

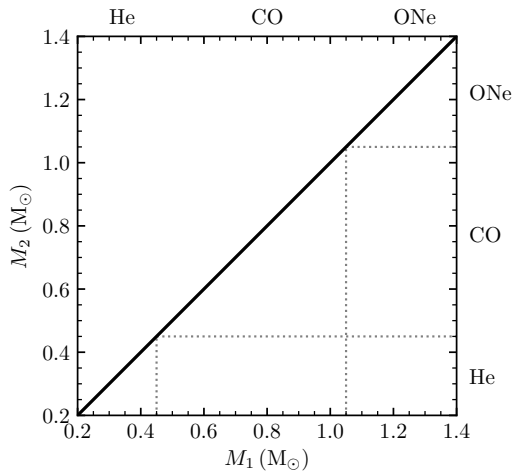
The Long-lived Remnants of Massive WD Mergers

Josiah Schwab

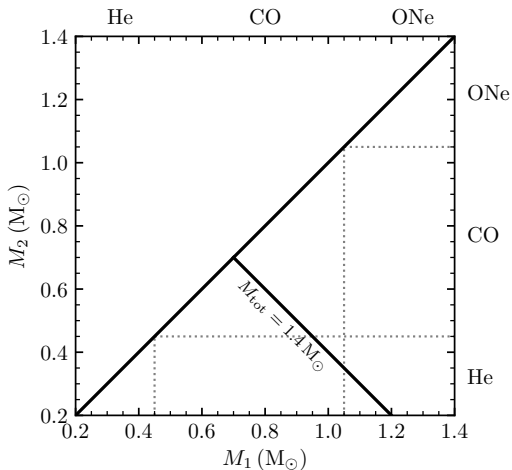
Hubble Fellow, UC Santa Cruz

05 July 2019

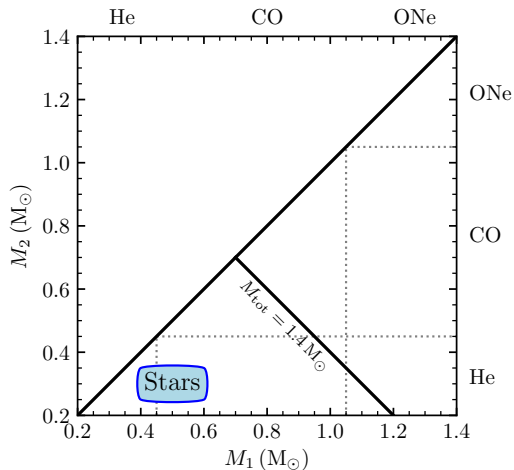
What is the final fate of two WDs that merge?



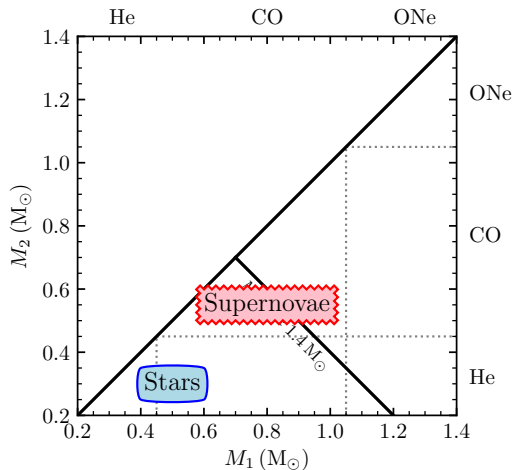
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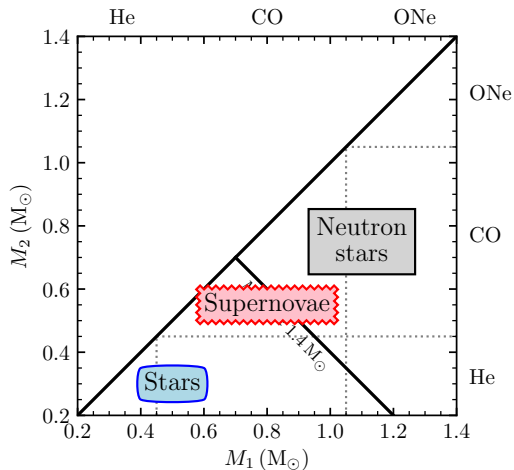
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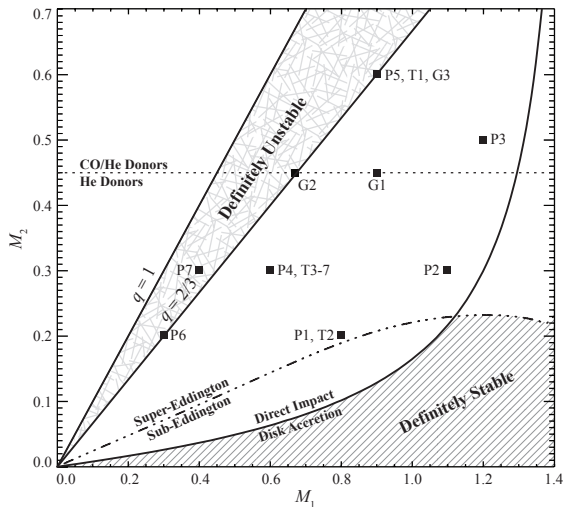
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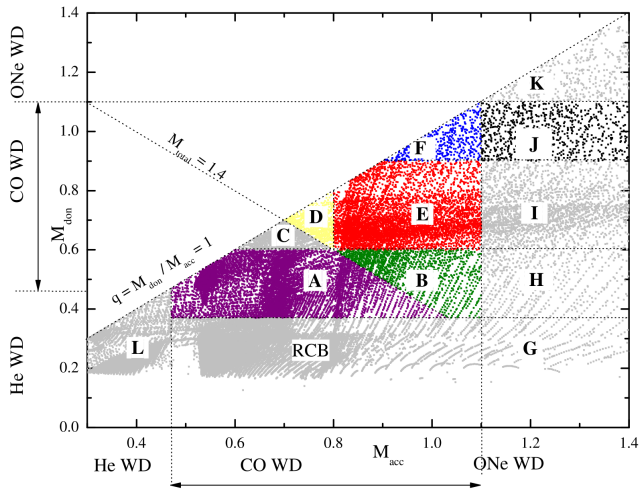


But which binaries become unstable and merge?



e.g., Marsh et al. (2004); Fig. from Dan et al. (2011)

And what populations of close double WDs form?



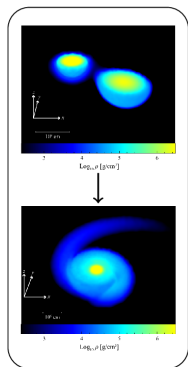
Yungelson & Kuranov (2017)

Review of white dwarf mergers

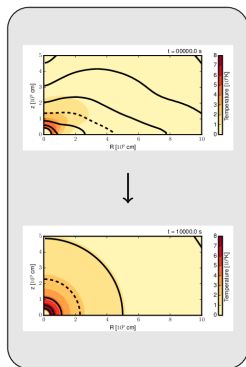
From super-Chandra merger to a neutron star

Summary

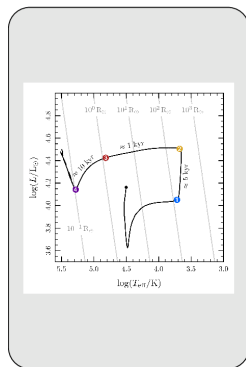
On the way to their final fates, double WD systems evolve through multiple phases.



~ minutes

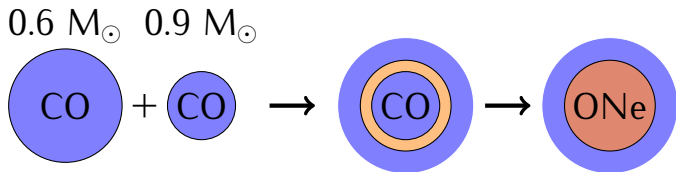


~ hours



~ $10^4 - 10^8$ yr

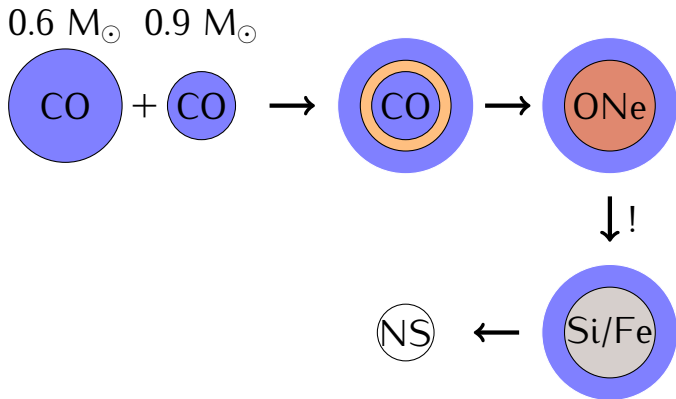
A super-Chandrasekhar total mass
does not imply a thermonuclear supernova.



Nomoto & Iben (1985); Saio & Nomoto (1985)

Schwab et al. (2016)

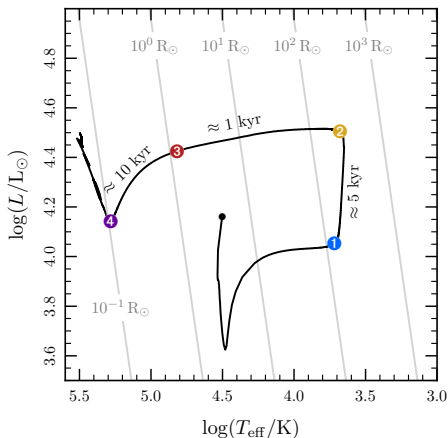
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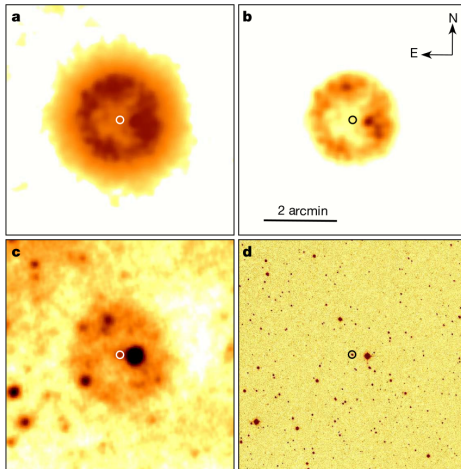
Schwab et al. (2016)

Collapse to an NS is essentially never prompt;
it typically requires $\gtrsim 10^4$ years.

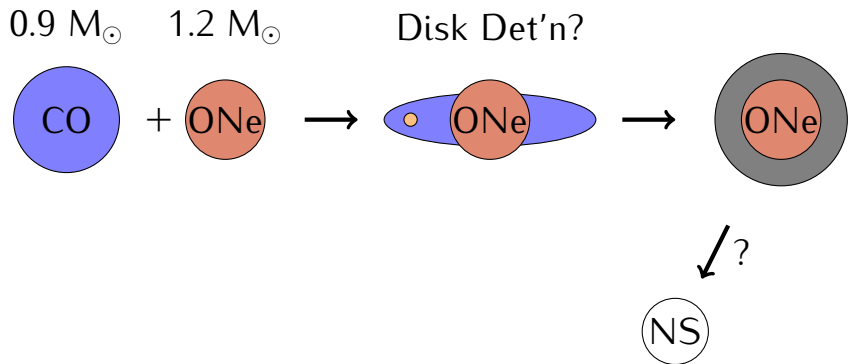


Schwab et al. (2016); see Gvaramadze et al. (2019)

A good candidate for such a merger remnant was recently found.



A CO + ONe merger may also produce an NS.



Lyutikov & Toonen (2018); Kashyap et al. (2018)

Review of white dwarf mergers

From super-Chandra merger to a neutron star

Summary

- ▶ Massive WD mergers make $\approx M_{\text{Ch}}$ metal cores, whose evolution may mirror that of "low mass massive stars".
- ▶ Delay ($\sim 10^4$ yr) between merger and collapse, so look for systems on the way to becoming a NS.

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