

ePESSTO+: survey and scientific output

(the advanced Public ESO Spectroscopic
Survey of Transient Objects)

Janet Ting-Wan Chen on behalf of ePESSTO+ collaboration
Alexander von Humboldt Fellow
Max Planck Institute for Extraterrestrial Physics



MAX-PLANCK-INSTITUT
FÜR EXTRATERRESTRISCHE PHYSIK

ePESSTO+

- The advanced (extended) **P**ublic **E**SO **S**pectroscopic **S**urvey of **T**ransient **O**bjects
- **137** scientists in ePESSTO+ collaboration
- From P103 to P105: 140 nights on ESO New Technology Telescope (**NTT**) at La Silla Observatory, Chile
- We use **EFOSC2** (ESO Faint Object Spectrograph and Camera) & **SOFI** (Son of ISAAC, a large field Infra-Red (1-2.5 micron) spectro-imager)



ePESSTO+ organisation

Science Board

- **Cosimo Inserra** (survey PI)
- **Joe Anderson** (SB Chair)
- **Ting-Wan Chen** (TAT Chair)
- **Maria Teresa Botticella**
- **Michel Dennefeld**
- **Avishay Gal-Yam** (ZTF point of contact)
- **Giorgos Leloudas**
- **Kate Maguire**
- **Rubina Kotak**
- **Stephen Smartt** (ATLAS and PS point of contact)
- **Stefan Taubenberger**
- **Lukasz Wyrzykowski** (OGLE point of contact)

27 science
projects

ePESSTO+ organisation

Operating team

- **Dave Young** - Marshall operations
- **Ofer Yaron** - data distribution via WISeREP
- **Shane Moran-Kelly & Eleonora Parrag**
- coordinators of photometric resources
- **Tomás Muller Bravo** - pipeline manager and phase 3 data manager
- **Erkki Kankare, Mariusz Gromadzki & Matt Nicholl**
- phase 3 and data release responsible(s)



@PESSTOsurvey

- **Charlotte Angus, Chris Frohmaier & Phil Wiseman**

Ombudspersons

- **Patricia Shady & Lluís Galbany**

Classification machine

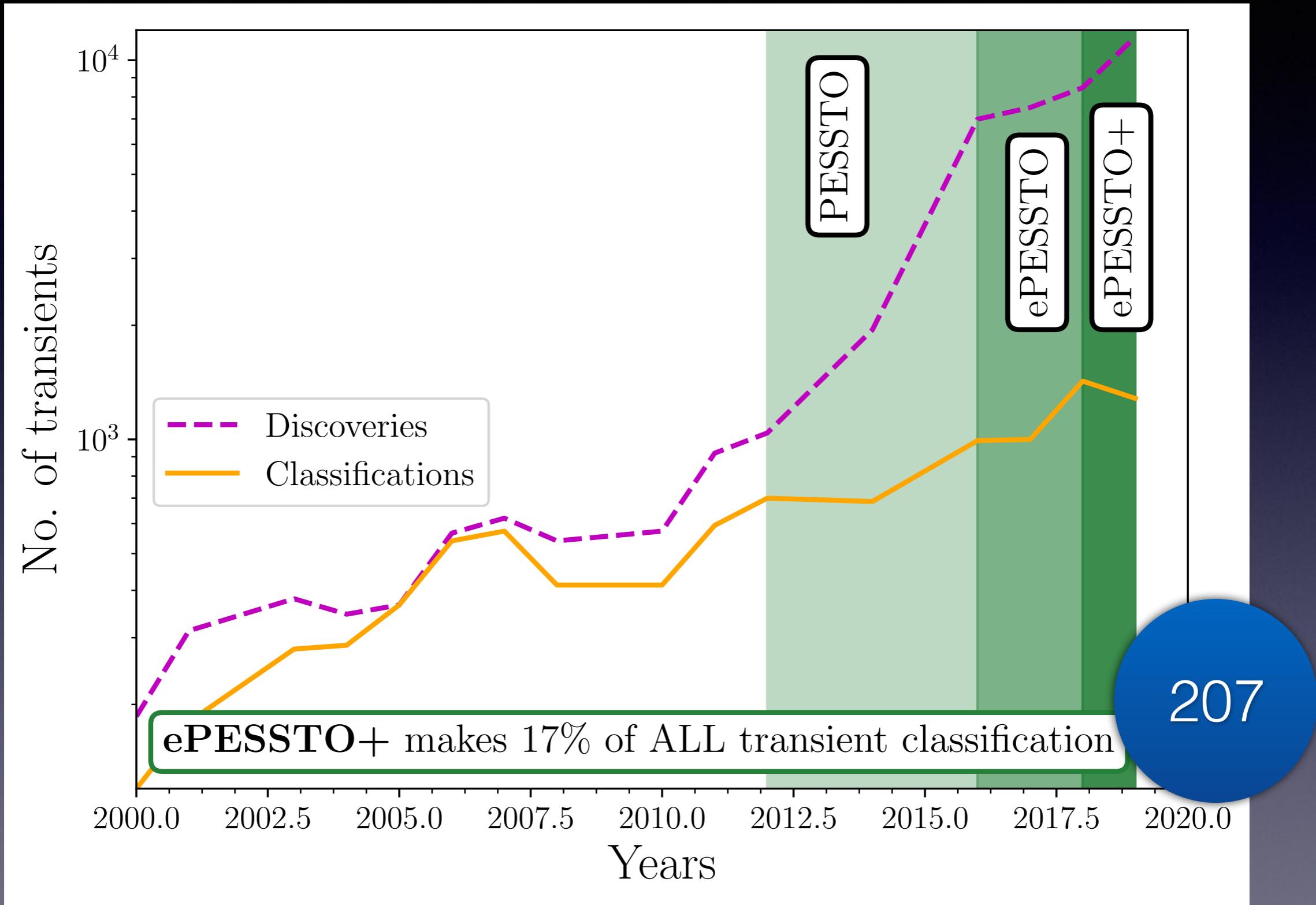


figure from Cosimo Inserra

Classification machine

Breakdown of the source of targets for ePESSTO+

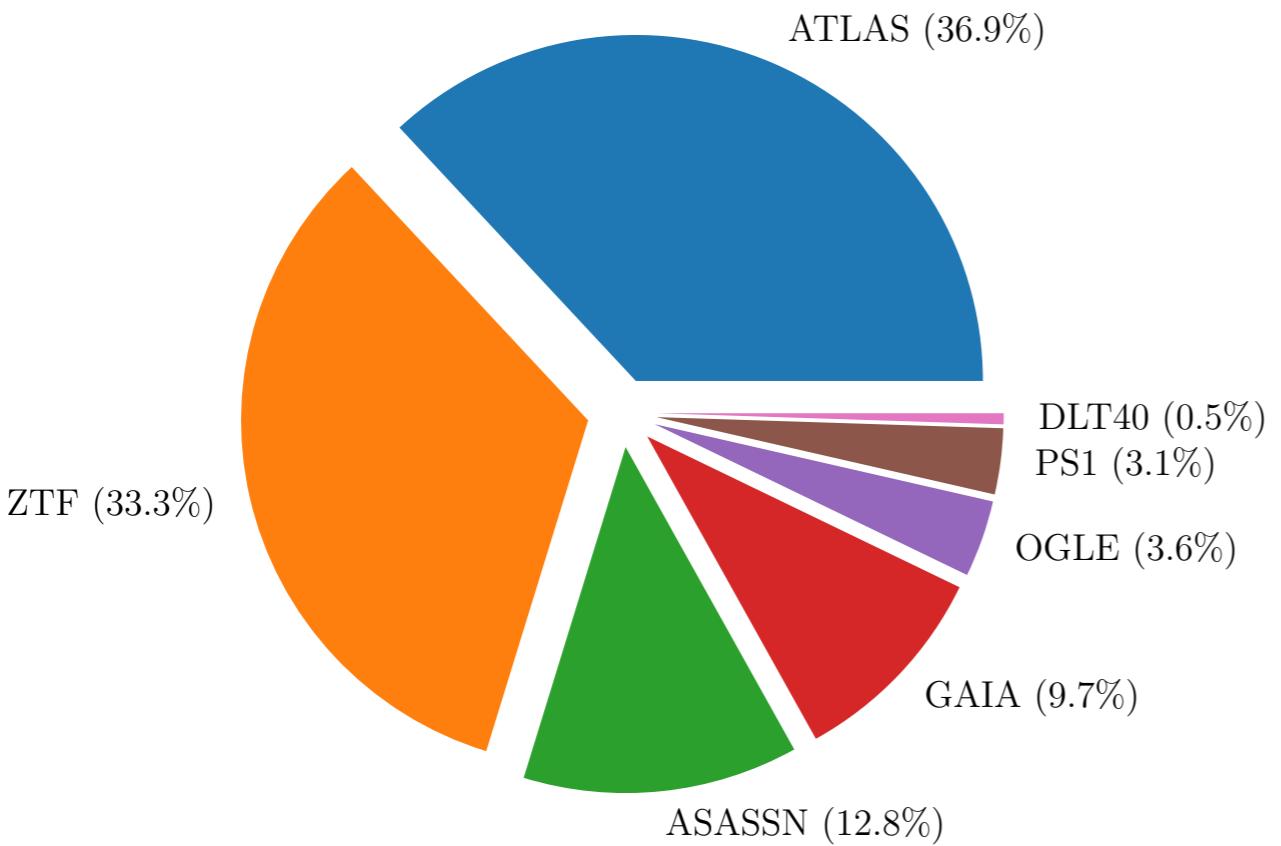
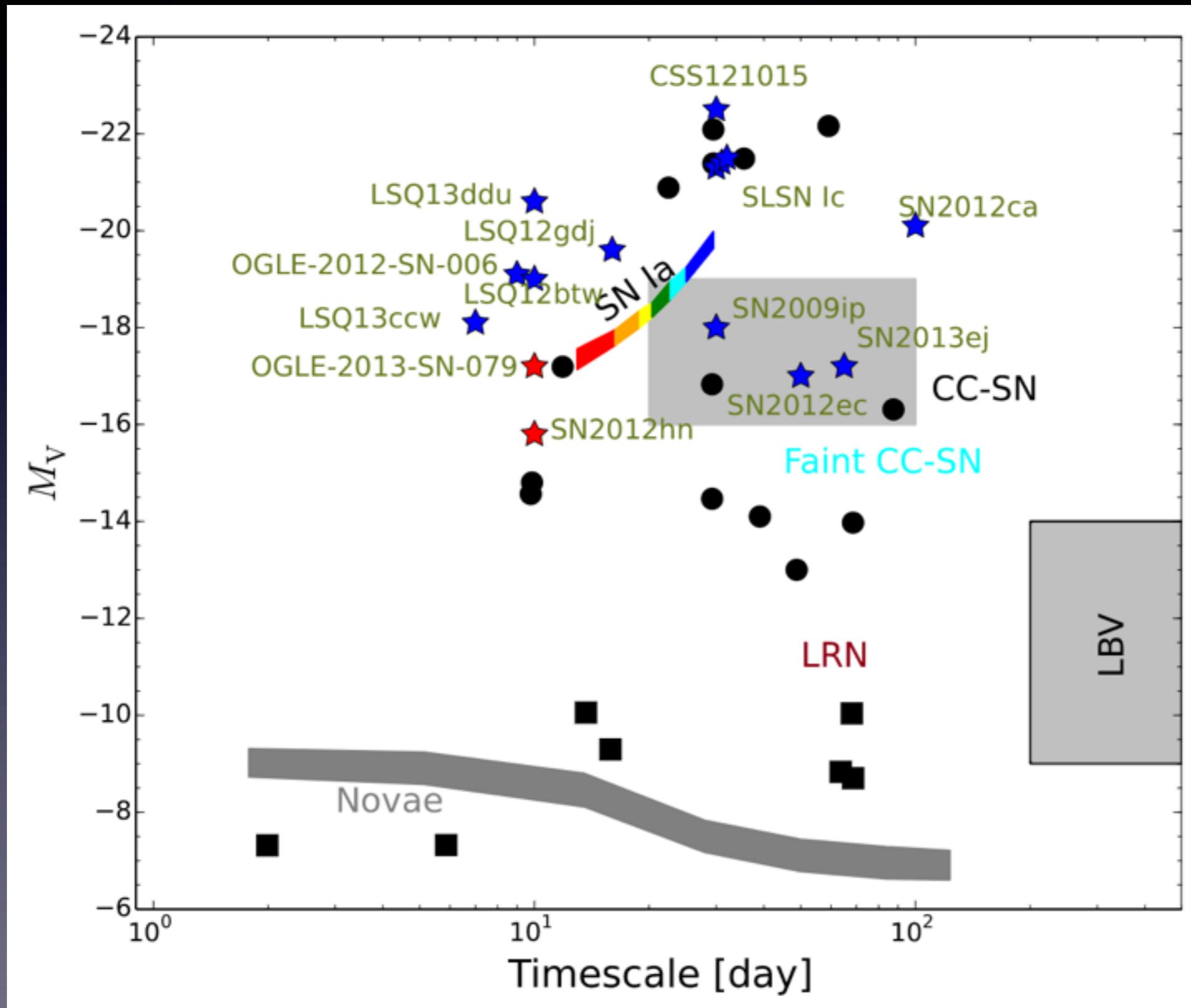


figure from Cosimo Inserra

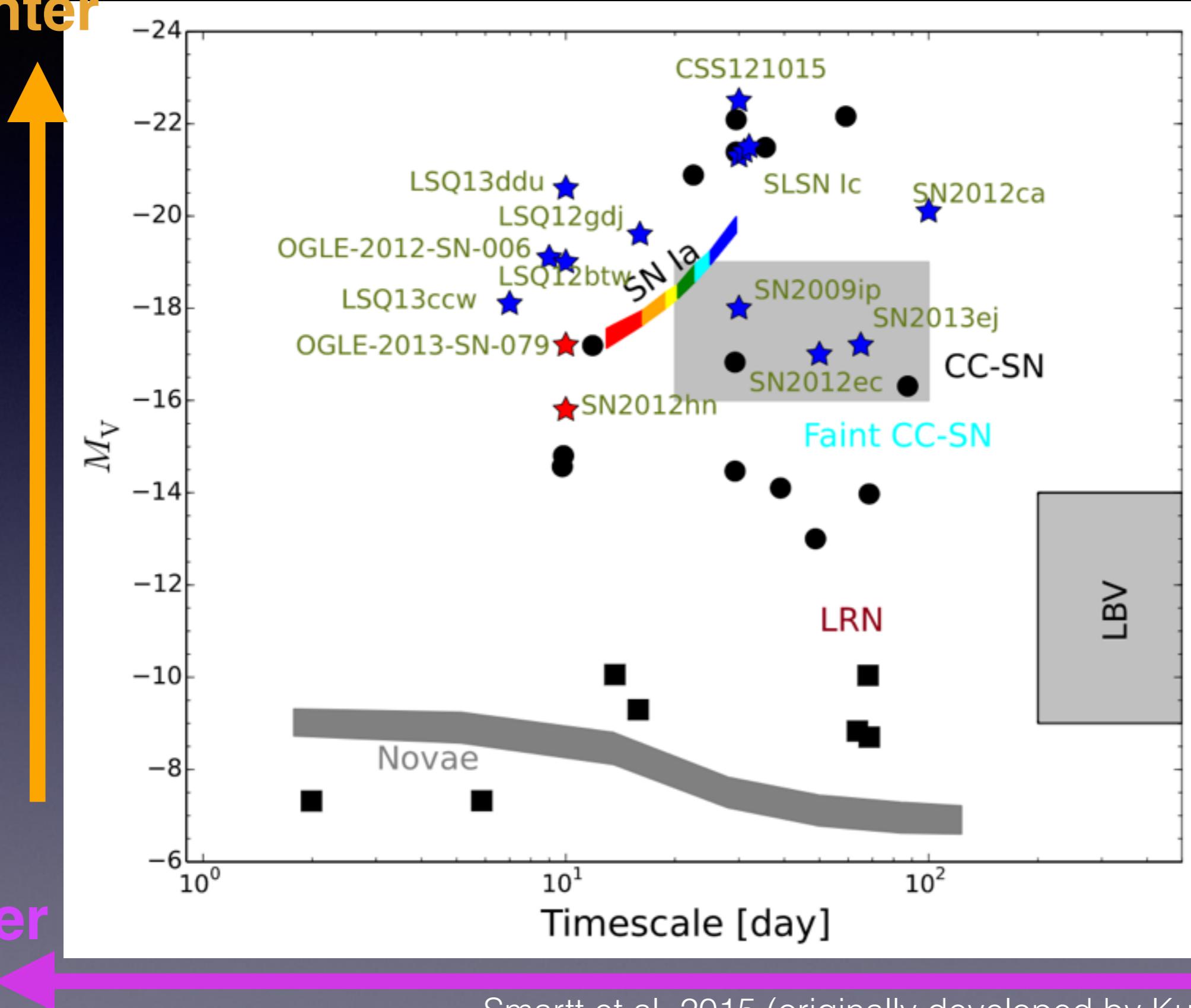
Type	Number	Per Cent
Ia	95	46
Ia-pec	8	4
II	35	17
Ib/c	6	3
Ic-BL	4	2
IIIn	7	3
IIb	3	2
Ibn	1	1
SLSNe	3	1
TDE	3	1
AGN	3	1
CV	11	5
Others	18	9
Not visible	10	5
Total	207	100

Transient space phase



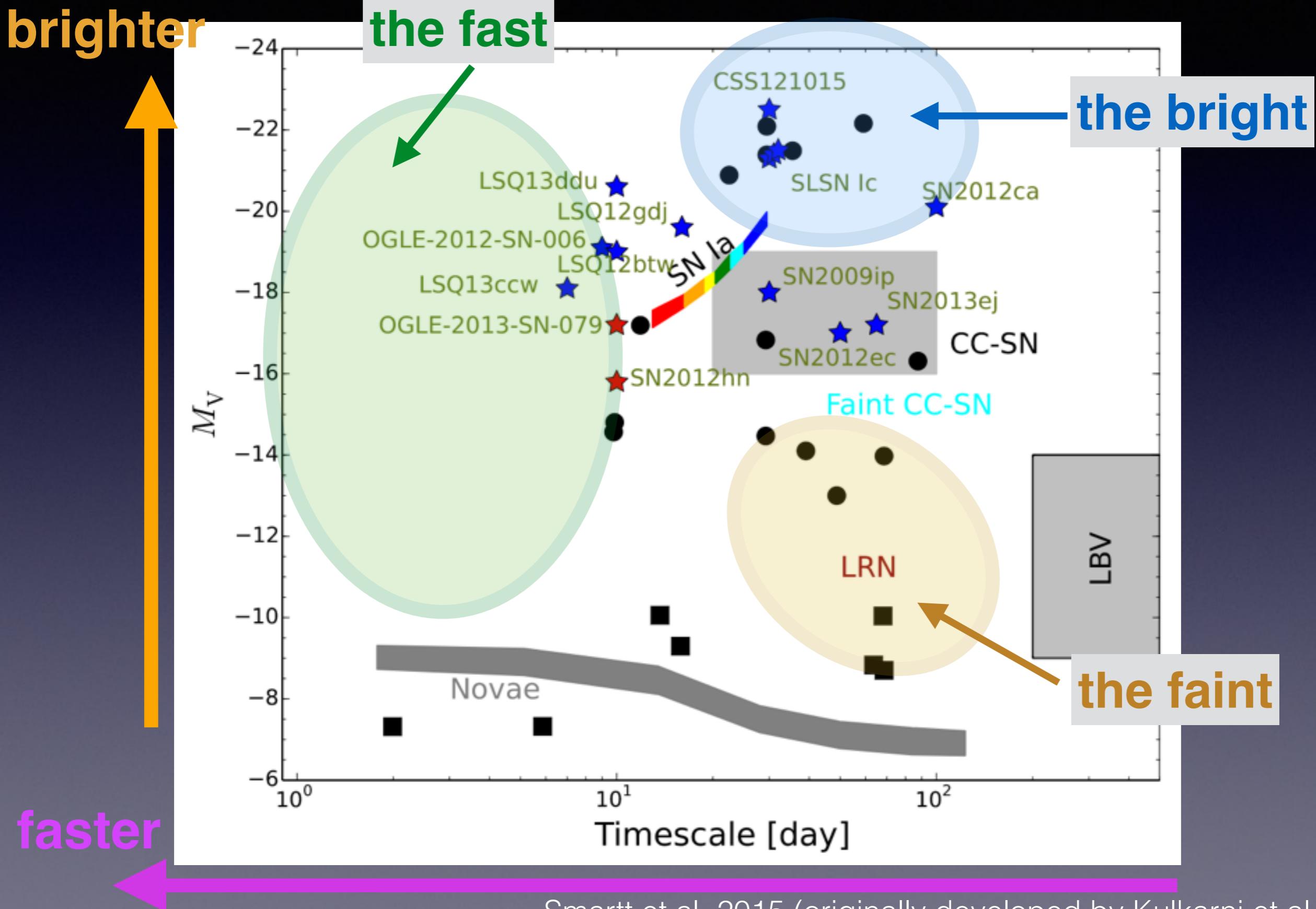
Transient space phase

brighter



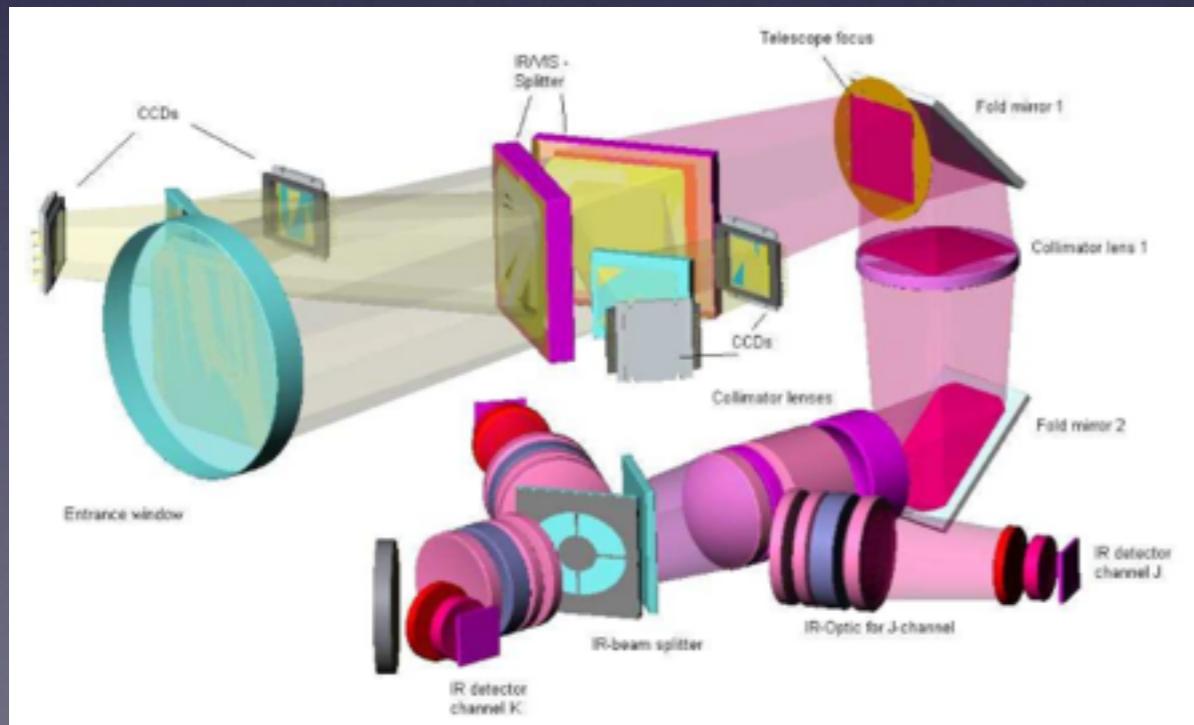
faster

Transient space phase



ePESSTO+ × GROND

- GROND: **G**amma-**R**ay burst
Optical/**N**ear-infrared **D**etector
- 7-channel imaging
(simultaneous in g'r'i'z'JHK_s)
mounted at 2.2m MPG
telescope at ESO La Silla
observatory



Greiner et al. 2008

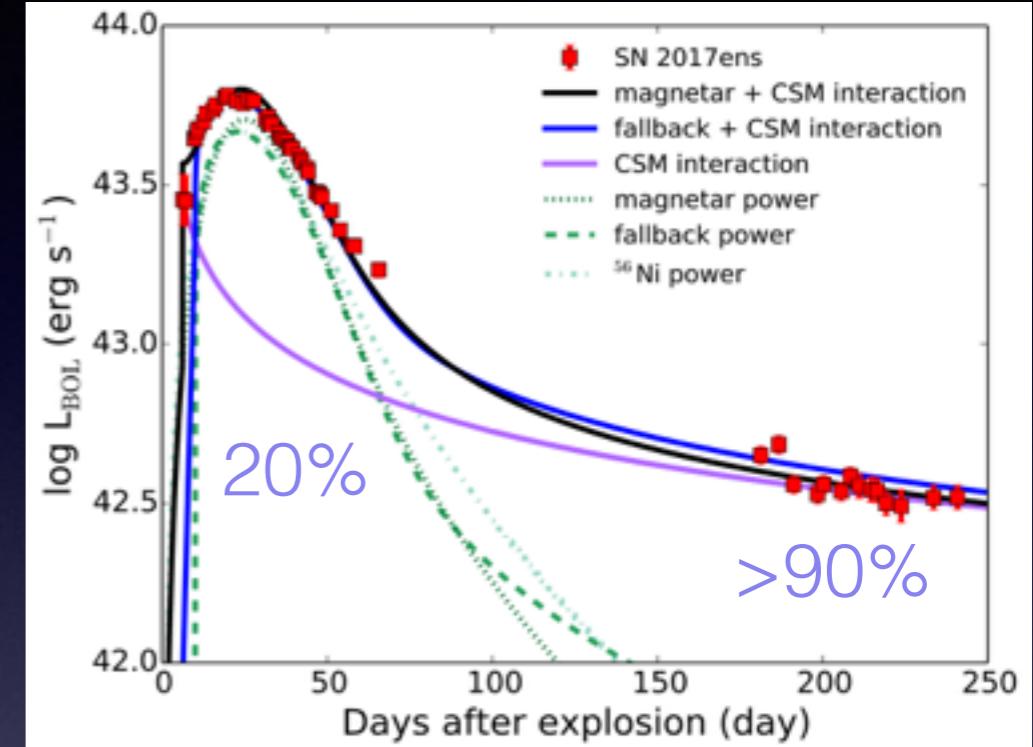
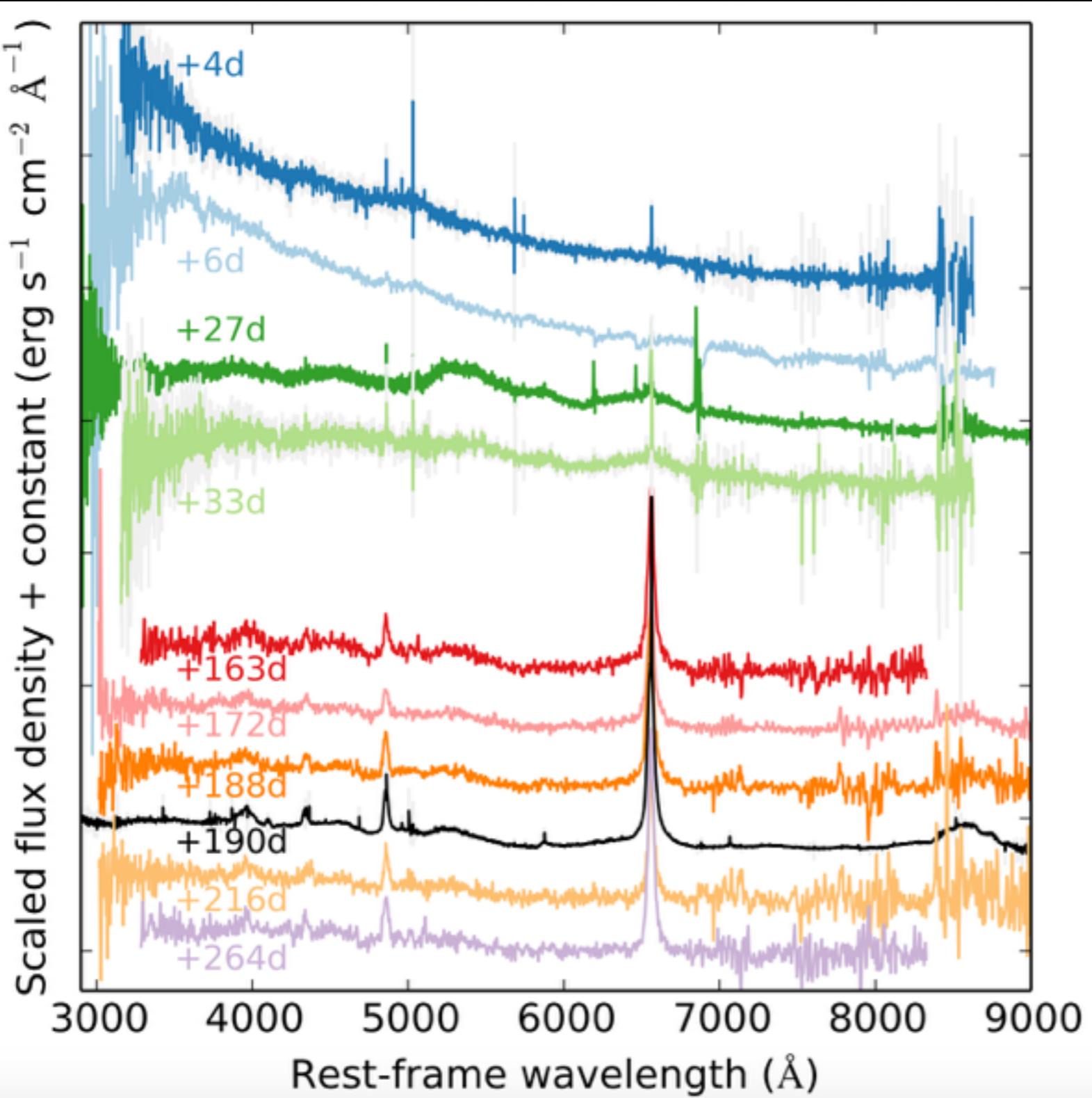
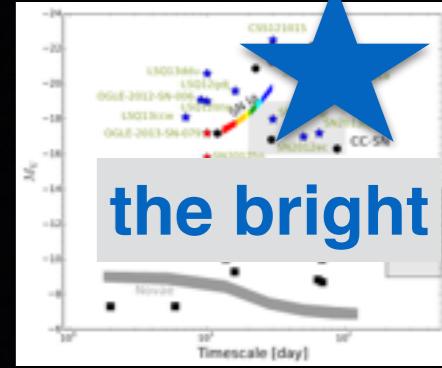
at the same site with NTT!

- GROND SN programme (PI: Chen)
15% telescope time sharing with
other projects (e.g. GRB, X-ray
binary...)
- GROND follows up 60 SNe
collaborating with ePESSTO(+)
since Nov. 2016
- GREAT (**GR**OND-**e**PESSTO-
ATLAS) survey identifies
2 SLSN IIn and 1 peculiar object

SN 2017ens

redshift = 0.1086
 M_g = -21.1 mag

(from luminous broad lined-Ic to IIn)

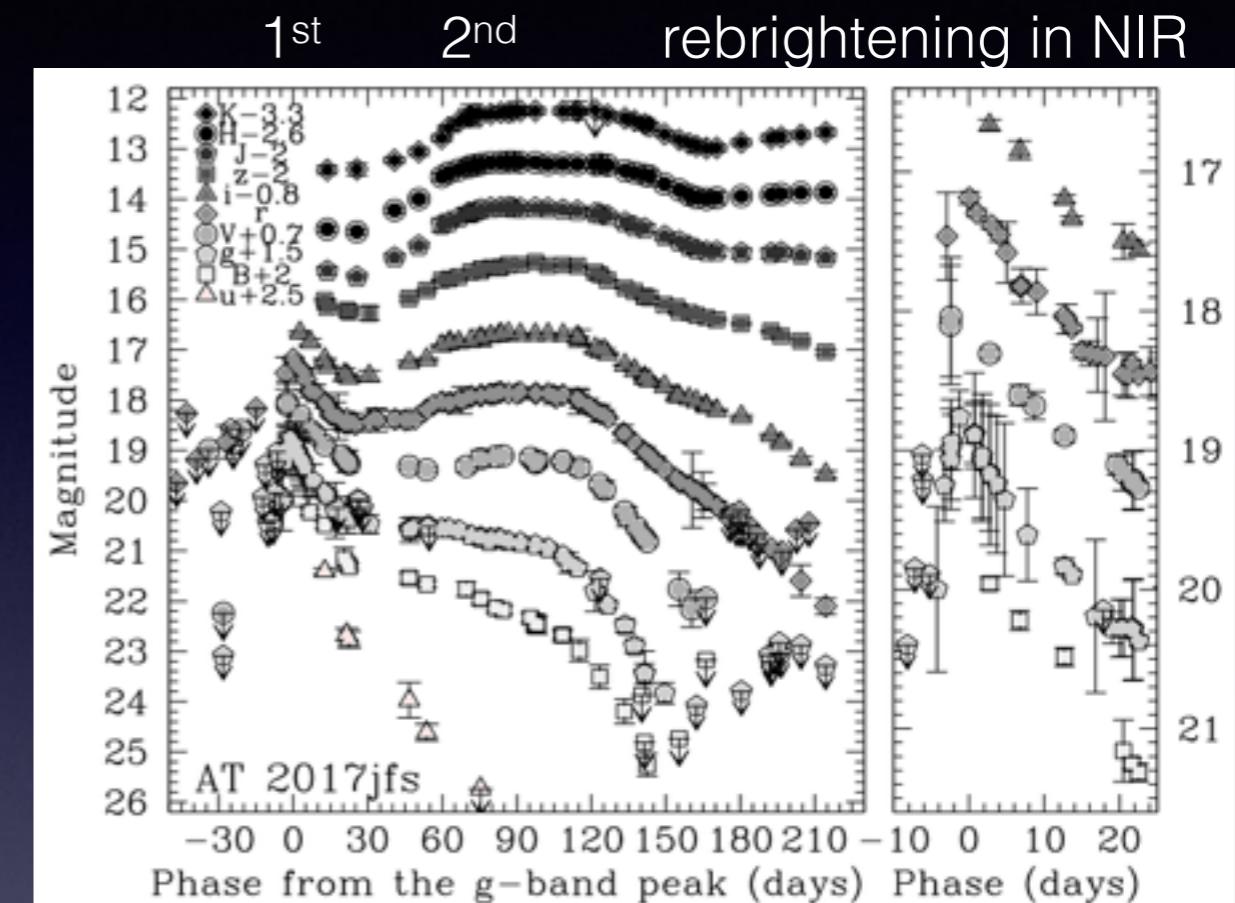
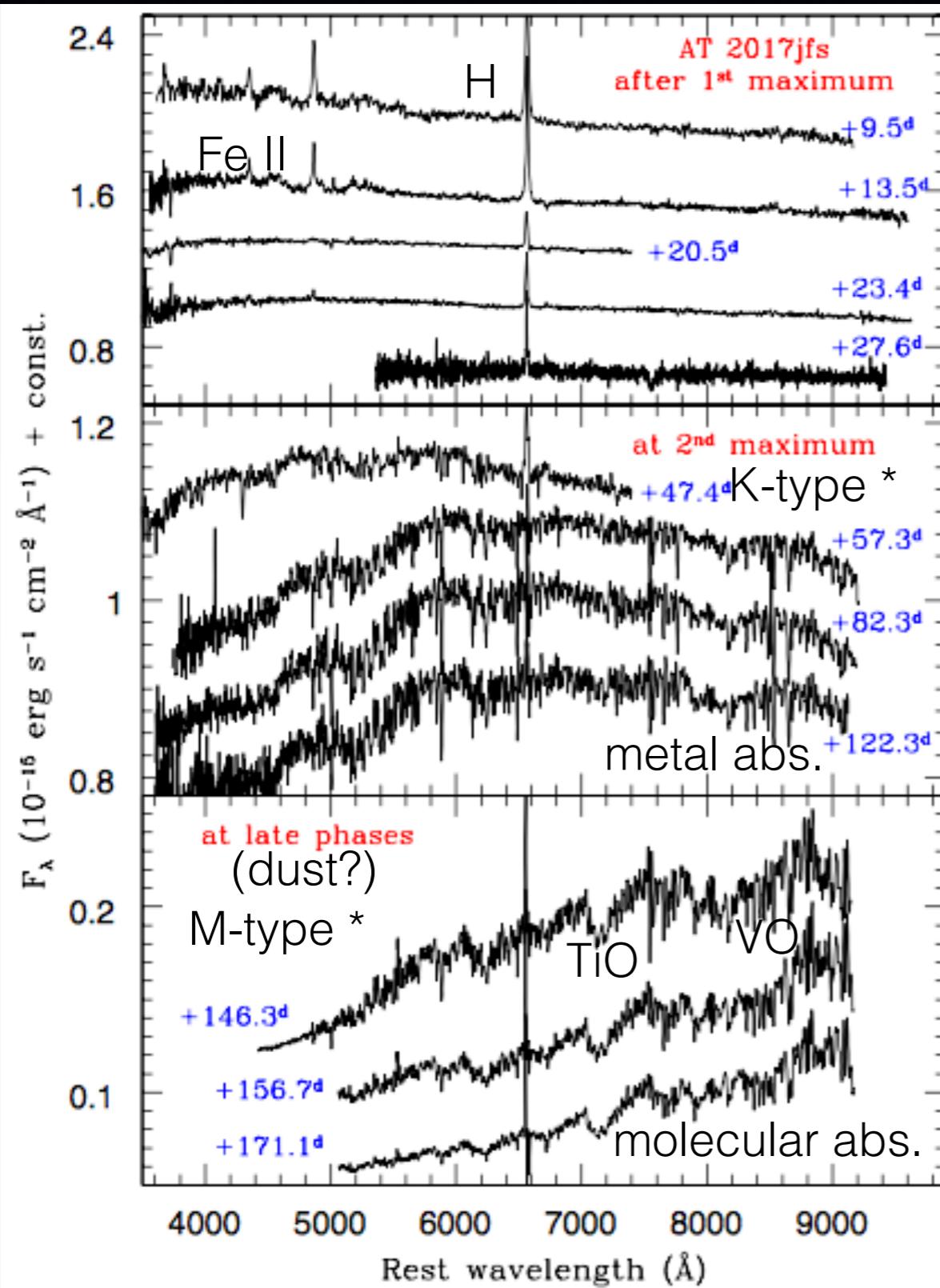
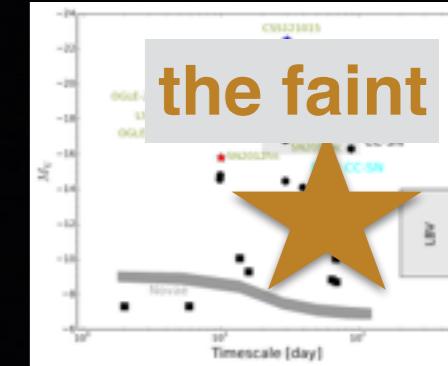


Chen et al. 2018, ApJ, 867L, 31

- Two components: central power and interaction.
- The mass loss rate of $5 \times 10^{-4} M_{\odot}/\text{yr}$ with the constant wind of 50 km/s

AT 2017jfs (luminous red nova)

distance = 35 Mpc
 $M_g = -15.5$ mag



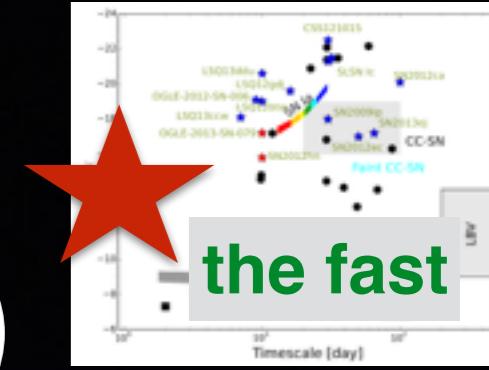
Pastorello, Chen et al. 2019,
A&A, 625L, 8

- common-envelope transient, possibly the outcome of a merging event in a massive binary

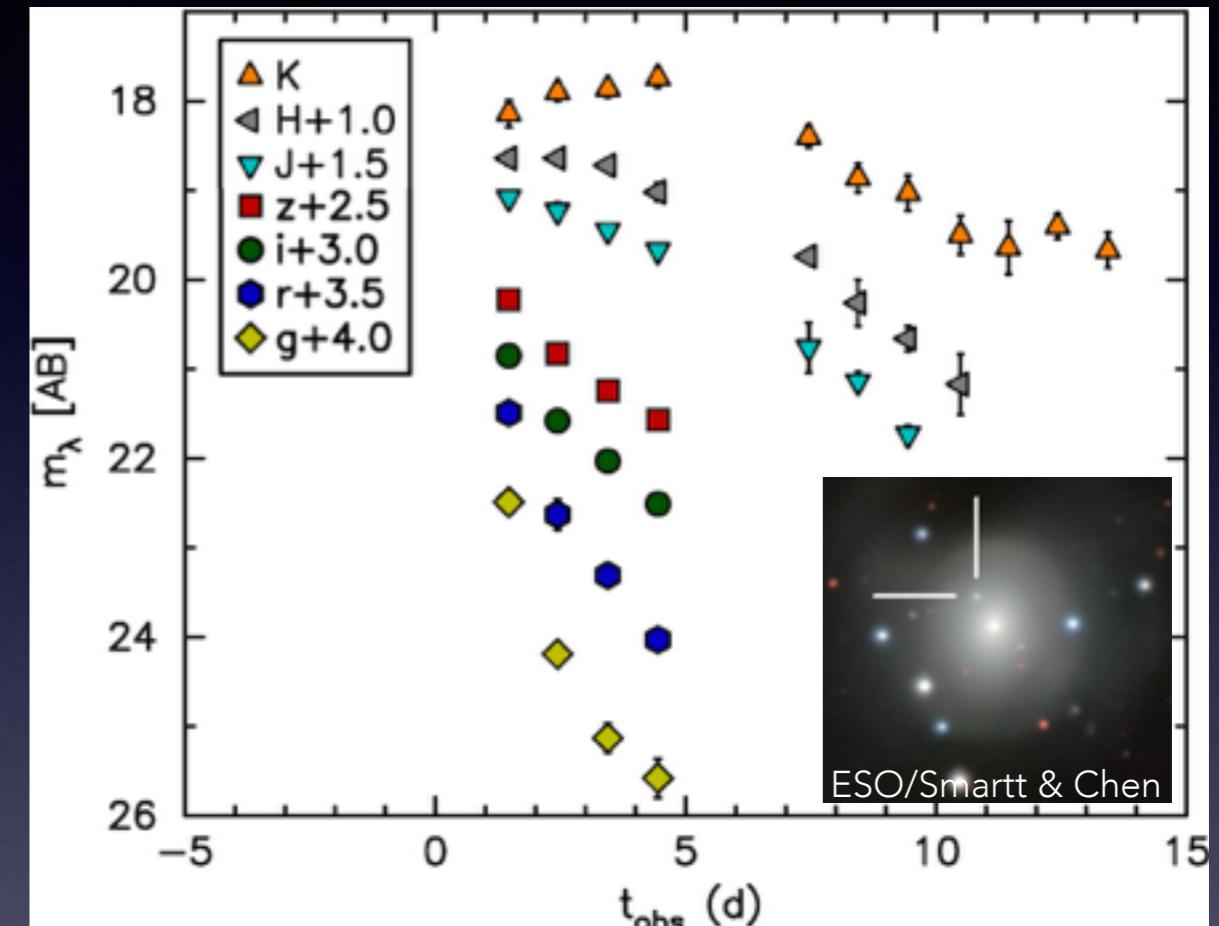
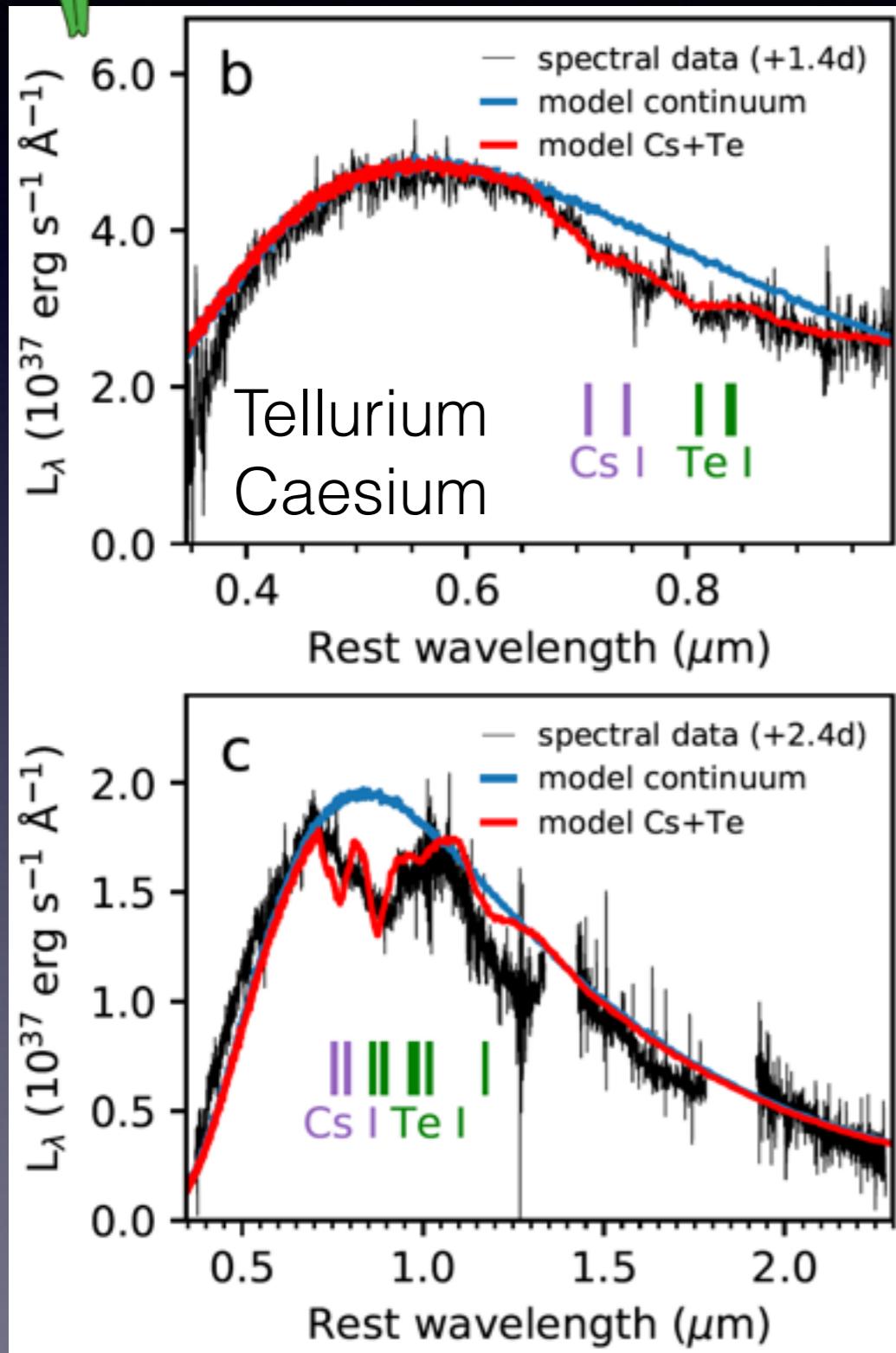


AT 2017gfo (kilonova, GW 170817)

distance = 44 Mpc
 $M_r = -15.8$ mag



the fast



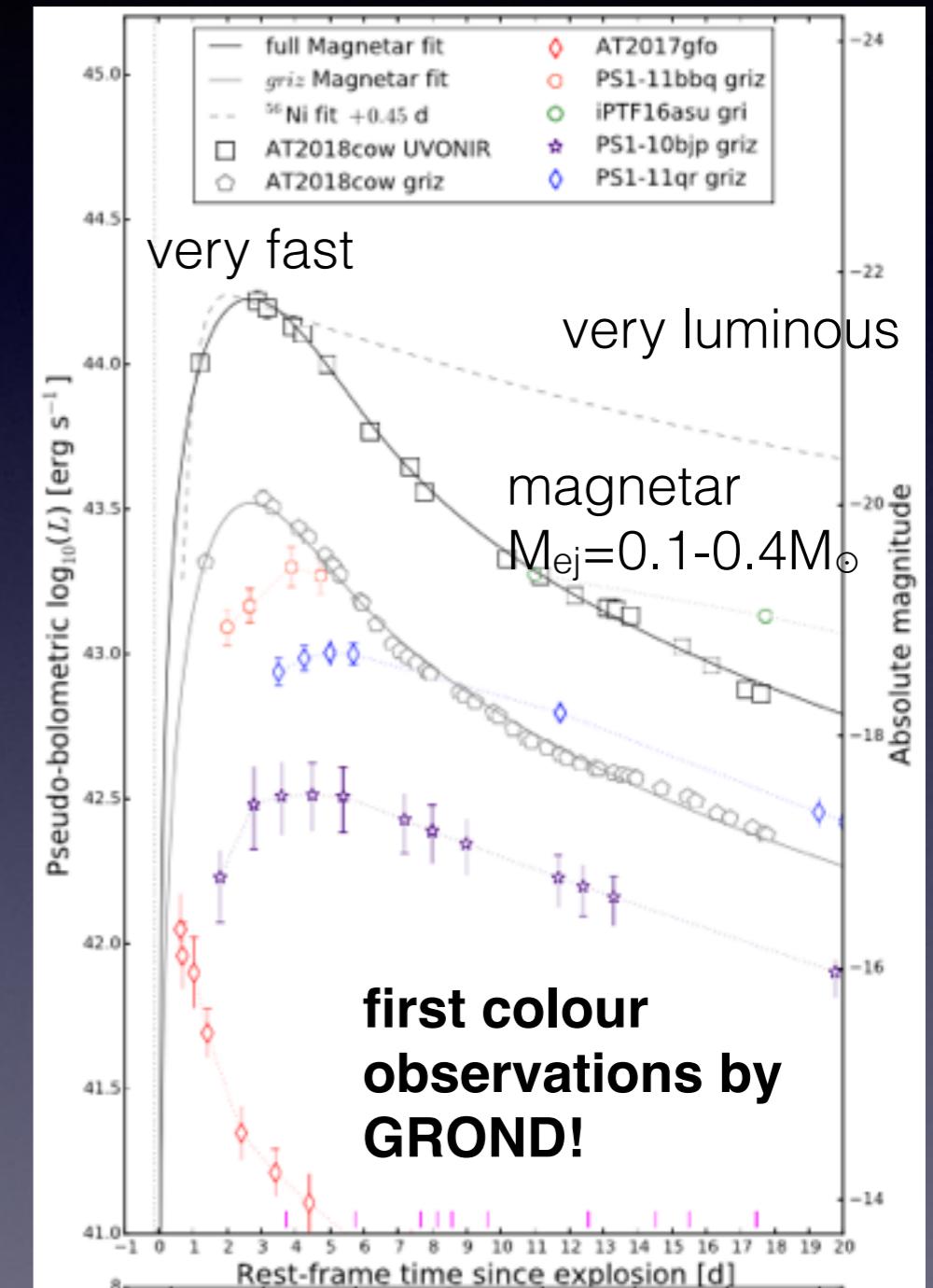
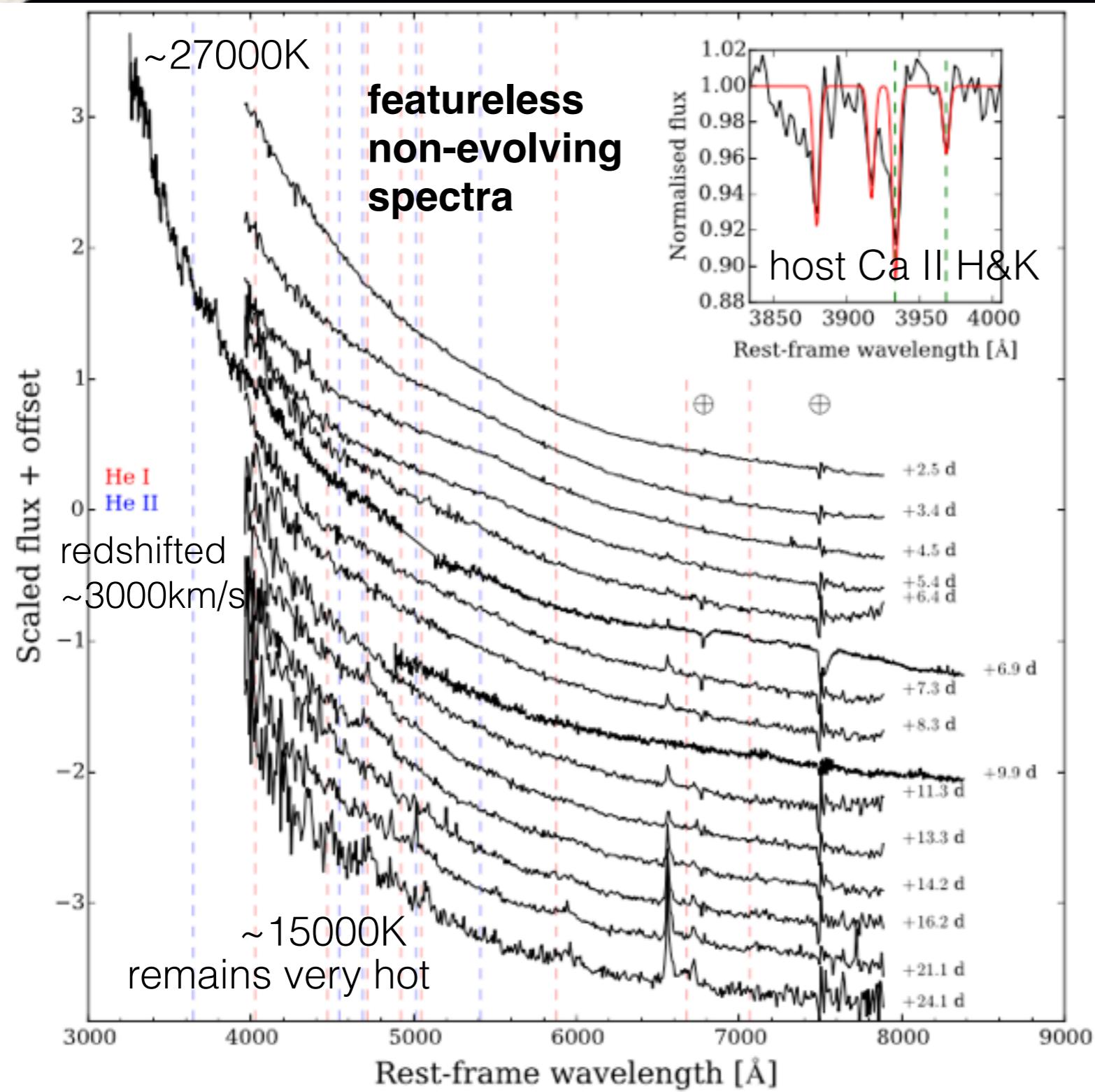
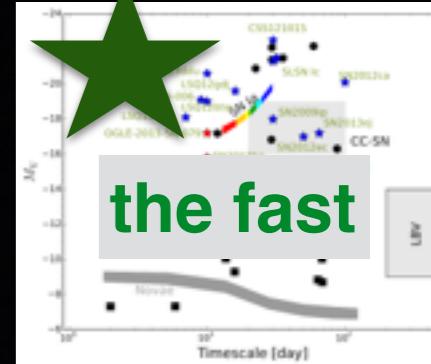
Smartt, Chen et al. Nature, 2017, 551, 75

The power source is constrained to have a power law slope of $\beta = -1.2^{+0.3}_{-0.3}$ consistent with radioactive powering from r-process nuclides

AT 2018cow (SN? NS merger? TDE?)

distance = 66 Mpc
 $M_g = -20.4$ mag

distance = 66 Mpc
 $M_g = -20.4$ mag

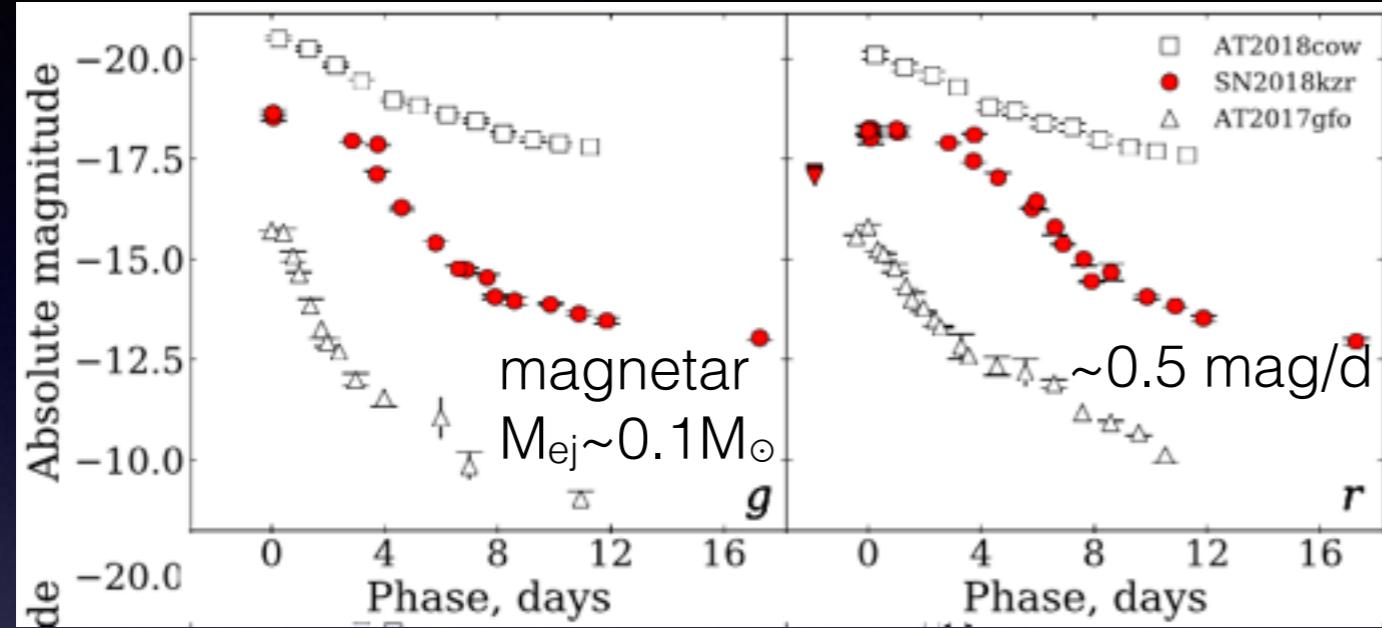
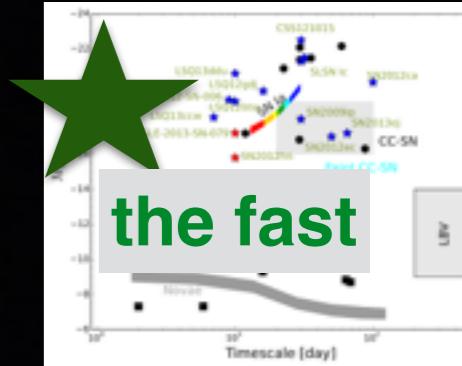


Prentice et al. 2018, ApJ, 865L, 3

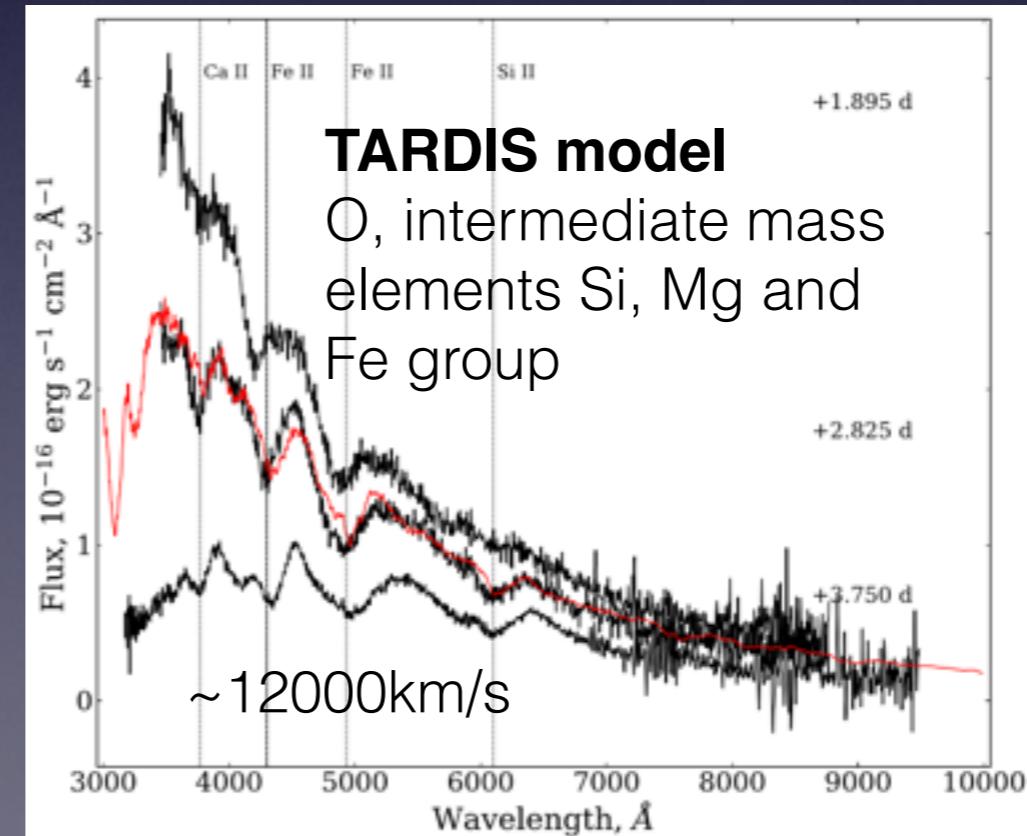
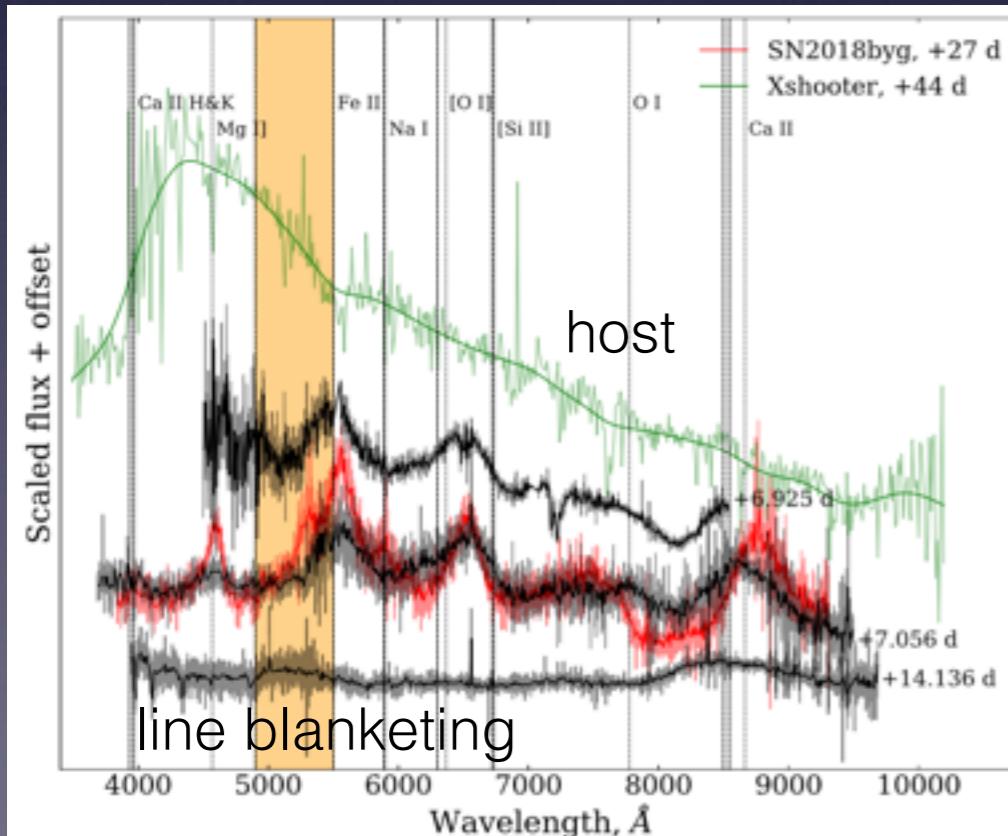
SN 2018kzr

(the fastest declining SN-like transient)

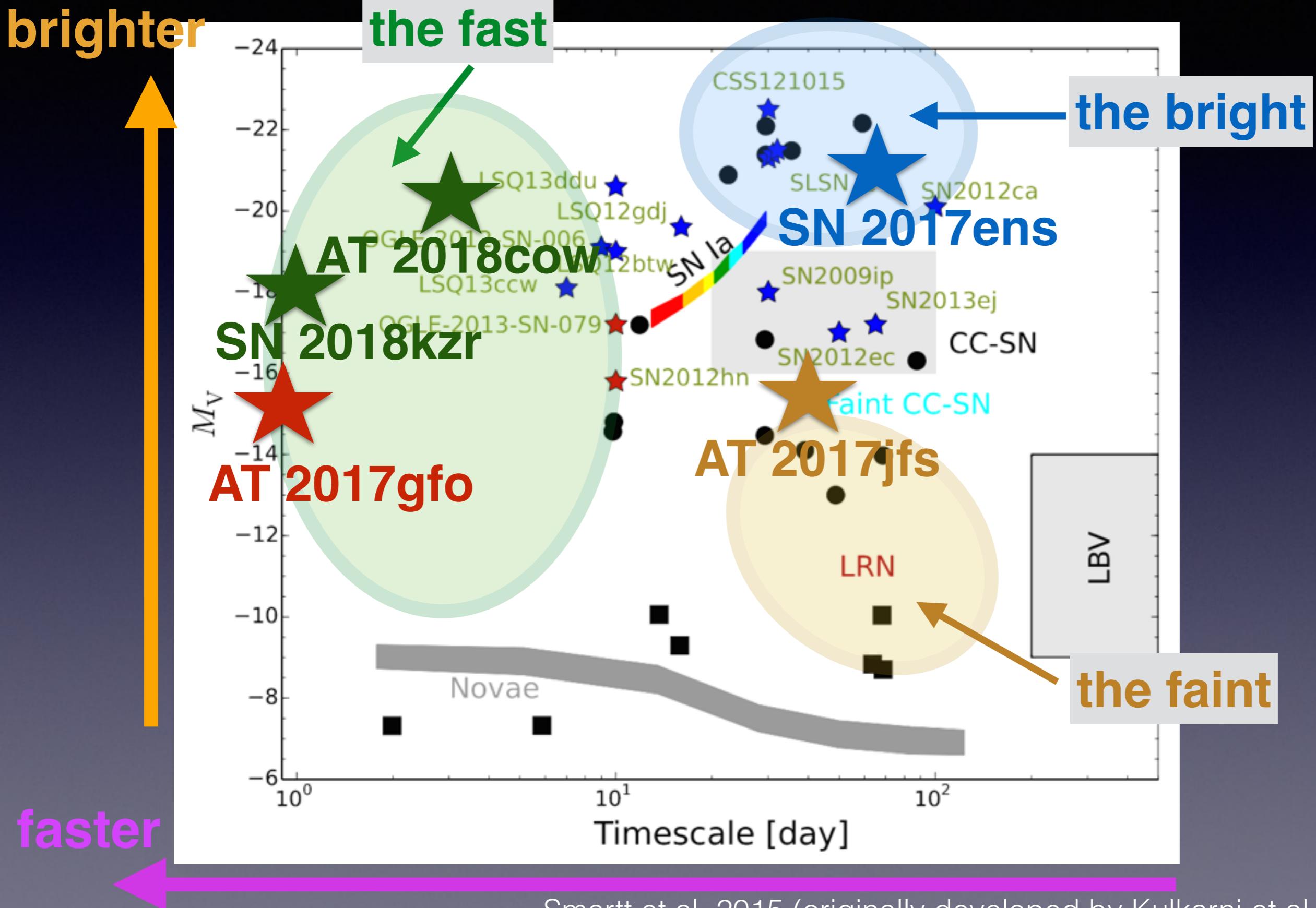
distance = 44 Mpc
 $M_r = -18$ mag



- alternative input energy source: the accretion induced collapse with magnetar powering or a WD - NS merger with energy from disk wind shocks
- further quantitative modelling: Gillanders et al. in prep.



Transient space phase



Transient space phase

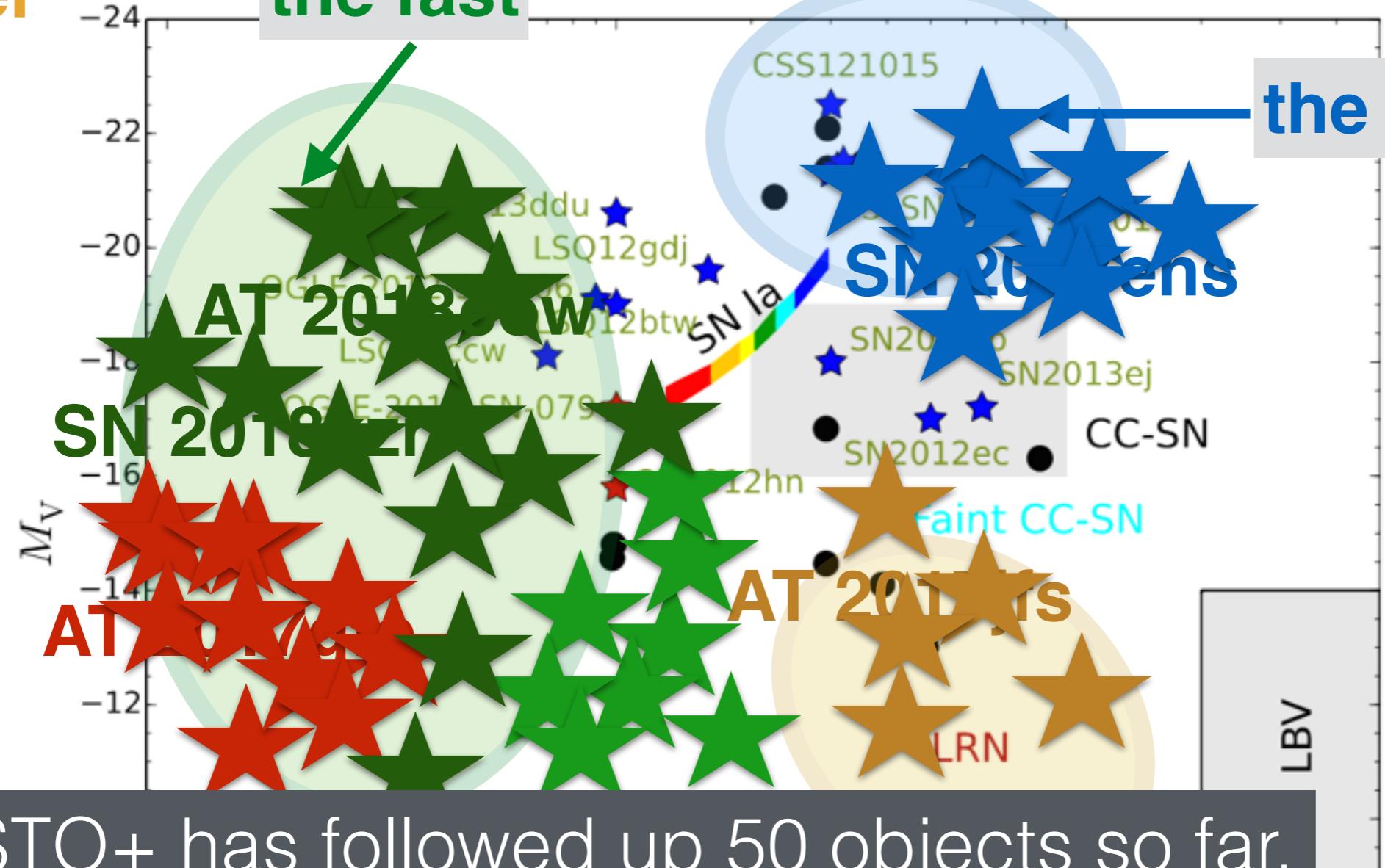
brighter

the fast

the bright

the faint

M_V



ePESSTO+ has followed up 50 objects so far,
22 of those were classified by ourselves!

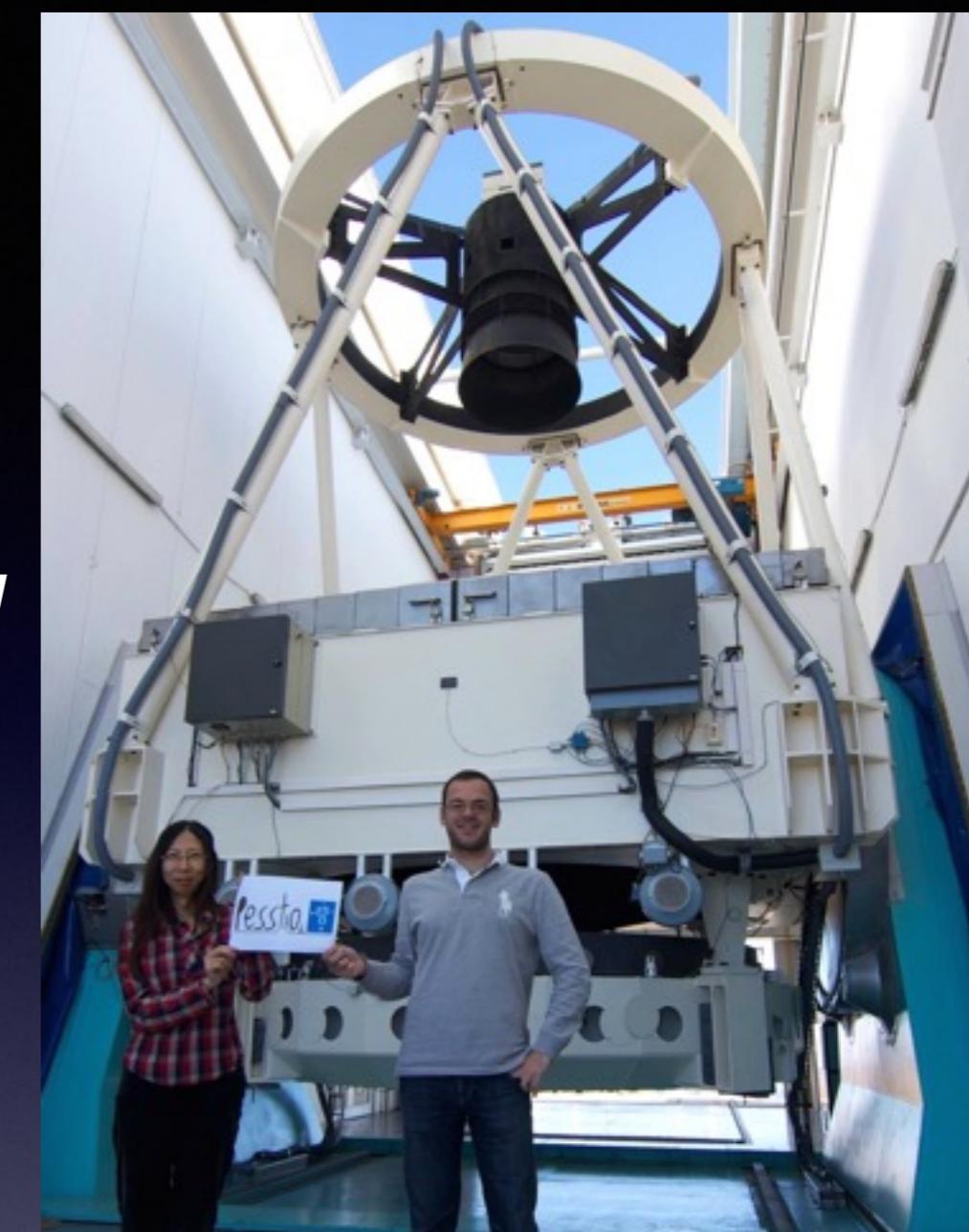
faster

Timescale [day]



Dankeschön!

<http://www.pessto.org/>



@NTT, La Silla
09/2013