

```
import numpy as np
import matplotlib.pyplot as plt
from scipy.optimize import minimize
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```
# 1. REAL OBSERVATIONAL DATA: Cosmic
    Chronometers (CC) H(z) Data
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```
cc_observational_data = np.array([
    ,[12.0 ,69.0 ,0.090] ,[19.6 ,69.0 ,0.070]
    ,[26.2 ,68.6 ,0.120]
```

,[4.0 ,75.0 ,0.179] ,[8.0 ,83.0 ,0.170]  
,[5.0 ,75.0 ,0.199]  
,[2.6 ,79.6 ,0.240] ,[29.6 ,72.9 ,0.200]  
,[14.0 ,77.0 ,0.270]  
,[14.0 ,83.0 ,0.352] ,[36.6 ,88.8 ,0.280]  
,[13.5 ,83.0 ,0.3802]  
,[10.2 ,77.0 ,0.4004] ,[17.0 ,95.0 ,0.400]  
,[11.2 ,87.1 ,0.4247]  
,[50.0 ,89.0 ,0.470] ,[12.9 ,92.8 ,0.4497]  
,[9.0 ,80.9 ,0.4783]  
,[13.0 ,104.0 ,0.593] ,[62.0 ,97.0 ,0.480]  
,[8.0 ,92.0 ,0.680]  
,[17.0 ,125.0 ,0.875] ,[12.0 ,105.0 ,0.781]  
,[40.0 ,90.0 ,0.880]  
,[20.0 ,154.0 ,1.037] ,[23.0 ,117.0 ,0.900]  
,[17.0 ,168.0 ,1.300]  
,[18.0 ,177.0 ,1.430] ,[33.6 ,160.0 ,1.363]  
,[14.0 ,140.0 ,1.530]  
[50.4 ,186.5 ,1.965] ,[40.0 ,202.0 ,1.750]  
([

z\_obs = cc\_observational\_data[:, 0]  
H\_obs = cc\_observational\_data[:, 1]

```
sigma_H = cc_observational_data[:, 2]
```

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## # 2. THEORETICAL MODELS AND FIXED PARAMETERS

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```
t0 = 4.352e17
```

```
km_s_Mpc_to_SI = 3.24077929e-20
```

```
# H0 is derived directly from the age of the  
universe and  $\beta=1.0$  (zero free parameters)
```

```
beta_RPT_fixed = 1.0
```

```
H0_RPT_SI = beta_RPT_fixed / t0
```

```
H0_RPT = H0_RPT_SI / km_s_Mpc_to_SI
```

```
def H_RPT_fixed(z):
```

```
    """RPT with fully locked  $\beta=1.0$  — no
```

```

        freedom whatsoever"""
    denominator = 1.0 - (1.0 + z)**(-1.0 /
                                beta_RPT_fixed)
    denominator = np.where(denominator
                           == 0, 1e-10, denominator)
    return H0_RPT * (z / denominator)

def H_LCDM(z, H0, Om0):
    """Standard  $\Lambda$ CDM with two free
        parameters"""
    return H0 * np.sqrt(Om0 * (1.0 + z)**3 +
                        (1.0 - Om0))

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### # 3. CHI-SQUARED CALCULATION FOR MODEL COMPARISON

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```
def chi2_LCDM(params):
```

```
    H0, Om0 = params
```

```
    H_pred = H_LCDM(z_obs, H0, Om0)
```

```
    return np.sum((((H_obs - H_pred) /  
                    sigma_H)**2)
```

```
    # Fitting  $\Lambda$ CDM to the data
```

```
    res_LCDM = minimize(chi2_LCDM,
```

```
    x0=[70.0, 0.3], bounds=[(50, 90), (0.1, 0.5)])
```

```
    best_H0_lcdm, best_Om0_lcdm =  
    res_LCDM.x
```

```
    # Computing  $\chi^2$  for RPT (without any  
    fitting)
```

```
    H_pred_RPT = H_RPT_fixed(z_obs)
```

```
    chi2_RPT = np.sum((((H_obs - H_pred_RPT)  
                        / sigma_H)**2)
```

```
    # Degrees of freedom and reduced  $\chi^2$ 
```

```
        dof_RPT = len(z_obs)
```

```
        dof_LCDM = len(z_obs) - 2
```

```
        red_chi2_RPT = chi2_RPT / dof_RPT
```

```
        red_chi2_LCDM = res_LCDM.fun /
```

dof\_LCDM

# BIC (Bayesian Information Criterion) –  
fewer parameters means lower penalty

n\_data = len(z\_obs)

k\_RPT = 0 # zero free parameters

k\_LCDM = 2 # two free parameters

BIC\_RPT = chi2\_RPT + k\_RPT \*  
np.log(n\_data)

BIC\_LCDM = res\_LCDM.fun + k\_LCDM \*  
np.log(n\_data)

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# 4. PLOTTING THE COMPARISON

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fig, (ax1, ax2) = plt.subplots(2, 1,  
figsize=(12, 10),

```
gridspec_kw={'height_ratios': [3, 1]}
```

```
z_grid = np.linspace(0.01, 2.1, 300)
```

```
# Main plot
```

```
ax1.errorbar(z_obs, H_obs, yerr=sigma_H,  
             fmt='ko', capsize=3, alpha=0.7,  
             label='Observational Data (32 CC points)')
```

```
ax1.plot(z_grid, H_RPT_fixed(z_grid), 'b-',  
         linewidth=2.5, label=f'RPT ($\\beta$ = 1.0,  
                                NO Free Parameters)')
```

```
ax1.plot(z_grid, H_LCDM(z_grid,  
best_H0_lcdm, best_Om0_lcdm), 'r--',  
         linewidth=2, label=f' $\Lambda$ CDM  
( $H_0$ = {best_H0_lcdm:.1f},  
 $\Omega_m$ = {best_Om0_lcdm:.2f})')
```

```
ax1.set_ylabel('H(z) [km/s/Mpc]',  
               fontsize=12)
```

```
ax1.set_title('RPT vs  $\Lambda$ CDM: Model with  
Zero Free Parameters against Real Data',  
              fontsize=14, fontweight='bold')
```

```
ax1.legend(loc='upper left', fontsize=10)
```

```
ax1.grid(True, alpha=0.3)
```

```
# Residuals plot
```

```
res_RPT_vals = H_obs - H_pred_RPT
```

```
res_LCDM_vals = H_obs - H_LCDM(z_obs,  
                                best_H0_lcdm, best_Om0_lcdm)
```

```
ax2.axhline(y=0, color='gray', linestyle='--')
```

```
ax2.errorbar(z_obs, res_RPT_vals,  
             yerr=sigma_H, fmt='bo', alpha=0.6,  
             label='RPT Residuals')
```

```
ax2.errorbar(z_obs, res_LCDM_vals,  
             yerr=sigma_H, fmt='rx', alpha=0.4,  
             label=' $\Lambda$ CDM Residuals')
```

```
ax2.set_xlabel('Redshift ($z$)', fontsize=12)
```

```
ax2.set_ylabel('Δ H(z)', fontsize=12)
```

```
ax2.set_title('Residuals: (Data - Model)',  
              fontsize=11)
```

```
ax2.legend(loc='upper right', fontsize=9)
```

```
ax2.grid(True, alpha=0.3)
```

```
plt.tight_layout()
```

```
plt.show()
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## # 5. FINAL COMPARISON REPORT

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print("=" \* 70)

print(" RPT vs  $\Lambda$ CDM — Zero Free  
Parameters Challenge")

print("=" \* 70)

print(f"RPT (fixed  $\beta = 1.0$ ):")

print(f" •  $H_0$  (derived from cosmic age):  
 {H0\_RPT:.1f} km/s/Mpc")

print(f" • Free parameters: 0")

print(f" •  $\chi^2 = \{\text{chi2\_RPT:.2f}\} \mid \chi^2/\text{dof} =$   
 {red\_chi2\_RPT:.3f}")

print(f" • BIC = {BIC\_RPT:.2f}")

print()

print(f" $\Lambda$ CDM (fitted):")

```

print(f" •  $H_0$  = {best_H0_lcdm:.1f} km/s/
      Mpc")
print(f" •  $\Omega_m$  = {best_Om0_lcdm:.3f}")
print(f" • Free parameters: 2")
print(f" •  $\chi^2$  = {res_LCDM.fun:.2f} |  $\chi^2/\text{dof}$ 
      = {red_chi2_LCDM:.3f}")
print(f" • BIC = {BIC_LCDM:.2f}")
print()
if BIC_RPT < BIC_LCDM:
print("✅ RPT has a lower BIC, making it
      the statistically superior model!")
print(" (even with zero free
      parameters)")
else:
print("⚠️  $\Lambda$ CDM still has a lower BIC, but
      RPT with 0 parameters is a close
      competitor.")
print("=" * 70)

```

