# A water budget dichotomy of rocky protoplanets from <sup>26</sup>Alheating

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### **Tim Lichtenberg (U Oxford)** Gregor J. Golabek (BGI Bayreuth) Remo Burn (U Bern)

Michael R. Meyer (U Michigan) Yann Alibert (CSH/U Bern) Taras V. Gerya (ETH Zurich) Christoph Mordasini (CSH/U Bern)

# INTRO

<sup>26</sup>Al powered internal heating and chemical differentiation of rocky bodies during planet formation in the early Solar System, but is likely inhomogeneously distributed across forming planetary systems.

# **METHODS**

We coupled models of <sup>26</sup>Al-driven planetesimal dehydration and a planet population synthesis code to investigate its effects on exoplanet population statistics and potential observability.

## **RESULTS & CONCLUSIONS**

In the planetesimal-based accretion framework, <sup>26</sup>Al displays first-order control on the distribution of planet water abundance, and thus exoplanet radii, possibly traceable with TESS or PLATO.





# Varying <sup>26</sup>Al abundances across planetary systems induce systemic changes in rocky exoplanet water content

 $\geq {}^{26}Al_{\odot}$  during planet formation: ~ 10–100× water depletion Up to 10% smaller Rtransit Intra-system size correlation



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# No/few <sup>26</sup>Al present:

- Water worlds dominant type of rocky exoplanets
- Stable against hydrodynamic escape



# **Planet population**



Final bulk planet water mass fraction  $f_{HoO}$  [wt%]

# Mean transit radius shift

