



KamLAND-Zen and SNO+

Christopher Grant

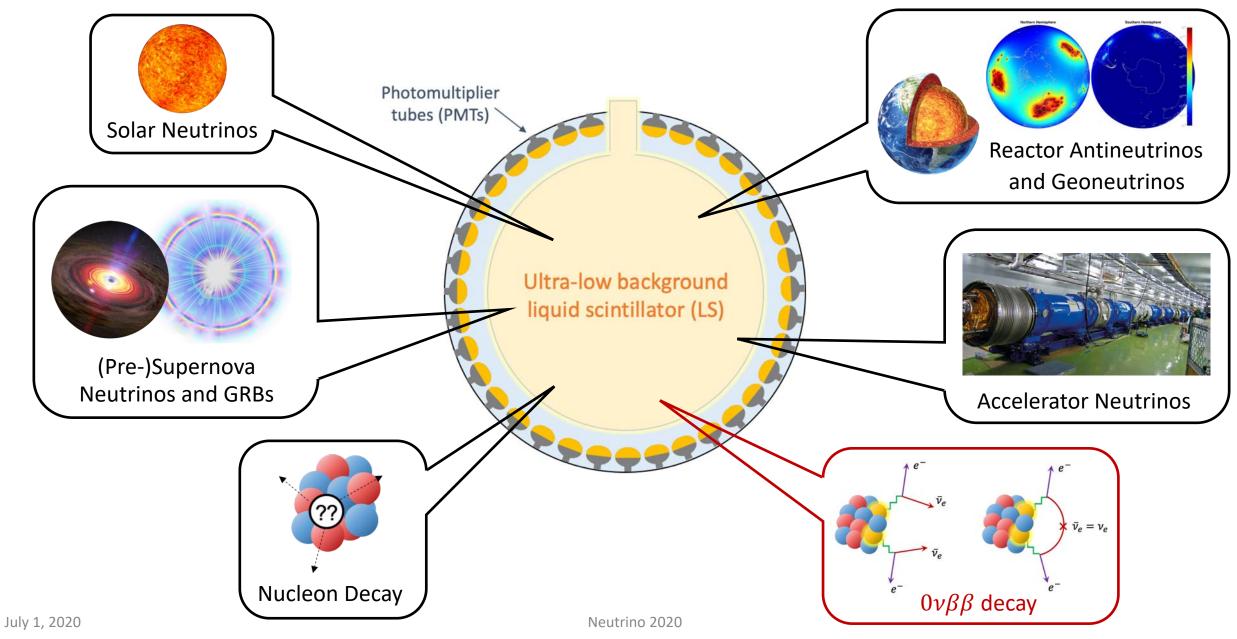
Boston University



On behalf of the SNO+ and KamLAND-Zen Collaborations



Pursuing Crosscutting Science with Big Detectors



1

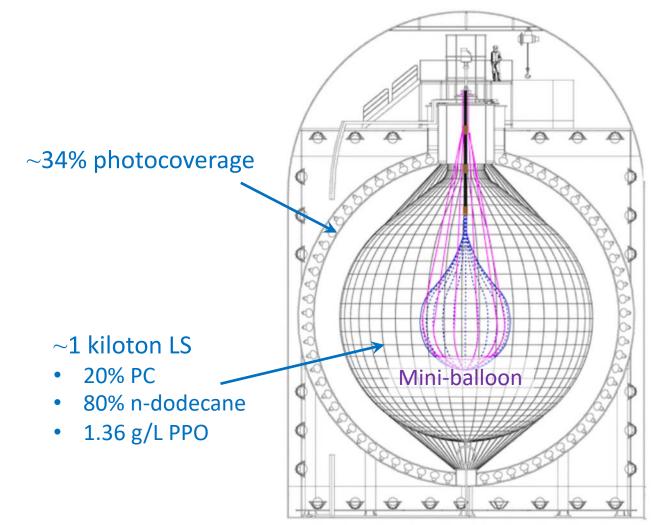
KamLAND-Zen

Mini-balloon:

- 25- μ m-thick nylon film (durable)
- Fabricated in class-1 clean room
- Highly transparent (~99% at 400 nm)

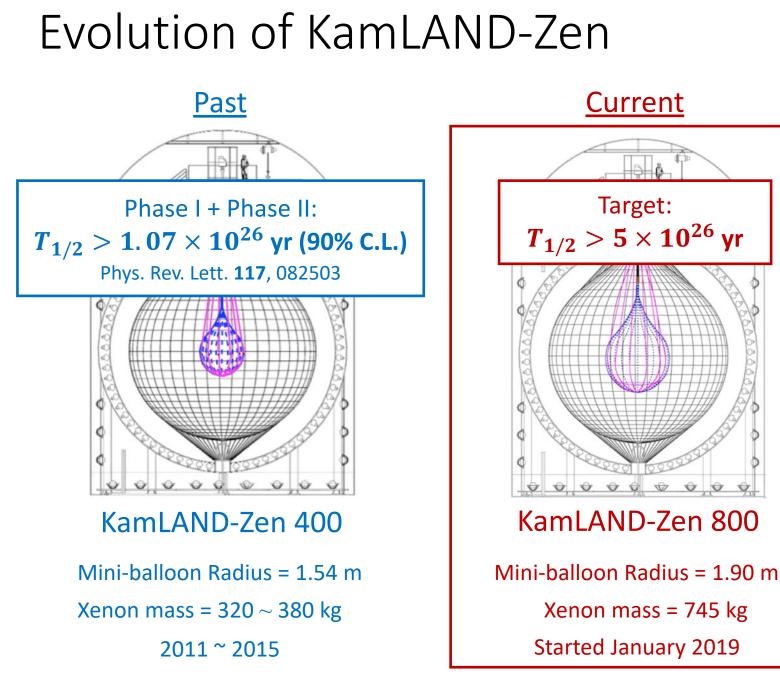
Xenon loading:

- Chemically stable (noble gas)
- Good solubility (3.2% wt in LS)
- Removable from LS
- Purification is well-established

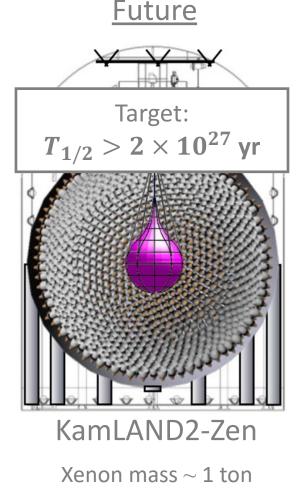


91% enriched ¹³⁶Xe loaded in LS inside mini-balloon (Q value = 2.4578 MeV)

2



"Future Neutrinoless Double Beta Decay Experiments" Jason Detwiler (next Session)

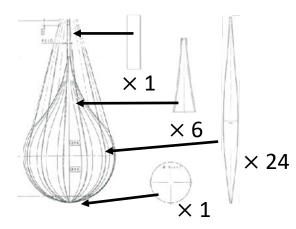


× 5 increase in light collection Scintillation balloon film

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Mini-Balloon Fabrication in Sendai (2017 - 2018)

All work performed inside a Class 1 clean room in Sendai





1 Film Washing



(2) Seam Welding



3 He leak test + repairs

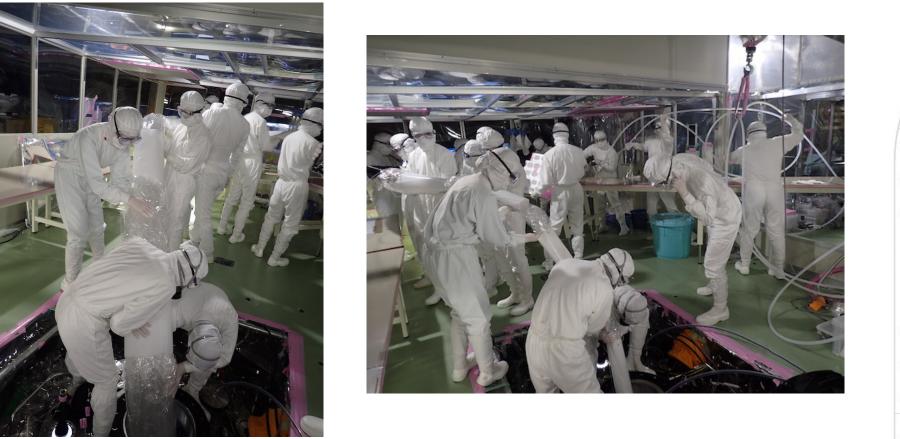


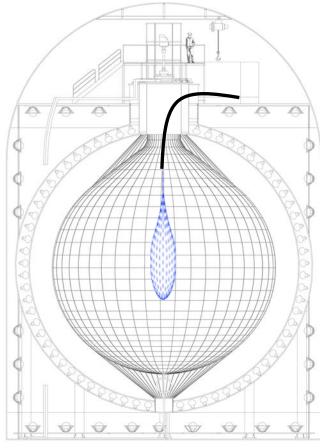




(5)Packaging

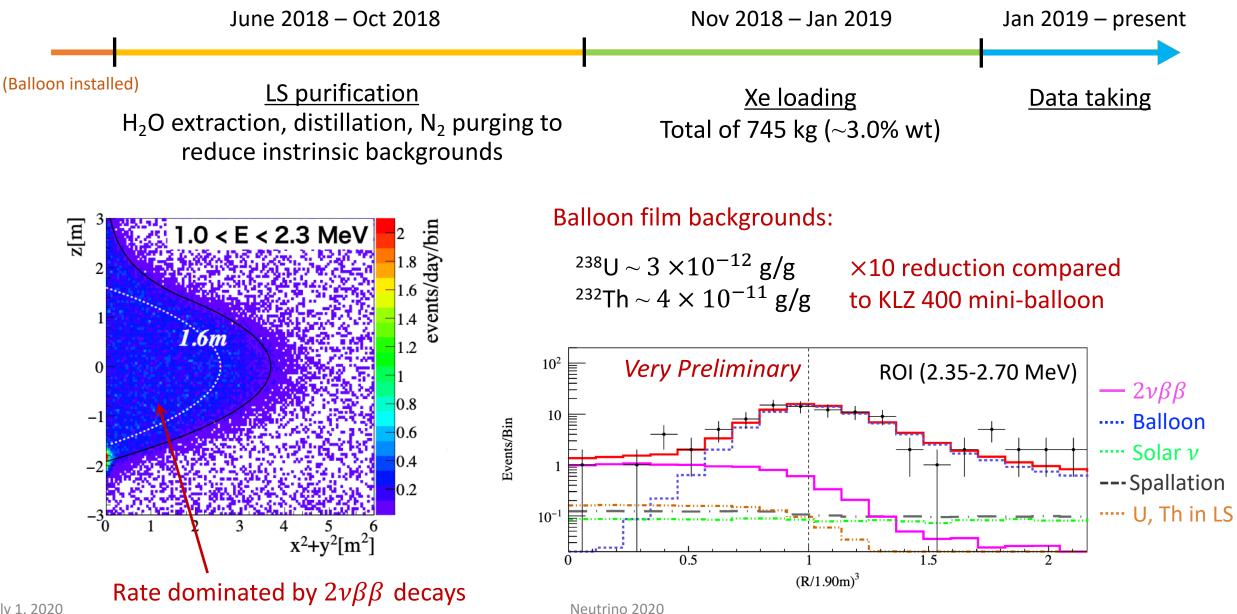
Mini-Balloon Installation (May 2018)





...after nearly 1.5 years, new mini-balloon fabrication and installation was finished

LS Purification and Xe Loading



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July 1, 2020

Event Selection (the following was presented at TAUP 2019)

Very Preliminary 10^{4} Events/0.05MeV 10^{3} 214**B** 208**T** 10^{2} 10 0.5 1.5 3.5 4.5 2.5 2 3 4 Visible Energy (MeV)

Total livetime of 132.7 days

Cuts used to reduce backgrounds:

r < 240 cm

Select events inside and just outside of the mini-balloon

Rn cut

Delayed coincidence cut for ²¹⁴Bi – ²¹⁴Po and ²¹²Bi – ²¹²Po

Fiducial volume cut

Further reduce backgrounds going from r < 240 cm to r < 157 cm

Spallation Cut

Remove events correlated with muons

Fitting the Data

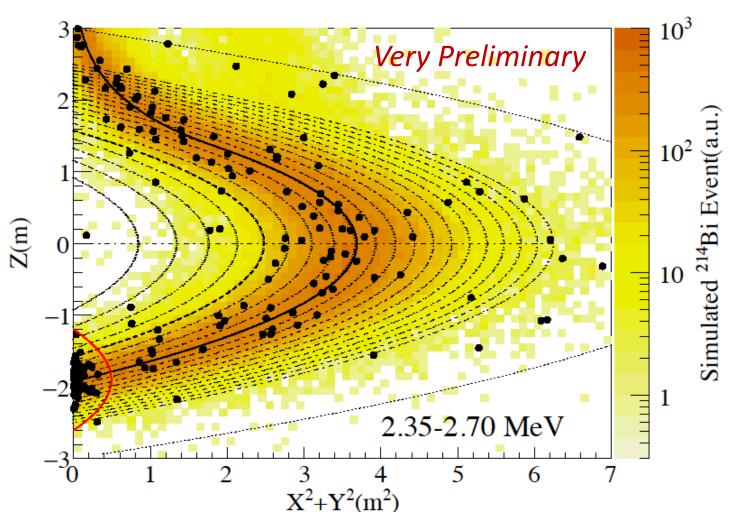
 Simultaneously fit 40 equal volume bins inside of r < 2.5 m

> **Outer region** \rightarrow more sensitive to backgrounds on mini-balloon film (²¹⁴Bi, etc.)

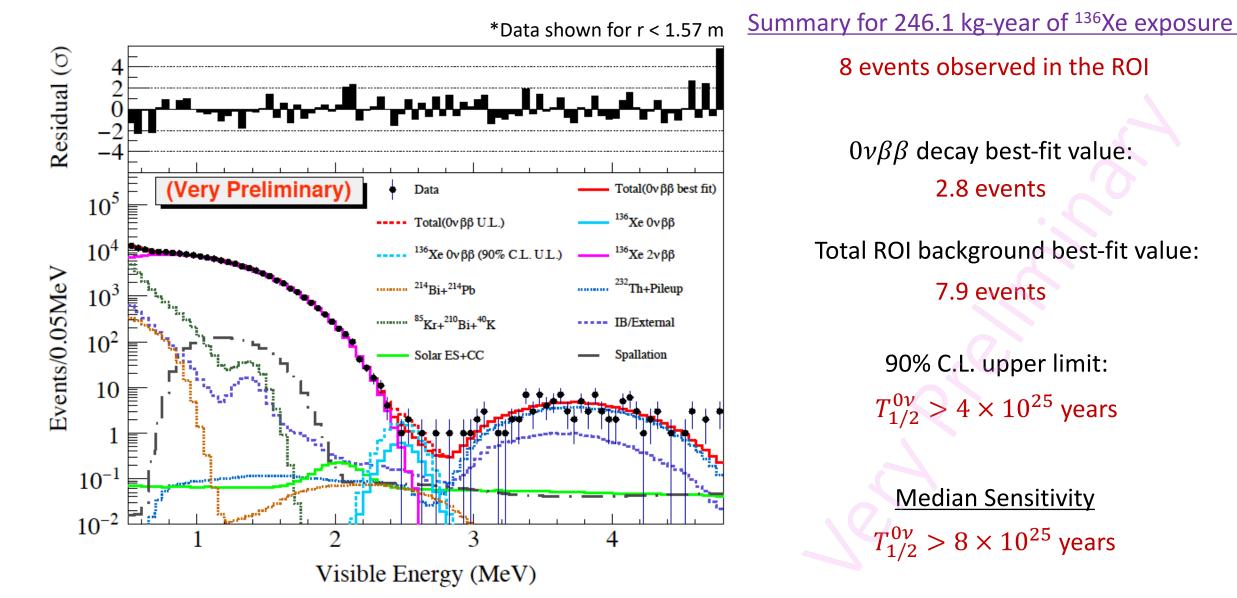
Inner region \rightarrow more sensitive to $0\nu\beta\beta$ decay

- Excess events at bottom of the miniballoon removed using spatial cut
- Separate frequentist and Bayesian fitting analyses performed with 31 free parameters and 40 volume bins



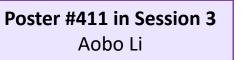


Frequentist Fitting Results



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Bayesian Fitting Results





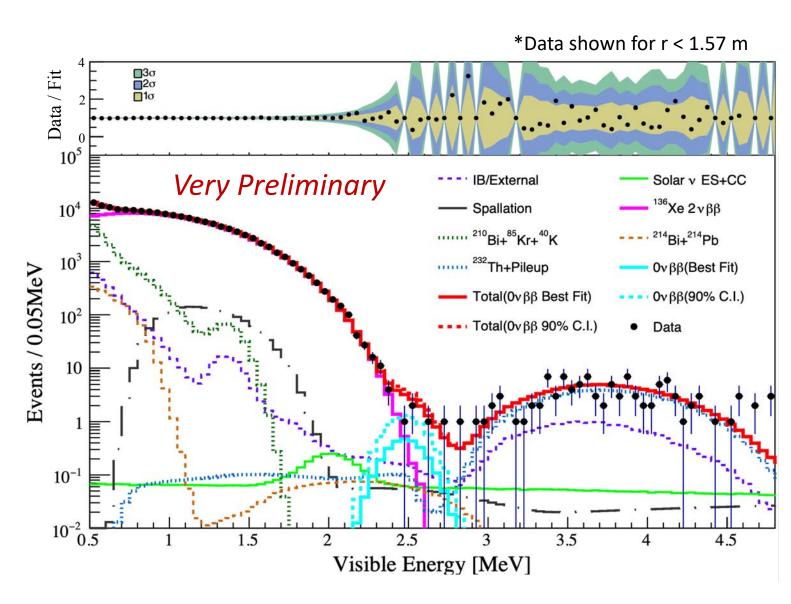
 $0\nu\beta\beta$ decay best-fit value: 2.6 events

Total ROI background best-fit value: 8.1 events

> 90% C.L. upper limit: $T_{1/2}^{0\nu} > 4.3 \times 10^{25}$ years

Frequentist and Bayesian analyses are in reasonable agreement.

We now have nearly twice the additional exposure for a total of 482 kg-year which is currently being analyzed.

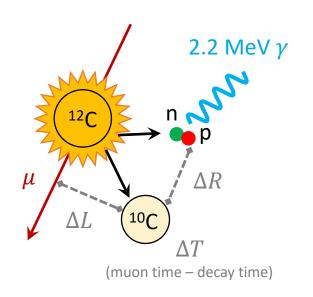


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Analysis Improvements Since TAUP 2019

Muon spallation on ¹³⁶Xe is being studied extensively

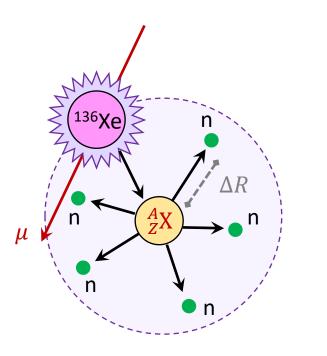
Spallation on ¹²C



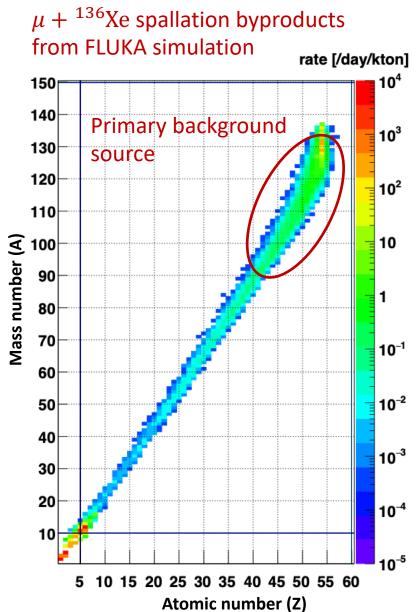
Isotopes with small *A* are rejected based on timing and spacial distance information (triple coincidence)

> Poster #106 in Session 3 Yuto Kamei

(NEW) Spallation on ¹³⁶Xe



Can a similar technique be applied to high A isotopes with high neutron multiplicity?



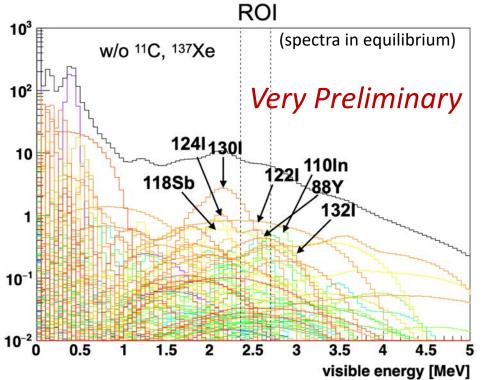
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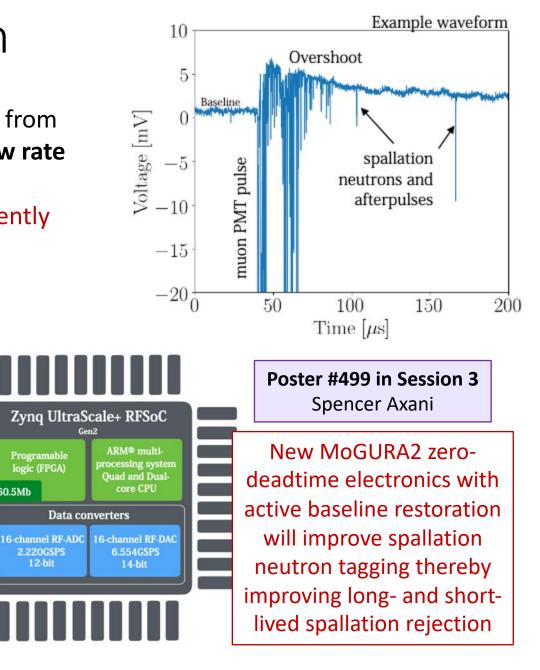
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Long-Lived Spallation Rejection

High A spallation isotopes in the ROI have half-lives ranging from several hours to several days – they also occur at a very low rate

We've developed a special day-long spallation veto to efficiently remove at least 50% of these events from the analysis





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Programable

logic (FPGA)

2.220GSPS

12-bit

60.5Mb

Background Rejection with Deep Learning

Background decays with ~MeV gamma-rays typically have energy deposits (Compton scatters) spread over distances of tens of centimeters.

Decays only containing ~MeV electrons are more localized.

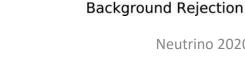
Spherical CNN applied to subset of KLZ data

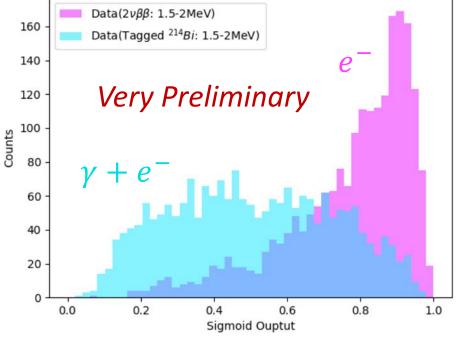
CNN applied to MC of ¹⁰C and $0\nu\beta\beta$ events in a KLZ-like detector 10 1.0 Xe-0vBB Center Normalized Counts 10-5 10-3 ³⁶Xe-0vBB 3m Sphere ¹⁰C 3m Sphere A. Li et al, NIM 947, 162604 (2019) Li et al, NIM 947, 162604 (20: Status PC(%) QE(%) Rejection 42.0 36.2 Ratio Upgrade 0.813 19.6 23.0 0.616 Current 0.0 0.6 0.7 0.8 0.9 1.0 0.5 10 20 30 40

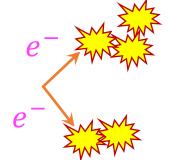
July 1, 2020

Time Since First Photon(ns)

1/







Poster #414 in Session 2 Zhenghao Fu

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Now moving from Kan



SNO+

- 12m diameter acrylic vessel with 780 tons of LS
- ~9300 inward facing PMTs with light concentrators and ~90 outward facing PMTs for tagging cosmic rays

New:

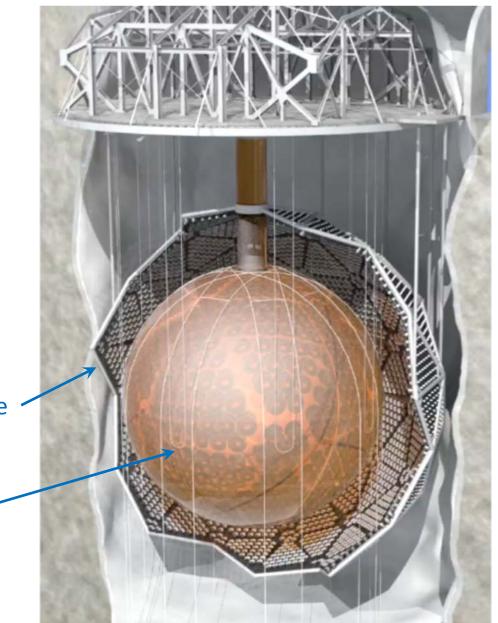
- Addition of hold-down ropes to counter buoyancy of LS
- Upgraded electronics for higher data-taking rates
- New calibration systems for LS

~50% photocoverage

780 tons of LS

- LAB
- 2 g/L PPO

Located at 6800 ft depth in SNOLAB (6000 m.w.e.)



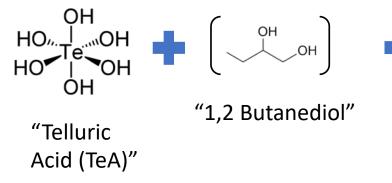
Neutrino 2020

¹³⁰Te Loading in Liquid Scintillator

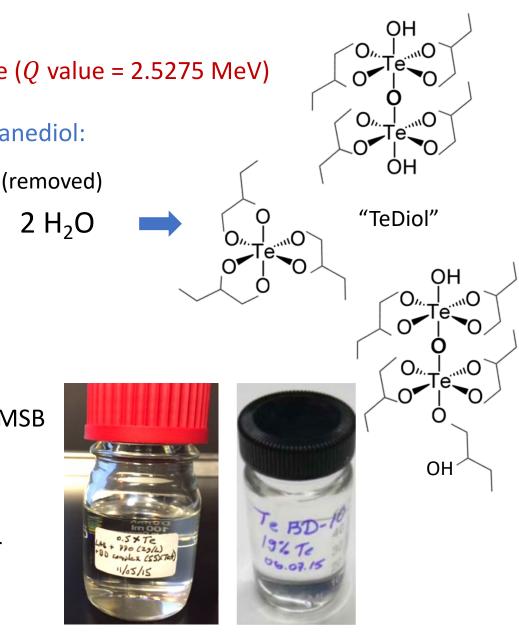
¹³⁰Te makes up 34% of the natural Te abundance (*Q* value = 2.5275 MeV)

Forming an organometallic compound from telluric acid and butanediol:

HO,



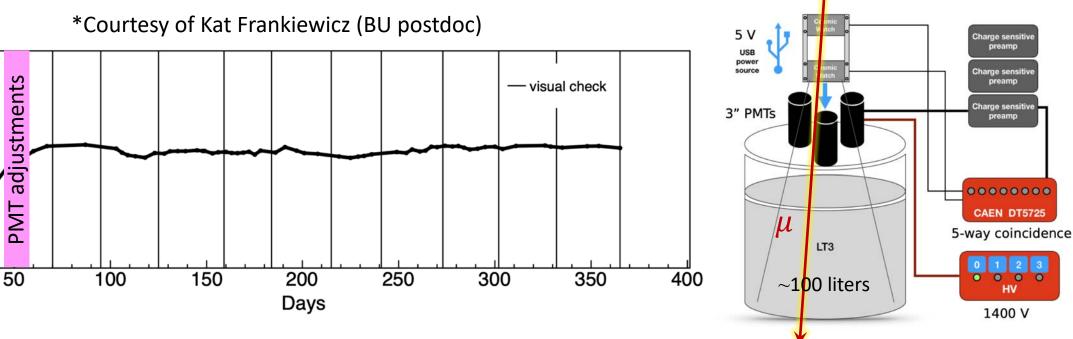
- TeDiol (TeBD) is mixed directly into SNO+ LS with 15 mg/L bis-MSB and a stabilizer called Dimethyldodecylamine (DDA)
- Optical transparency and light yield of the final Te-loaded LS cocktail are expected to produce \sim 460 p.e. / MeV in SNO+ for 0.5% ^{nat}Te loading by weight



¹³⁰Te Loading in Liquid Scintillator

- Chemistry and long-term stability of Te-loaded LS has been studied extensively by SNO+
- One example: Light yield of ~100 liters 0.5% ^{nat}Te-loaded LS was measured by Long-Term Test Tank (LT³) at BU. Light yield been stable for more than a year and counting.





1600

1400

1200

1000 귣

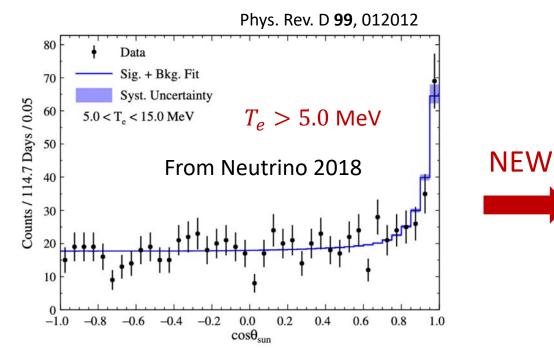
800

600^[]

ADC count

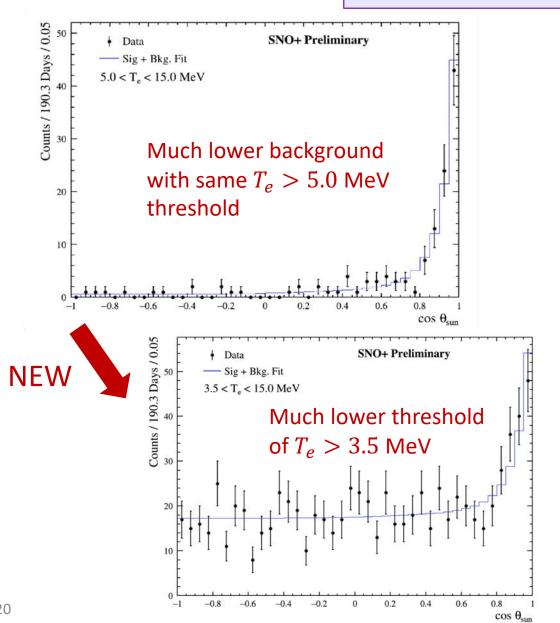
Improved Background Suppression in Water

Poster #424 in Session 4 Brian Krar



Radon ingress down the neck of the detector was mitigated with a N_2 cover gas, further suppressing the already low backgrounds and enabling a much lower threshold on the solar neutrino analysis

Same strategy is being used during scintillator filling



LS Purification and Filling

Purification and Filling Systems

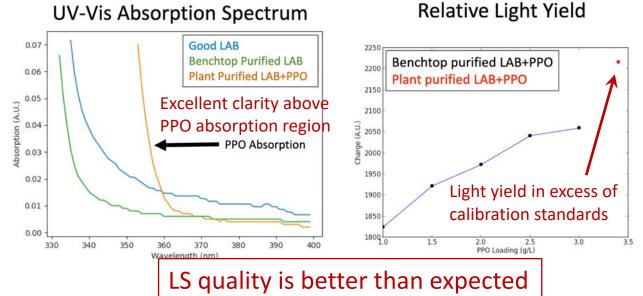


Transfer of LAB from surface to underground in tank railcars at SNOLAB

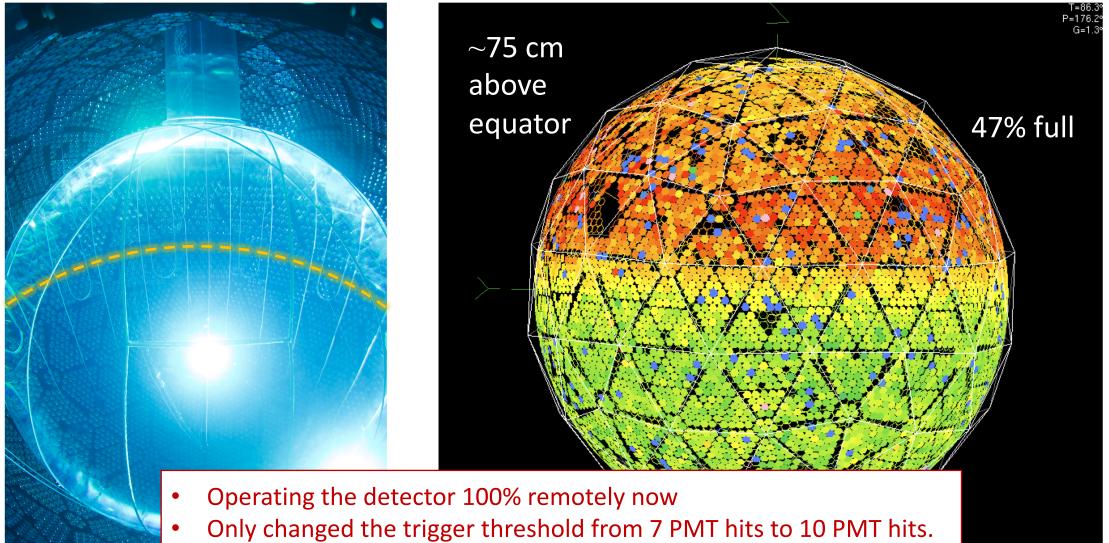


SNOLAB is leading the scintillator filling campaign

- Commissioned underground systems and began filling SNO+ in 2019.
- Filling was temporarily halted in April 2020 due to the pandemic, but as operations begin to fully resume, the scintillator fill will soon be completed.
- Right before pandemic we filled 75 tons in one week.



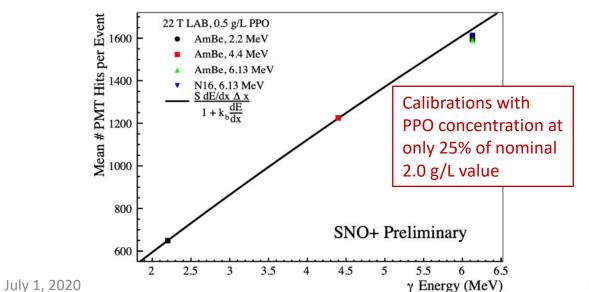
Nearly half-filled with LS since April 2020

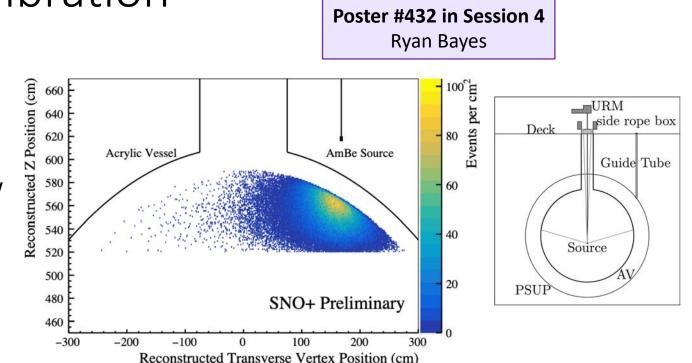


Previously triggered at \sim 1 MeV in pure water, but now at \sim 35 keV!

Partial LS Fill Detector Calibration

- Detector response during LS fill was measured with optical and radioactive source calibrations
- Source deployments through guide tubes in regions outside the acrylic vessel leave the new LS undisturbed and avoid contamination
- Demonstrated the capability to reconstruct events in a hybrid LS/water detector



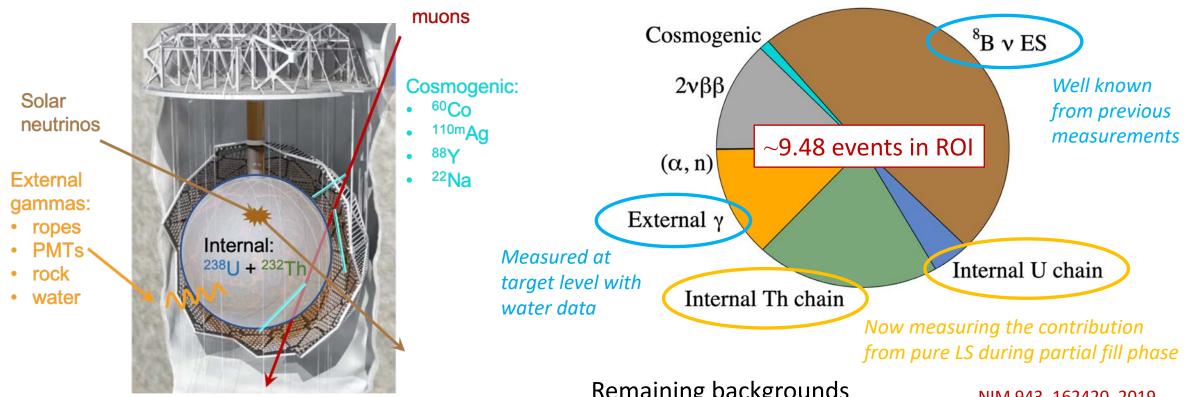


With a PPO concentration of only 0.5 g/L (25% of the nominal value) we see a light yield equivalent to ~300 p.e. / MeV

Extrapolates to ${\sim}650$ p.e. / MeV at 2.0 g/L PPO

$0\nu\beta\beta$ Background Predictions

ROI: 2.42 - 2.56 MeV $[-0.5\sigma - 1.5\sigma]$

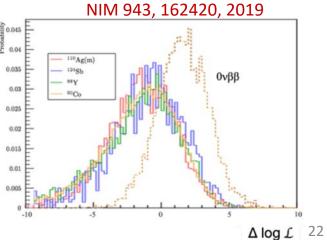


Target concentrations of less than 10^{-15} g/g U and 10^{-16} g/g Th in pure LS are required for $0\nu\beta\beta$ decay.

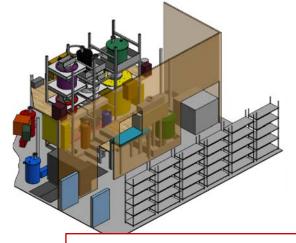
We're currently below our targets for U and Th in the partial LS phase.

Remaining backgrounds will be measured during Te-loading!

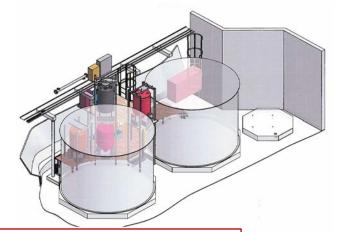
Cosmogenic backgrounds will be verified by multisite analysis



Commissioning Tellurium Plants



- Construction and installation of the purification and loading plants is finished
- Preparing for the first test batch of Te purification and synthesis when activities resume in the lab





~8 tons of telluric acid has been "cooling" underground for several years.

Ton-scale underground purification of telluric acid for further background reduction.



SNO+ Projected Sensitivity

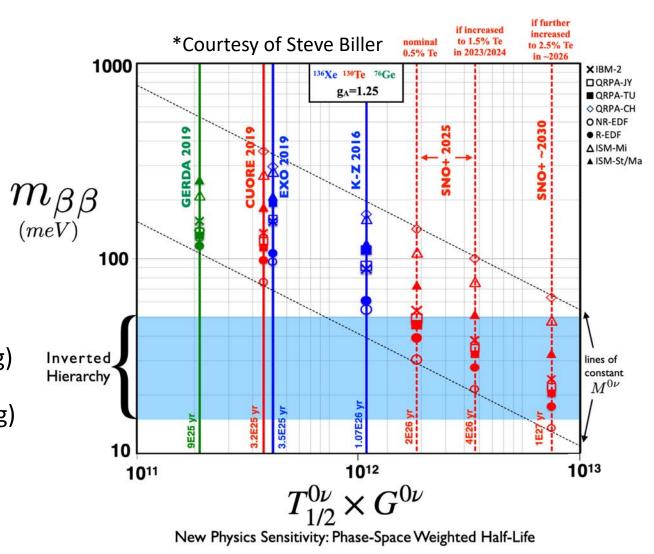
$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu} |M|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$

Likelihood analysis after 3 years at 0.5% Te loading:

 $T_{1/2}^{0\nu} > 2 \times 10^{26}$ years (90% C.L.)

Same analysis and same SNO+ detector for increased Te loading scenarios:

$$T_{1/2}^{0\nu} > 4 \times 10^{26}$$
 years (90% C.L.) (1.5% loading)
 $T_{1/2}^{0\nu} > 1 \times 10^{27}$ years (90% C.L.) (2.5% loading)



Future Outlook



- KLZ-800 is aiming for $T_{1/2}^{0\nu} > 5 \times 10^{26}$ years after 5 years of data taking.
- New analysis tools, including deep learning, and new MoGURA2 electronics upgrade will further improve background rejection.
- A future detector upgrade to KamLAND2-Zen is aiming for $T_{1/2}^{0\nu}>2 imes10^{27}$ years.
- Stay tuned for new exciting results from KLZ-800 to be released soon!



- LS filling will finish soon!
- We're preparing for tellurium loading (initially 0.5% ^{nat}Te by weight)
- Tellurium loading technique is highly scalable in LS and cost is relatively very low (< \$2M per ton of $0\nu\beta\beta$ decay isotope)
- Increased loading to 2.5% could allow SNO+ to reach a half-life of $T_{1/2}^{0\nu} > 1 \times 10^{27}$ years after 4 years of data taking

KamLAND-Zen Collaboration



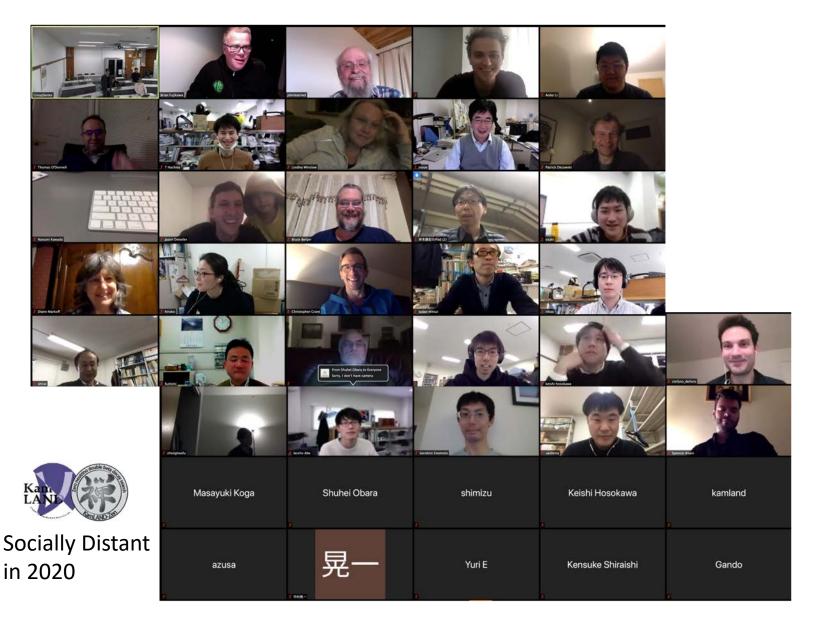
RCNS, Tohoku Univ. Kavli-IPMU Univ. of Tokyo Osaka Univ. Tokushima Univ. Kyoto Univ.



Lawrence Berkeley National Lab Univ. of Tennessee Triangle Univ. Nuclear Lab Univ. of Washington Massachusetts Institute of Technology Virginia Polytechnic Institute and State Univ. Univ. of Hawaii Boston Univ.



Nikhef, Univ. of Amsterdam



SNO+ Collaboration



Univ. of Alberta UC Berkeley / Lawrence Berkeley National Lab Boston Univ. Brookhaven National Lab Univ. of Chicago UC Davis Technical Univ. of Dresden Lancaster Univ. Laurentian Univ. LIP Lisbon and Coimbra Univ. of Liverpool Univ. National Autonoma de Mexico King's College London Norwich Univ. Univ. of Oxford Univ. of Pennsylvania Queen's Univ. Queen Mary Univ. of London SNOLAB Univ. of Sussex TRIUMF





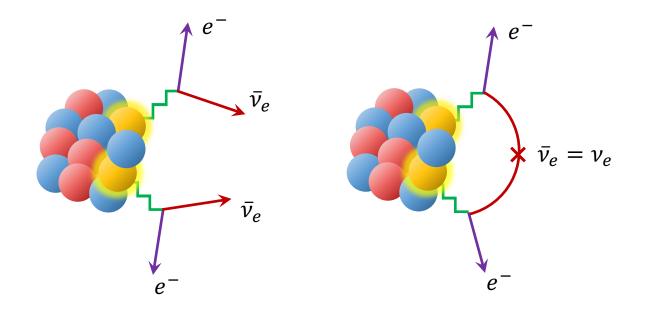












Thank you!



