

# 1 Controlling competing photochemical reactions 2 stabilises perovskite solar cells

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20

21 **Abstract**

22 **Metal halide perovskites have been successfully applied as optically active layers in**  
23 **photovoltaics and in various optoelectronic devices. Long-term reliability must**  
24 **however be assured. Instabilities are manifested as light-induced ion migration and**  
25 **segregation, which can eventually lead to material degradation. Discordant reports**  
26 **have reported a beneficial role of ion migration under illumination, leading to**  
27 **defect healing. By combining *ab initio* simulations with photoluminescence**  
28 **measurements under controlled conditions, we demonstrate that photo-instabilities**  
29 **are related to light-induced formation and annihilation of defects acting as carrier**  
30 **trap states. We show that these phenomena coexist and compete. In particular,**  
31 **long-living carrier traps related to halide defects trigger photo-induced material**  
32 **transformations, which drive both processes. On short scales, defect annihilation**  
33 **can prevail over defect formation, which occurs on longer scales. Nevertheless,**  
34 **defect formation can be controlled by blocking under-coordinated surface sites,**  
35 **which act as a defect reservoir. By a specific surface passivation strategy we are**  
36 **thus able to stabilize the perovskite layer towards photo-induced instabilities,**  
37 **leading to improved optoelectronic material quality and enhanced photo-stability**  
38 **in a working solar cell.**

39       The interest in metal halide perovskites has been growing continuously in recent  
40 years due to the remarkable evolution of their application in photovoltaics – that have  
41 reached power conversion efficiencies exceeding 23%<sup>1</sup> – as well as promising  
42 performance in light emitting diodes<sup>2</sup>, photodetectors, and lasers<sup>3,4</sup>. This success owes  
43 much to their optimal optoelectronics properties, *i.e.* high absorption coefficients and  
44 long carrier lifetimes<sup>5-7</sup>. The soft nature of the lead-halide lattice induces relatively low  
45 defect formation energies, which implies a significant probability of defects formation<sup>8</sup>

46 through combined thermal- and light-induced phenomena. While the impact of defects  
47 seems not to be detrimental with respect to the figures of merit of perovskite-based  
48 devices,<sup>9</sup> their activity poses a challenge to the widespread uptake of perovskite-based  
49 optoelectronic devices in terms of material stability under electrical and optical  
50 stimuli.<sup>10-14</sup> Monitoring the photoluminescence (PL) and dynamics of photogenerated  
51 charge carriers is an effective approach for gathering information about the opto-  
52 electronic properties of a semiconductor. The PL efficiency is extremely sensitive to a  
53 combination of parameters such as recombination rate constants, defect densities, carrier  
54 lifetimes, and non-radiative decay paths; as such PL efficiency is usually taken as a  
55 measure of the material optoelectronic quality.

56         Since the primary function of lead halide perovskites in solar cells is to absorb  
57 light, the material response to light exposure and the possible associated phenomena are  
58 fundamental properties to enable reliable devices. A PL intensity decrease (PLID) has  
59 been observed in lead halide perovskite thin-films of different composition as a result of  
60 photo-induced trap formation in inert atmosphere<sup>15</sup>. Enhanced photo-induced ion  
61 migration<sup>16</sup> was interpreted as a defect density increase under light irradiation<sup>17</sup>. In  
62 direct contradiction, PL intensity enhancement (PLIE) under illumination has also been  
63 observed<sup>18,19</sup> which was explained by a light-induced trap annihilation mechanism.<sup>20</sup> A  
64 dependence of the PL efficiency variation on defect density<sup>21</sup> and excitation intensity<sup>22</sup>,  
65 with both PLIE and PLID, was reported. The atmosphere also has a strong impact on PL  
66 efficiency, with strong PLIE observed by exposing the sample to an oxygen-rich  
67 environment<sup>15,23-27</sup>.

68         The presence of transient phenomena upon electrical and photo-excitation and a  
69 lack of understanding of the mechanisms behind such phenomena raise skepticism  
70 about the successful integration of lead halide perovskites in a reliable technology.

71 From a more fundamental point of view, no general consensus exists on whether the  
72 observed transient behaviors can eventually lead to an improved material quality or they  
73 can actually destroy the material.

74 Here we provide a comprehensive picture of the PLIE and PLID processes in  
75 methylammonium lead iodide ( $\text{MAPbI}_3$ ) and bromide ( $\text{MAPbBr}_3$ ) perovskites, showing  
76 that both phenomena can happen and compete in both materials by varying the  
77 illumination conditions, although with different characteristic times and temperature  
78 response.

79 Combining our observations with first-principles calculations we propose a  
80 model based on lead halide defect chemistry for both PLID and PLIE which  
81 consistently explains the overall phenomenology. We find that long-living carrier traps  
82 related to halide defects trigger photo-induced transformations which drive both PLIE  
83 and PLID. As long as the density of traps is relatively low or they are scarcely  
84 populated, photo-induced PLIE will be dominant. On the other hand, photo-induced  
85 PLID will take over in the presence of a high density of populated traps located close to  
86 thin film surface and grain boundaries, which provide a reservoir of binding sites for  
87 photo-generated species eventually leading to material degradation. Consistent with the  
88 developed model, we show that surface passivation of  $\text{MAPbI}_3$  thin films allows us to  
89 switch off the PLID mechanism and make the PLIE process dominant, which translates  
90 into an associated solar cell efficiency increase and enhanced temporal stability.

91

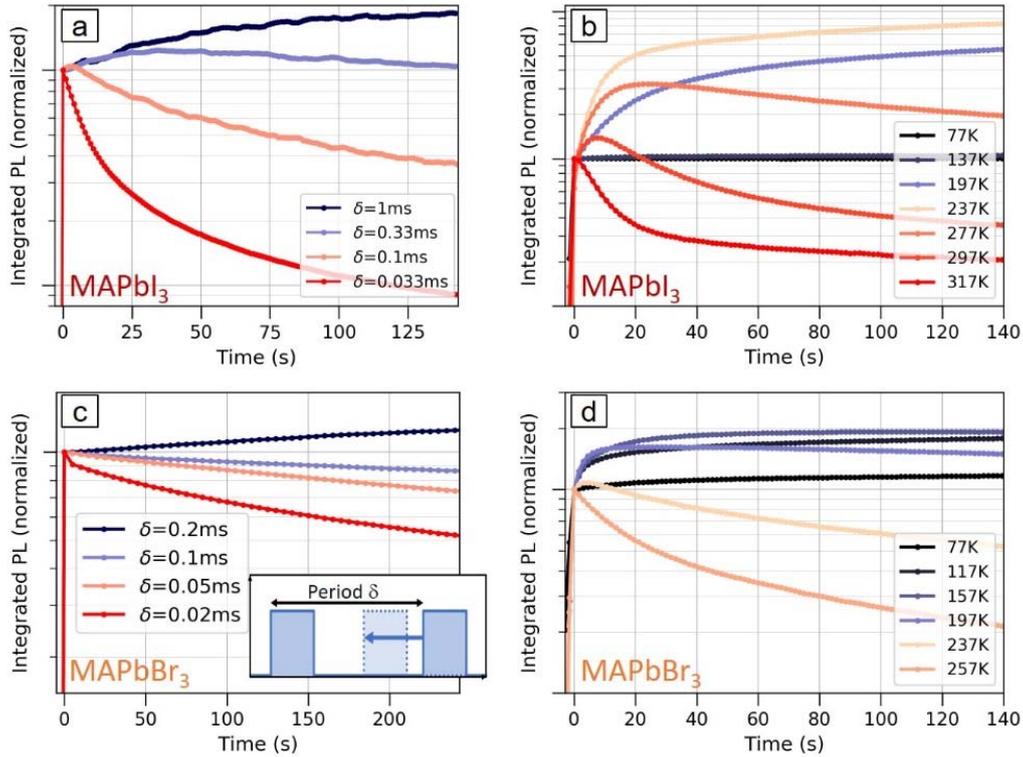
## 92 **Results**

93 We monitored the PL intensity of  $\text{MAPbI}_3$  and  $\text{MAPbBr}_3$  perovskite thin-films used in  
94 efficient solar cell devices as a function of illumination time. Figure 1a shows the room  
95 temperature integrated PL over time of  $\text{MAPbI}_3$  polycrystalline films deposited on

96 glass. The sample was excited with a laser beam incident on the film surface. The laser  
97 was modulated with a fixed pulse width of 200 ns and varying repetition rate, *i.e.*  
98 varying the period ( $\delta$ ) between two excitation pulses. The pulse energy was kept  
99 constant to guarantee the same carrier generation density at each excitation pulse (in  
100 Table S2 of SI the average excitation intensity as a function of  $\delta$  are reported). To  
101 exclude the effect of oxygen and moisture<sup>15,24,25</sup> all measurements were performed  
102 under active vacuum (pressure  $<10^{-5}$  mbar, under constantly running pump). We  
103 observe that the PL intensity changes over time, showing both PLIE and PLID,  
104 depending on the repetition rate for both MAPbI<sub>3</sub> and MAPbBr<sub>3</sub> (Figures 1a,c). At low  
105 repetition rates -that is long  $\delta$ , corresponding to the sample being in the dark for ~ms  
106 time – the PL intensity cumulatively grows with subsequent illumination periods.  
107 Increasing the repetition rate - short  $\delta$ , corresponding to longer light exposure and less  
108 time in the dark- a quenching process kicks in, which becomes eventually dominant  
109 (Figures 1a,c). These observations clearly indicate that *light absorption can lead to*  
110 *competing PLIE and PLID phenomena*, depending on the excitation conditions. We thus  
111 performed additional experiments at varying temperatures to disentangle the factors  
112 underlying the individual phenomena involved in PLIE and PLID. Figure 1b shows the  
113 time evolution of the integrated PL intensity of a MAPbI<sub>3</sub> film excited with continuous-  
114 wave (CW) light, analogous to a very high repetition rate, *i.e.* short  $\delta$ , at different  
115 temperatures. Each curve was taken on a fresh spot of the sample. For MAPbI<sub>3</sub>, at 77 K  
116 the PL is stable for the entire duration of the experiment, and the same holds when  
117 heating the sample to 137 K. Further raising the temperature to 197 K we observe a  
118 strong PLIE over time which is further boosted up to 237 K, leading to a factor of ~8 PL  
119 increase compared to the initial value. At still higher temperature (277 K) an initial  
120 PLIE is superseded by PLID at longer exposure times, which eventually dominates at

121 room temperature and above. Similar behavior was observed for MAPbBr<sub>3</sub> films upon  
122 varying the temperature, though with lower thresholds for PLID and PLIE (Figure 1c,d),  
123 making us safely exclude any role of structural phase transition in the observed  
124 dynamics (see also S27 in SI).

125         The PL intensity variations with increasing temperature and the transition from  
126 stable PL, to PLIE to PLID, is clearly suggestive of competing thermally activated  
127 processes that either improve or hinder the efficiency of radiative recombination. The  
128 dominant contribution of PLID at higher temperature suggests a higher energy demand  
129 for this process compared to PLIE. Also, the enhanced stability range of MAPbI<sub>3</sub>  
130 compared to MAPbBr<sub>3</sub> at low temperatures (cf. Figure 1b and 1d, with PLID starting at  
131 277 and 197 K respectively) suggests that the threshold for switching between the two  
132 phenomena is determined by the precise material composition, in particular by the  
133 nature of the halide.

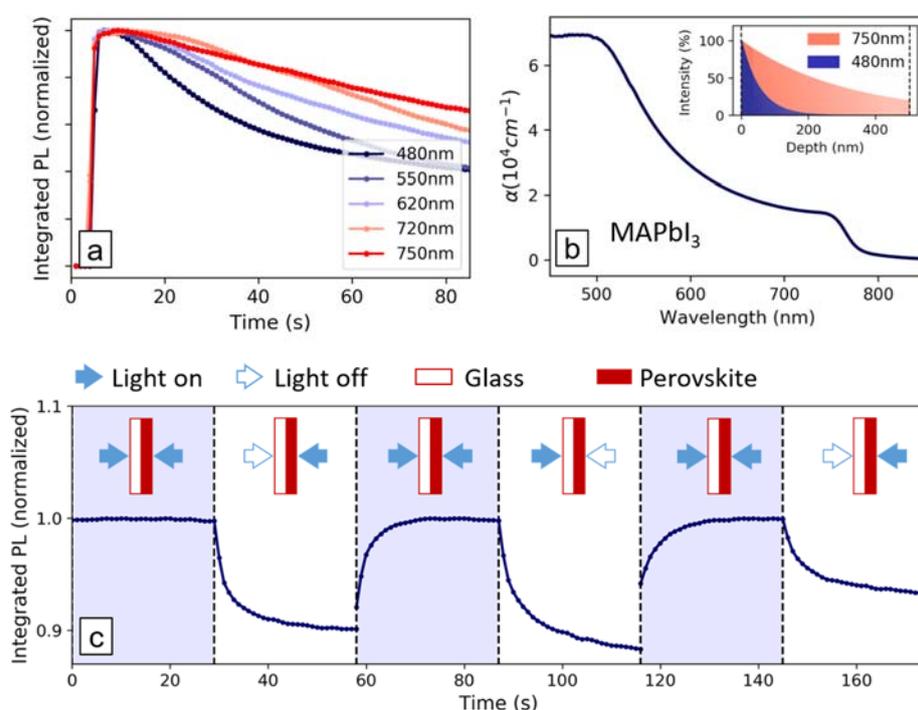


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135 **Figure 1** a,c) Integrated PL intensity, normalized at time  $t=0$ , over time (excitation  
 136 density  $\sim 10^{16} \text{ cm}^{-3}$ ) on fresh spots of polycrystalline films of MAPbI<sub>3</sub> (a) and  
 137 MAPbBr<sub>3</sub> (c) with increasing repetition rate (decreasing period  $\delta$ ) of the excitation  
 138 light. The pulse width was fixed at 200 ns with constant pulse energy of 0.2 nJ (fluence  
 139  $\sim 0.5 \mu\text{J}\cdot\text{cm}^{-2}$ , average intensity from 0.6 to 20 mW/cm<sup>2</sup>, see also data with 10  $\mu\text{s}$  pulse  
 140 width shown in Fig S5). b-d) Integrated PL over time of MAPbI<sub>3</sub> (b) and MAPbBr<sub>3</sub> (d)  
 141 at different temperatures under CW excitation (excitation intensity  $\sim 50 \text{ mW/cm}^2$ ).  
 142 Excitation wavelength: 450 nm for MAPbI<sub>3</sub> and 405nm for MAPbBr<sub>3</sub>.

143 To identify what additional factors affect the PLID and PLIE processes, we investigated  
 144 the PL time evolution of MAPbI<sub>3</sub> as a function of the excitation wavelength, as shown  
 145 in Figure 2a. The incident photon density over the illuminated spot area was kept  
 146 constant, and each curve was taken on a fresh spot of the film. We observe that PLID is  
 147 significantly enhanced at shorter excitation wavelengths, while band-edge excitation  
 148 results in more stable PL. Besides the additional thermal energy provided by above-gap

149 excitation, shorter wavelength light also has significantly shorter penetration depth  
 150 (Figure 2b). This results in a higher carrier generation close to the film surface, with a  
 151 factor  $\sim 2$  increase in the first 30 nm film away from the surface in case of 480 vs. 750  
 152 nm excitation. While we cannot exclude direct excitation of  $\text{PbI}_2$  as the cause of PLID  
 153 and PLIE,<sup>27</sup> the observation of both phenomena for  $\text{MAPbBr}_3$  is consistent with a  
 154 minor role of this phenomenon in our conditions, as excitation of the higher band-gap  
 155 ( $\sim 4$  eV)  $\text{PbBr}_2$  phase only occurs at much shorter excitation wavelength than those used  
 156 in our experiments. Our observations suggest that the competition between PLID and  
 157 PLIE is strongly related to the illumination conditions of the thin film which may  
 158 include factors such as proximity to the reactive surface, uniformity of the excitation  
 159 profile, and excitation density, with possible contribution of thermal dissipation effects  
 160 following photo-excitation.



161  
 162 **Figure 2.** (a) Integrated PL over time obtained using excitation light of different  
 163 wavelengths. (b) Absorption spectrum of  $\text{MAPbI}_3$  thin film (see absorbance in S28) and  
 164 photoexcitation profiles when using 480 nm or 750 nm light (inset). (c) Integrated PL

165 intensity over time of a MAPbI<sub>3</sub> film on glass. Excitation is performed at 560 nm with  
166 symmetrical beams of equal intensity impinging on both the perovskite and glass side of  
167 the sample. Each section of the time trace is normalized by its maximum intensity for  
168 visualization purpose (See Figure S6 in SI for unnormalized data).

169

170 To complement our phenomenological consideration, we investigated whether the  
171 experiment geometry could affect the relative weight of PLID vs. PLIE. We find a  
172 reproducible PLIE when illuminating the film from two sides (the open surface and the  
173 glass side) with two identical beams covering a large area of the thin film, suggesting  
174 that a uniform illumination may reduce the impact of PLID (Figure 2c). Similar  
175 observations, *i.e.* a switch from PLID to a more stable PL is also observed by varying  
176 the lateral illumination gradient (see Supplementary Information Figure S7).

177 To provide a rationale for these observations we first consider the possible role  
178 of trap-state filling, followed by de-trapping, as a possible cause determining the  
179 observed phenomenology. Despite the long lifetimes of trapped carriers in both MAPbI<sub>3</sub>  
180 and MAPbBr<sub>3</sub><sup>9,15,28</sup>, extending up to several  $\mu$ s, the timescales spanned by the repetition  
181 rates in Figures 1a,c are too long to be solely explained by trapping/de-trapping  
182 mechanisms. The system can reach almost a stationary density of charge carriers within  
183 the 200 ns pulse (see Supplementary Information, Figure S2-S4 and S22-S25, for  
184 simulations of charge carrier dynamics under CW or modulated illumination. Please  
185 note that the change in modulation mainly results in a change in the averaged excitation  
186 intensity and an enhancement of the PL absolute value at time  $t=0$ , due to an associated  
187 enhancement of counts during the integration, see Fig S26 in SI ); if no concomitant  
188 processes take place, though, the ms delay time between two consecutive pulses allows  
189 the system to return to the ground state before a second pulse comes, in contrast to the

190 cumulative effect of pulsed illumination observed in Figure 1a,c. Likewise,  
191 charge/lattice interaction through formation of large polarons<sup>29</sup> show lifetimes  
192 comparable to those of the charge carriers, inconsistent with the cumulative effect  
193 observed with ms delay. Similarly, we can rule out a role of light induced release of  
194 lattice strain<sup>30</sup> in PLE, since electronic effects are too fast to provide a cumulative effect  
195 on the ms repetition time scale of our experiments.

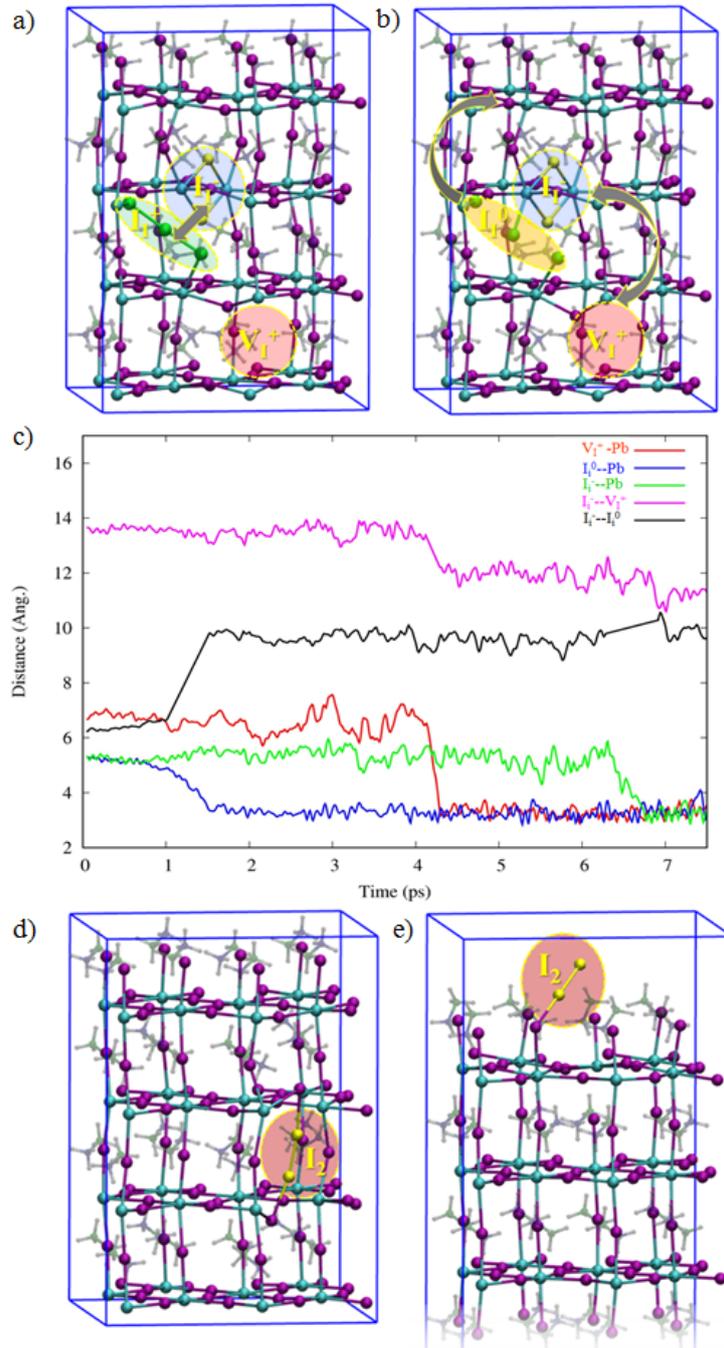
196         The time scales involved in PLID and PLIE are consistent with the reported  
197 ionic activities, such as ion/defect annihilation and migration rates, in both MAPbI<sub>3</sub> and  
198 MAPbBr<sub>3</sub>,<sup>20,31</sup> and with the time scale of photo-induced transformations in free standing  
199 MAPbI<sub>3</sub> films.<sup>32</sup> Comparably, in mixed halide (I, Br) perovskites ion segregation was  
200 shown to take place via halide defects, mainly vacancies and interstitials, driven by the  
201 gradient in carrier generation through the thickness of these strongly absorbing  
202 materials<sup>33-35</sup>. Importantly, the observation of typical spectral features associated with  
203 trap carriers and defects in the lattice that are enhanced under illumination (see  
204 Supplementary Information Figure S8 and S9) is suggestive of photo-induced defect  
205 formation and annihilation being respectively related to PLID and PLIE.

206         With the aid of first-principles calculations we propose a model based on lead-  
207 halides defect chemistry which allows us to consistently explain the observations of the  
208 data set presented above. MAPbI<sub>3</sub> and MAPbBr<sub>3</sub> show a remarkably similar defect  
209 chemistry (see Supplementary Information for comparative analysis based on hybrid  
210 DFT calculations including spin-orbit coupling) which is dominated by lead vacancies  
211 ( $V_{\text{Pb}}$ ) and interstitial halogen ( $I_i$  or  $\text{Br}_i$ ) defects. We henceforth use the notation for  
212 defects in MAPbI<sub>3</sub> (*e.g.*  $I_i$ ), but the mechanism applies similarly to MAPbBr<sub>3</sub>. We  
213 notice, however, that for the latter the involved defects are somehow shallower than in  
214 MAPbI<sub>3</sub> (see Supplementary Information Figure S10). Both halide interstitials and lead

215 vacancies show thermodynamic ionization levels in the band gap. At the calculated  
216 native Fermi level  $V_{Pb}$  is stable in the 2- charge state, thus this defect could trap holes  
217 by the (-/2-) transition level whose energy falls 0.13 eV above the VB, while  $I_i^-$  exhibits  
218 a (0/-) charge transition 0.29 eV above the VB. The partial shallow character of  $V_{Pb}$   
219 transition, less deep than  $I_i$ , and the high migration barriers associated to the diffusion of  
220 the defect ( $E_a \sim 1.0$  eV)<sup>36</sup> highlight that this defect is only moderately active as a trap,  
221 while  $I_i$  can significantly trap both electrons and holes (see Supplementary Information  
222 Figure S10) through the (+/0) and (0/-) transitions. Hole trapping/de-trapping at  $I_i^-$  is a  
223 relatively fast process, while electron trapping at  $I_i^+$  induces long living states due to the  
224 strong associated geometrical relaxation decreasing the kinetics of electron  
225 detrapping<sup>9,28</sup>. The promoter of possible material transformations must be a long-living  
226 species, whose formation may trigger subsequent photo-chemical transformations  
227 leading to both PLID and PLIE. The most natural candidates are thus the filled electron  
228 traps with their associated  $\mu s$  lifetime.

229         Recombination of  $I_i^-$  with  $V_i^+$  was proposed as a mean of trap annihilation<sup>20</sup>,  
230 which restores the pristine material and decreases the concentration of traps, thus it can  
231 be associated to PLIE. Here we further disclose that long living electrons at trapped  $I_i^+$   
232 defects may additionally mediate Frenkel defect annihilation.

233



234

235 **Figure 3.** a) Optimized structure of the interacting (as signaled by the arrow)  $I_i^+/I_i^-$   
 236 defect pair (shaded green and blue areas, respectively) and a distant  $V_I^+$  defect (shaded  
 237 red area). The formation of the  $I_i^---V_I^+$  defects in the presence of an interacting  $I_i^+$  is  
 238 endothermic by 0.55 eV. b) Electron trapped at  $I_i^+$  to form  $I_i^0/I_i^-$  defect pair (shaded  
 239 orange and blue areas, respectively). The arrows indicate the direction of  $I_i^0$  migration  
 240 and  $I_i^---V_I^+$  annihilation. The formation of the  $I_i^---V_I^+$  defects in the presence of a

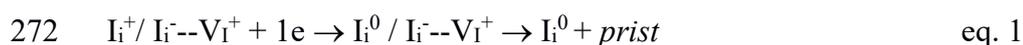
241 neutral  $I_i^0$  is endothermic by 0.86 eV. c) Dynamical evolution of the system after  
242 electron trapping at  $I_i^+$ . A sequence of ion reorganization and migration events can be  
243 visualized: 1) following electron trapping at  $I_i^+$ , the system rapidly reaches the  
244 minimum energy structure of  $I_i^0$  (blue line) increasing the  $I_i^0/I_i^-$  distance (black line); 2)  
245  $V_I^+$  migrates in the direction of  $I_i^-$  (red and magenta lines); 3)  $I_i^-$  further migrates in the  
246 direction of  $V_I^+$  (green and magenta lines). d) Optimized structure of an unstable bulk  $I_2$   
247 molecule (dark red shaded area), which migrates to the perovskite (001) surface in e) to  
248 form a stable surface bound  $I_2$  molecule (dark red shaded area). Binding of an  $I_2$   
249 molecule to the perovskite surface (bulk) is exothermic (endothermic) by 0.48 (0.40) eV.

250

251 Briefly, the formation of a  $I_i^-$ -- $V_I^+$  Frenkel pair in the presence of a neighboring  $I_i^+$  is  
252 endothermic by 0.55 eV since this process requires to break a Pb-I bond and displacing  
253 the  $I_i^-$ -- $V_I^+$  defects away, Figure 3a. Upon electron trapping at  $I_i^+$ , which is preferentially  
254 located next to  $I_i^-$ , the system is further destabilized by 0.31 eV, Figure 3b, adding a  
255 significant driving force to restore the pristine material. The long-living nature of the  
256 formed  $I_i^0$ , which mirrors the long electron de-trapping time, may allow sufficient time  
257 for  $I_i^-$ -- $V_I^+$  annihilation to effectively take place. Further insight on such process was  
258 obtained by *ab initio* molecular dynamics. We start from the minimum energy geometry  
259 of  $I_i^-$ -- $V_I^+$  Frenkel pair interacting with  $I_i^+$  (Figure 3a) and add one electron to the  
260 system, following the dynamical evolution of this globally neutral system. A sequence  
261 of ion reorganization and migration events can be visualized: 1) following electron  
262 trapping at  $I_i^+$ , confirmed by spin localization at the defect site, the system rapidly  
263 reaches the minimum energy structure of  $I_i^0$  (blue line in Figure 3c) increasing the  $I_i^0/I_i^-$   
264 distance (black line); 2)  $V_I^+$  migrates in the direction of  $I_i^-$  (red and magenta lines); 3)  $I_i^-$   
265 further migrates in the direction of  $V_I^+$  (green and magenta lines). The results are

266 consistent with electron trapping at  $I_i^+$  promoting  $I_i^-$ -- $V_I^+$  annihilation, as inferred from  
267 the two oppositely charged defects migrating towards each other twice within less than  
268 10 ps. A simulation starting from the same structure but with no added electron at  $I_i^+$  did  
269 not show any shortening of the  $I_i^-$ -- $V_I^+$  distance, though the  $I_i^+$  /  $I_i^-$  defects jointly  
270 migrated in the investigated time, see Supporting Information Figure S11.

271 The light-induced PLIE process can be schematized as follows:



273 where *prist* represents the pristine material. Annihilation of the  $I_i^-$ -- $V_I^+$  Frenkel pair is  
274 entropically disfavored, with a small energy barrier related to the migration of  $I_i^-$  and  
275  $V_I^+$  ( $\sim 0.1$  eV)<sup>37</sup>. This reaction is thus favored at low temperature but it still requires a  
276 sufficient thermal energy to accomplish defect migration.

277 The detrapping reaction:



279 instead has a significant activation energy (0.29 eV)<sup>9</sup>, so the annihilation process of  
280 eq.1, mediated by the high ion mobility of iodine defects, can effectively prevail over  
281 the thermally activated detrapping of eq.2. Such PLIE mechanism is fully consistent  
282 with the data of Figure 1b,d, showing a preferential PLIE at low temperature but no  
283 PLIE if the temperature is too low to hamper defect migration. In the absence of other  
284 (PLID) processes this PLIE mechanism is cumulative since it requires a long time after  
285 de-trapping to restore the initial density of Frenkel defects due to the endothermic  
286 nature of the process, with the possible associated entropic gain hindered at low  
287 temperature. By definition, in a Frenkel defect an ion leaves its place in the lattice,  
288 creating a vacancy, and it becomes an interstitial by lodging in a nearby location, thus  
289 we expect to observe their annihilation happening even when only short-range  
290 migration is allowed (*e.g.* at low temperature).

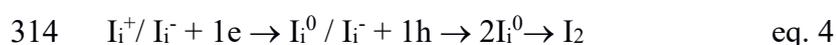
291 The PLID process is likely more complex than PLIE. We know that PLID is  
292 favored by high illumination repetition rates. This is in turn associated with a larger  
293 fraction of traps being filled at any time, and possibly to a higher effective temperature  
294 due to heating of the sample. We also know from the wavelength dependence of PLID  
295 that this process is more efficient when the photo-generation happens close to the  
296 surface where a larger trap density can be envisaged. It is also amplified when ion  
297 diffusion is enhanced, *i.e.* by the presence of gradient across dark/light regions.

298 A possible PLID mechanism could thus be reasonably associated to a  
299 bimolecular reaction – boosted by increasing the encountering probability of the  
300 reactants, *i.e.* by increasing their density and/or their mobility – occurring among the  
301 filled, long living traps, in a film region close to the surface. Following the initiating  
302 trap filling reaction of eq. 2, two filled traps could react to form an I<sub>2</sub> molecule:



304 This radical quenching reaction, involving two coordinated I<sub>2</sub><sup>-</sup> molecules,<sup>38</sup> takes place  
305 virtually with no energy barrier (Supplementary Information Figure S12) being it only  
306 limited by diffusion of the reactants. Such bimolecular reaction is favored by a high trap  
307 concentration in a small volume that increases the probability of a bimolecular  
308 encounter, as realized by high irradiation intensity of short wavelength, as well as by  
309 allowing for long-range I<sub>i</sub><sup>0</sup> migration. The direct recombination of two interacting I<sub>i</sub><sup>-</sup>/I<sub>i</sub><sup>+</sup>  
310 defects to give I<sub>2</sub> has instead a fairly high activation energy (0.33 eV, when calculated  
311 among surface-adsorbed species, see Supporting Information Figure S13).

312 Sequential electron and hole trapping could alternatively take place at the I<sub>i</sub><sup>+</sup>/I<sub>i</sub><sup>-</sup>  
313 defect pair:



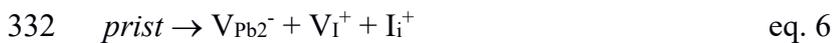
315 The last reaction step leads to the same I<sub>2</sub> product as eq. 3 but its reaction probability is  
316 likely limited by the short living trapped hole at I<sub>i</sub><sup>-</sup>.

317 A filled trap may also take a second electron to form I<sub>i</sub><sup>-</sup>:



319 the bielectronic reaction in eq. 5 is favored by high irradiation intensity. Formation of I<sub>2</sub>  
320 is possibly favored over capture of a second electron, but both mechanisms can coexist.  
321 I<sub>2</sub> loss into toluene solution following MAPbI<sub>3</sub> light irradiation was recently reported<sup>17</sup>,  
322 consistent with eq.3 and 4. I<sub>2</sub> can be trapped as a complex in the bulk (formally I<sub>4</sub><sup>2-</sup>,  
323 Figure 3d) or it may disproportionate to reform the I<sub>i</sub><sup>+</sup>/ I<sub>i</sub><sup>-</sup> pair<sup>37,39</sup>. Importantly, both  
324 coordinated I<sub>2</sub> and the I<sub>i</sub><sup>+</sup>/ I<sub>i</sub><sup>-</sup> pair tend to migrate to the surface, being stabilized by ~0.4  
325 eV. A surface-coordinated I<sub>2</sub> molecule represents the final reaction product (Figure 3e).  
326 Such I<sub>2</sub> molecule can remain bound to the surface or to grain boundaries. Notably I<sub>2</sub> can  
327 act as an electron trap, as I<sub>i</sub><sup>+</sup> does.

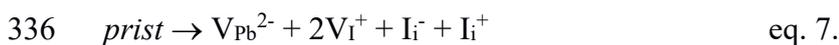
328 The iodine imbalance between surface and bulk can trigger a series of  
329 compensating reactions that regenerate the starting equilibrium distribution of I<sub>i</sub><sup>+</sup>/ I<sub>i</sub><sup>-</sup>  
330 defects to compensate their transformation to surface bound I<sub>2</sub>. This can be  
331 accomplished by disrupting the pristine lattice as



333 or through formation of Frenkel pairs, according to eq.1, here simplified as



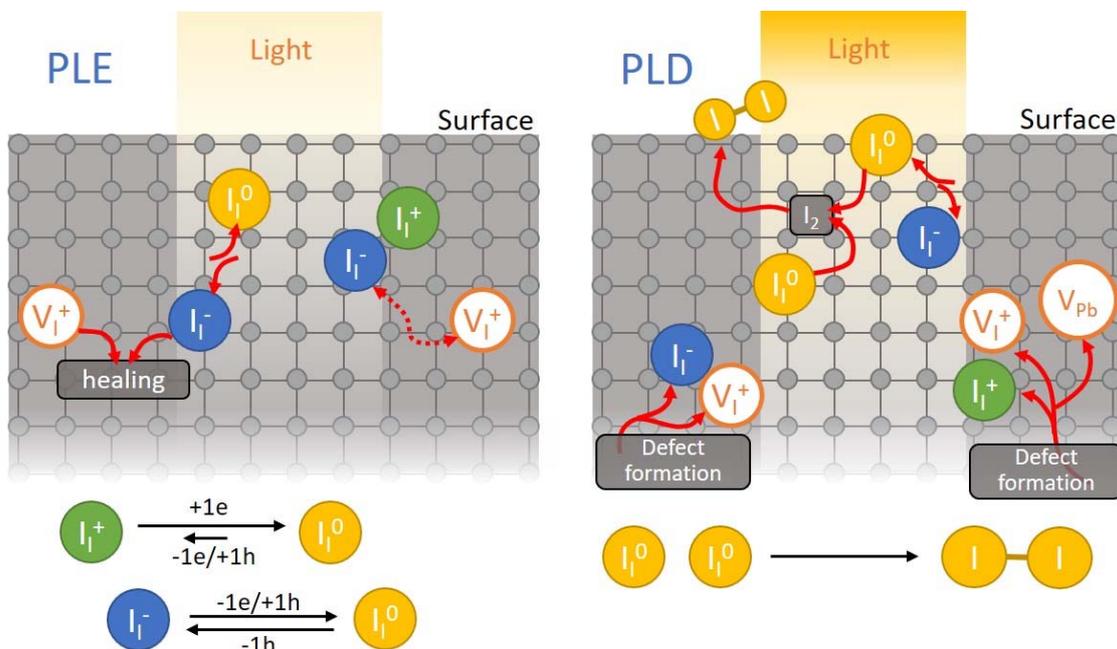
335 or the net reaction:



337 Note that V<sub>Pb</sub><sup>2-</sup> + 2V<sub>I</sub><sup>+</sup> corresponds to the loss of PbI<sub>2</sub>, which may segregate as a separate  
338 phase, with consequent material degradation<sup>40</sup> and associated PLID. Alternatively,  
339 metallic Pb could nucleate on the material surfaces<sup>41</sup>.

340 The trigger to PLID is thus the surface (or grain boundary) stabilization of  
 341 coordinated  $I_2$ . Treating the thin film with appropriate passivating agents and *blocking*  
 342 *undercoordinated surface Pb atoms is thus the key to avoid PLID*<sup>42</sup>. Surface passivation  
 343 will result not only in an improvement of the absolute PL quantum yield of the  
 344 semiconductor thin film (surface-bound  $I_i^-$  is destabilized when the MAPbI<sub>3</sub> surface is  
 345 capped by aliphatic ethers, see Supplementary Information Figure S14) but also in  
 346 hampering the PLID phenomenon, leaving room for PLIE to happen.

347 The key to block PLID and likely to significantly slow down perovskite  
 348 degradation is the passivation of surface sites to which  $I_2$  can bind in its various forms.  
 349 These undercoordinated Pb surface sites, not directly involved in trapping/detrapping  
 350 events, act as a reservoir for  $I_2$ , leading to iodine imbalance in the bulk and stimulating  
 351 the production of additional defects. The process continues until surface sites are  
 352 saturated, after which the material may start releasing  $I_2$  and possibly form a segregated  
 353  $PbI_2$  phase. Scheme 1 summarizes the PLIE and PLID mechanisms.



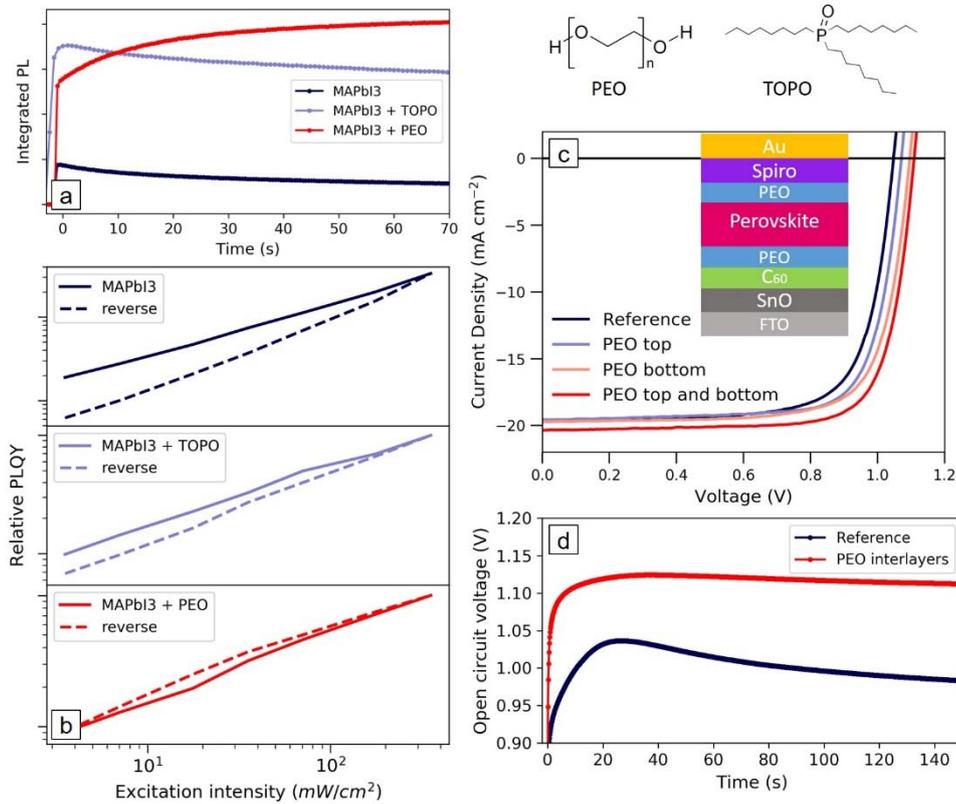
Scheme 1. Ion migration promoting PLID and PLIE in MAPbI<sub>3</sub> thin film.

357 To check the importance of the surface in PLID, we compared a pristine  
358 MAPbI<sub>3</sub> film with one coated with tri-n-octylphosphine oxide (TOPO) or polyethylene  
359 oxide (PEO) (Figure 4a). TOPO has been applied for the passivation of perovskite films  
360 resulting in enhanced PL efficiencies and lifetimes.<sup>7</sup> PEO is a polymer containing polar  
361 ether groups and hydroxyl terminations, representing a moisture barrier and a possible  
362 passivating agent<sup>43,44</sup>. In Figure 4a, by illuminating a fresh spot of the samples (560 nm,  
363 CW excitation, fluence of ~0.1 μJcm<sup>-2</sup>) and monitoring the integrated PL signal over a  
364 few minutes we observed a slow quenching for the bare MAPbI<sub>3</sub> thin film and an initial  
365 enhancement followed by a slower quenching for the TOPO passivated thin film. On  
366 the other hand, in the presence of PEO we observe that the mechanism leading to PLID  
367 is switched off even at room temperature (Figure 4b). Despite TOPO and PEO are both  
368 characterized by the presence of oxygen atoms which interact with undercoordinated  
369 surface Pb atoms, (see Supplementary Information Figure S14) the increased PLIE  
370 observed with PEO is suggestive of a more effective passivation by the latter, probably  
371 because of the larger fraction of oxygen atoms against aliphatic carbon chains in PEO.  
372 Figure 4b shows the relative PL quantum yield, calculated as the integrated PL intensity  
373 normalized by the excitation intensity, of MAPbI<sub>3</sub> films. Each data point was taken after  
374 a hold time of ~30 seconds. When measuring from low to high excitation intensities  
375 (solid lines) and then in reverse order (dashed lines), we observe a hysteretic behavior  
376 resulting from photoinduced trap formation, *i.e.* PLID and a concomitant PL intensity  
377 reduction from bare perovskite thin films. By illuminating the film from the passivated  
378 side, both the TOPO and PEO coated samples show a higher PL signal, indicating a  
379 decrease in the density of native defects on the thin film surface. Thus, also in  
380 agreement with the data shown in Figure 4a, the sample coated with TOPO shows a less  
381 pronounced hysteretic behavior compared to the uncoated film, while the PEO

382 passivated thin film presents an opposite hysteretic behavior of the relative PL as a  
383 function of excitation intensity, as a consequence of a dominant PLIE process. This  
384 clearly shows that efficient surface passivation can reduce the density of defect states  
385 and photo-stabilize the perovskite thin film. To connect our experimental and  
386 computational model, in particular the central role of surface-stabilized I<sub>2</sub> in PLID, we  
387 investigated the I<sub>2</sub>-releasing reaction in hexane solution from MAPbI<sub>3</sub> films considering  
388 three cases: i) the bare thin film; ii) MAPbI<sub>3</sub> capped by an inert polystyrene film; and  
389 iii) a PEO-passivated film. In line with the results of Kim et al.<sup>17</sup> prolonged irradiation  
390 of the bare MAPbI<sub>3</sub> and polystyrene-capped thin films produced the same amount of I<sub>2</sub>,  
391 while a significant reduction in I<sub>2</sub> production was observed for the PEO-passivated thin  
392 film (Supplementary Figures S18-S20).

393         The passivating PEO layer has a significant technological relevance as it allows  
394 for the formation of high quality interlayer thin films which can be easily implemented  
395 in perovskite solar cells. In Figure 4c we show the J-V characteristic of MAPbI<sub>3</sub> based  
396 solar cells. Here we add PEO interlayers between the perovskite active layer and the  
397 selective charge extracting layers. The PEO layer is seen to improve the open circuit  
398 voltage (V<sub>OC</sub>) when applied to either the electron- or hole-extracting interface, with the  
399 best performance given by the device with PEO applied to both sides (see the devices  
400 figure of merit in Table S1 and their statistics in Figure S15 of the SI). In Figure 4.d we  
401 show the evolution of the V<sub>OC</sub> of solar cells with and without the PEO interlayers under  
402 one sun illumination (see Figure S17 of SI for monochromatic photo-excitation). The  
403 devices were encapsulated in inert atmosphere. Despite the fact that the perovskite layer  
404 in the reference device is already interfaced with materials which have, to some extent a  
405 passivating role, such as C<sub>60</sub><sup>45,46</sup>, we still see competing dynamics in the Voc which  
406 mirror the PLID and PLIE shown in Figure 4.a. However, the relative contribution of

407 the PLID is drastically reduced in the PEO-treated device, resulting in a higher and  
 408 more stable Voc.  
 409



410  
 411  
 412 **Figure 4.** (a) PL intensity over time of the MAPbI<sub>3</sub> films bare (blue) and coated with  
 413 TOPO (light blue) or PEO (red), taken on fresh spots (560 nm, CW, 40mW cm<sup>-2</sup>); (b)  
 414 Relative PLQY, taken with increasing (solid lines) or decreasing (dashed lines)  
 415 excitation intensities (CW illumination, 560 nm, hold time between data points  
 416 ~30 seconds) of MAPbI<sub>3</sub> thin films bare (dark blue) and coated with TOPO (light blue)  
 417 or PEO (red); (c) Current density voltage (J-V) curves measured under 1 sun of  
 418 MAPbI<sub>3</sub> devices with and without PEO interlayers; (d) open-circuit voltage (Voc)  
 419 transient measured from solar cells with and without PEO interlayers under one sun  
 420 illumination.  
 421

422 Most notably, the PEO-passivated device shows also significantly longer stability under  
423 light soaking at the maximum power point, retaining more than 90% of its initial  
424 efficiency for 15h; while the bare MAPbI<sub>3</sub> device loses 35% of its initial efficiency  
425 within the first 6h, see Supplementary Figure S21.

426 It is worth to mention that LEDs may operate at a relatively high density of carriers,  
427 thus they may suffer more from the PLID mechanism. As we have demonstrated, PEO  
428 interlayers are capable of passivating the perovskite surfaces minimizing the formation  
429 of degradation products (*i.e.* I<sub>2</sub>) without blocking the charge injection in the device.  
430 Therefore we believe PEO could be beneficial to LED stability too.

431 In conclusion, by monitoring the evolution of PL intensity under systematic  
432 control of the experimental parameters we identified two distinct processes behind the  
433 instabilities observed in lead halide perovskites, excluding effects related to the  
434 atmosphere. We demonstrate that such processes are simultaneously happening within  
435 the thin films under photo-excitation, and that depending on the conditions, one can  
436 overcome the other, conciliating the conflicting reports in literature and providing a  
437 consistent understanding of photoinduced phenomena in this class of materials. Based  
438 on simple yet comprehensive experimental evidence we are able to provide a clear  
439 model, in which long-living trapped carriers – typical of halide photochemistry -  
440 mediate photo-induced ionic dynamics which can lead to a dominant PL enhancement  
441 or decrease. The latter stems from a bimolecular reaction which due to the encounter of  
442 two I<sup>0</sup> species to form I<sub>2</sub>, thus it is favored in the presence of high trap density  
443 concentration and/or high temperature. The stabilization of coordinated I<sub>2</sub> at the surface  
444 (or grain boundary) is a fundamental factor for the efficiency loss. On such grounds, we  
445 show that passivating the thin film surface prevents the detrimental process leading to  
446 PL decrease, favoring the defect healing. This allows us to target the relevant defect

447 sites by passivating the semiconductor surface with PEO interlayers, improving device  
448 performance and stability.

449

#### 450 ACKNOWLEDGEMENTS

451 This work has been funded by the European Union project PERT PV under grant  
452 agreement n° 763977, and ERC project SOPHY under grant agreement N 771528

453 S.G.M. thanks the CNPq (Conselho Nacional de Desenvolvimento Científico e  
454 Tecnológico - Brasil) for the scholarship [206502/2014-1]. The authors thank Dr  
455 Giuseppe Paternò for his support to set-up the transient Voc characterization.

456

#### 457 AUTHOR CONTRIBUTIONS

458 S.G.M. performed the PL measurements, D.M. and E.M. performed the first-principle  
459 calculations, A.R.P., J.M.B, M.G. and M.K. were responsible for the thin films  
460 fabrication, M.K. fabricated the solar cell devices and M.K and A.J.B characterized the  
461 solar cell. A.P., S.M, A.J.B., D.M. and F.D.A. analyzed the data. S.G.M, F.D.A. and  
462 A.P. wrote the first draft of the manuscript and all authors contributed with the  
463 discussions and finalized the manuscript. AP supervised the project.

464

#### 465 COMPETING INTERESTS

466 The authors declare no competing interests.

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647 SUPPLEMENTARY Info

648

## 649 Controlling competing photochemical reactions 650 stabilises perovskite solar cells

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## 672 Supplementary Information

673

### 674 **Sample preparation**

675 Lead(II) bromide ( $\text{PbBr}_2$ ,  $\geq 98\%$ ), N,N-dimethylformamide (DMF, anhydrous,  
676 99.8%), Chlorobenzene (anhydrous, 99.8%), and dimethyl sulfoxide (DMSO, anhydrous,  
677  $\geq 99.9\%$ ) were purchased from Sigma-Aldrich; methylammonium bromide (MABr) and  
678 methylammonium iodide (MAI) were purchased from Dyesol; and lead (II) iodide ( $\text{PbI}_2$ ,  
679 99.9985%, CAS No. 10101-63-0) was purchased from Alfa Aesar. All chemicals were  
680 used without any further purification. Glass substrates were cleaned in acetone and  
681 isopropyl alcohol (IPA) for 10 minutes by sonication. The cleaned glass substrates were  
682 treated with Oxygen plasma for 10 minutes before any further deposition.

683

684 *MAPbBr<sub>3</sub> thin films.* These films could be fabricated by an adapted Nanocrystal-  
685 Pinning technique [Ref. Cho H. et al., Science (2015), 350-6265, 1222-5]. In this case,  
686 two steps of spin-coating speed were used (500 rpm for 7 seconds, 3000 rpm 90 seconds).  
687 After spin-speed acceleration, a solution of MABr and  $\text{PbBr}_2$  (molar ratio 1.05:1) in  
688 DMSO was spin-coated onto the clean glass substrate. After 60 seconds, the pinning  
689 occurred by dropping 300  $\mu\text{l}$  of chlorobenzene on the spinning sample. The samples were  
690 then baked at 90°C for 10 minutes.

691

692 *MAPbI<sub>3</sub> thin films.* These films were fabricated by quenching a precursor solution  
693 with an antisolvent during spin coating [M. Xiao et al., Angewandte Chemie vol. 126,  
694 p10056 (2014)], in a nitrogen filled glovebox. A 1.45 M precursor solution of  
695  $\text{PbI}_2$ :MAI:DMSO in a molar ratio of 1:1:1 was prepared in DMF. This solution was spin  
696 coated onto the glass substrate at 4000 rpm, with an acceleration of 4000 rpm/s, for 15 s.  
697 After 6 s toluene, an antisolvent to the precursor solution, was dropped onto spinning  
698 sample by pipette. The samples were then annealed at 100°C for 10 minutes.

699

700 *MAPbI<sub>3</sub> thin films with passivation.*

701 tri-n-octylphosphine oxide (TOPO) or polyethylene oxide (PEO) were dissolved in  
702 anhydrous chlorobenzene at concentration of 10  $\text{mg}\cdot\text{ml}^{-1}$ . Solutions were sonicated for  
703 10 minutes for complete dissolution before deposition on top of the perovskite film by  
704 spin-coating at 4000 rpm.

705        **Device fabrication**

706        *Film fabrication:* ITO on glass substrate was etched with 2M aqueous HCl solution  
707 and zinc powder (Sigma Aldrich). The ITO substrates were sequentially washed with 2%  
708 Hellamanex in deionized water, deionized water, iso-propanol, acetone, iso-propanol in  
709 a sonication bath, followed by O<sub>2</sub> plasma cleaning for 10 min. A cleaned ITO substrate  
710 was covered with a SnO<sub>2</sub> layer (~20 nm) by spin-coating of a diluted SnO<sub>2</sub> nanoparticle  
711 solution (Alfa Aesar) and annealed at 180 °C for 1 h. On the SnO<sub>2</sub> layer, C<sub>60</sub> layer was  
712 deposited from a C<sub>60</sub> solution (10 mg·ml<sup>-1</sup>) in 1,2-dichlorobenzene by spin-coating at  
713 3000 rpm, followed by thermal annealing at 100 °C for 10 min. The perovskite layer was  
714 fabricated through applying anti-solvent quenching method. For hole transport layer, a  
715 spiro-MeOTAD solution was spin coated on the perovskite layer at 4000 rpm for 30 s.  
716 Spiro-MeOTAD solution was prepared by dissolving 73 mg of spiro-MeOTAD in 1 mL  
717 chlorobenzene (99.8%; Sigma–Aldrich), to which were added 28.8 μL of 4-tert-  
718 butylpyridine (96%, Sigma-Aldrich), 17.5 μL lithium bis(trifluoromethanesulfonyl)imide  
719 (LiTFSI) solution (520 mg LiTFSI in 1 mL acetonitrile, 99.8%, Sigma-Aldrich). This  
720 fabrication process was carried out under controlled conditions in a glove-box  
721 atmosphere. Finally, 75 nm gold was thermally evaporated on top of the device at a  
722 pressure of  $1 \times 10^{-6}$  mbar to form the top-contact electrode.

723        **Photovoltaic characterization:**

724        All devices were measured under simulated AM1.5 solar illumination using a class  
725 AAA solar simulator (Oriel Sol3A, Newport). The illumination intensity was calibrated  
726 using an unfiltered certified Si reference diode (area = 4 cm<sup>2</sup>, Newport) to be 100 mW/cm<sup>2</sup>  
727 (typical spectral mismatch factor of 1.01 for solar cells based on CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>). The  
728 current density-voltage (J-V) characteristics were recorded with a Keithley 2440. No light  
729 soaking or pre-biasing was applied before the J-V measurements. The illuminated  
730 electrode area, defined with a holed black anodized aluminium mask, was 0.0935 cm<sup>2</sup>.  
731 Devices were measured in the ambient atmosphere at 23 ± 2°C and 40-60% relative  
732 humidity. The scan rates of J-V sweep were 0.2 V s<sup>-1</sup>. The forward scan started from 0 V  
733 (the short circuit condition) to 1.4 V, while backward scan from 1.4 V to 0 V. The stability  
734 of the current/power output with time was recorded for 120 s. The EQE was measured  
735 with a home-built setup. EQE spectra were recorded using the monochromated (Bentham)  
736 output from a tungsten halogen lamp calibrated with a Newport UV-818 photodiode.

737

738 **Transient open circuit voltage measurements**

739

740 The excitation source was a simulated AM1.5 solar illumination using a class AAA  
741 solar simulator (Oriel Sol3A, Newport) or an unfocused beam of a 450 nm CW diode  
742 laser (Oxxius). The beam was expanded to match the active area of device,  $\sim 0.1 \text{ cm}^2$ , and  
743 the excitation power density were around  $100 \text{ mW cm}^{-2}$ . The devices were encapsulated  
744 using epoxy and glass, and then were mounted perpendicular to the excitation. Keithley  
745 2401 was used to measure the transient open circuit voltage over a duration of time,  $\sim 1500$   
746 s.

747

748

749

750 **Modulation frequency dependent Photoluminescence**

751 Excitation was provided with a continuous wave (CW) diode laser (Oxxius  
752 laserboxx). The excitation wavelength was 405 nm for MAPbBr<sub>3</sub> films and 450 nm for  
753 MAPbI<sub>3</sub>. Modulation was accomplished by a waveform generator connected directly to  
754 the laser. Pulse widths were kept fixed to exclude the effects of varying illumination times  
755 so we could probe only the dependence on the intervals in the dark. MAPbI<sub>3</sub> films were  
756 illuminated with pulses of 200 ns at frequencies of 1 kHz, 3 kHz, 10 kHz, and 30 kHz.  
757 MAPbBr<sub>3</sub> films were illuminated with pulses of 10  $\mu\text{s}$  and frequencies of 50 Hz, 500 Hz  
758 and 5 kHz. Excitation fluences used were around  $0.5 \mu\text{Jcm}^{-2}$ , corresponding to an  
759 excitation density of  $\sim 10^{16} \text{ cm}^{-3}$ . Samples were mounted inside a vacuum chamber  
760 mounted on a translation stage that allowed fresh spots to be probed between every curve  
761 by changing the sample height without damaging the collection alignment. PL was  
762 collected in reflection mode and focused into a fiber coupled to a spectrometer (Ocean  
763 Optics Maya Pro 2000).

764

765 **Temperature dependent Photoluminescence**

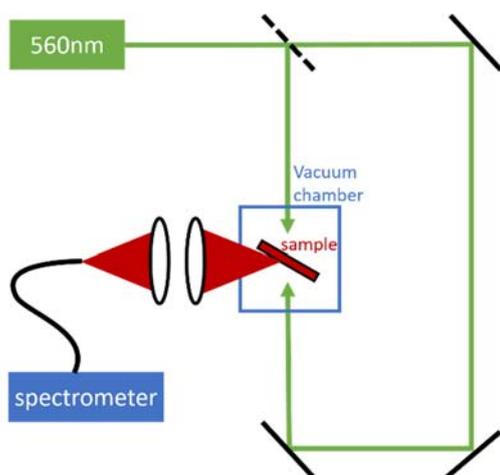
766 Excitation was provided with a CW diode laser (Oxxius laserboxx). The excitation  
767 wavelength was 405 nm for MAPbBr<sub>3</sub> films and 450 nm for MAPbI<sub>3</sub>. Excitation  
768 intensity was around  $50 \text{ mW/cm}^2$ , corresponding to an estimated excitation density  
769  $\sim 10^{16} \text{ cm}^{-3}$ . Films were mounted in vacuum in a cold finger cryostat cooled with liquid  
770 N<sub>2</sub> (Oxford Instruments) perpendicular to the excitation line. The excitation beam was

771 directed to the sample by a dichroic mirror and PL was collected in reflection mode and  
772 focused into a fiber coupled to a spectrometer (Ocean Optics Maya Pro 2000).

773

### 774 **Double sided illumination**

775 The beam of a 560 nm CW diode laser (Oxxius) was expanded and collimated with a  
776 telescope to a diameter of ~3 mm. A beam splitter was used to obtain the two excitation  
777 lines and attenuators were used to balance the intensities. Samples were mounted inside  
778 a vacuum chamber (pressure below  $10^{-5}$  mbar) and PL was collected at a right angle from  
779 the excitation line (Figure S1) and focused into a fiber coupled to a spectrometer (Ocean  
780 Optics Maya Pro 2000).



781

782 Figure S1. Schematics of the experimental setup used for double side illumination on  
783 thin films.

784

### 785 **Photoluminescence on passivated samples**

786

787 The excitation source was an unfocused beam of a 560 nm CW diode laser (Oxxius).  
788 Samples were mounted inside a vacuum chamber (pressure below  $10^{-5}$  mbar) and PL was  
789 collected in reflection mode at a right angle from the excitation line and focused into a  
790 fiber coupled to a spectrometer (Ocean Optics Maya Pro 2000). Integrated PL over time  
791 was measured on fresh spots of the film with intensity  $\sim 40$  mW/cm<sup>2</sup>.

792 For relative PLQY measurements, the integrated PL was measured at varying  
793 excitation intensities and plotted as:

794

$$\text{Relative PLQY} = \frac{I_{PL}}{I_{pump}}$$

795

796 **Excitation wavelength dependence**

797 The excitation source was a supercontinuum laser (SuperK Extreme, NKT Photonics),  
798 which is spectrally filtered by an acousto-optic modulator (SuperK Select, NKT  
799 Photonics). The selected output is directed to a home-built transmission microscope in a  
800 confocal configuration. The sample is placed on top of a piezoelectric translation stage.  
801 The PL transients (Figure 2a) were taken tuning the wavelength of the acousto-optic  
802 modulator output and monitoring the PL over time with a spectrometer (Ocean Optics  
803 Maya Pro 2000) coupled to the microscope for detection.

804

805 **Simulation of PL variations over time**

806

807 A system of rate equations based on the model proposed by Stranks et al.<sup>1</sup> was used  
808 to simulate the evolution of the populations of free electrons, free holes and trapped  
809 carriers (Figure S2-S4). The model considers one single type of trap, assuming that in the  
810 presence of both hole and electron traps the more abundant type dominates the dynamics.  
811 In our case we consider the electron traps as dominant leading to the following system of  
812 equations:

813

814 
$$\frac{dn_e}{dt} = G - \beta_{rad}n_en_h - k_t n_e(N_t - n_t) - \gamma n_e n_h^2$$

815 
$$\frac{dn_h}{dt} = G - \beta_{rad}n_en_h - R_t n_h n_t - \gamma n_e n_h^2$$

816 
$$\frac{dn_t}{dt} = k_t n_e(N_t - n_t) - R_t n_h n_t$$

817

818 where  $n_e$  is the density of free electrons,  $n_h$  is the density of free holes,  $n_t$  is the  
819 density of trapped carriers,  $G$  is the photogeneration rate,  $\beta_{rad}$  is the radiative  
820 recombination rate,  $\gamma$  is the Auger recombination rate,  $k_t$  is the trapping rate,  $N_t$  is the  
821 density of available traps, and  $R_t$  is the rate of recombination of a trapped electrons with  
822 free holes.

823 The PL intensity at a given time is

824

825 
$$I_{PL} = \beta_{rad}n_en_h$$

826

827 The parameters used in our simulations for modulated illumination were:

$$\begin{aligned} \beta_{rad} &= 5 \cdot 10^{-10} \text{ cm}^3/\text{s} \\ k_t &= 1 \cdot 10^{-10} \text{ cm}^3/\text{s} \\ R_t &= 1 \cdot 10^{-12} \text{ cm}^3/\text{s} \\ \gamma &= 1 \cdot 10^{-26} \text{ cm}^3/\text{s} \\ N_t &= 1 \cdot 10^{16} \text{ cm}^{-3} \end{aligned}$$

834 These values were chosen based on fits of experimental data, including the low rate  
835 for trapped carrier recombination, which is consistent with a fit of hole dynamics (Figure  
836 S2) to our previous results from transient absorption measurements<sup>2,3</sup>.

837 We obtain the evolution of the carrier populations in time switching the generation  
838 rate on and off to simulate the modulated illumination. The PL intensity can change over  
839 time because of pulse piling up and trap filling effects, but in the conditions similar to the  
840 experiments performed in this work, these effects should reach a steady state condition  
841 on a scale shorter than milliseconds. Several variations of the parameters above were  
842 tested in order to rule out that longer timescales could result from different combinations  
843 of values, but in all the possible scenarios we considered, stable PL is established after no  
844 more than a second. In the section below we discuss the results of these simulations in  
845 the interpretation of our experimental observations.

### 847 **Role of electronic effects in PL dynamics in Fig 1 of the Manuscript**

848 We can exclude the role of pure electronic processes, i.e. trap filling effects, on the  
849 observed PL dynamics in the milliseconds/seconds time domain. This can be visualized  
850 by the simulations reported in Figure S22-25. The simulations have been performed as  
851 described in the section “Simulation of PL variations over time” above. Here the  
852 simulations are run by testing up to 500 pulses coming at different frequencies between  
853 1 kHz and 30 kHz. For each defined frequency we plot the electron and hole dynamics,  
854 the trap sites filling and the corresponding PL dynamics. It is clear that even when trap  
855 state filling start piling up, above 10 kHz, the PL reach a steady state condition well before  
856 the ms-s time regime, where we probe the PL quenching and enhancement transients.  
857 Moreover, it is worth noticing that increasing the pulse frequency mainly leads to pile up  
858 effects, i.e. growing population upon trap filling and this should lead to an enhancement  
859 of the PL, while we observe the PL quenching kicking in.

860 The main effect correlated to an enhancement of the average flux density is an  
861 enhancement of the PL absolute value at time  $t=0$ , due to an associated enhancement of  
862 counts during the integration (see Figure S26).

Frequency	$\delta$ (time between pulses)	Pulse energy	Average Intensity
1 kHz	1 ms	0.2 nJ	0.6 mW/cm <sup>2</sup>
3 kHz	0.33 ms	0.2 nJ	2 mW/cm <sup>2</sup>
10 kHz	0.1 ms	0.2 nJ	6 mW/cm <sup>2</sup>
30 kHz	0.033 ms	0.2 nJ	20 mW/cm <sup>2</sup>

863 Table S2. Average intensity as a function of the change in the excitation repetition  
864 rate.

865

866

### 867 **Computational details**

868 All calculation have been carried out with Quantum Espresso<sup>4</sup> program package.

#### 869 Defects formation energies and ionization levels calculations:

870 DFT calculations have been performed in the tetragonal phases of MAPbI<sub>3</sub> and  
871 MAPbBr<sub>3</sub>. In all cases the cell parameters have been fixed to the experimental values, i.e.  
872  $a=b=8.849$  Ang,  $c = 12.642$  Å for MAPbI<sub>3</sub><sup>5</sup>;  $a=b=8.345$ ,  $c = 11.802$  Å for MAPbBr<sub>3</sub>.

873 Defects structures have been calculated in the 2x2x1 supercells for both MAPbI<sub>3</sub> and  
874 MAPbBr<sub>3</sub> by using the PBE functional<sup>6</sup> and ultrasoft pseudopotentials with a cutoff on  
875 the wavefunctions of 40 Ryd (320 Ryd on the charge density) and 1x1x2 k-point grids in  
876 the Brillouin zone (BZ). Defects calculations have been thus refined at the hybrid level  
877 by using the HSE06 functional<sup>7</sup> ( $\alpha=0.43$ ) by including spin-orbit corrections and  
878 dispersions interactions a posteriori within the DFT-D3 scheme.<sup>8</sup> Single point hybrid  
879 calculations have been performed at the calculated PBE structures by using norm  
880 conserving pseudopotentials and a cutoff energy on the wavefunctions of 40 Ryd and  
881 1x1x2 k-points in the BZ.

882 Defects formation energies and thermodynamic ionization levels have been calculated  
883 following the approach reported in Ref.<sup>9</sup> Chemical potentials have been set by imposing  
884 thermodynamic equilibrium between perovskites and the relative lead precursors, i.e.

$$885 \quad \mu(\text{MAPbI}_3) = \mu(\text{MA}) + \mu(\text{Pb}) + 3\mu(\text{I}), \quad \mu(\text{Pb}) + 2\mu(\text{I}) = \mu(\text{PbI}_2)$$

$$886 \quad \mu(\text{MAPbBr}_3) = \mu(\text{MA}) + \mu(\text{Pb}) + 3\mu(\text{Br}), \quad \mu(\text{Pb}) + 2\mu(\text{Br}) = \mu(\text{PbBr}_2)$$

887 Halides medium conditions have been modelled by using intermediate chemical  
888 potentials between halide-rich and halide-poor conditions. For halide rich conditions I  
889 and Br chemical potentials have been fixed to the values of the respective gas molecules,  
890 i.e.  $\mu(\text{I}) = \frac{1}{2} \mu(\text{I}_2^{\text{gas}})$  and  $\mu(\text{Br}) = \frac{1}{2} \mu(\text{Br}_2^{\text{gas}})$ , while in halide poor conditions the chemical  
891 potentials of lead has been set to the metallic bulk Pb. Defects formation energies have  
892 been corrected by including potential alignment and Makov-Payne corrections<sup>10</sup> (ionic  
893 dielectric constants  $\epsilon=24.0$  and  $\epsilon=20.0$  for MAPbI<sub>3</sub> and MAPbBr<sub>3</sub>, respectively).

#### 894 2x2x2 Supercell and slabs calculations

895 Geometry optimizations are carried out at  $\Gamma$  point using experimental cell parameters  
896 along with plane-wave basis set cutoffs for the smooth part of the wave functions and  
897 augmented electronic density expansions of 25 and 200Ry, respectively. Electrons-ions  
898 interactions were described by ultrasoft pseudo-potentials with electrons from I 5s, 5p;  
899 N, C 2s, 2p; O 2s 2p; H 1s; Pb, 6s, 6p, 5d; shells explicitly included in calculations. For  
900 the slab simulation we take into account both the 001 MAI- and PbI<sub>2</sub>-terminated surface  
901 with a thickness of three and five Pb layers, respectively. A 10 Å of vacuum was added  
902 along the c non periodic direction.

#### 903 Car-Parrinello Molecular Dynamics

904 Car-Parrinello Molecular Dynamics<sup>11,12</sup> have been carried out using a 2x2x2 supercell  
905 MAPbI<sub>3</sub> model generated by duplicating the 1x1x1 optimized structure using the  
906 experimental cell parameters. Electrons-ions interactions were described by ultrasoft  
907 pseudo-potentials with electrons from I 5s, 5p; N, C 2s, 2p; H 1s; Pb 6s, 6p, 5d; shells  
908 explicitly included in calculations. Plane-wave basis set cutoffs for the smooth part of the  
909 wave functions and augmented electronic density expansions of 25 and 200Ry,  
910 respectively. A time step of 10 a. u., electronic mass of 1000 a.u. and masses equal to 5.0  
911 a.m.u have been used for all atoms. The simulation has been carried out at 400 K for a  
912 simulation time of 14 ps comprising the initial thermalization.

913

## 914 **I<sub>2</sub> production under illumination**

915 We have identified hexane as a good solvent for I<sub>2</sub> and poor solvent for PEO, allowing  
916 us to investigate I<sub>2</sub> formation free of other solvent induced degradation processes.

917 We prepared three thin films, a pristine MAPbI<sub>3</sub>, and two twin samples coated by  
918 PEO and polystyrene (PS), respectively. The use of PS serves as blank, as this system  
919 covers the perovskite surface without significant chemical interaction (Energy Environ.  
920 Sci., 2018, 11, 2609-2619) thus representing a simple capping agent.

921 In figure S18 we show the I<sub>2</sub> absorption spectrum in hexane, characterized by a broad  
922 band peaking at about 520 nm.<sup>14</sup>

923 We kept the uncoated, PS- and PEO-coated samples under illumination for 36h in a  
924 vial full of hexane and during this time window we monitored the absorption spectrum of  
925 the ensuing solution, Figure R4a-c. In line with results by Kim et al.<sup>15</sup>, we see that under  
926 photo-excitation I<sub>2</sub> molecules reach the perovskite surface and are then dissolved in the  
927 solvent, as demonstrated by the growing 520nm band over time. Importantly, in the  
928 presence of the inert PS coating, the concentration of I<sub>2</sub> in the hexane solution is the same  
929 as the one produced by the uncoated thin film. On the other hand, the presence of PEO  
930 strongly reduces the I<sub>2</sub> production. This can be clearly visualized Figure S19 where we  
931 show the integrated area of the band at 520nm versus time. Notably, leaving the three  
932 twin samples in contact with a hexane solution in the dark, as a control experiment, no  
933 trace of I<sub>2</sub> could be probed for either uncoated, PS- and PEO-coated samples (Fig S19).

934 This set of observations represent an experimental evidence of our model,  
935 demonstrating that chemically active surface passivation, not simple coating, provides a  
936 mean to reduce I<sub>2</sub> loss from the perovskite thin film under photo-excitation, pointing at  
937 the surface as the preferential site for defect nucleation and formation.

938 It is also worth to notice that a small amount of I<sub>2</sub> is observed also in the PEO-coated  
939 sample, which we tentatively ascribe to some degradation of the thin film coating over  
940 time, as we observe through AFM analysis of the surface (Figure S20)

941

## 942 **PL temporal evolution in T and structural phase transition**

943 To further exclude any direct dependence from the crystalline phase, in Figure S27,  
944 we show the PL spectra of MAPbI<sub>3</sub>, over time, at certain temperatures where the thin film  
945 is undergoing the phase transition, therefore one can observe at the same time the  
946 emission from the orthorhombic and tetragonal phases.<sup>16,17</sup> Both peaks showed a slight  
947 enhancement, in a similar way, demonstrating that the phenomenon is not related to the  
948 specific crystalline phases but rather to the specific thermal energy available.

949 On the computational side, DFT simulations of defects in MAPbI<sub>3</sub> and MAPbBr<sub>3</sub> have  
950 been carried out in the related tetragonal cells in order to obtain comparable defect  
951 formation energies and thermodynamic ionization levels of the two systems, by avoiding  
952 possible variation associated to the different band edges estimate of the tetragonal and  
953 cubic phases. Such variations may affect the calculation of defects formation energies and  
954 ionization levels.

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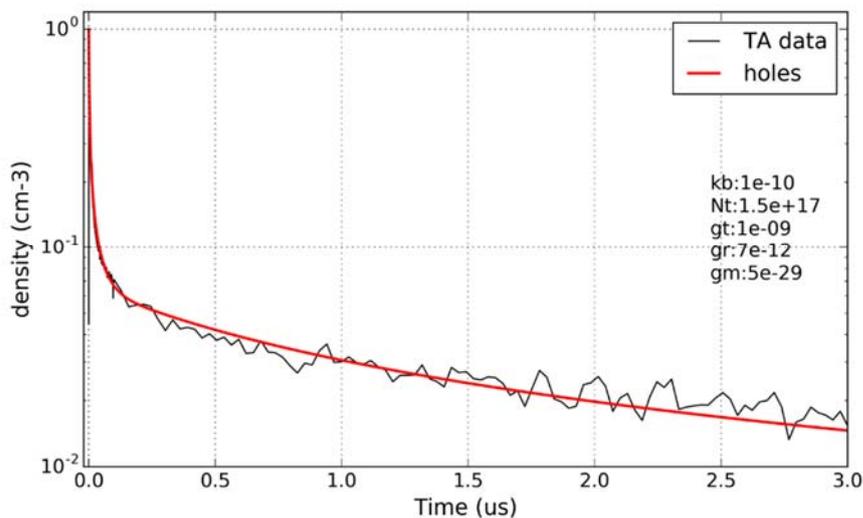
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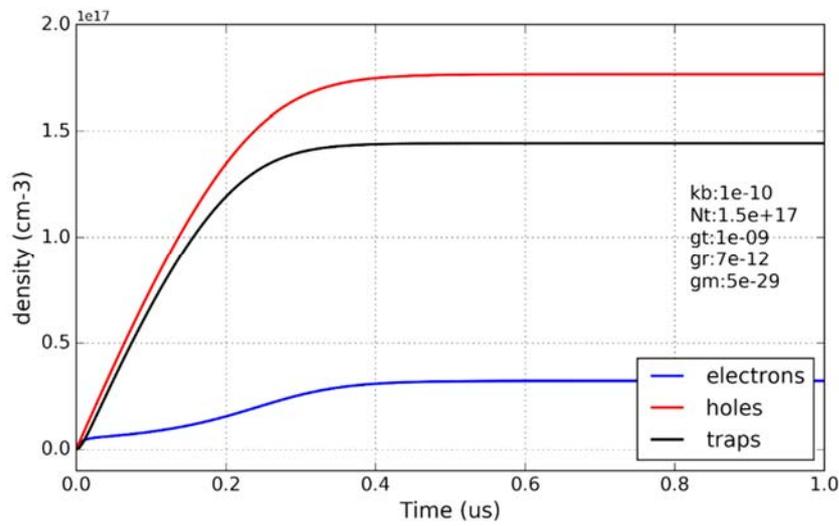
## 963 **Supplementary figures**

964

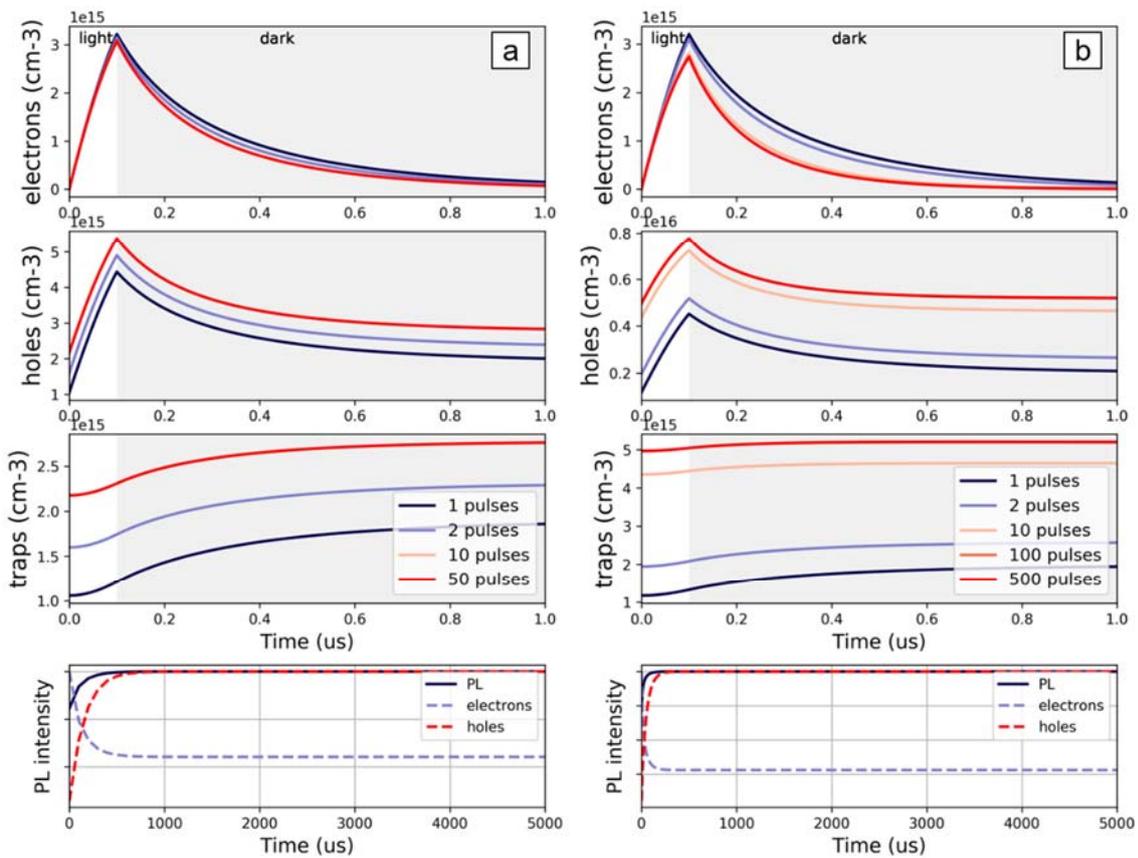


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966 Figure S2. Transient Absorption at the band edge of a MAPbBr3 film (525nm) and a fit  
 967 of the hole dynamics resulting from the rate equation model described.

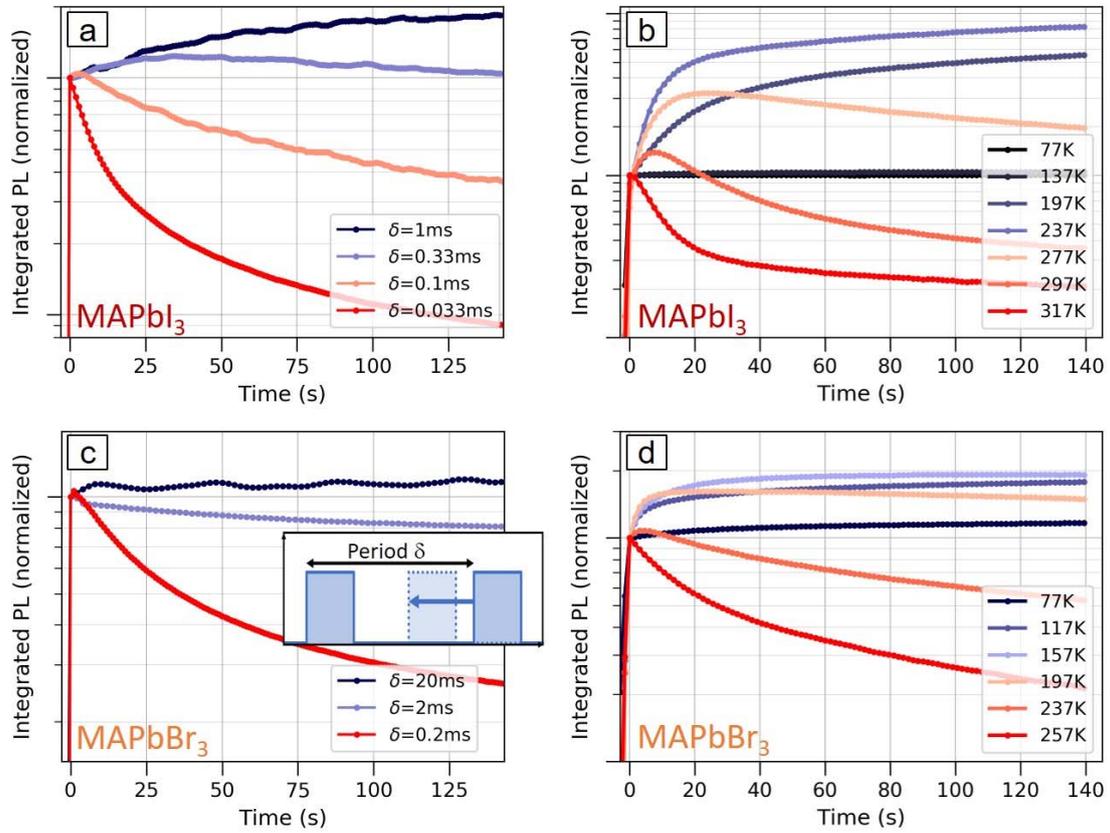


968  
 969 Figure S3. Simulation of the evolution of carrier populations in time under CW  
 970 illumination until steady state conditions are reached.

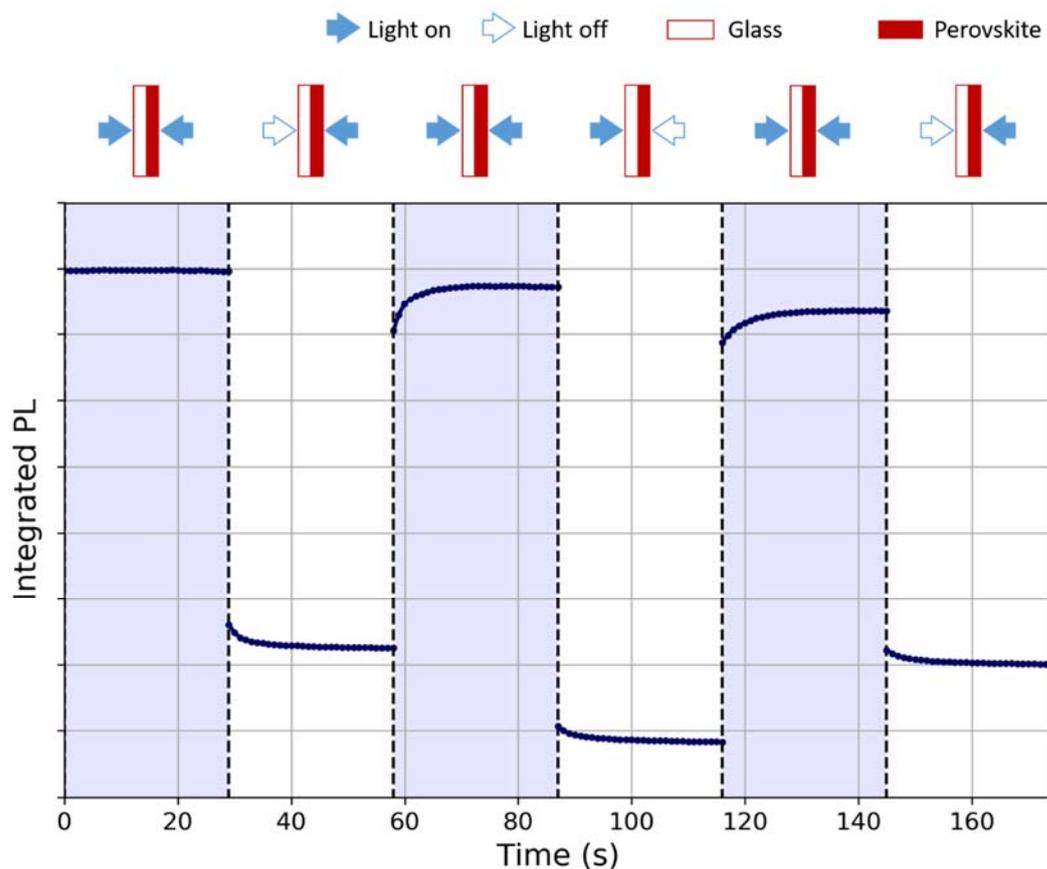


971

972 Figure S4. Simulation of the modulated illumination with pulses of 100ns and frequencies  
 973 of (a) 10 kHz and (b) 100 kHz.  
 974



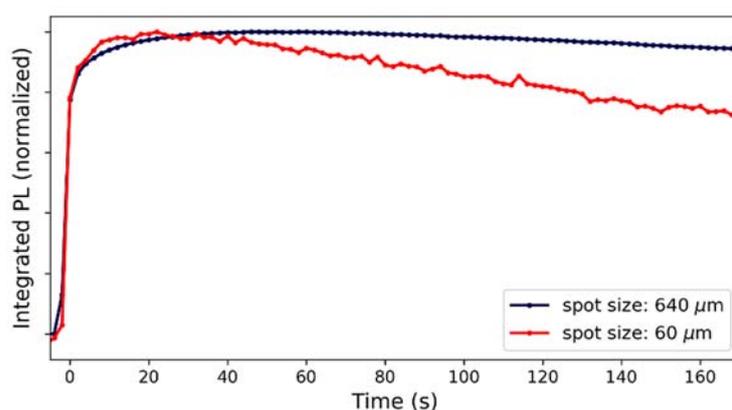
975  
 976 Figure S5. Reproduction of Fig 1 of the manuscript with the pulse width for MAPbBr<sub>3</sub>  
 977 in c) fixed at 10  $\mu$ s.



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979 Figure S6. Integrated PL intensity over time of a MAPbI<sub>3</sub> film on glass. Excitation is  
 980 performed with symmetrical beams of equal intensity (CW illumination, 560 nm)  
 981 impinging on both the perovskite and glass side of the sample.

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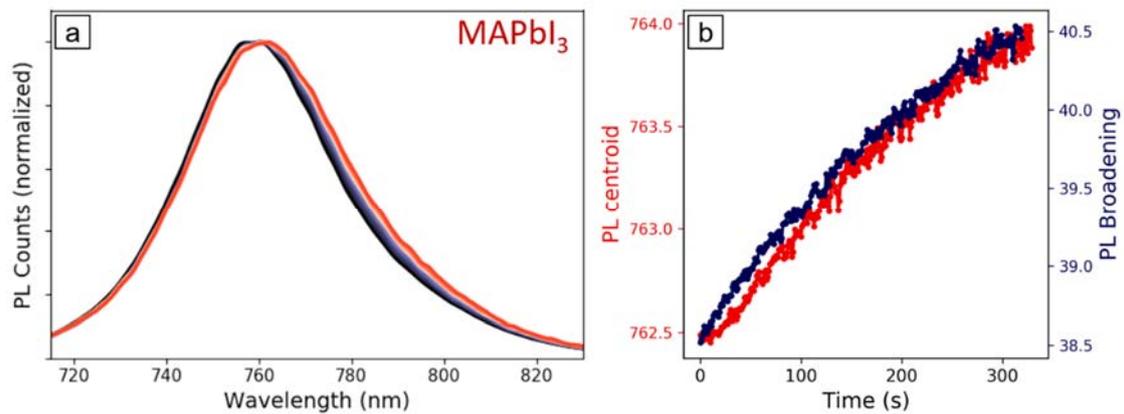


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984 Figure S7. Normalized integrated PL over time of a MAPbBr<sub>3</sub> film obtained using a  
 985 450 nm CW diode laser focalized into 640 μm or 60 μm spot size, with fixed intensity of  
 986 21 mW/cm<sup>2</sup>.

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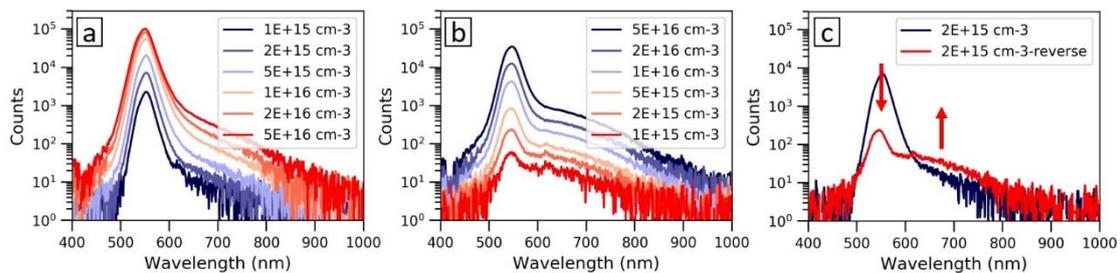
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990 Figure S8. a) PL spectrum of a MAPbI<sub>3</sub> film in PLD conditions (100 kHz, 200 ns pulses),  
991 from t=0 s (dark blue) to t=300 s (red) under illumination, and b) PL broadening and shift  
992 of center of mass over time. When we monitor the spectrum of the emission during the  
993 PL quenching process, we see a small but consistent transformation to broader linewidths,  
994 consistent with a heating scenario. The PL position also slightly red shifts, which is the  
995 opposite behavior from the usual temperature dependence on this material.<sup>13</sup> We suggest  
996 these spectral changes result from the increase of traps in the film, increasing the disorder  
997 in the lattice and density of subgap states.

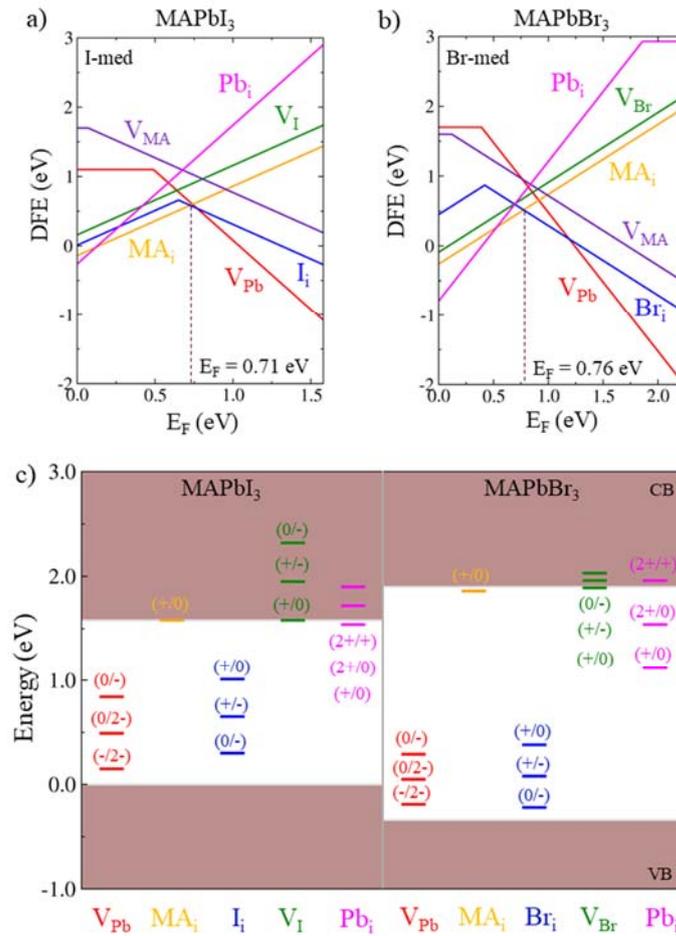
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999

1000 Figure S9. PL spectra of a MAPbBr<sub>3</sub> thin film with a) increasing and b) decreasing  
1001 excitation intensity, and c) comparison of the PL spectra at a single excitation intensity  
1002 before and after the intensity cycle.

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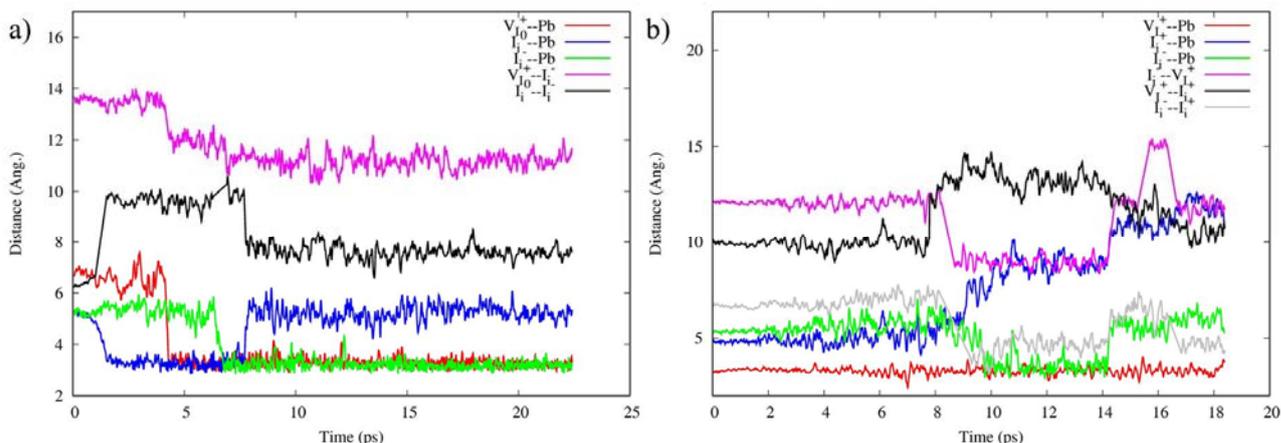


1004 Figure S10. Defects formation energies of a) MAPbI<sub>3</sub> and b) MAPbBr<sub>3</sub> calculated in  
 1005 halide medium conditions by using the HSE06 functional ( $\alpha=0.43$ ) by including spin orbit  
 1006 coupling and a posteriori dispersion corrections; c) Associated thermodynamic ionization  
 1007 levels.

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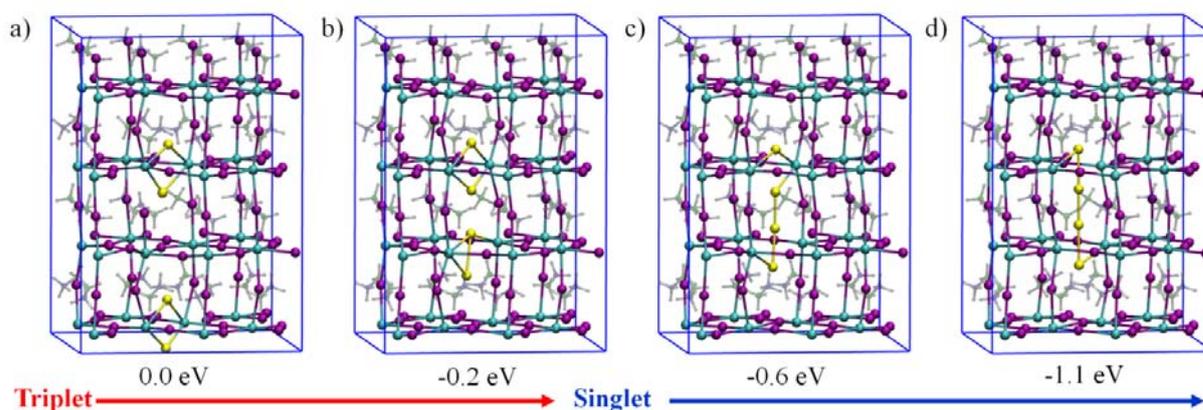
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1012 Figure S11. (a) Dynamical evolution of the system after electron trapping at  $I_i^+$ . A  
 1013 sequence of ion reorganization and migration events can be visualized: 1) following  
 1014 electron trapping at  $I_i^+$ , the system rapidly reaches the minimum energy structure of  $I_i^0$   
 1015 ( $I_i^0$  /  $I_i^-$  distance (black curve) increasing the  $I_i^0$  /  $I_i^-$  distance (black curve); 2)  $V_i^+$  migrates in the direction  
 1016 of  $I_i^-$  (red and magenta lines); 3)  $I_i^-$  further migrates in the direction of  $V_i^+$  (green and  
 1017 magenta lines). (b) Dynamical evolution of the system  $I_i^+ / I_i^-$  in presence of a  $V_i^+$ . A  
 1018 sequence of ion reorganization and migration events can be visualized. In particular the  
 1019 migration of  $I_i^+$  (blue) appears simultaneously with the migration of  $I_i^-$  (green), while the  
 1020  $V_i^+$  remain in its original position (red). The  $I_i^+ / I_i^-$  migrates coupled and the defects  
 1021 separation is not found (gray) and it does not move towards to the  $V_i^+$  (black).

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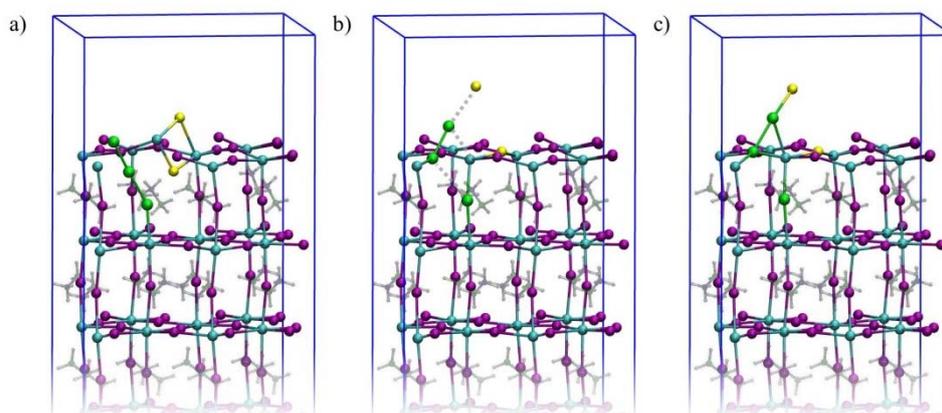


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1024 Figure S12. Radical quenching reaction of two  $I_i^0$  species with associated energetics. The  
 1025 initial stage features a global triplet state corresponding to the two non interacting  
 1026 radicals. Upon shortening the inter-radical distance, a singlet state is stabilized by the

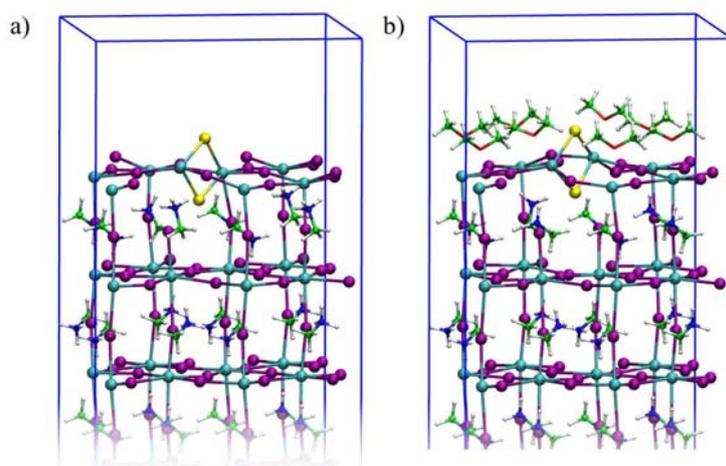
1027 electron pairing to form the I-I bond. The system spontaneously evolves towards the  
1028 formation of a coordinated I<sub>2</sub> molecule.

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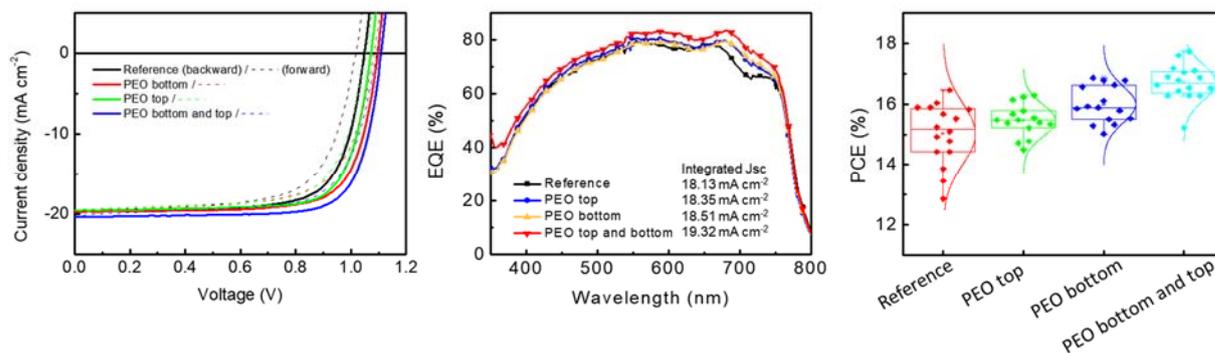
1031 Figure S13. Direct I<sub>i</sub><sup>-</sup>/I<sub>i</sub><sup>+</sup> (a) recombination to move through the transition state (b) towards  
1032 to the final I<sub>2</sub> product (c). This is calculated on the (001) PbI<sub>2</sub>-terminated surface with I<sub>2</sub>  
1033 adsorbed (c). The transition state is calculated 0.33 eV above the reagent.



1034

1035 Figure S14. I<sub>i</sub><sup>0</sup> defect at the PbI<sub>2</sub>-terminated surface bare (a) and with di-methyl ether (b).  
1036 The formation energy of the I<sub>i</sub><sup>0</sup> defect in the presence of the di-methyl ether capping layer  
1037 is less favorable by 0.1 eV with respect the bare surface.

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1039  
 1040 Figure S15. (a) Forward scan (dash lines) and backward scan (solid lines) J-V curves, (b) External  
 1041 quantum efficiency (EQE) spectra, and (c) Performance parameters (from 16 cells) of the  
 1042 perovskite devices with and without PEO passivation.

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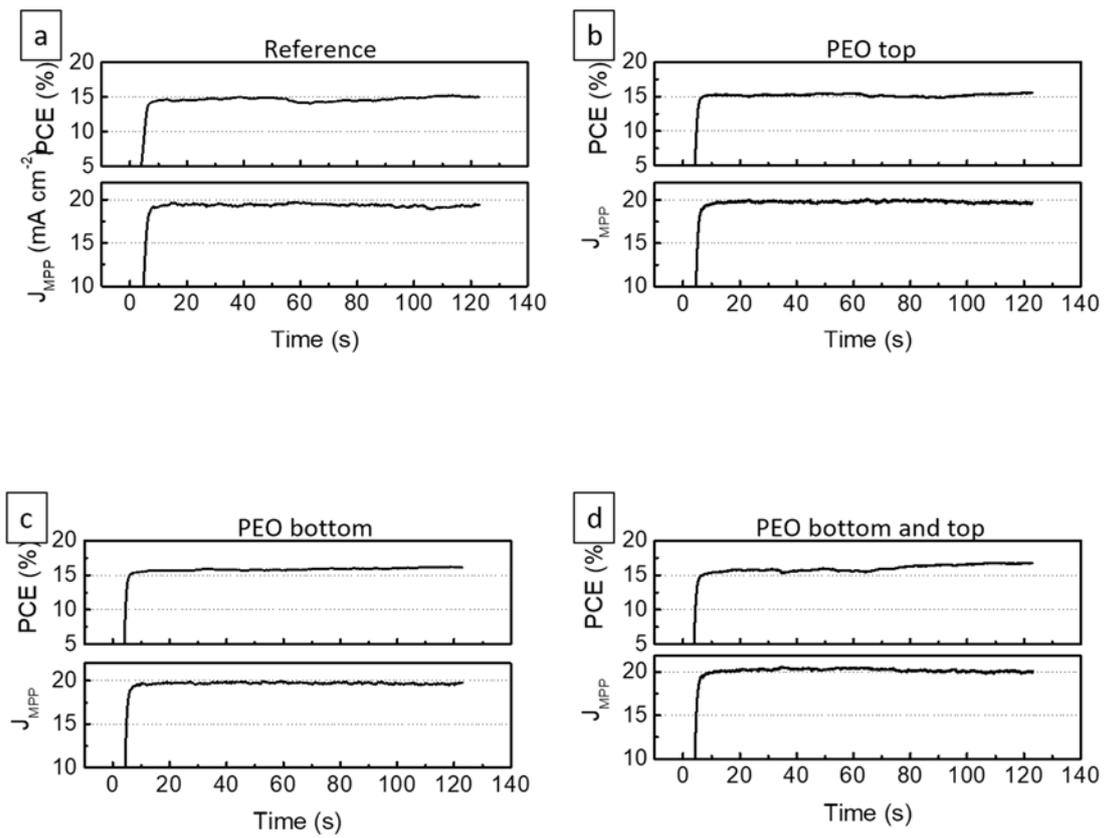
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	Scan direction	J <sub>sc</sub> (mA/cm <sup>2</sup> )	V <sub>oc</sub> (V)	PCE (%)	FF
Reference	Backward	19.72	1.049	15.05	0.73
	Forward	19.66	1.017	13.55	0.68
Bottom	Backward	19.70	1.097	16.14	0.75
	Forward	19.64	1.082	14.46	0.68
Top	Backward	19.57	1.071	15.86	0.76
	Forward	19.51	1.054	14.47	0.70
Bottom and top	Backward	20.35	1.109	17.23	0.76
	Forward	20.29	1.092	15.50	0.70

1049 Table S1. Device parameters of the perovskite devices with and without PEO passivation.

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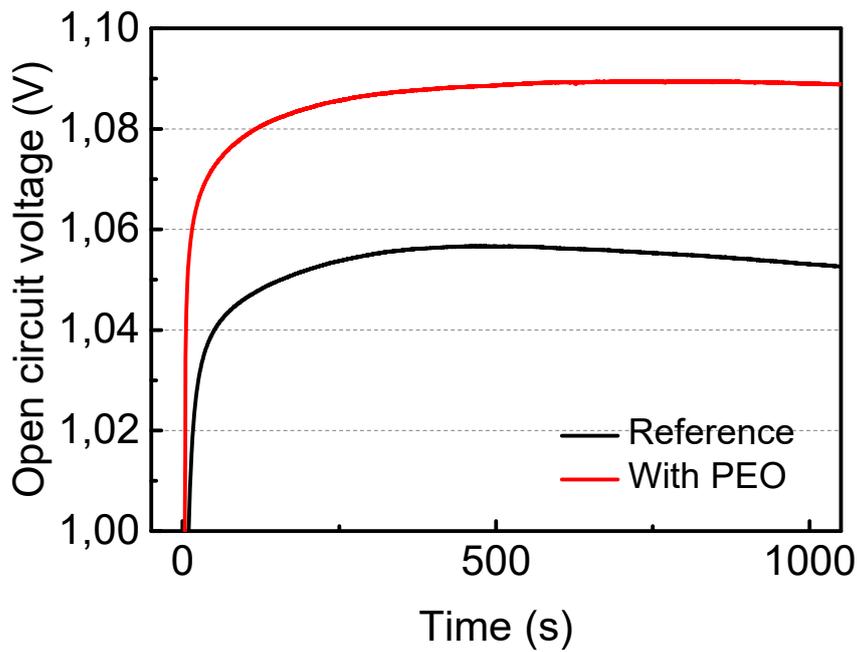
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1053 Figure S16. Time evolution of the maximum power point and with the photocurrent at  
 1054 maximum power point.

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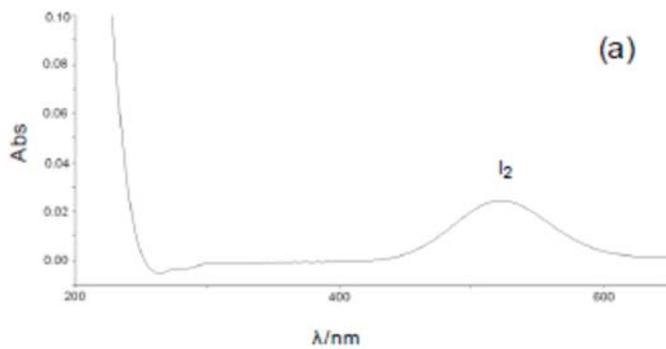


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Figure

1057 S17. open-circuit voltage ( $V_{oc}$ ) transient measured from solar cells with and without PEO

1058 interlayers under an unfocused beam of a 450 nm CW diode laser (Oxxius).



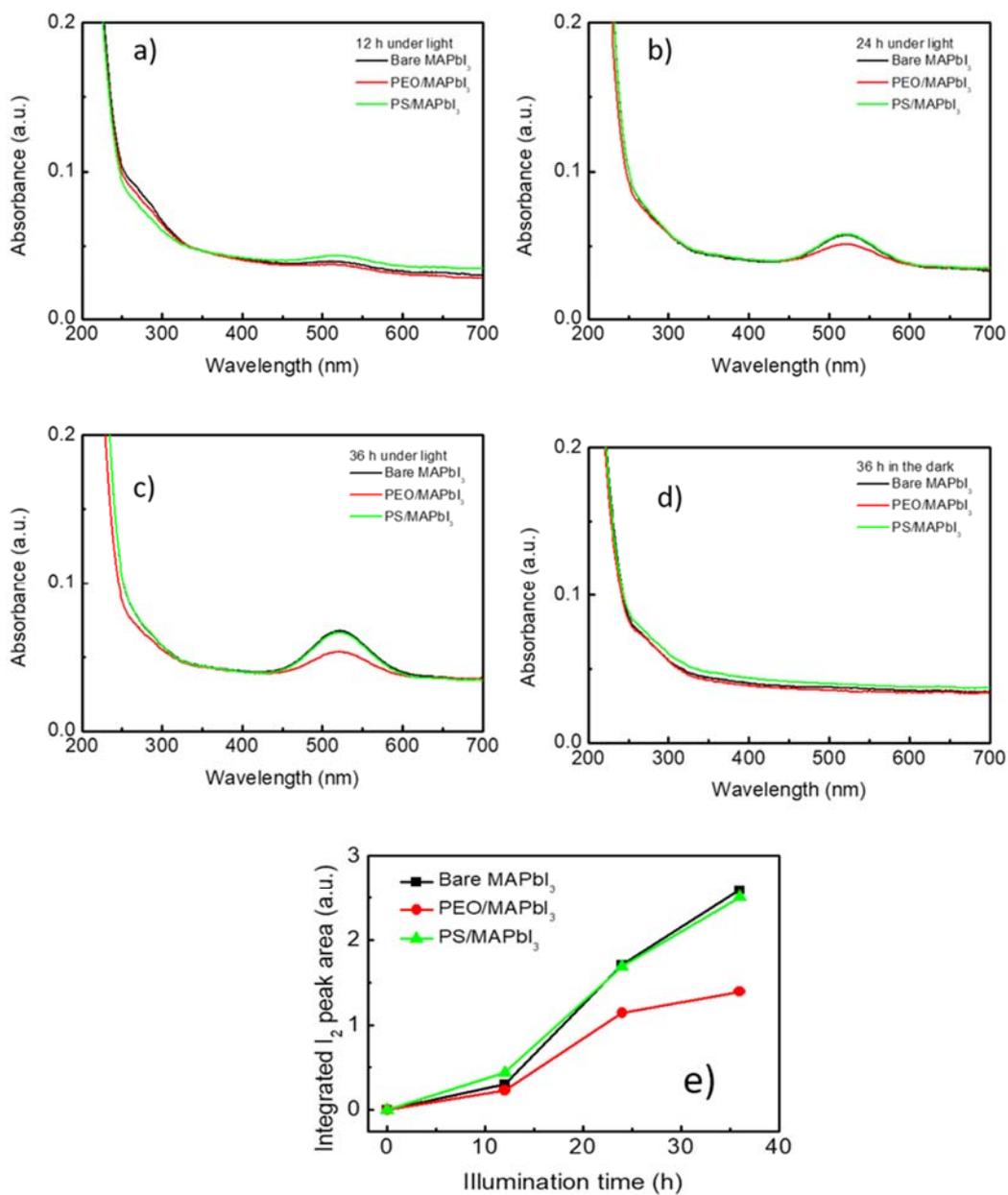
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1060 Figure S18. Absorption spectrum of  $I_2$  in hexane, taken from Atmos. Meas. Tech., 3, 177–

1061 185, 2010.

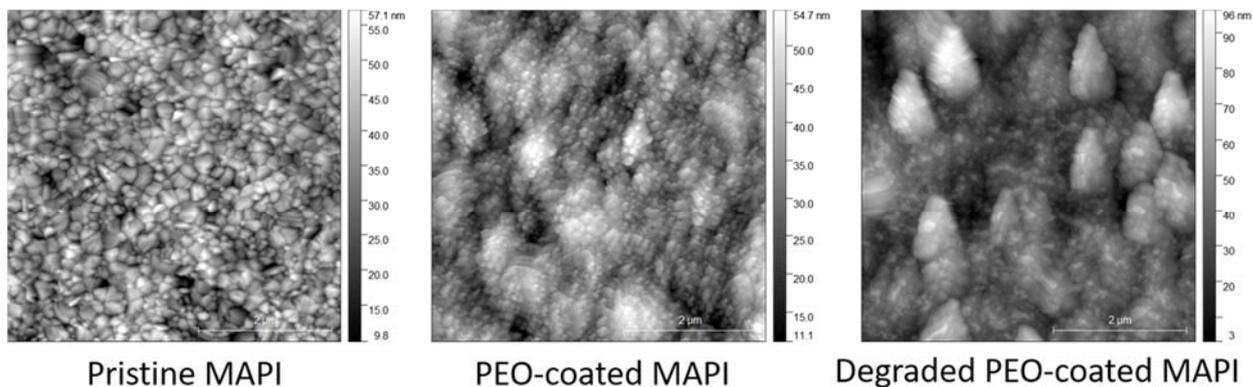
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1065 Figure S19. The absorption spectra of the hexane solution contained in the vials with  
 1066 uncoated, PS- and PEO-coated samples under illumination for a) 12h, b)24h, c) 36h and  
 1067 d) in dark for 36h. e) Integrated I<sub>2</sub> peak area over time as measured in a)-c).



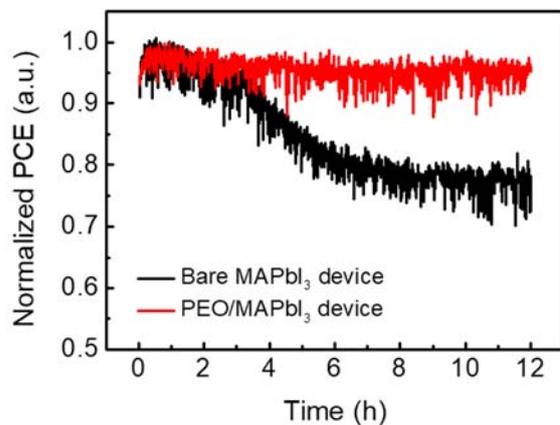
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1070 Figure S20. AFM top images of the bare and PEO-coated MAPbI<sub>3</sub> thin films measured  
 1071 in Figure R4, showing the partial decomposition of the PEO layer by the hexane.

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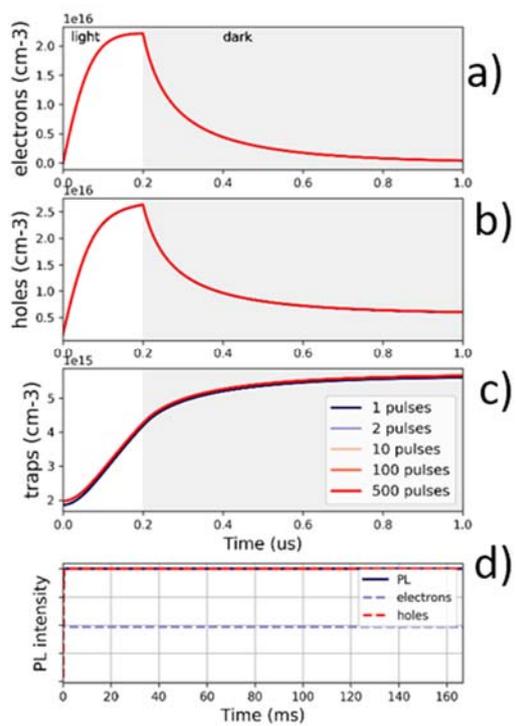
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1076 Figure S21. Solar cells based on bare MAPbI<sub>3</sub> and PEO sandwiched MAPbI<sub>3</sub> tested at  
 1077 the maximum power point, under full solar spectrum illumination (1SUN). The devices  
 1078 were encapsulated to exclude environmental effects.

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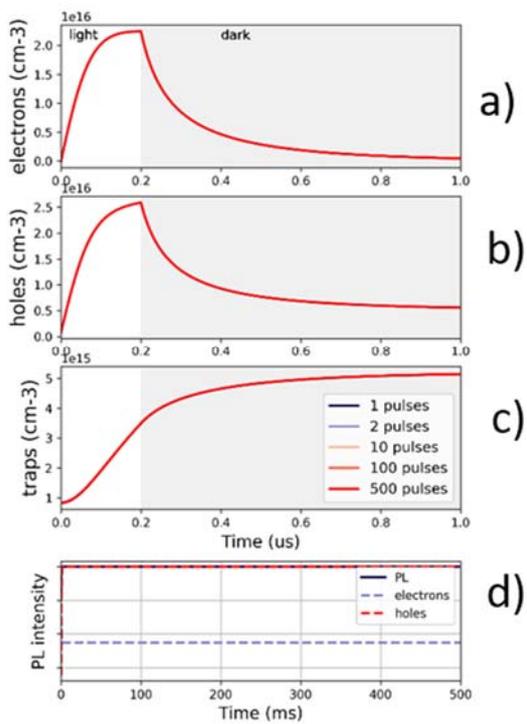
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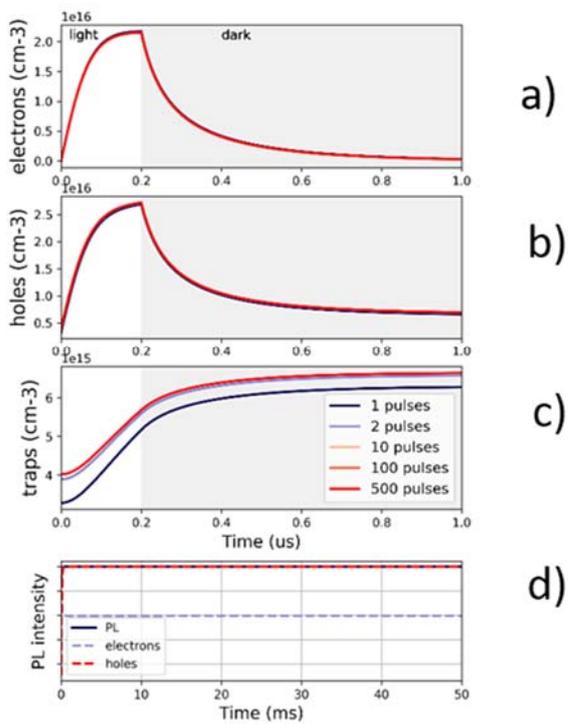
1083 Figure S22. Simulation of a) electron, b) hole, c) trap population dynamics and d) the  
 1084 corresponding photoluminescence intensity over time for pulse length of 200 ns, pulse  
 1085 energy 0.2 nJ, repetition rate 1 kHz, average power 0.6 mW/cm<sup>2</sup>.



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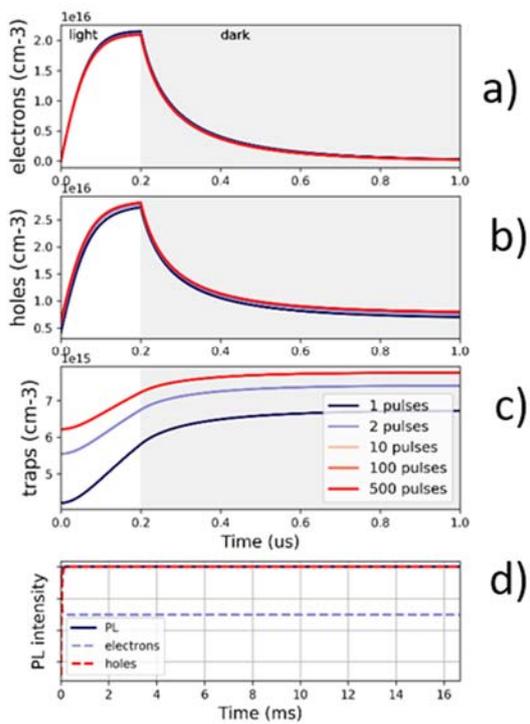
1087 Figure 23. Simulation of a) electron, b) hole, c) trap population dynamics and d) the  
 1088 corresponding photoluminescence intensity over time for pulse length of 200 ns, pulse  
 1089 energy 0.2 nJ, repetition rate 3 kHz, average power 2 mW/cm<sup>2</sup>.

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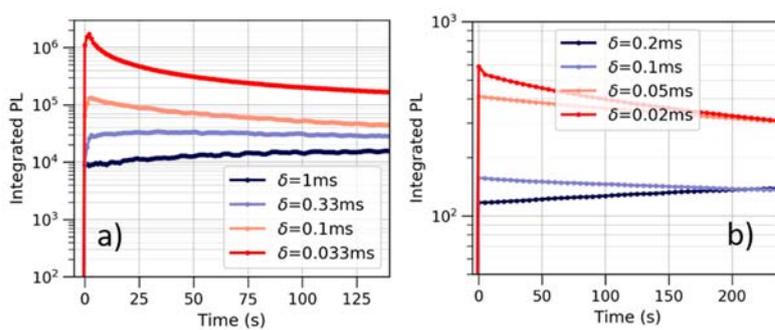
1092 Figure S24. Simulation of a) electron, b) hole, c) trap population dynamics and d) the  
 1093 corresponding photoluminescence intensity over time for pulse length of 200 ns, pulse  
 1094 energy 0.2 nJ, repetition rate 10 kHz, average power 6 mW/cm<sup>2</sup>.



1095

1096 Figure S25. Simulation of a) electron, b) hole, c) trap population dynamics and d) the  
 1097 corresponding photoluminescence intensity over time for pulse length of 200 ns, pulse  
 1098 energy 0.2 nJ, repetition rate 30 kHz, average power 20 mW/cm<sup>2</sup>.

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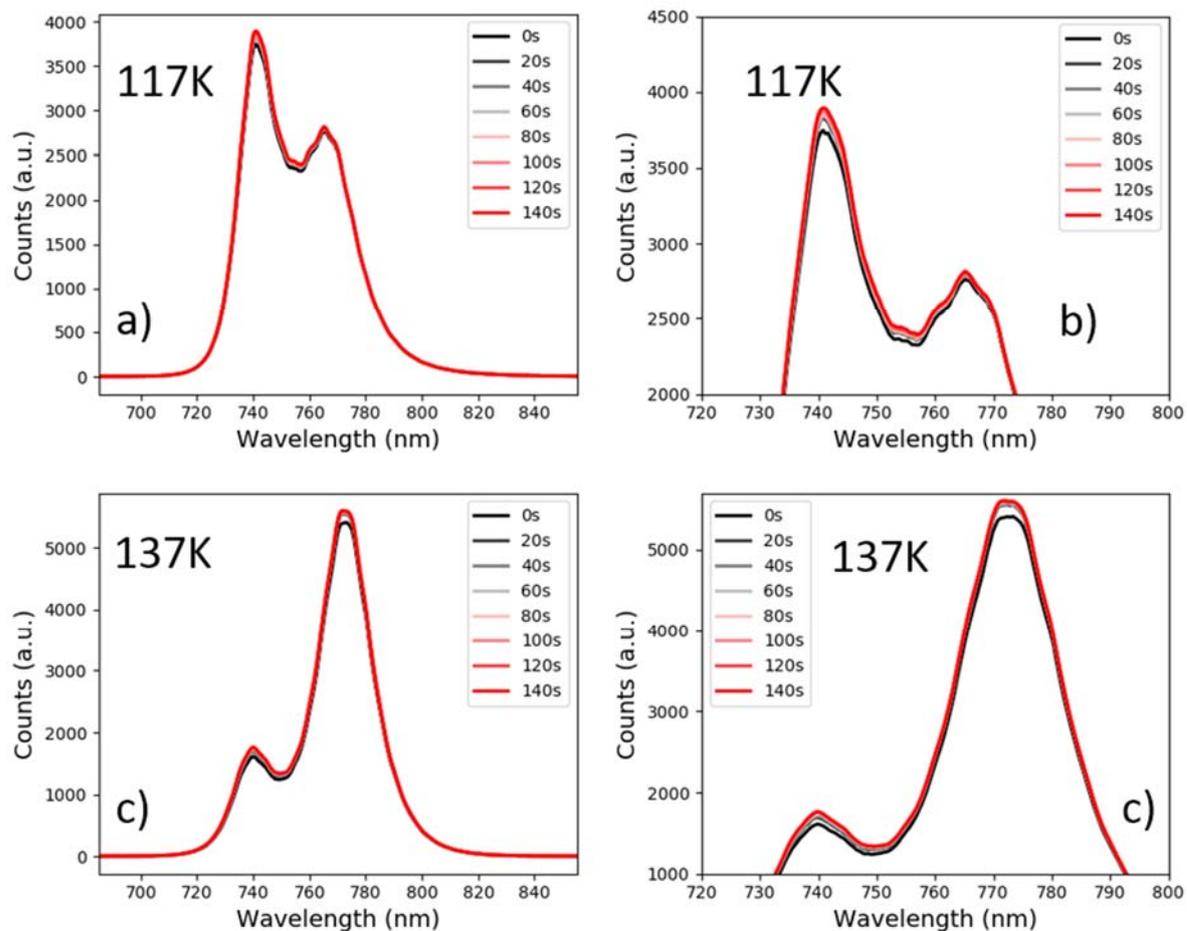


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1101 Figure S26. not normalized integrated PL dynamics at different repetition rates.

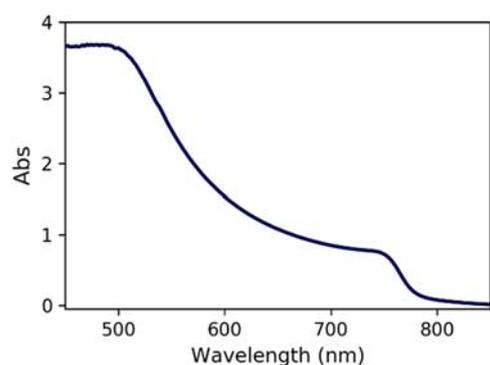
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1105 Figure S27. Temporal evolution of MAPbI<sub>3</sub> PL spectra at temperatures across the phase  
 1106 transitions. Figure b) and c) are a zoom in of the spectra in Fig a) and b).



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1108 Figure R28. Absorption spectrum of MAPbI<sub>3</sub>.

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