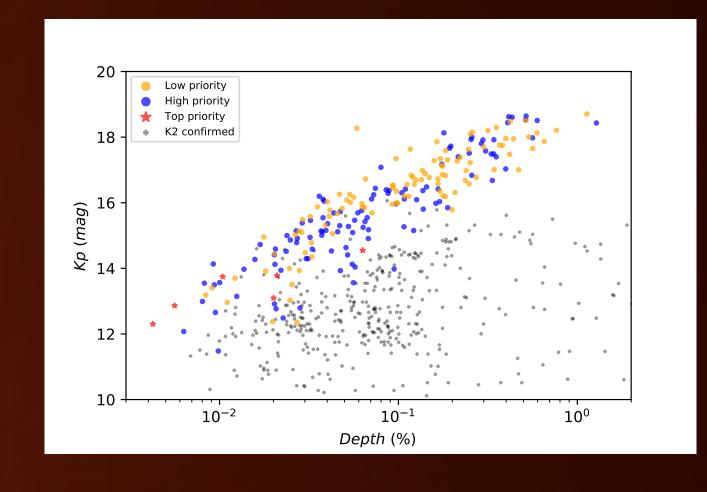
- [1] D. del Ser and O. Fors. In: *MNRAS* 498.2 (Oct. 2020), pp. 2778–2797.
- [2] D. del Ser, O. Fors, and J. Núñez. In: *A&A* 619, A86 (Nov. 2018), A86.
- [3] Kevin K. Hardegree-Ullman et al. In: *ApJS* 247.1, 28 (Mar. 2020), p. 28.[4] Michael Hippke and René Heller. In: *A&A* 623, A39 (Mar. 2019), A39.
- [5] G. Kovács, S. Zucker, and T. Mazeh. In: *MNRAS* 356 (Jan. 2005), pp. 557–567.
- [6] Rodrigo Luger et al. In: *AJ* 156.3, 99 (Sept. 2018), p. 99.[7] Timothy D. Morton. In: *ApJ* 761.1, 6 (Dec. 2012), p. 6.

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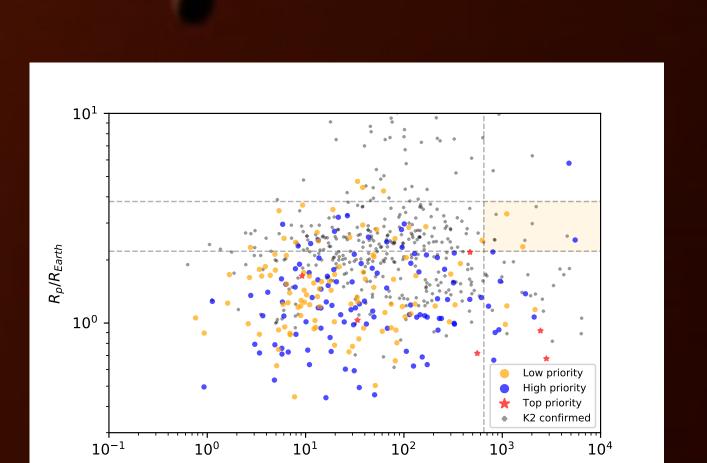
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STATISTICAL RESULTS

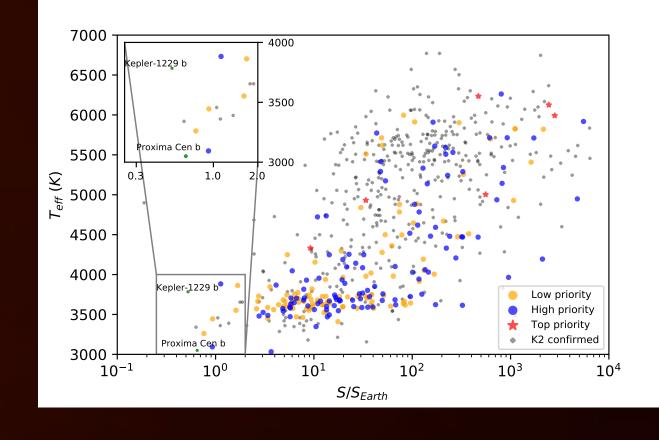
- ► We present a sample of **220 new planet candidates** detected with TFAW and TLS from K2 campaigns C1-C8 and C12-C18.
- All have passed our vetting procedure and validated by VESPA.
- Classified in three priority ranks: 6 "top" (red): detected in two or more K2 campaigns; 110 "high" (blue): vetted and validated; and 104 "low" (orange): vetted and validated but either have SDEs~9, FPP~2% or need further reanalysis.



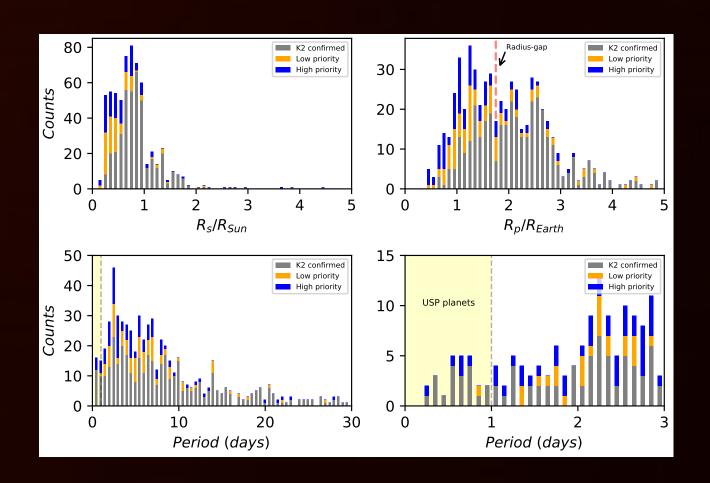
- TFAW opens a **new niche** for exoplanetary candidate search.
- It allows to detect transiting objects around fainter stars than any other previous survey.
- Most of our candidates orbit low-mass stars.
- Four candidates lie in the conservative HZ of their host stars.
- Two other lie in the optimistic HZ.



- We mainly populate the **small** planet radius range ($< 1.5R_{\oplus}$).
- We duplicate the sample of candidates within the small planet radius gap ($\sim 1.8R_{\oplus}$).
- We also add some candidates in the Ultra-Short-Period (USP) range.

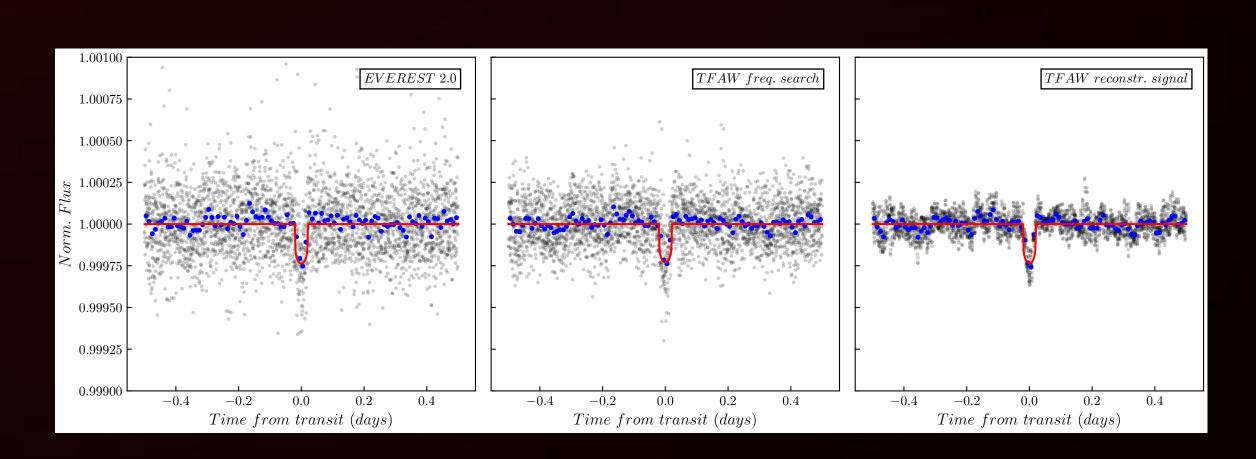


- No Jupiter-sized candidates found around M-dwarfs.
- Four candidates lie within the Hot-Super-Earth desert.



EPIC 248828117.02: A NEW EARTH-SIZED CANDIDATE FROM C14 FOUND BY TFAW

- We present EPIC 248828117.02 to illustrate the noise filtering efficiency of TFAW.
- ► EPIC 248828117.02 is a $1.3R_{\oplus}$ planet candidate orbiting a G3 star with a 2.896 day orbit.



Normalized phase-folded EPIC 248828117.02 light curve with TLS model data marked in red for EVEREST 2.0 (left), TFAW's frequency analysis step (middle) and for the TFAW-reconstructed signal (right)). Blue represents the binned data.