

Searching for dark matter signatures in 20 years of GPS atomic clock data

Archival data for fundamental and exotic physics

Benjamin M. Roberts

(University of Queensland, Australia)

G. Blewitt, C. Dailey, M. Pospelov, G. Panelli, and A. Derevianko

(UNR, Reno; Perimeter Institute)

P. Delva, A. Hees, E. Savalle, P. Wolf

(SYRTE, Observatoire de Paris)

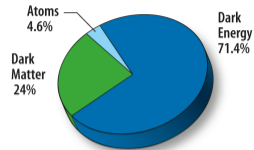
& the Optical clock, link, comb, & cavity teams of NPL, PTB, & SYRTE

CHEP2019, Adelaide

4–8 November 2019

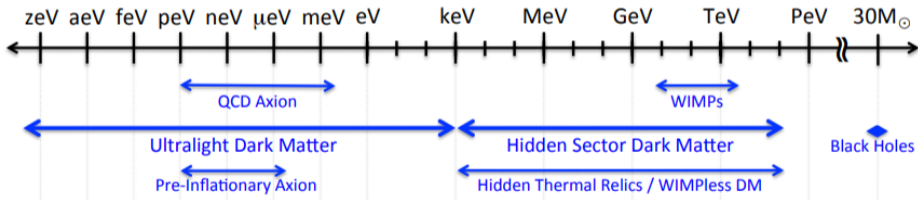
Dark Matter: What is it?

- $\sim 25\%$ of Universe energy budget (cf $\sim 5\%$ for “normal” matter)
- Possible mass range: ~ 90 orders-of-magnitude:



TODAY

[image: wmap.gsfc.nasa.gov]



[• US Cosmic Visions report, arXiv:1707.04591]

⇒ Wide range of possibilities: requires large range of experiments

- And the best kind of experiment is one that's already done

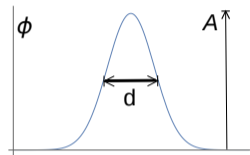
Dark Matter Clumps: (Topological Defects)

- Ultralight ($m_\phi \ll eV$) \implies high occupation number

Many possibilities: Here: TDs

Topological Defects

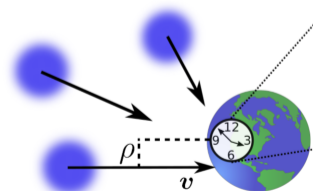
- monopoles, strings, walls,
- Defect width: $d \sim 1/m_\phi$
- Earth-scale object: $m_\phi \sim 10^{-14} eV$



Inside: $\phi^2 \rightarrow A^2$, Outside: $\phi^2 \rightarrow 0$

Dark matter: Gas of defects

- DM: galactic speeds: $v_g \sim 10^{-3} c$
- Collisions offer chance for lab detection



- Vilekin '85, Coleman '85, Lee '89, Kibble '80, ...
- Derevianko, Pospelov, Nature Phys. 10, **933** (2014).

Variation of fundamental constants

- Here: (quadratic) scalar: $\mathcal{L}_{\text{int}} \sim \phi^2(a\bar{\psi}\psi + bF_{\mu\nu}^2 + \dots)$
- Parameterised in with Λ “energy scale” (\sim inverse coupling strength)

\implies transient additions to *effective values* of fundamental constants

$$\alpha^{\text{eff}}(r, t) = \alpha \left(1 + \frac{\phi^2(r, t)}{\Lambda_\alpha^2} \right), \quad m_f^{\text{eff}}(r, t) = m_f \left(1 + \frac{\phi^2(r, t)}{\Lambda_f^2} \right),$$

\implies shifts in energy levels \implies shifts in clock frequencies

$$\frac{\delta\omega(r, t)}{\omega_0} = K_\alpha \frac{\delta\alpha(r, t)}{\alpha} = \phi^2(r, t) \frac{K_\alpha}{\Lambda_\alpha^2}$$

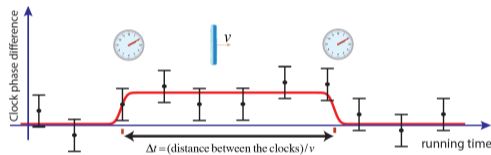
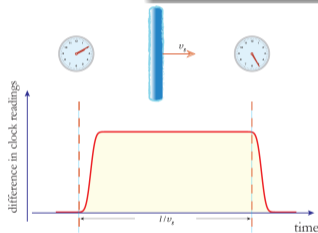
Monitor Atomic Clocks

- Clocks: lock frequency to atomic transition \implies Monitor atomic frequencies using atomic clocks
- Shift $\delta\omega$ occurs only when ϕ non-zero (inside DM object)

Shift in atomic clock frequencies

Monitor Atomic Clocks

- Temporary frequency shift \rightarrow bias (phase) build-up
- Initially synchronised clocks become desynchronised



Signal v. noise?

- Transient signal: looks essentially like any outlier
- So how to distinguish from noise? i.e. what is the specific DM signature?

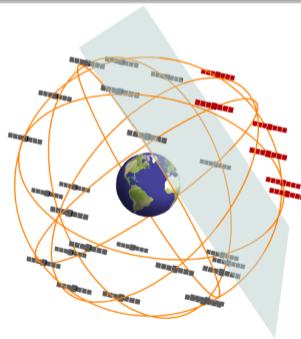
- Derevianko, Pospelov, Nat. Phys. 10, 933 (2014).

Global network of precision devices

- DM expected to move at \sim galactic speeds
- Correlated, directional signal: propagation through network, $v \sim 300$ km/s
- \vec{v} encoded in time-delay, ordering: $\Delta t \sim$ seconds – minutes
- Also: multiple clock-types in network, each has different K_α (prediction of theory)

GPS: 50,000 km DM observatory

- 32 satellite clocks (Rb/Cs),
 $\sim 16+$ years of high-quality data
- Also several H-maser ground-based clocks.
- Data from JPL:
(sideshow.jpl.nasa.gov/pub/jpligsac/)
 - 30s sampled data; 0.01–0.1 ns precision

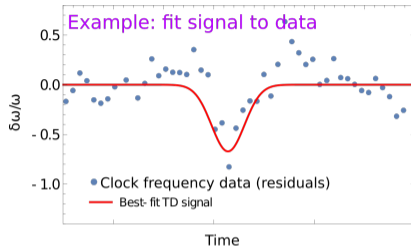


- Clocks: Derevianko, Pospelov, Nature Phys. 10, **933** (2014).
- Magnetometer: Pospelov, Pustelny, Ledbetter, Kimball, Gawlik, Budker, Phys.Rev.Lett. **110**, 021803 (13).

Search method

Fibre network

- Max-likelihood fit method¹
- Coherent $\delta\omega$ variations (with K_α)
- Signal template $s = s(t_0, d, \mathbf{v})$

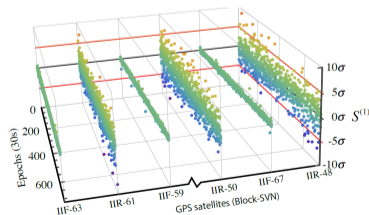


$$p(D_t|\delta\alpha, \theta) = C p(\theta) \exp\left(\frac{-1}{2}[d - \delta\alpha s]^T H[d - \delta\alpha s]\right),$$

d : data stream; $s = s(\theta)$ signal template, $\delta\alpha$: magnitude; $H = E^{-1}$ covariance

- Signal-to-noise $R >$ threshold? \implies event detection
- Largest $\delta\alpha$ that cannot be ruled out? \implies set limits

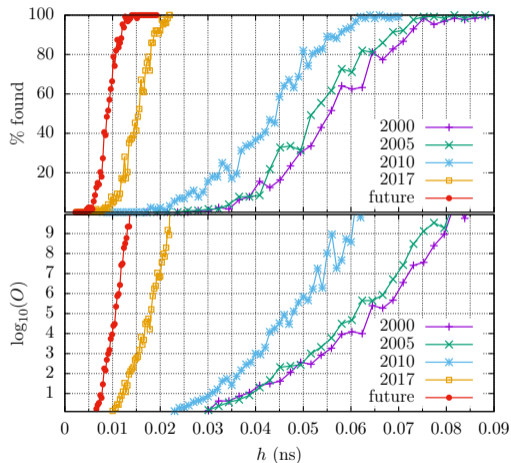
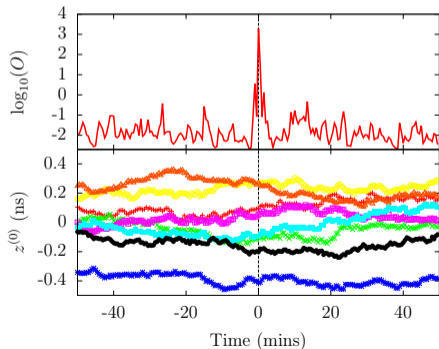
$$\delta\alpha^{\text{bf}} = dHs/sHs. \quad \implies \quad \delta\alpha < dHs/sHs + n(sHs)^{-1/2}.$$



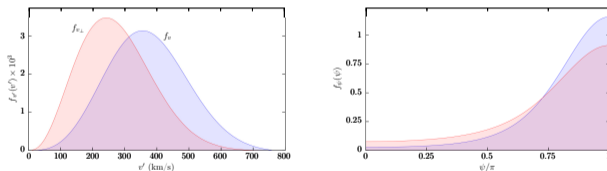
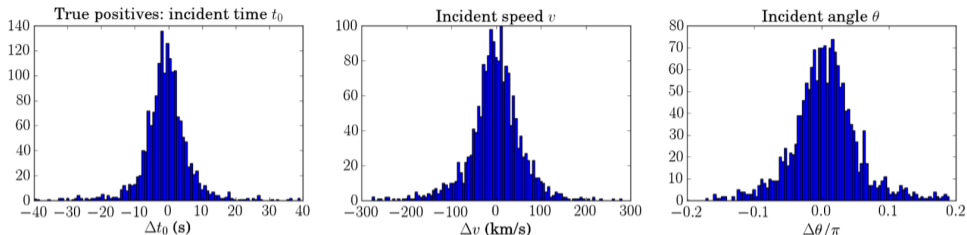
Testing: simulated GPS

Odds ratios: Marginalise all parameters

- Simulated GPS data (known power spectrum)
- Threshold: set by false positives
- True positives: inject fake events



Resolve speed + direction (simulation)

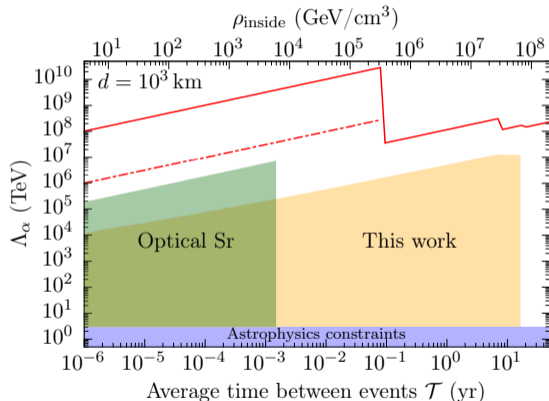


Resolution: simulation using GPS

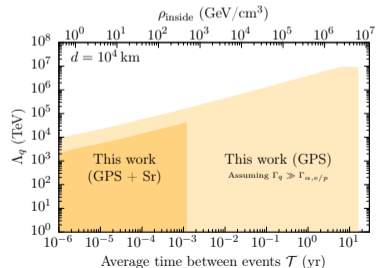
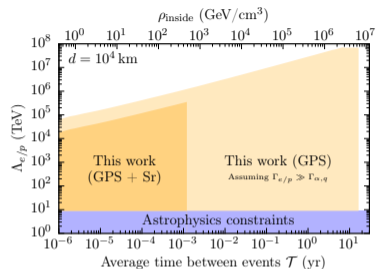
- Resolve v – DM vel. distro is “known” – reject false positives!
- Many clocks + High sampling frequency and/or **Large distances**

Results: Limits

- Initial search: simplified method (extra large events)
- Exclusion plot (assume coupling to photon dominates)



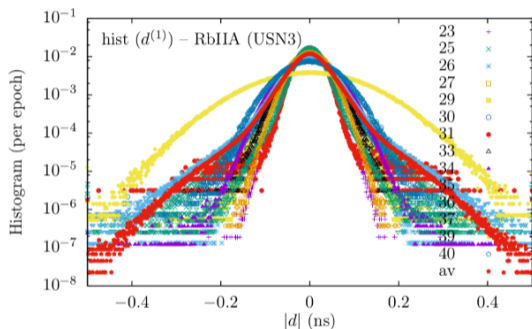
Sr: Wcislo, Morzynski, Bober, Cygan, Lisak, Ciurylo, Zawada, Nat. Astron. 1, 9 (2016).
 Astro: Olive, Pospelov, Phys. Rev. D. 77, 43524 (2008).



- BMR, Blewitt, Dailey, Murphy, Pospelov, Rollings, Sherman, Williams, Dereviako, Nature Comm. 8, 1195 (2017).

Aside: challenges of re-purposed data

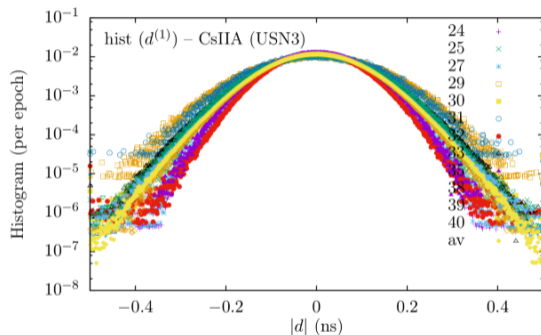
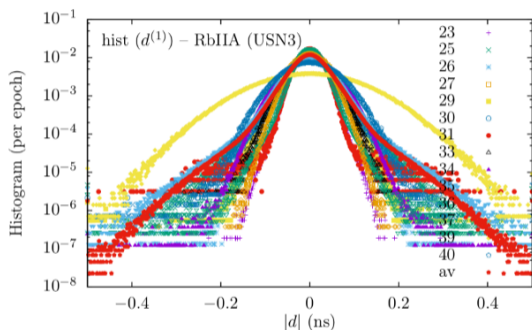
data from JPL: Histogram



- Possible that some clocks mis-identified (Here, one of the “Rb” clocks is probably Cs).
- Same discrepancy in autocorrelation function, Allan variance etc.

Aside: challenges of re-purposed data

data from JPL: Histogram



- Possible that some clocks mis-identified (Here, one of the “Rb” clocks is probably Cs).
- Same discrepancy in autocorrelation function, Allan variance etc.

European fibre-linked optical clock network

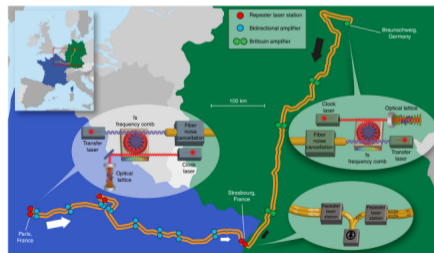


- Sensitivity: ✓ ($\delta\alpha, \Lambda$)
limited only by clocks:
Sr-Sr: $\delta\omega/\omega \sim 3 \times 10^{-17}$ at 1000s
- Long observation time: ✓ (T)
- Long-term stability: ✓ (d)

- Lisdat *et al.* (PTB, LNE-SYRTE), Nature Commun. **7**, 12443 (2016).
- Delva *et al.* (PTB, SYRTE, NPL, ..), Phys. Rev. Lett. **118**, 221102 (2017).

Fibre network

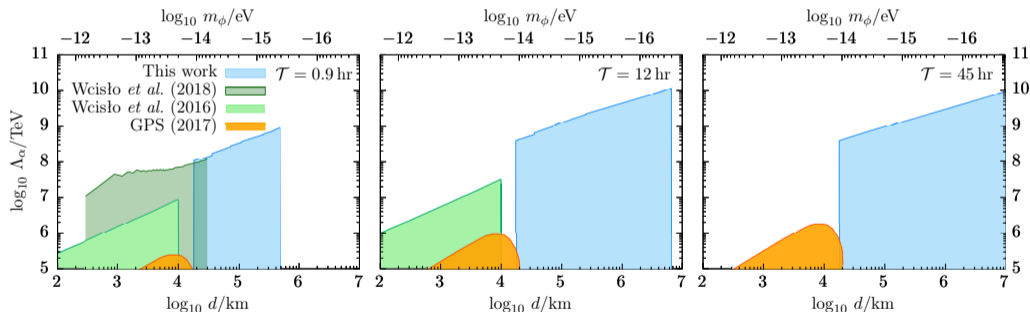
- High-accuracy long-distance clock-clock (atom-atom) comparisons
- Different clocks: Hg/Sr/Yb⁺
- ~ Days – weeks synchronous running



Topological defect dark matter

Very sensitive to different models, different range of parameters

$$\phi_0^2 = \hbar c \rho_{\text{DM}} v_g \mathcal{T} d, \quad \mathcal{T} = \frac{\rho_{\text{inside}}}{\rho_{\text{DM}}} \frac{d}{v_g} \implies \Lambda_\alpha^2(\mathcal{T}, d) > \frac{\hbar c \alpha \rho_{\text{DM}} v_g \mathcal{T} d}{|\delta\alpha_0(\mathcal{T}, \tau_{\text{int}})|}.$$



- nb: GPS results (orange): go up to $\mathcal{T} \sim 10 \text{ yrs} \sim 10^5 \text{ hrs}$

GPS: BMR, Blewitt, Dailey, Murphy, Pospelov, Rollings, Sherman, Williams, Derevianko, **Nature Comm.** **8**, 1195 (2017).

2016: Wcislo, Morzynski, Bober, Cygan, Lisak, Ciurylo, Zawada, **Nat. Astro.** **1**, 0009 (2016).

2018: Wcislo *et al.*, **Sci. Adv.** **4**, 4869 (2018).

Search for transient variations of the fine structure constant and dark matter using fiber-linked optical atomic clocks

B. M. Roberts,^{1,*} P. Delva,¹ A. Al-Masoudi,² A. Amy-Klein,³ C. Bærentsen,¹ C. F. A. Baynham,⁴
E. Benkler,² S. Bilicki,¹ S. Bize,¹ W. Bowden,⁴ J. Calvert,¹ V. Cambier,¹ E. Cantin,^{1,3} E. A. Curtis,⁴
S. Dörscher,² M. Favier,¹ F. Frank,¹ P. Gill,⁴ R. M. Godun,⁴ G. Grosche,² C. Guo,¹ A. Hees,¹ I. R. Hill,⁴
R. Hobson,⁴ N. Huntemann,² J. Kronjäger,⁴ S. Koke,² A. Kuhl,² R. Lange,² T. Legero,² B. Lipphardt,²
C. Lisdat,² J. Lodewyck,¹ O. Lopez,³ H. S. Margolis,⁴ H. Álvarez-Martínez,^{1,5} F. Meynadier,^{1,6} F. Ozimek,⁴
E. Peik,² P.-E. Pottie,¹ N. Quintin,⁷ C. Sanner,^{2,†} L. De Sarlo,¹ M. Schioppo,⁴ R. Schwarz,² A. Silva,⁴
U. Sterr,² Chr. Tamm,² R. Le Targat,¹ P. Tuckey,¹ G. Vallet,¹ T. Waterholter,² D. Xu,¹ and P. Wolf^{1,‡}



Conclusion

Fundamental physics with archival data

- Beyond standard model physics + dark matter
- Wide variety of possible models \implies wide range of experiments
- Archived time-stamped data from precision measurement devices

GPS: 50,000 km DM observatory

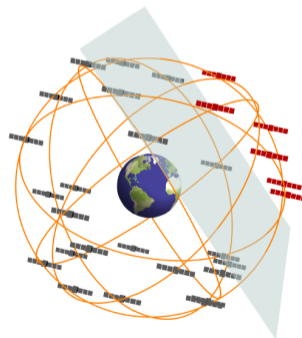
- Initial limits: orders of magnitude improvement
- Nature Comms. **8**, 1195 (2017)

Optical clock networks

- Also: European fibre-linked clocks [arXiv:1907.02661]

Other networks

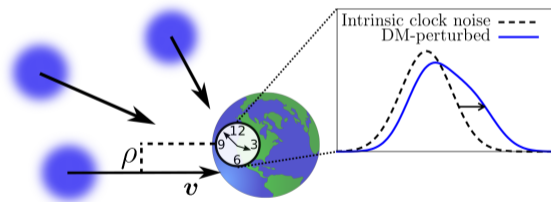
- Magnetometers, gravimeters, EDM searches
- complementary: sensitive to different models



Extra: Asymmetry

Small objects: no correlated signal

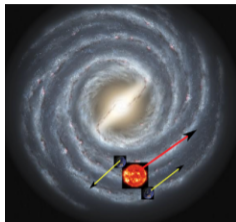
- small size $\sim \rightarrow$ large rate
- Shift in mean: unobservable (DM always present)
- Induce non-Gaussian features (such as an asymmetry)



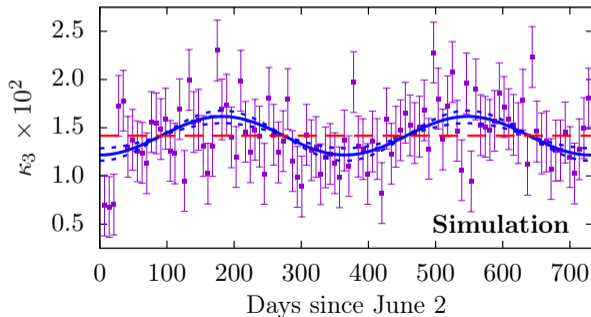
$$\mathcal{R} = \frac{1}{\mathcal{T}} = \frac{3\rho_{\text{DM}}v_g}{4\rho_{\text{inside}}R},$$

$$\kappa_3 \approx \frac{2\mathcal{R}\tau_0\chi_0^3}{5\sigma^3}.$$

Extra: Annual modulation

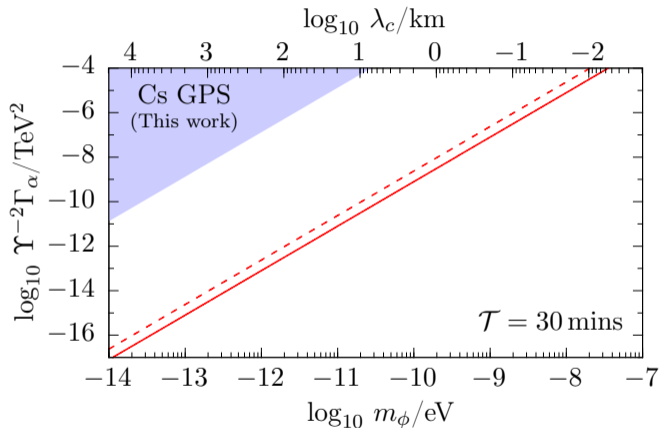


- Yearly change in event rate:
- Sun + Earth velocities add
- $R(t) = R_0 + R_m \cos(\omega t + \phi_{\text{June}2})$
- $\Delta\kappa_3/\kappa_3 = 10\%$
- $\Delta\kappa_4/\kappa_4 = 15\%$



Extra: Limits Q-balls: α (photon field)

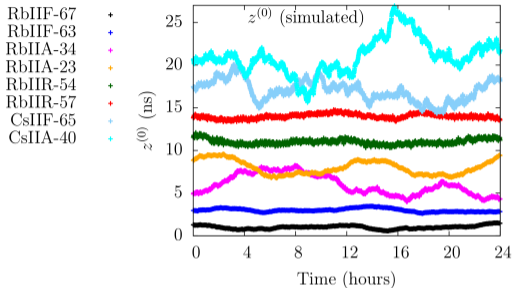
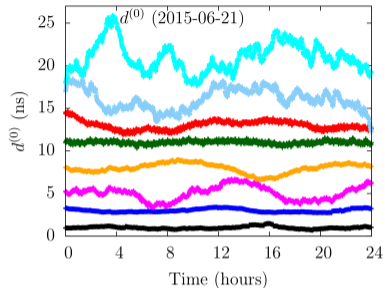
Proof of concept: Limits from GPS Cs/H $K_\alpha(\text{Cs}/H) \simeq 0.8$



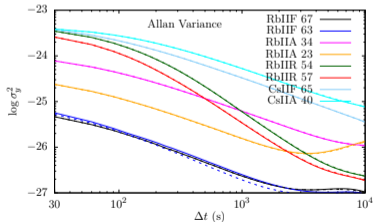
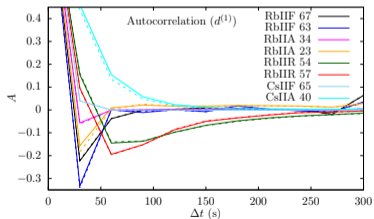
● BMR, Derevianko, arXiv:1803.00617

Red line: sensitivity estimate for 1 year of optical Sr

Test the method: simulated GPS



RbIIF-67 +
RbIIF-63 +
RbIIA-34 +
RbIIA-23 +
RbIIR-54 +
RbIIR-57 +
CsIIF-65 +
CsIIA-40 +



• Inject fake events: True positive rate

• Don't inject events: False positive rate