

Multiwavelength analysis of galaxies' structure to unveil quenching signatures

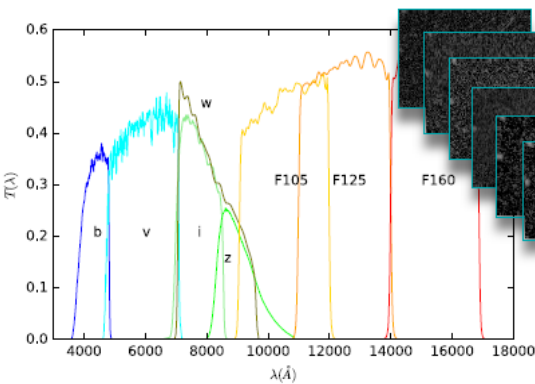
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Bulge-disk decomposition

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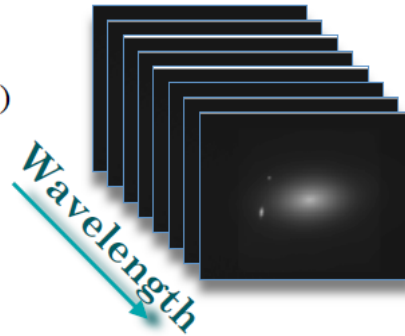
1) Modeling the surface brightness profile



7/4 GOODS/CANDELS
bands (400 - 1500 nm)

GalfitM
(Häussler et al. 2013)

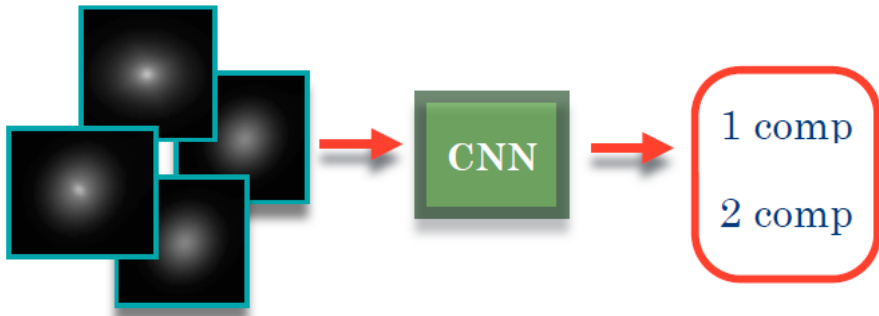
$Mag(\lambda)$
 $n(\lambda)$
 $Re(\lambda)$
 $q(\lambda)$



Sersic profile

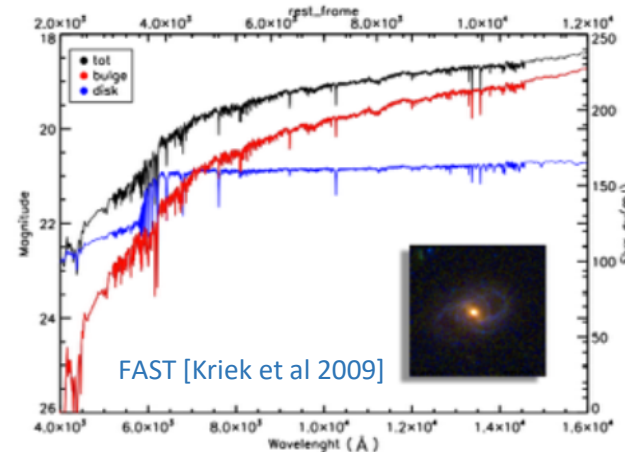
exp. disk

2) Best model selection



[Dimauro et al. 2018]

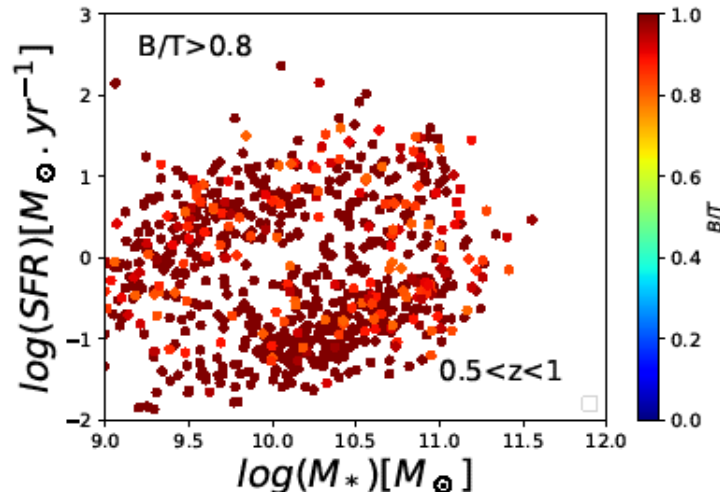
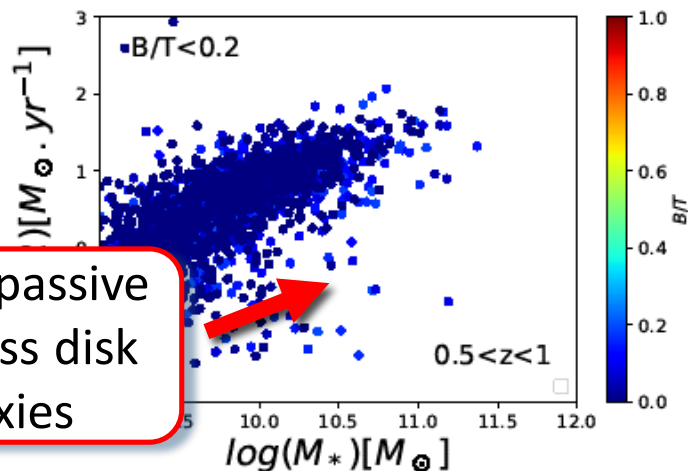
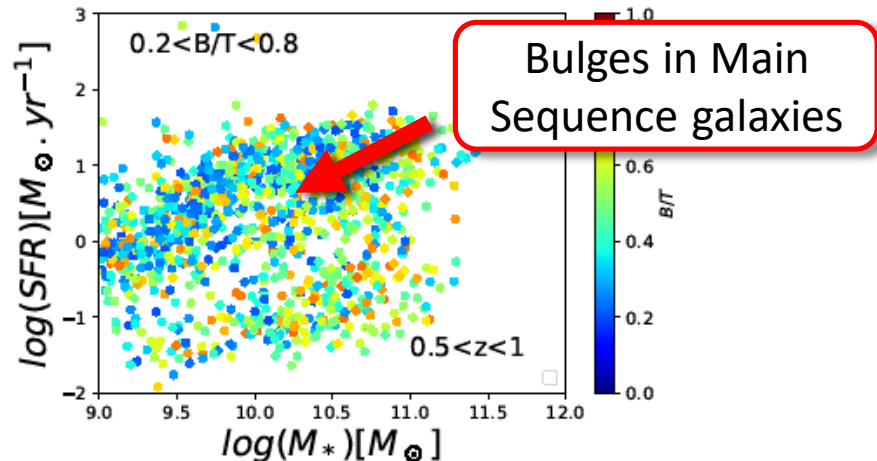
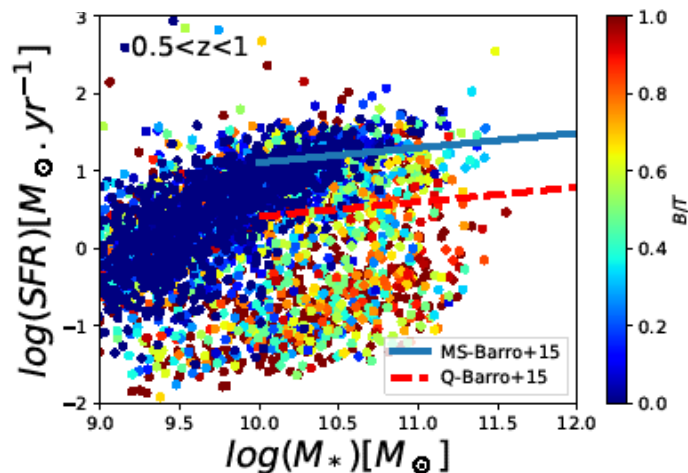
3) Spectral Energy Distribution



- Stellar masses
- rest-frame colors

Star formation and Bulge/Total mass ratio

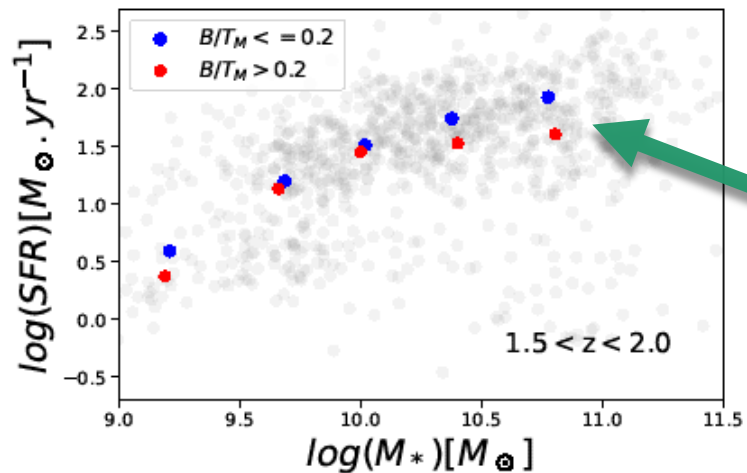
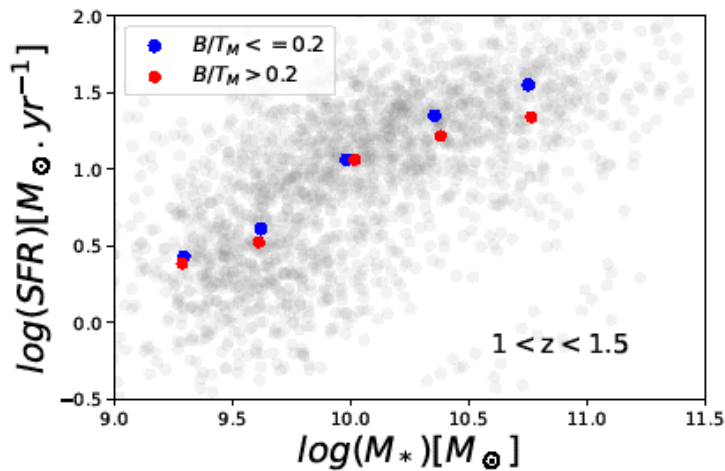
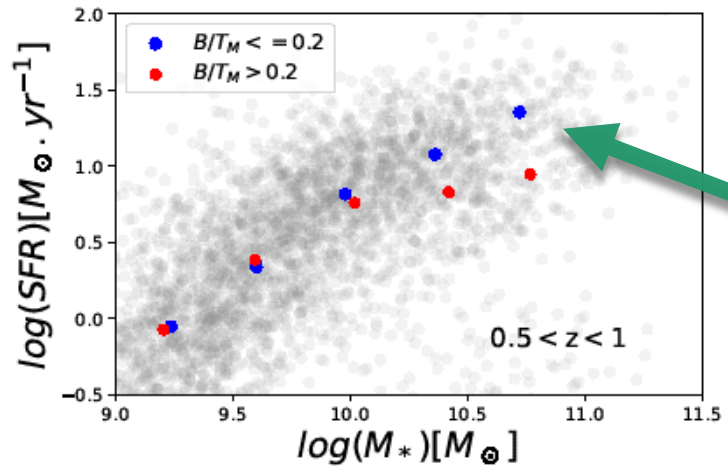
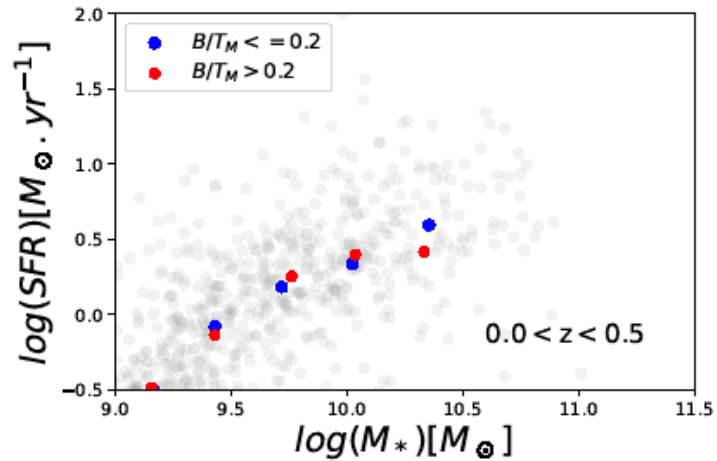
[Dimauro et al. 2020 in prep.]



Bulge is a required but not sufficient condition to quench

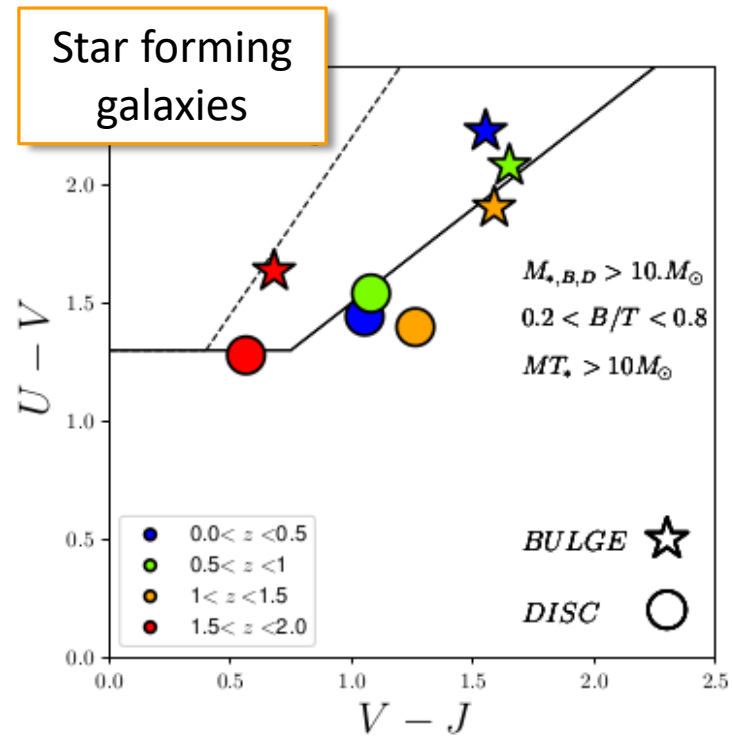
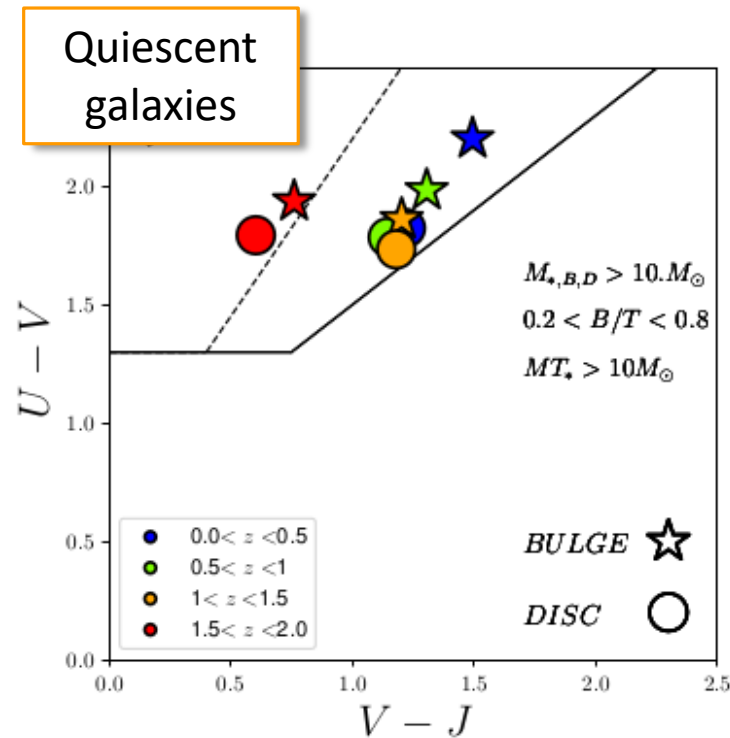
Bending of the Main Sequence

[Dimauro et al. 2020 in prep.]



Linked to the presence of massive bulges

Rest-frame colors



All the galaxies show a negative color gradient.
Possible signature of inside-out quenching

The star formation is localized in the disk

[Dimauro et al. 2020 in prep.]

Main results

The structure of internal components weakly depend on the star-formation activity neither on the morphology of the hosting galaxy
[Dimauro et al, 2019]

The bending of the main sequence is directly linked to the morphological gradient.

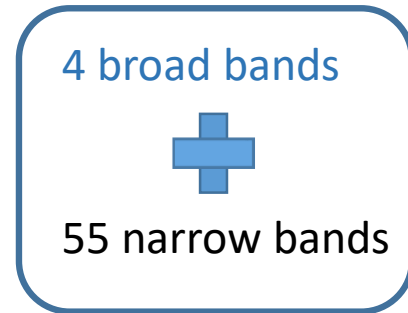
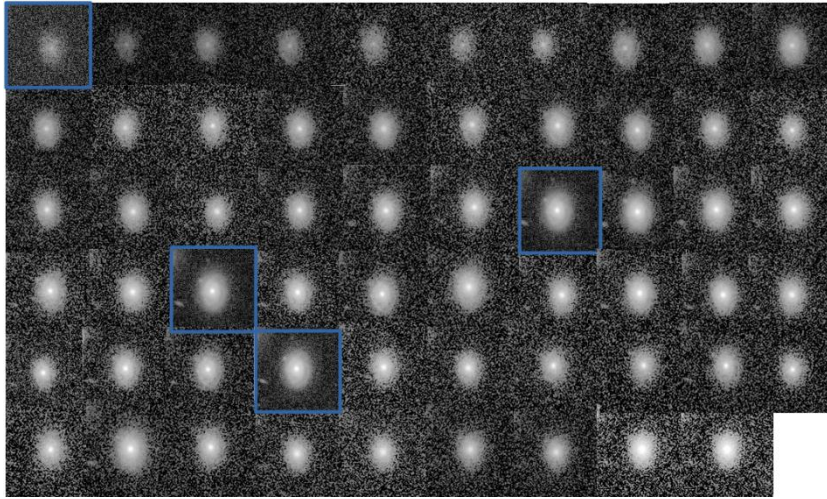
No passive bulgeless disk are observed: bulge growth is a required condition to quench.

Galaxies with a significant bulge populate the main sequence:
bulge formation starts while galaxies are still star forming.
Bulge is a required but not sufficient condition to quench.

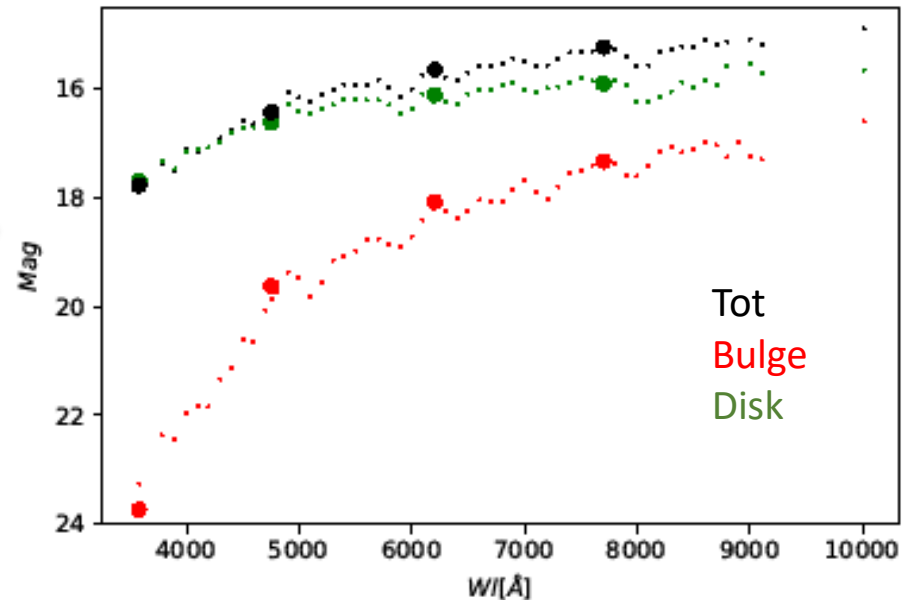
All the galaxies show a negative color gradient. Bulge are always quenched. The star formation is localized in the disk.

Morphological transformation as a main quenching path

Bulge-disk decomposition with J-PAS

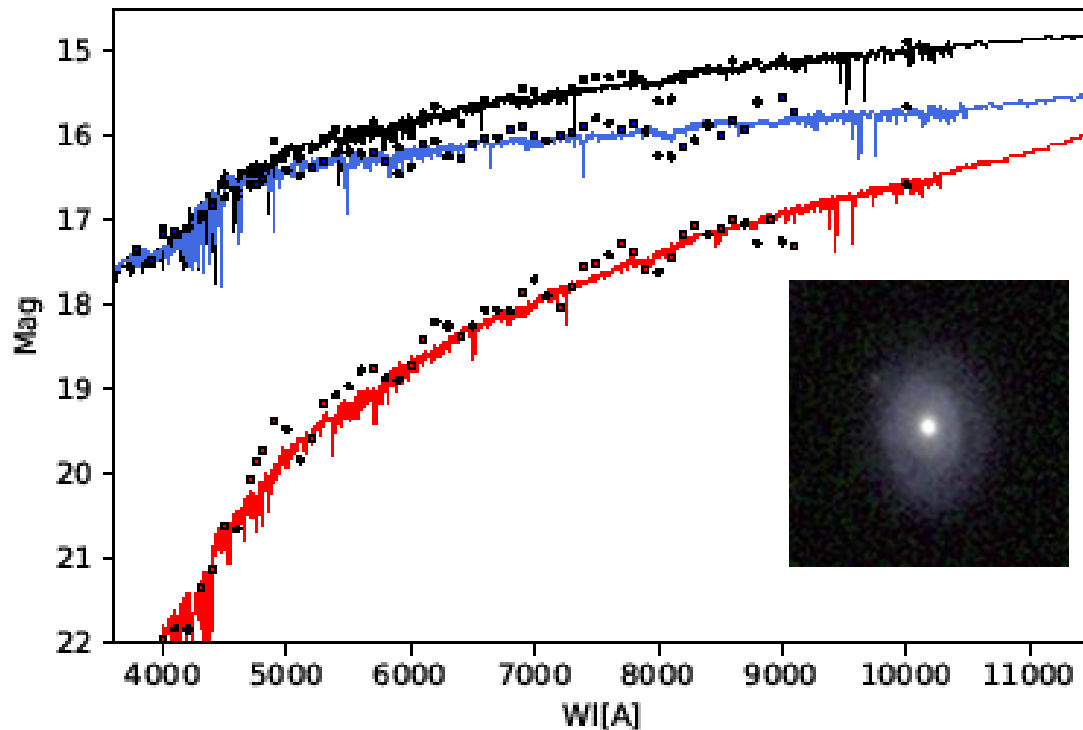


GalfitM
[Häußler et al. 2013]



Bulge-disk decomposition with J-PAS

Promising results
combining broad and
narrow bands!



Preliminary results

	Bulge	Disk
Mass	10.52	10.8
Isfr	-	0.42
Av	1.90	0
lage	10	9.60

SED fitting code: FAST [Kriek et al 2009]

What else:

- Morfometryka to automatically classify galaxies (Ferrari et al. 2015)
- Curvature method (Lucatelli, Ferrari et al. 2019) to predict number of components as well as to estimate starting values for GalfitM
- Compare final modes with the output of CICLE (Jiménez et al)

See the poster of Cortesi et al.