

High-energy monitoring of NGC 4593 with XMM-Newton and NuSTAR

Francesco Ursini



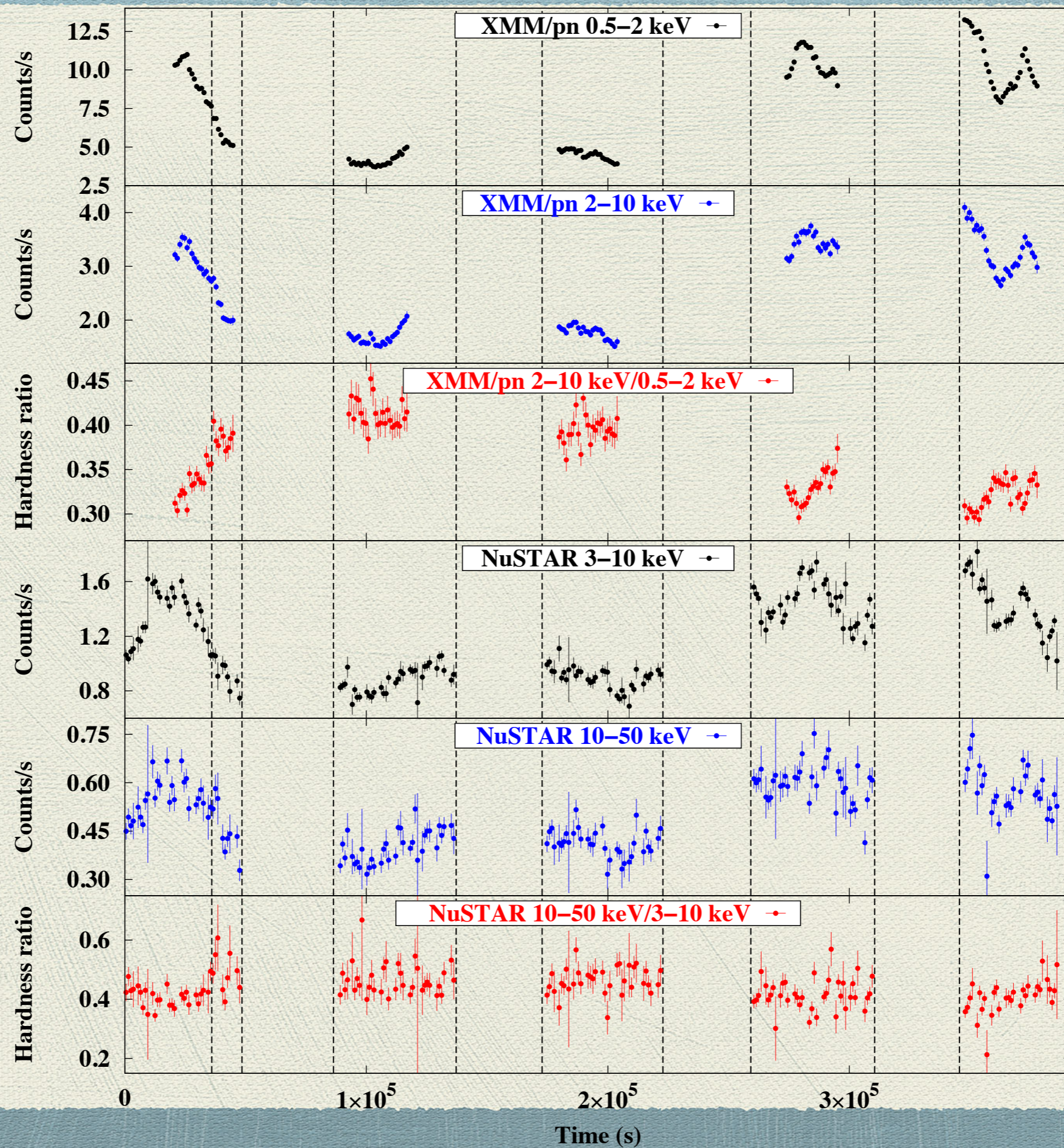
P.-O. Petrucci, G. Matt, S. Bianchi, M. Cappi, B. De Marco,
A. De Rosa, J. Malzac, A. Marinucci, G. Ponti, A. Tortosa

AGN 12

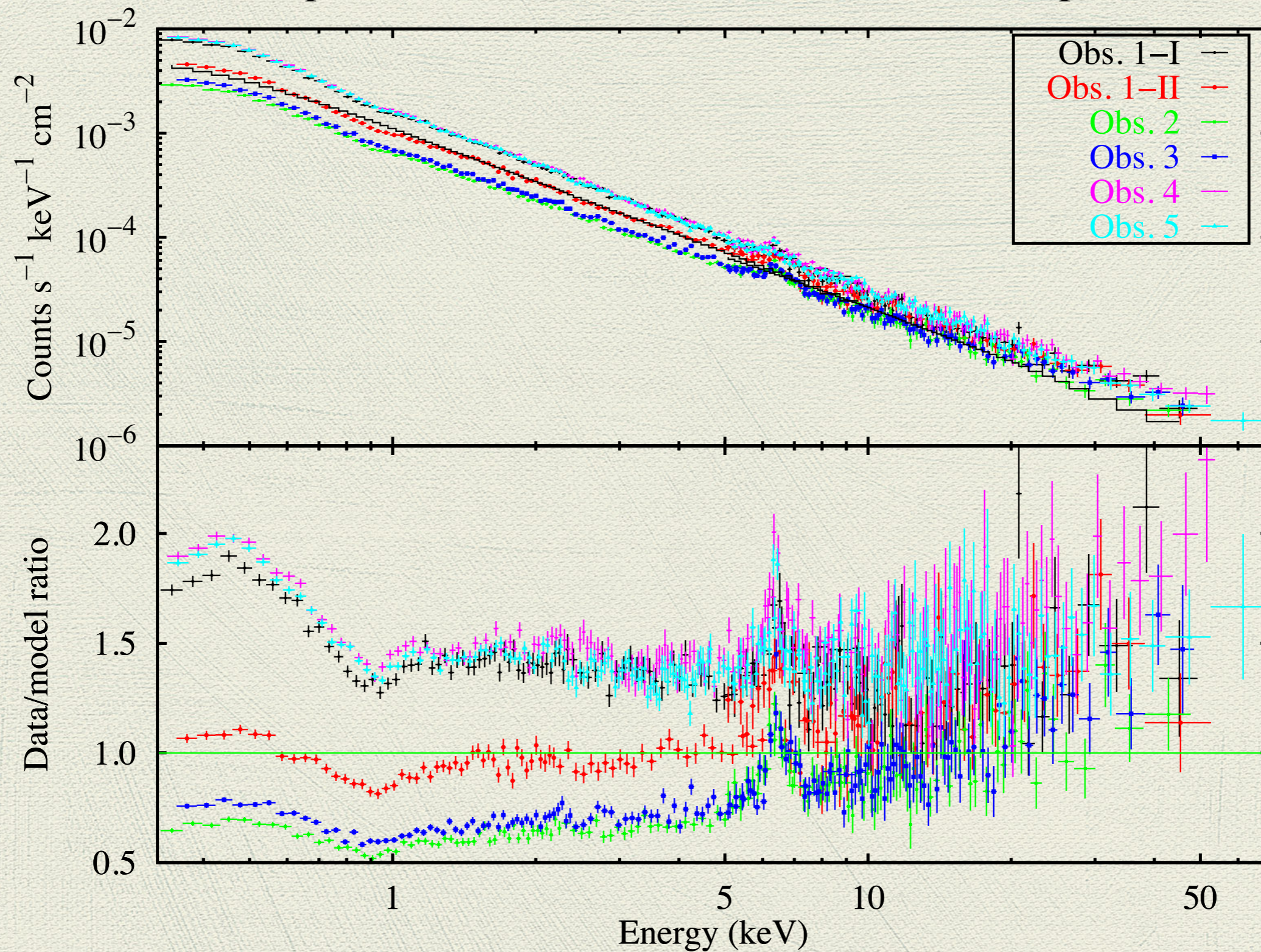
A joint XMM+NuSTAR monitoring

Obs.	Satellites	Obs. Id.	Start time (UTC) yyyy-mm-dd	Net exp. (ks)
1	<i>XMM-Newton</i> <i>NuSTAR</i>	0740920201 60001149002	2014-12-29	16 22
2	<i>XMM-Newton</i> <i>NuSTAR</i>	0740920301 60001149004	2014-12-31	17 22
3	<i>XMM-Newton</i> <i>NuSTAR</i>	0740920401 60001149006	2015-01-02	17 21
4	<i>XMM-Newton</i> <i>NuSTAR</i>	0740920501 60001149008	2015-01-04	15 23
5	<i>XMM-Newton</i> <i>NuSTAR</i>	0740920601 60001149010	2015-01-06	21 21

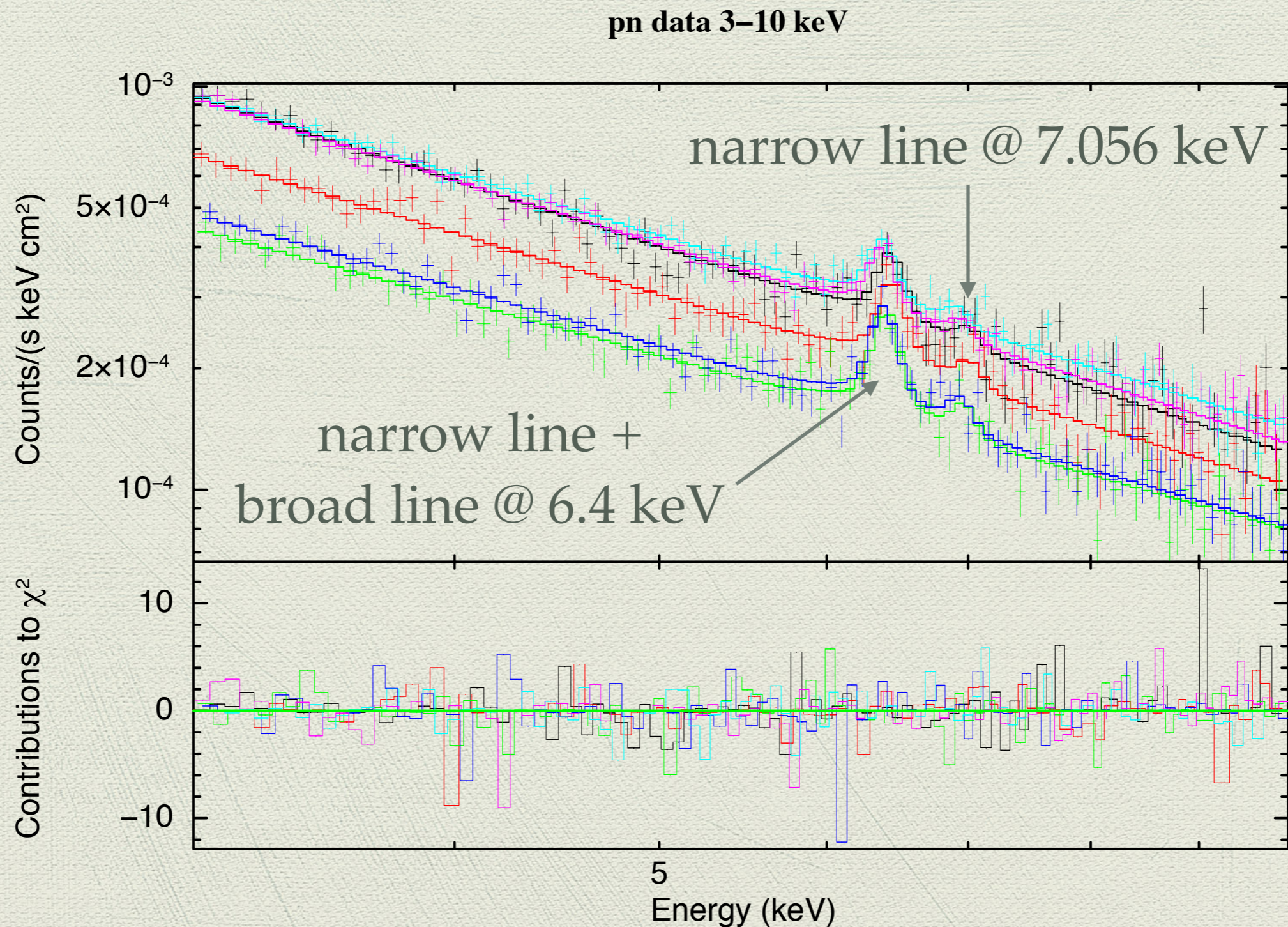
XMM/pn and NuSTAR/FPMA+FPMB light curves and hardness ratios



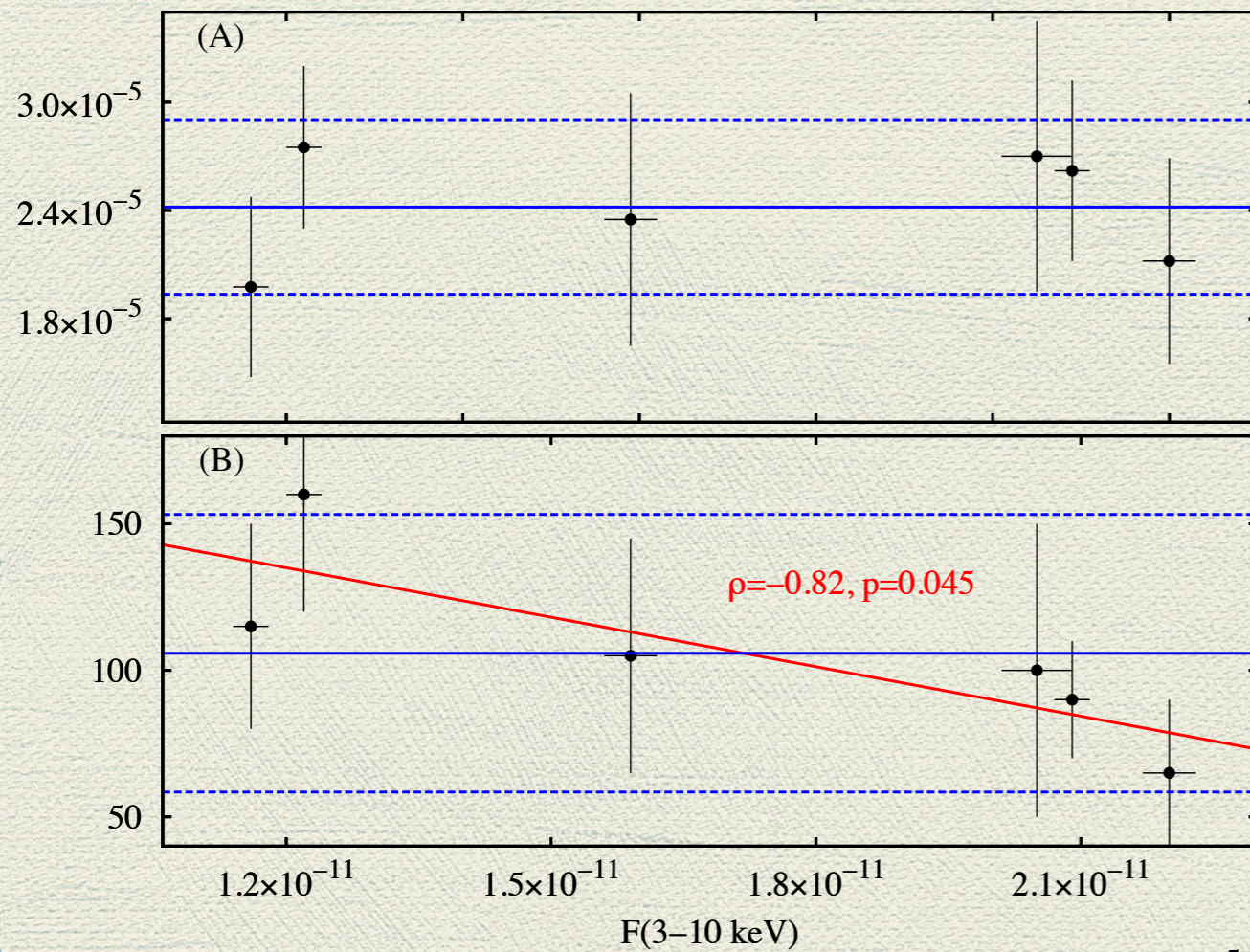
XMM/pn and NuSTAR/FPMA data fitted with a power law



The iron line(s)

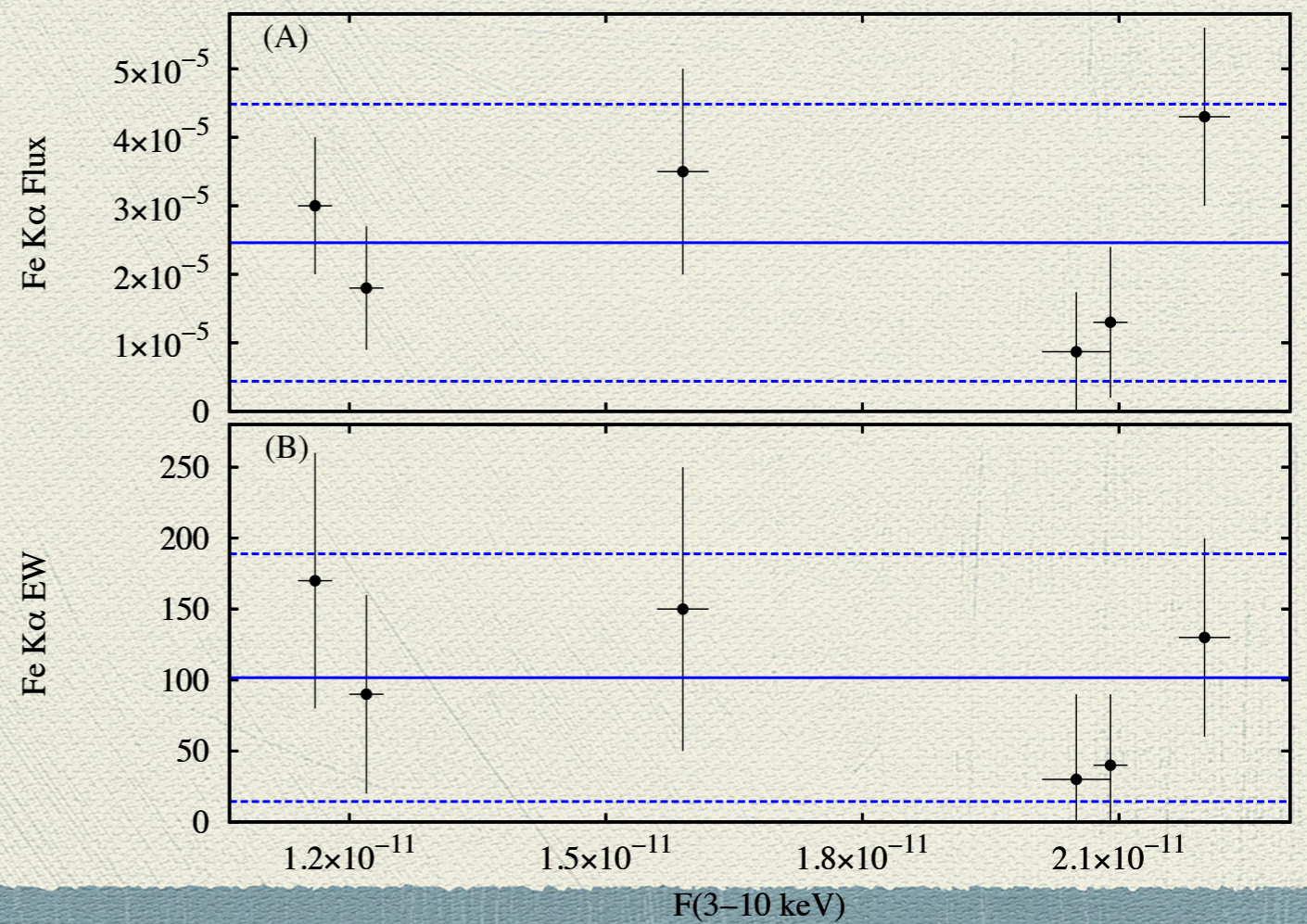


Narrow Fe K α line flux and EW versus primary flux



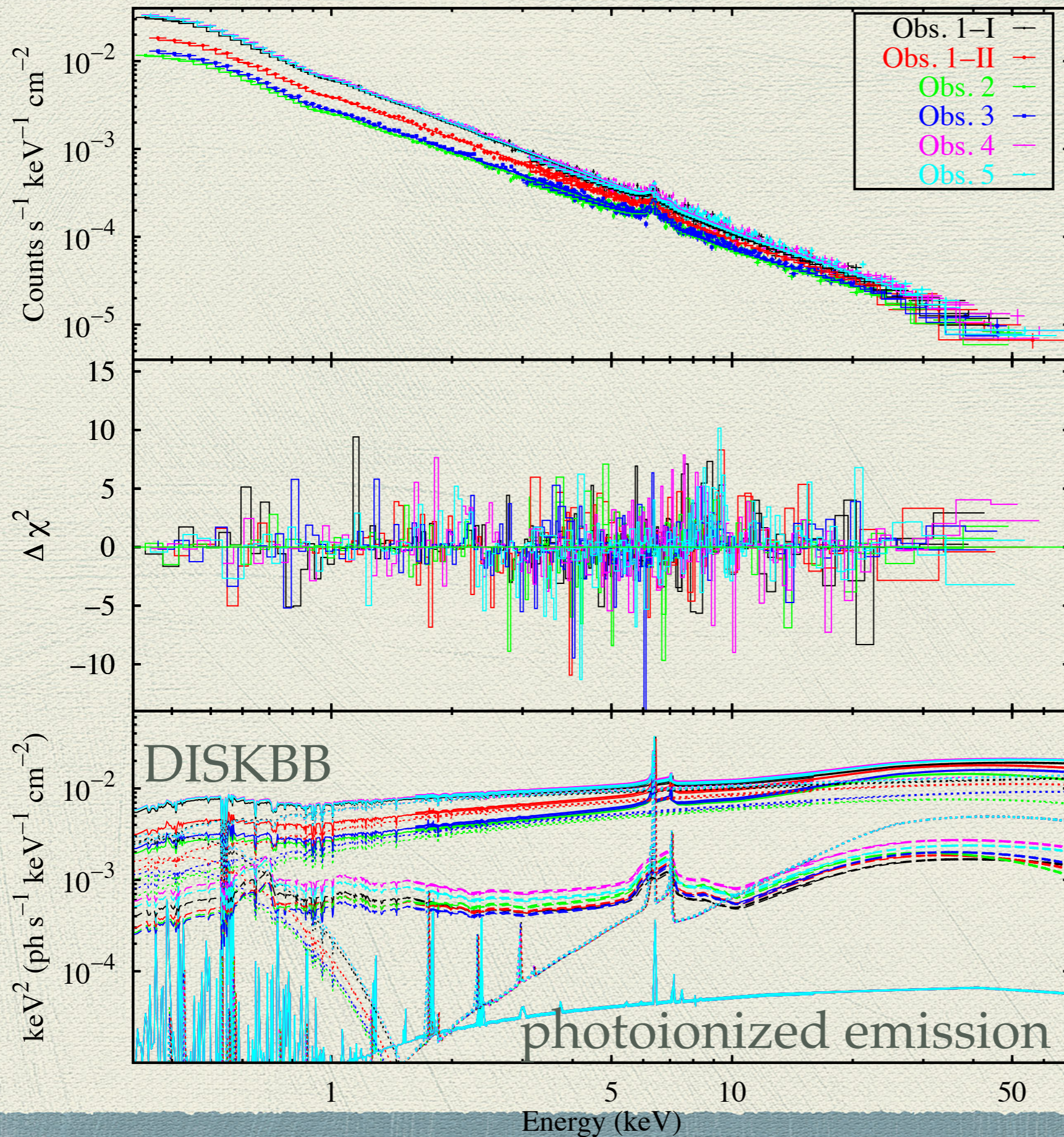
Constant narrow Fe K alpha line

Broad Fe K α line flux and EW versus primary flux



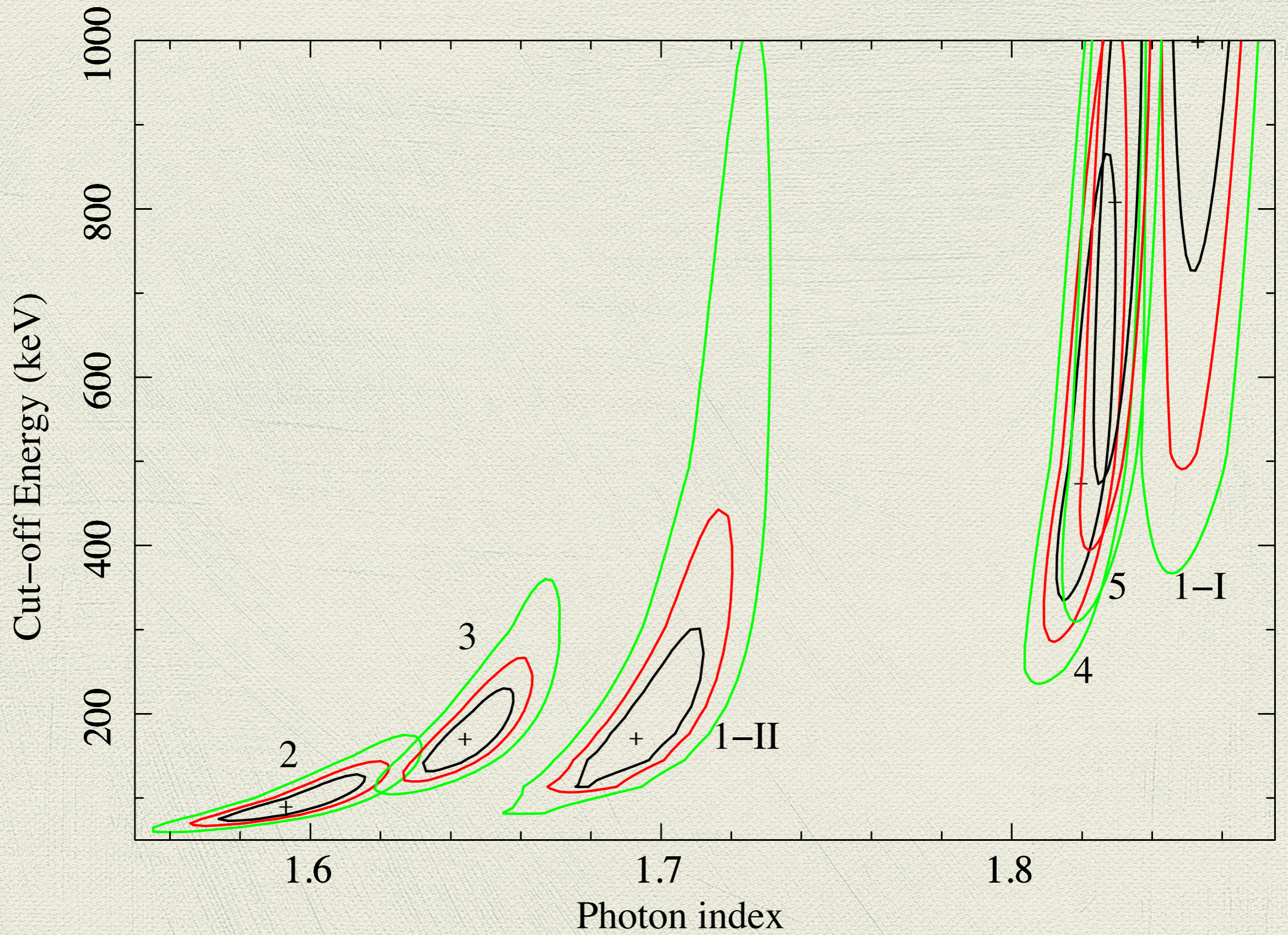
E (keV)	σ (eV)	average flux	average EW (eV)
6.4 (narrow)	0	2.42	106
6.4 (broad)	300^{+130}_{-70}	2.36	102
7.056	0	0.6	30

XMM/pn and NuSTAR data with best-fitting model

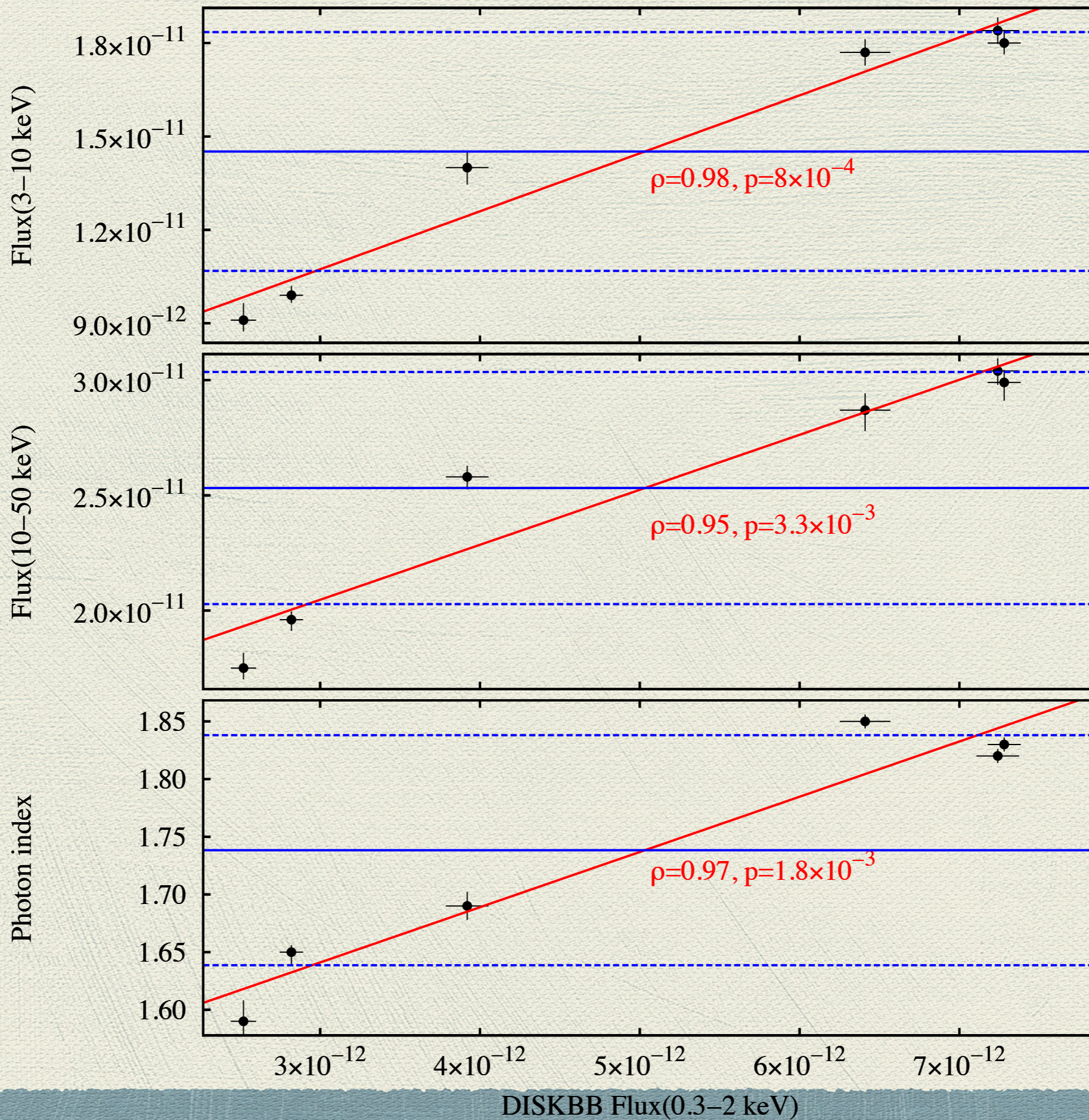


Variable primary
cut-off power law
+
2 reflection
components
+
soft excess

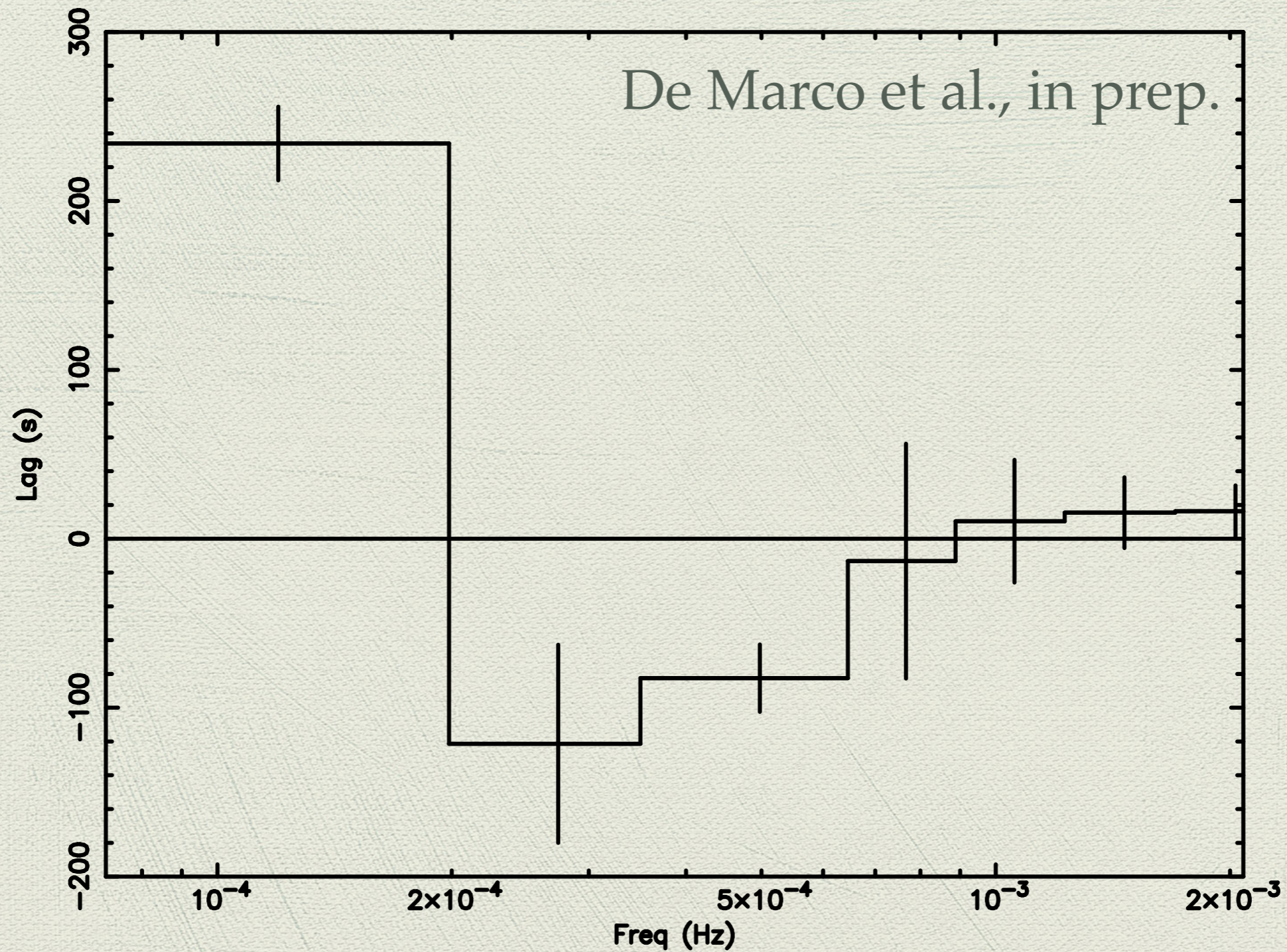
cut-off PL
XILLVER ($R_s \sim 0.3-0.6$)
RELXILL ($R_s \sim 0.2$)
($R_{in} = 40 R_g$)



Correlation between soft excess and primary emission



Timing



Main results

- ◆ Remarkable variability, both in flux and spectral shape over ~days and down to ~ks
- ◆ Significant variations of Gamma (1.6-1.8) and cut-off (~100 keV up to >500 keV): temperature / optical depth variations?
- ◆ 2 reflection components, giving rise to a narrow and a broad Fe K alpha lines. One (XILLVER) is from neutral and distant matter, one (RELXILL) from an ionized disc with $R_{in} \sim 40 R_g$
- ◆ Soft excess correlated with primary emission: warm Comptonization?
Link with the UV? (in progress) →
- ◆ See Ursini et al. 2016, MNRAS, 463, 382

