

# 3D Model of Magnetic Reconnection in Sigmoidal Regions

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# Background Info

- Controls all the activity that you can and cannot see on the sun
- Twisting magnetic field lines



Video: UCAR via SOHO (ESA & NASA)

# Outline

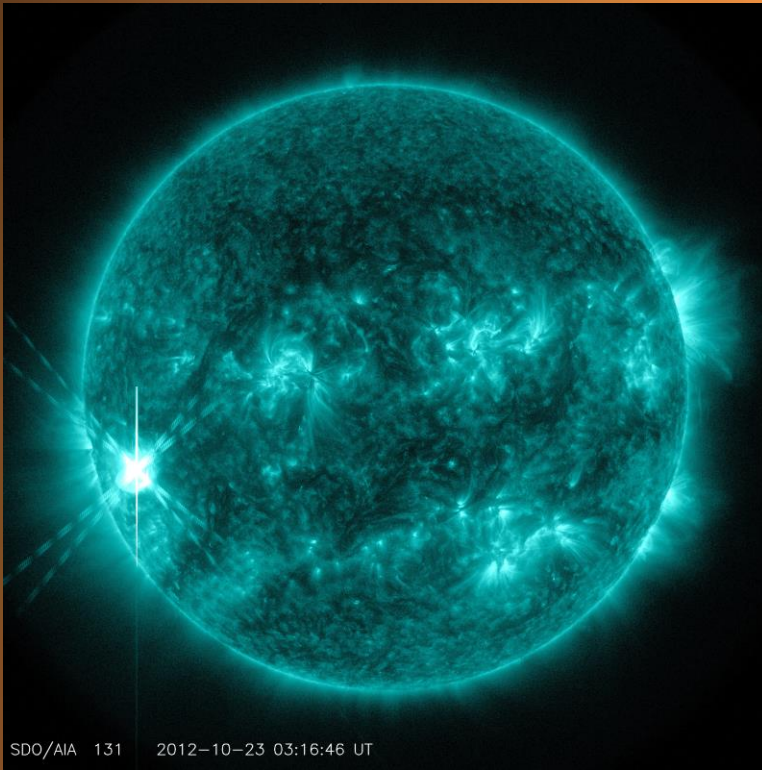
Jargon: Sigmoids,  
QSLs, Slip-Running  
Reconnection



Process: Coronal  
Modeling System  
(CMS) via IDL



Analysis: Compare  
slippage observations  
with 3D models



# JARGON YOU NEED TO KNOW TO MAKE MODELS



# Sigmoids

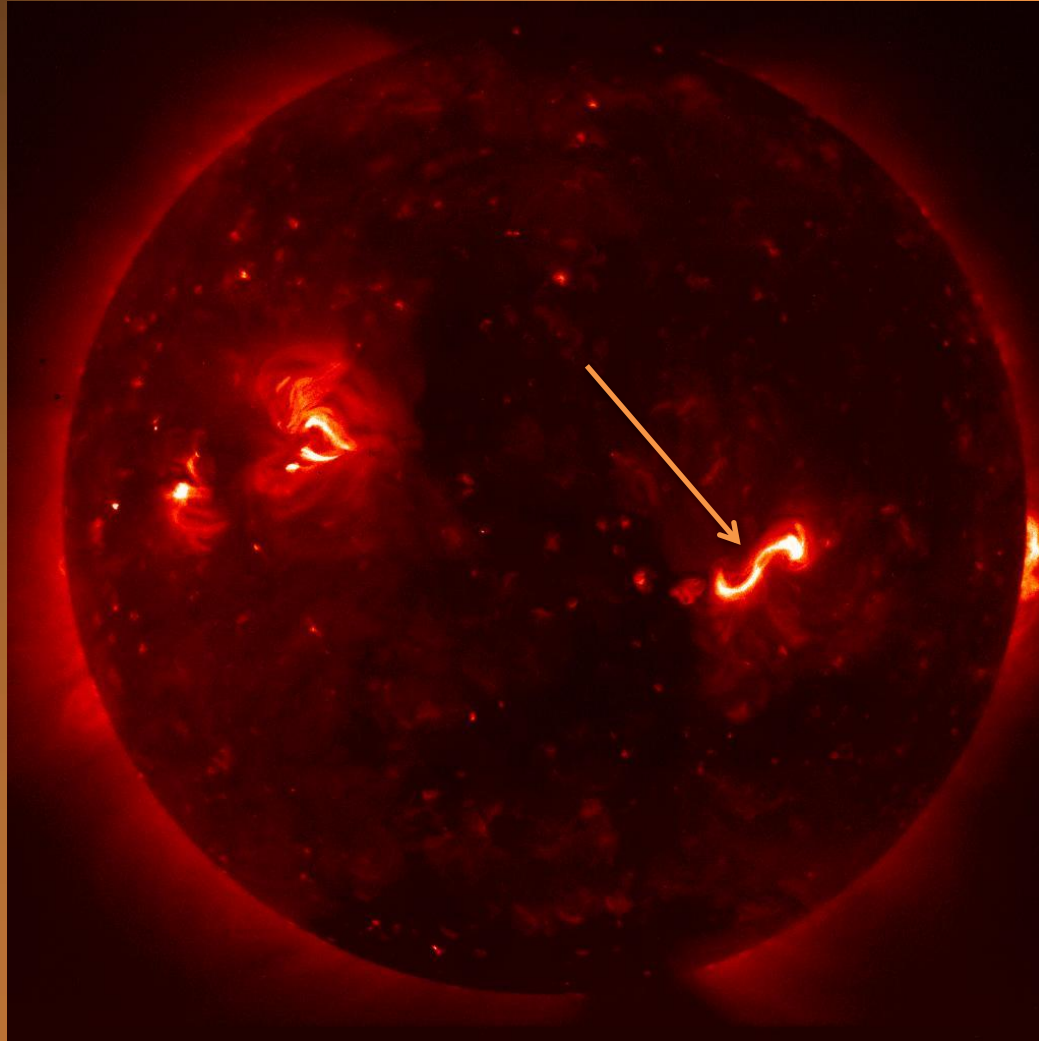
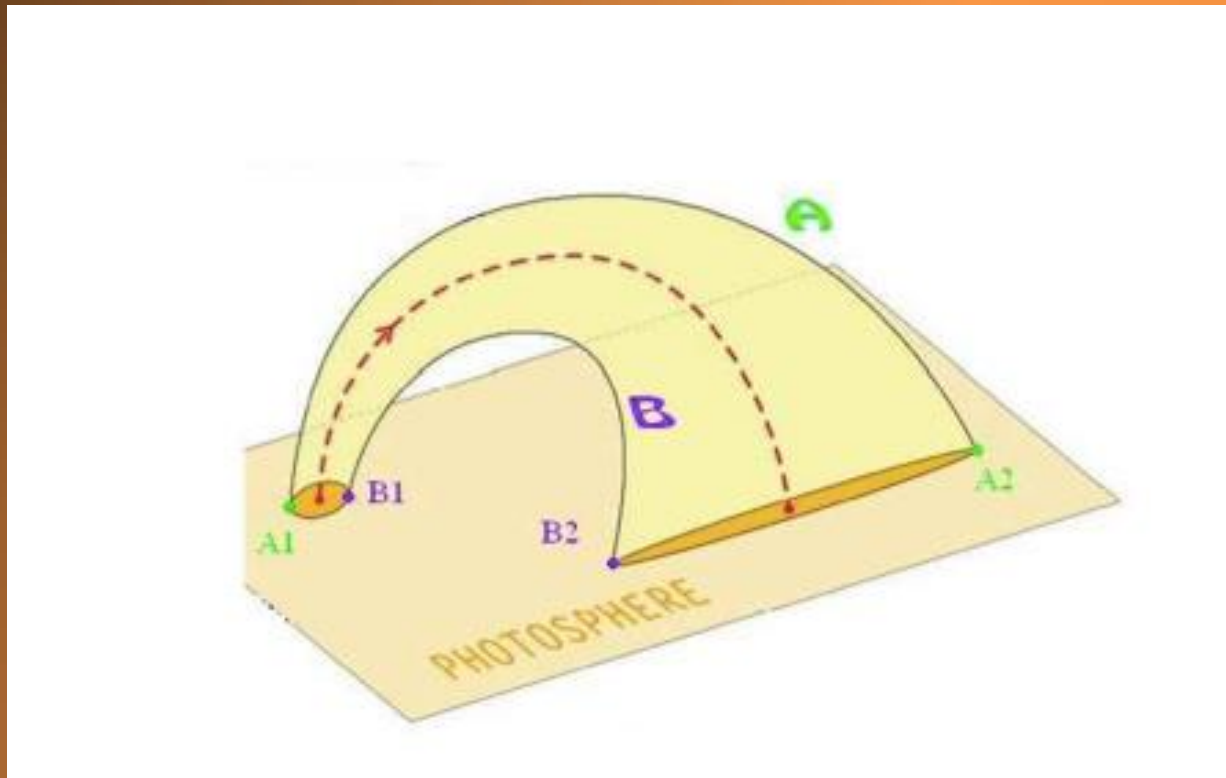


Image: NASA/STFC/ISAS/JAXA

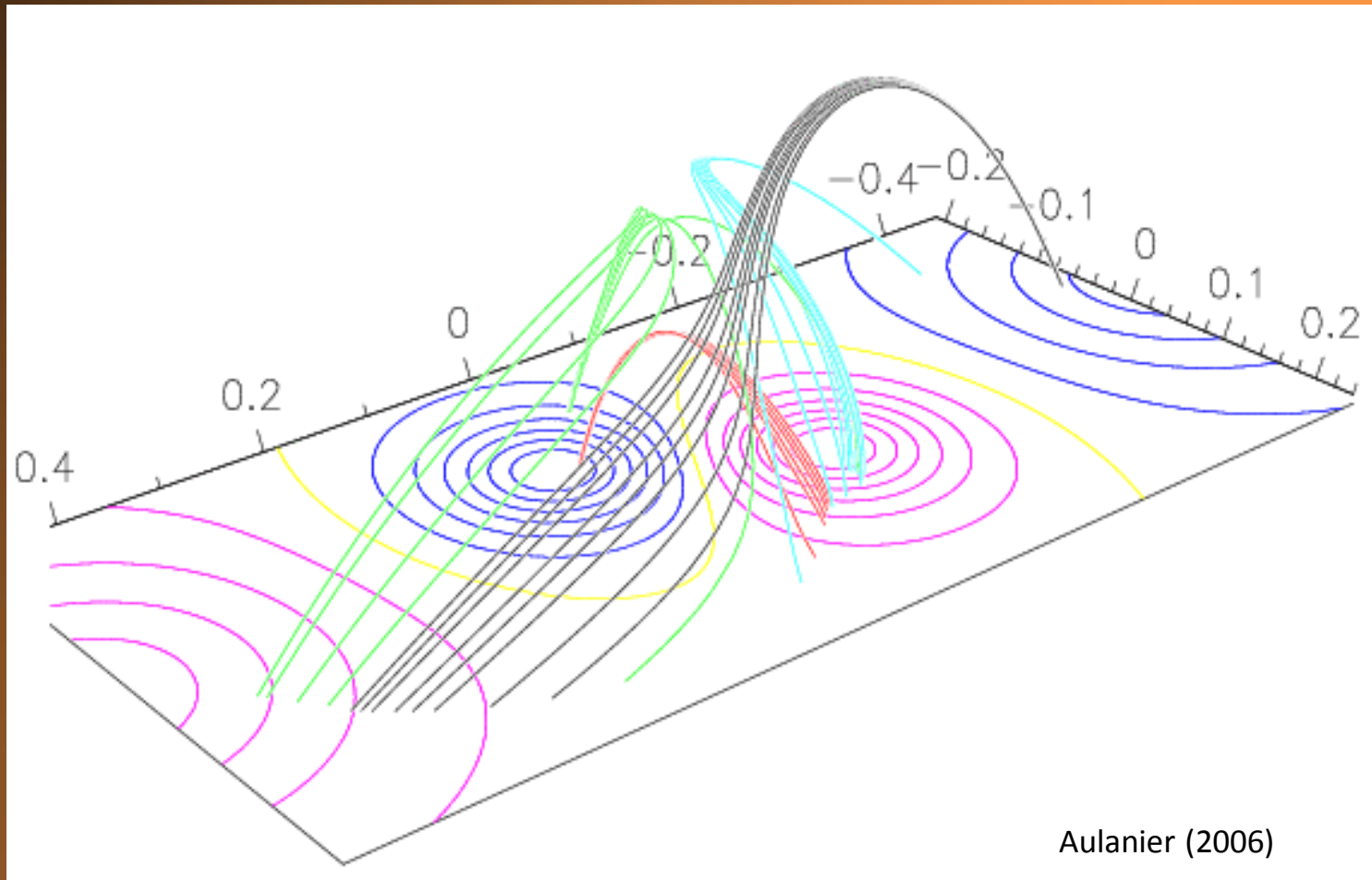
# Quasi Separatrix Layers (QSLs)

3D topological features



Pariat (2006)

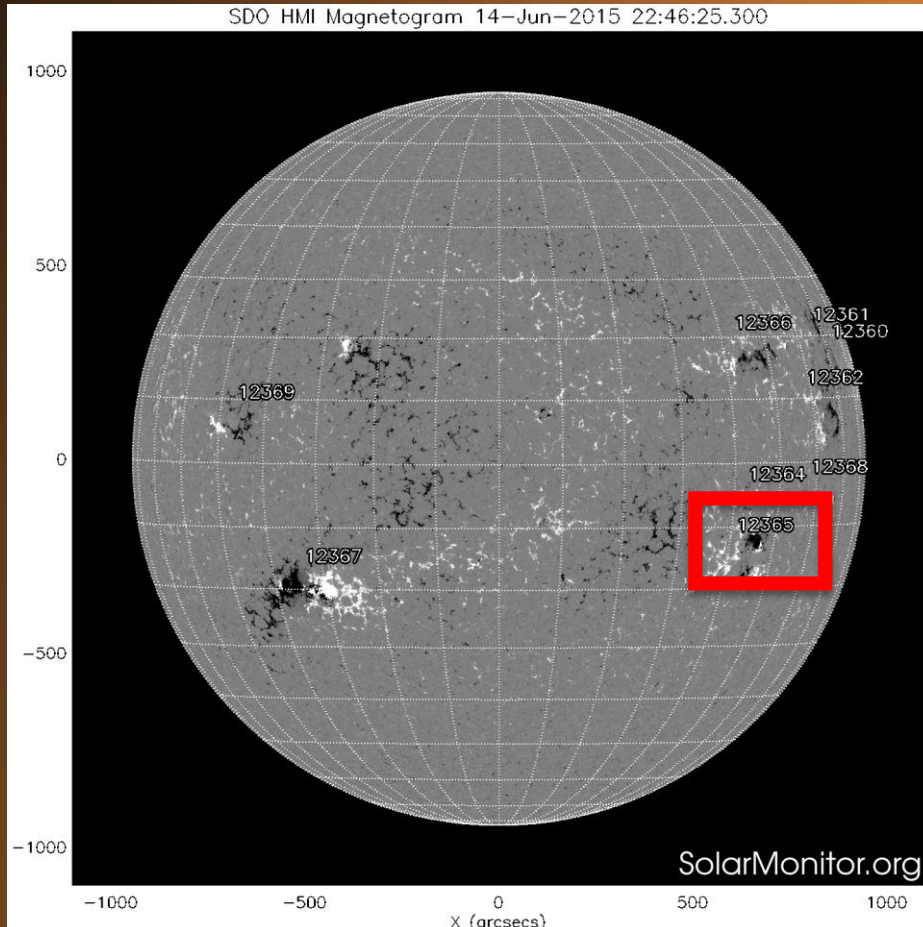
# Slip Running Reconnection



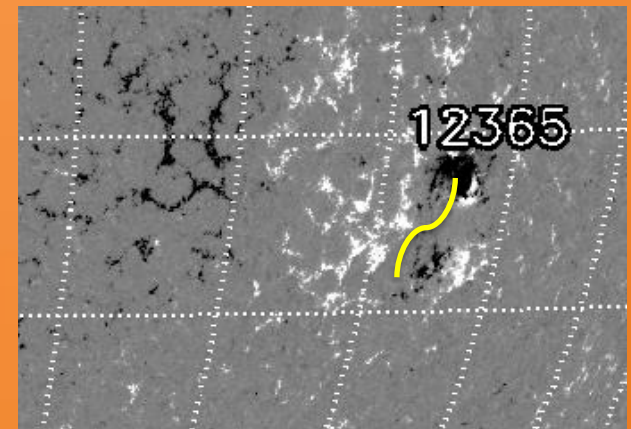
# CMS MODELING PROCESS



# CMS Modeling



3. Compute Potential field (axial flux and poloidal flux)
4. Relax the solution for  $\sim 30,000$  to  $60,000$  iterations

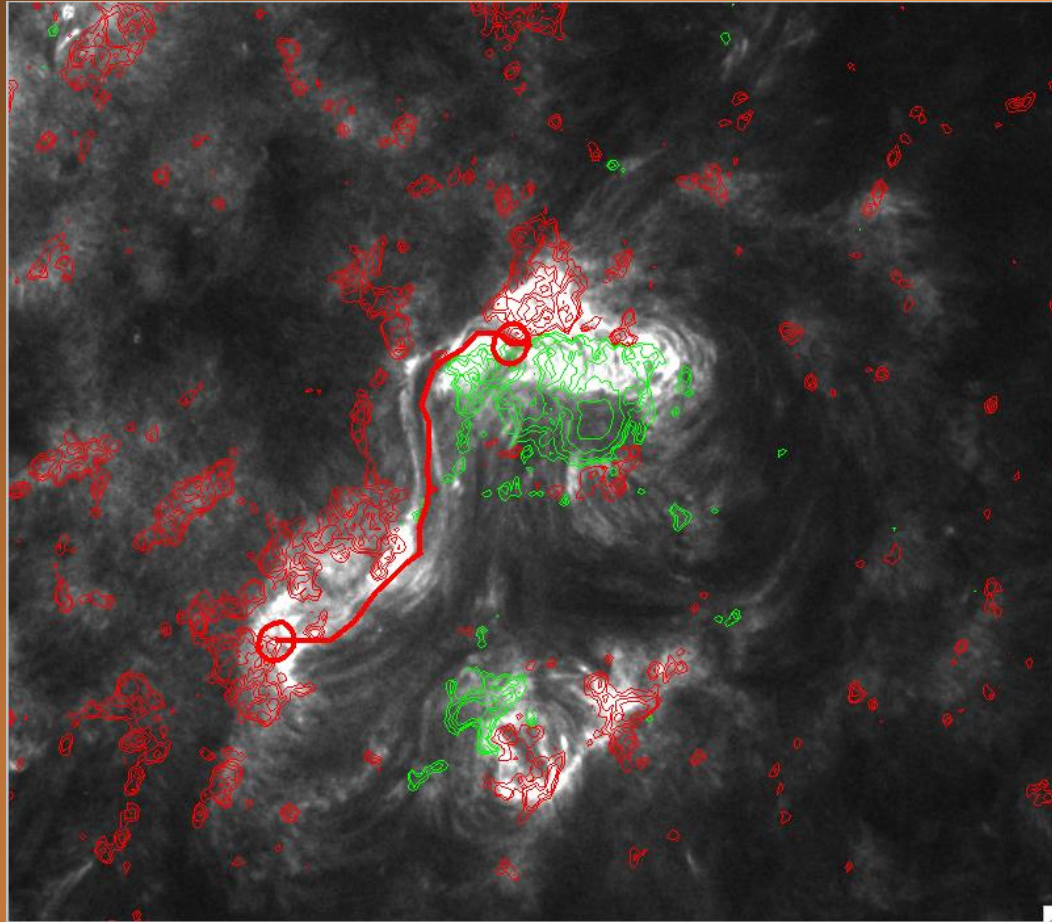


**Non-linear force free field (NLFFF)**  
– used for investigating the structure, dynamics and evolution of the coronae of solar active regions (DeRosa et al 2009)

# Non-linear Force Free Field

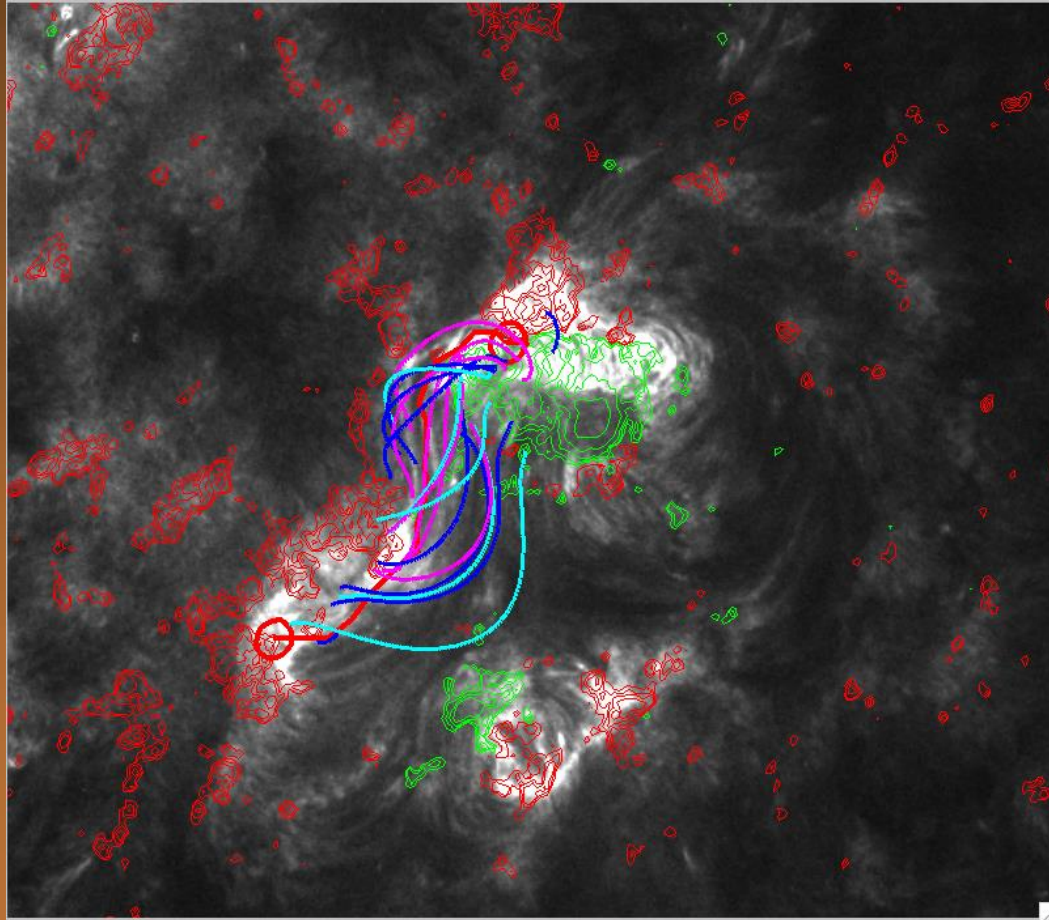
- Assume Lorentz force is zero
- $\mathbf{J} \times \mathbf{B} = (\nabla \times \mathbf{B}) \cdot \mathbf{B} = 0$  ,  $\nabla \cdot \mathbf{B} = 0$
- $\mathbf{J} = 0$  linear force (potential field)
- $\mathbf{J} \parallel \mathbf{B} = 0$  ,  $\mathbf{J} = \alpha * \mathbf{B}$  non-linear force
  - $\alpha = C$  linear force free field ( $\alpha = 0$  special linear field)
  - $\alpha \neq C$  non linear force free field
- Makes numerical simulations not as complex

# Flux Rope Path



AIA 304 Å: 14 June 2015

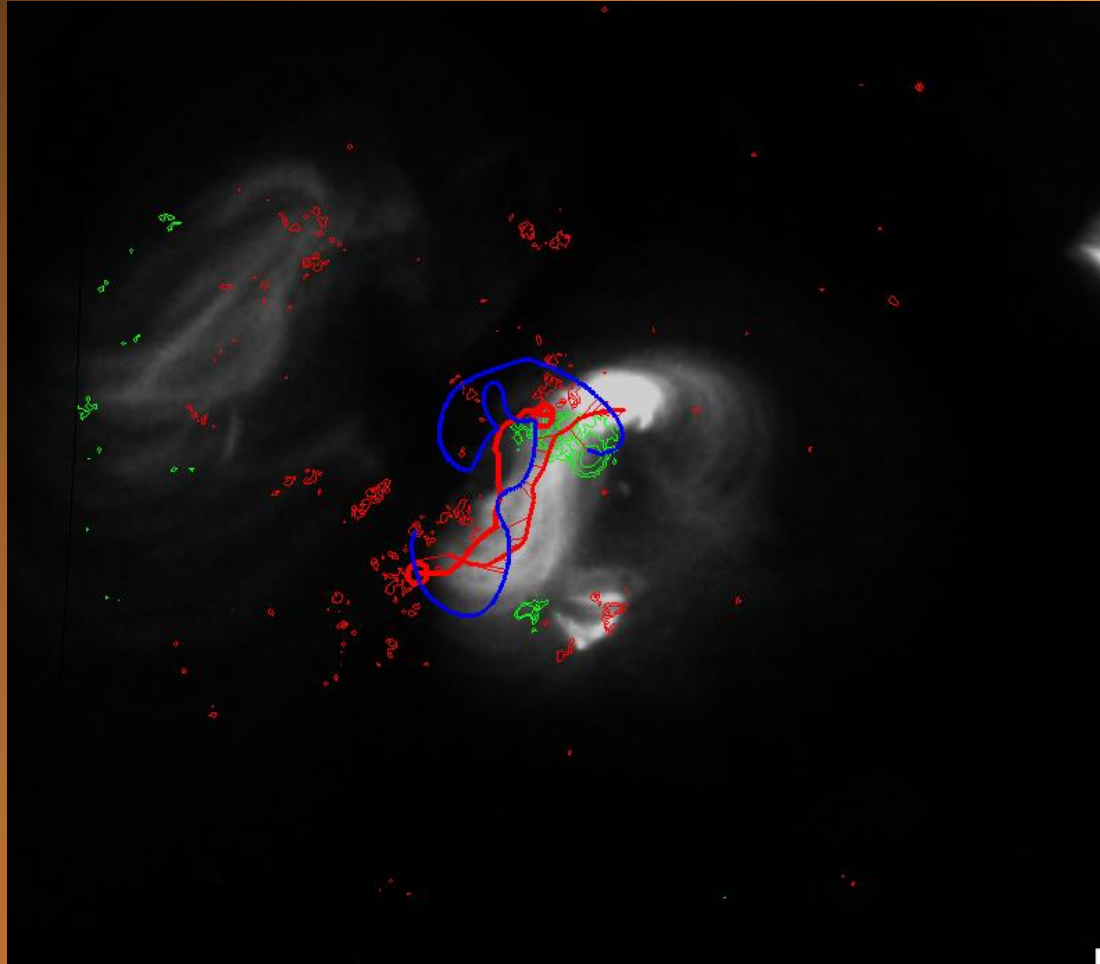
# Viewing Field Lines



AIA 304 Å: 14 June 2015



# Fitting the Data



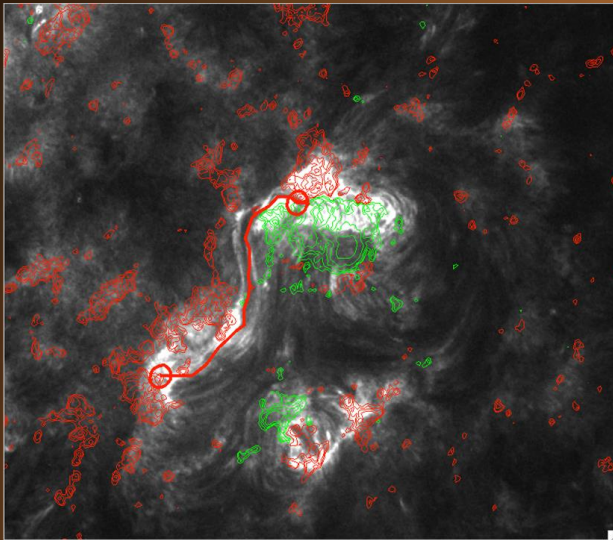
XRT: 14 June 2015

GoodFit (0.00525)

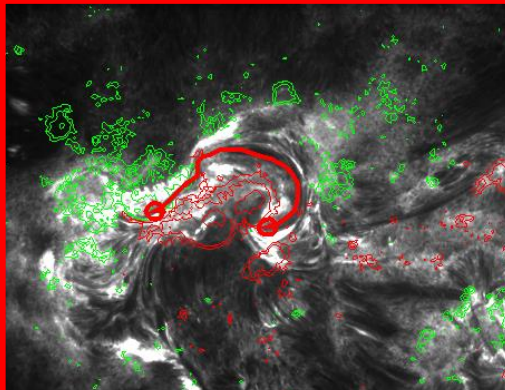
# ANALYSIS & RESULTS



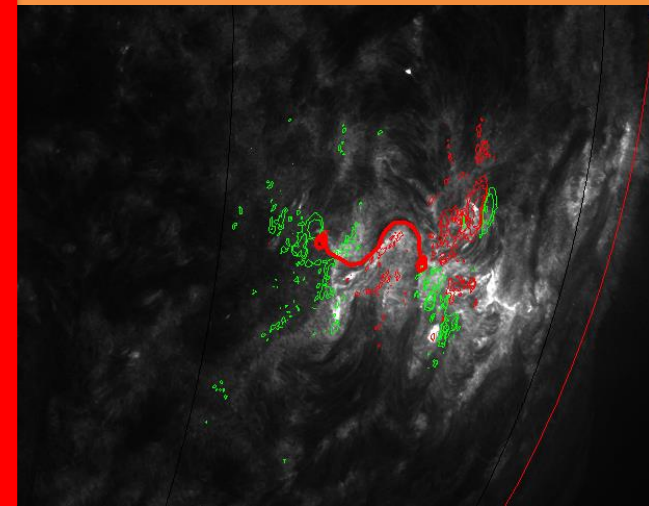
# 3 Regions Modeled



14 June 2015

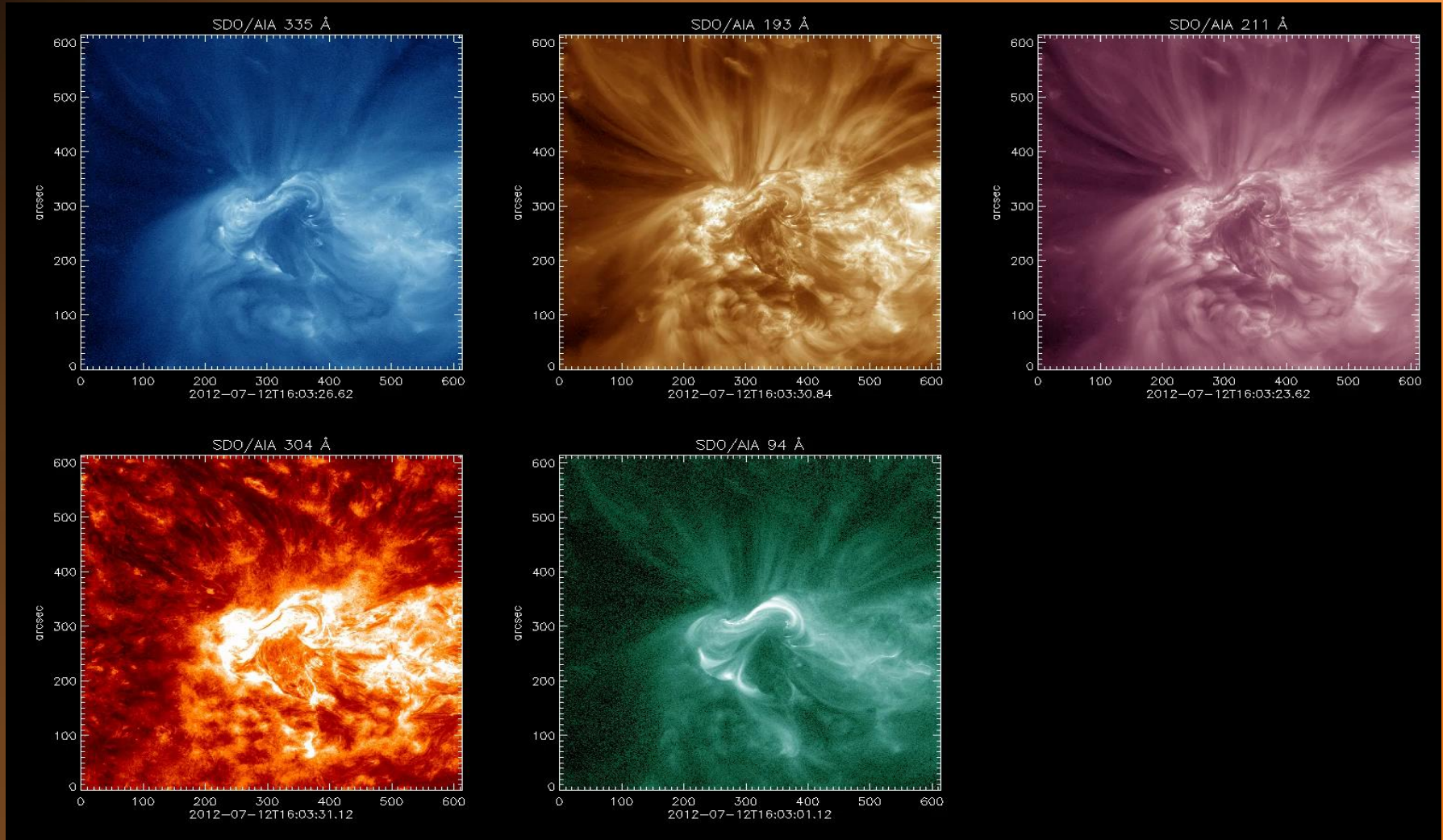


12 July 2012



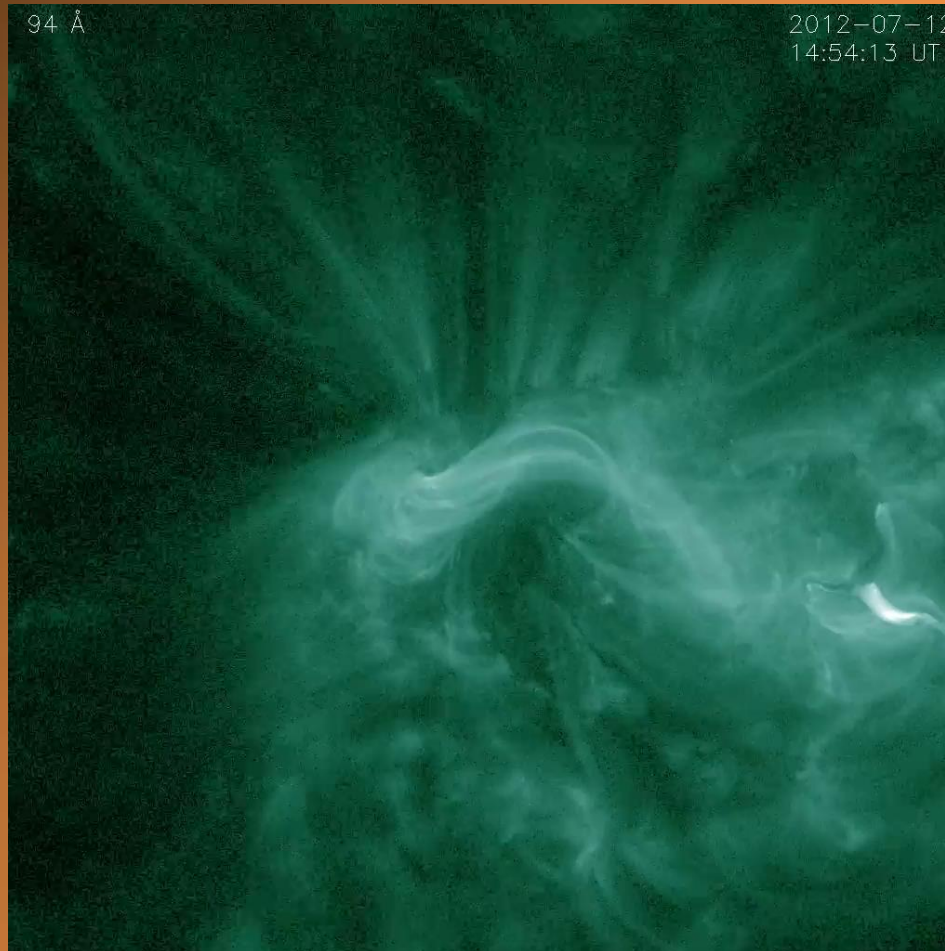
30 September 2014

# Comparing the Data



12 July 2012

# Comparing the Data

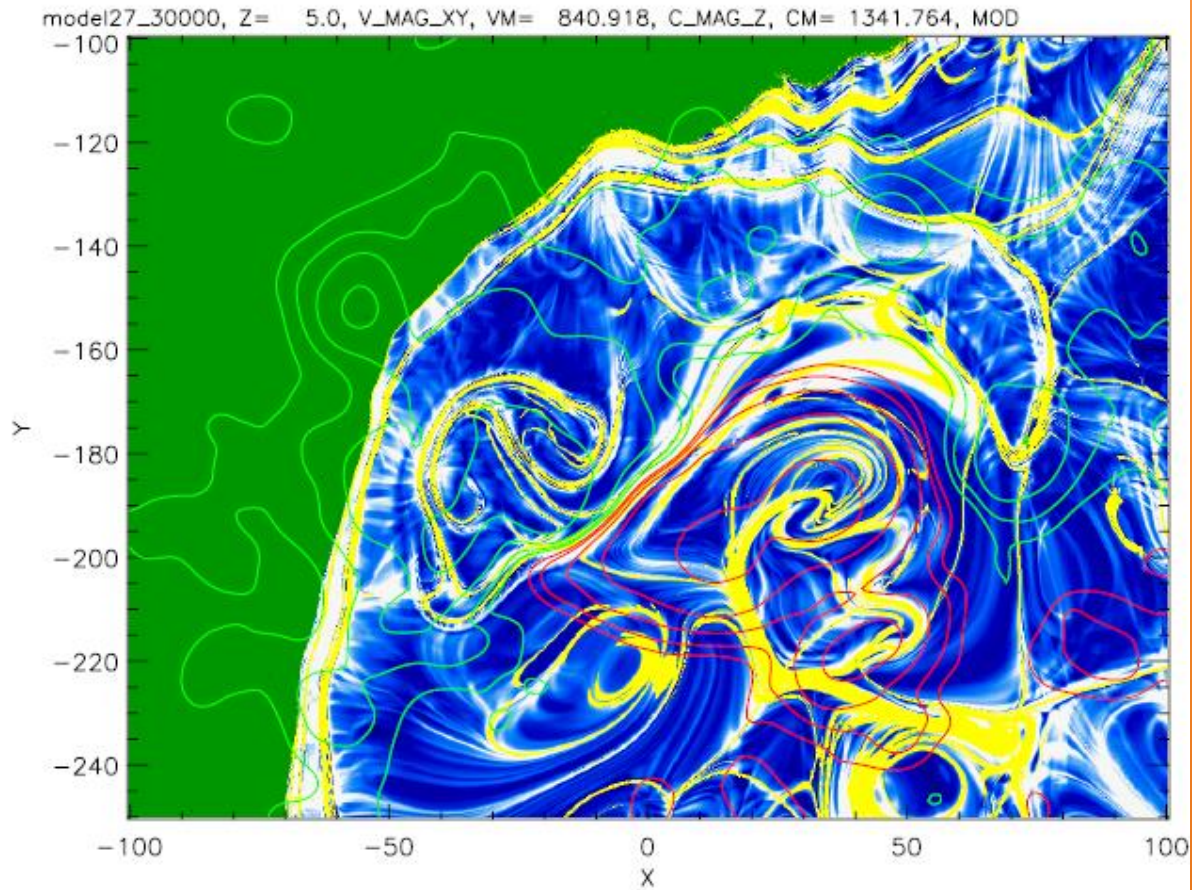


AIA 94 Å: 12 July 2012



# QSL Map

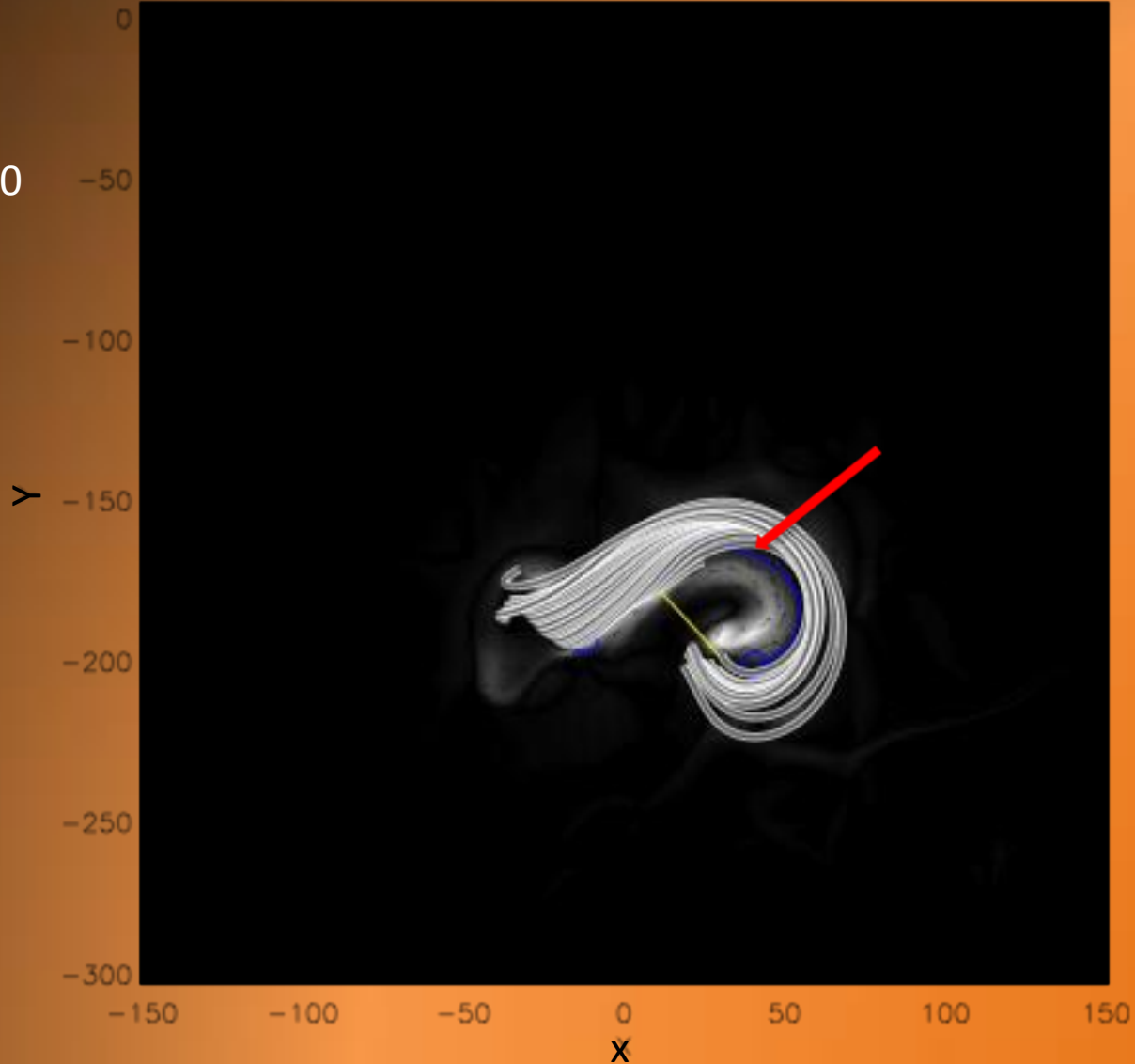
Model27\_30000  
QSL Map



# Plotting best fit data

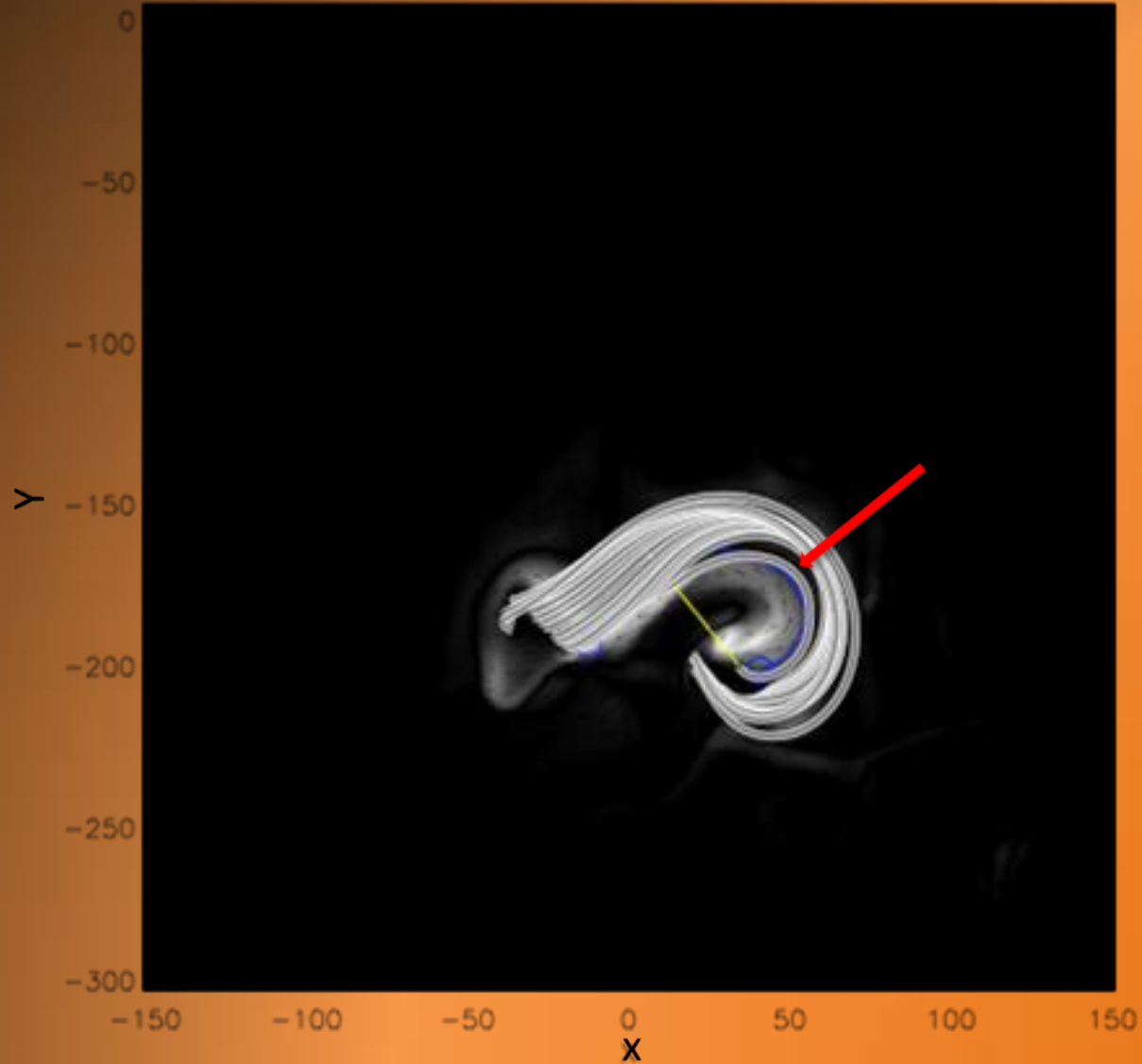
model6\_20000, Z= 5.0, V\_MAG\_XY, VM= 844.855, C\_CUR\_A, CM= 165383.5, MOD

Model 6\_20000  
XY- Plane



# Model6\_30000

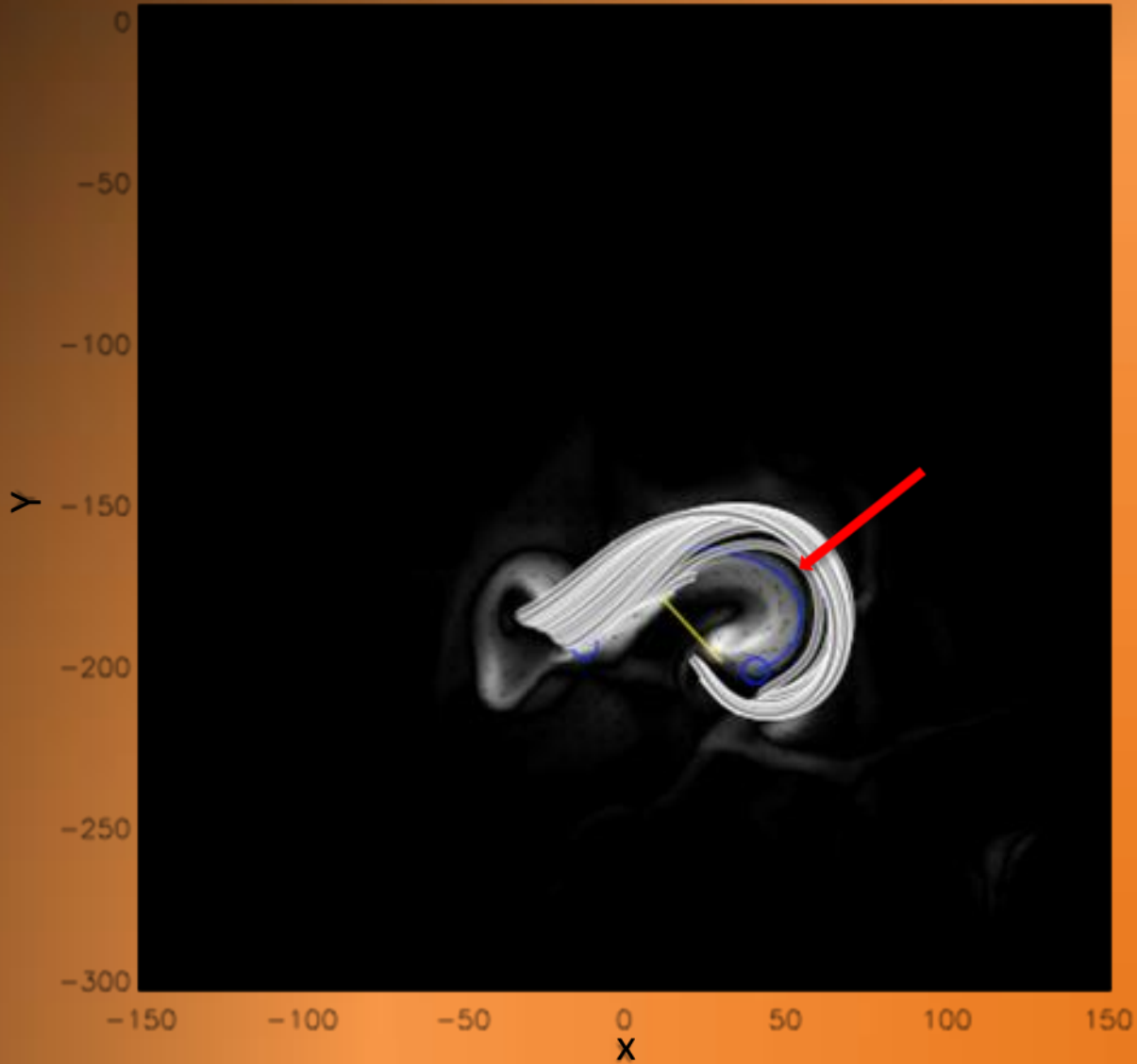
model6\_30000, Z= 5.0, V\_MAG\_XY, VM= 787.611, C\_CUR\_A, CM= 137847.8, MOD





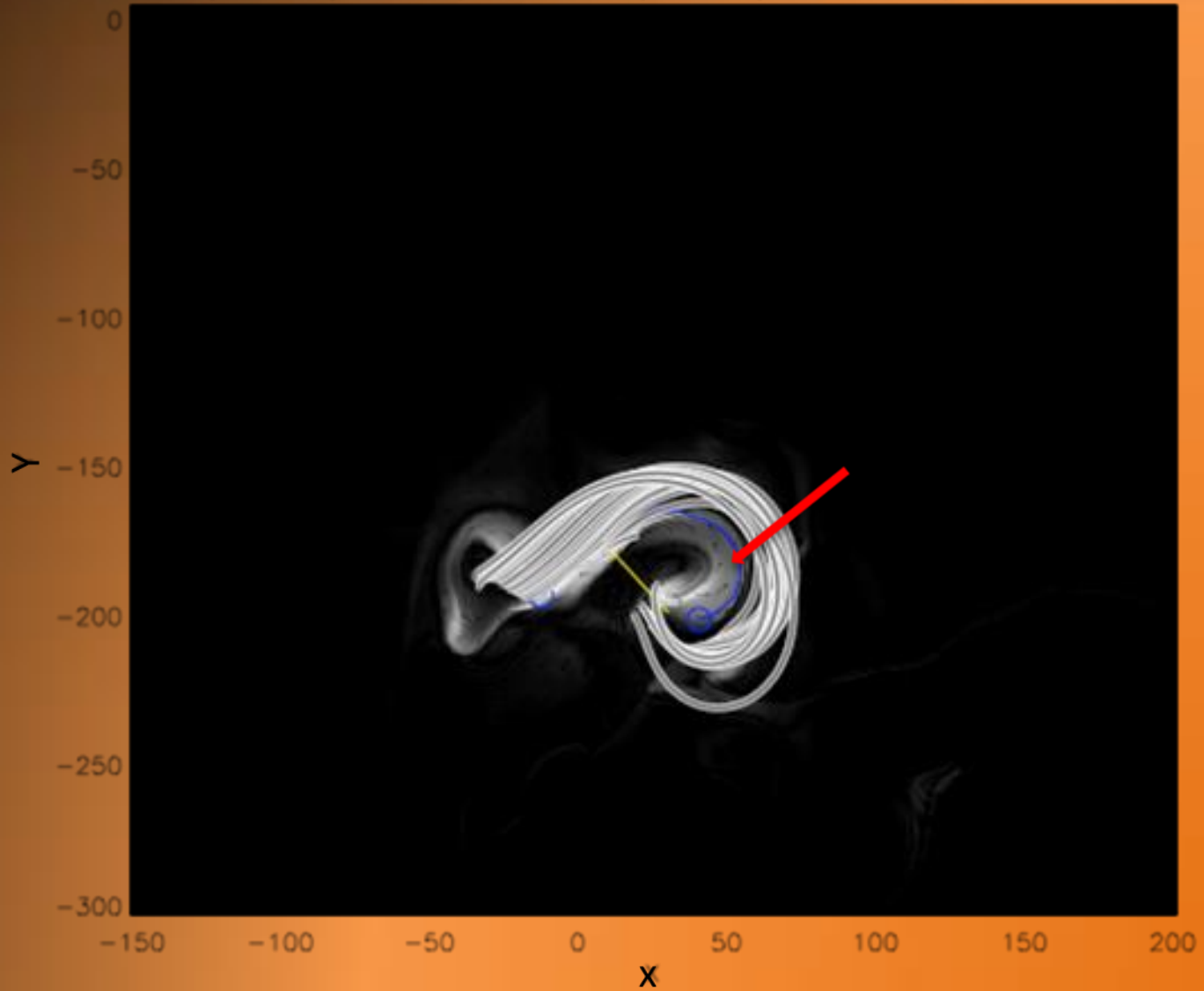
# Model6\_40000

model6\_40000, Z= 5.0, V\_MAG\_XY, VM= 780.585, C\_CUR\_A, CM= 120184.8, MOD

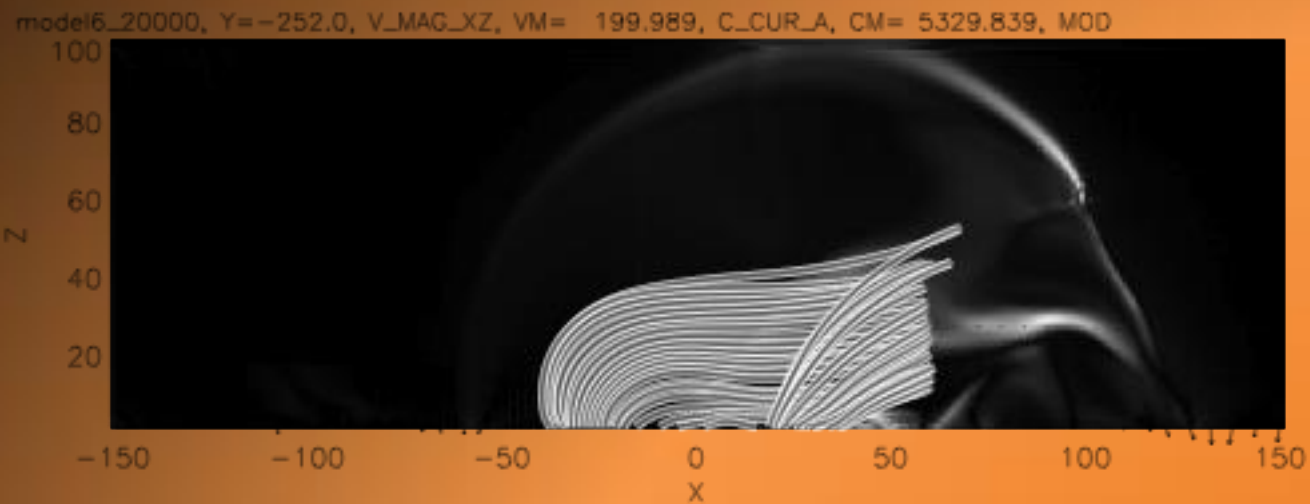


# Model6\_50000

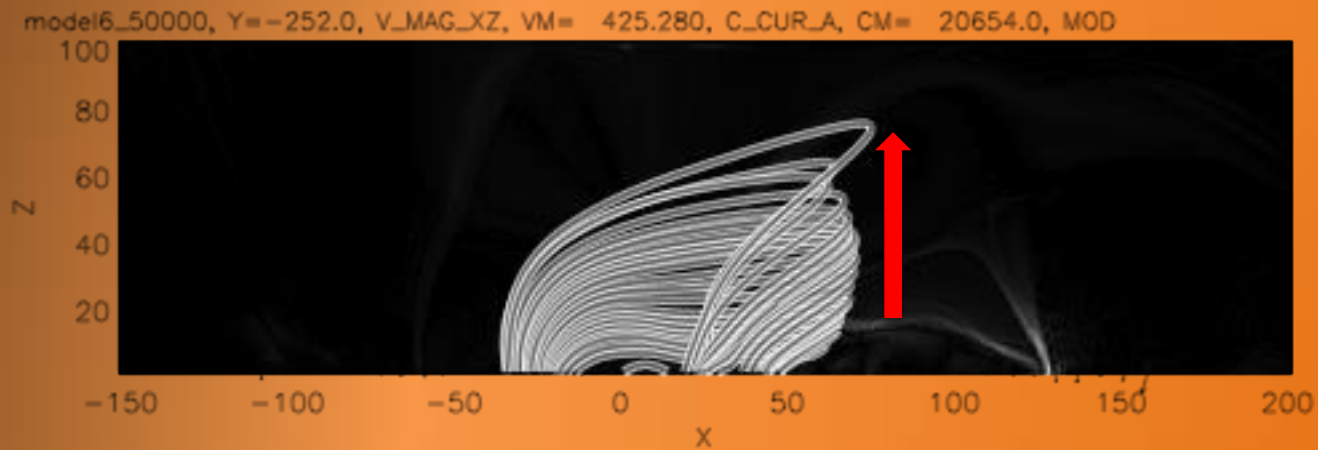
model6\_50000, Z= 5.0, V\_MAG\_XY, VM= 800.263, C\_CUR\_A, CM= 119581.1, MOD



### Model 6\_20000 SZ Plane



### Model 6\_50000 SZ Plane



# Summary

- Sigmoids are areas where eruptive events caused by change in connectivity can occur
- There is a correlation between slipping reconnection and the created models
- NLFFF modeling may be able to capture realistic 3D magnetic structures associated with slipping reconnection

# Thank You!

# Acknowledgements

- Mentors Antonia Savcheva & Ed DeLuca
- Professor John Johnson and Banneker Institute
- Center for Astrophysics faculty and staff
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# Questions?