What can we learn about stellar activity cycles from ZDI?

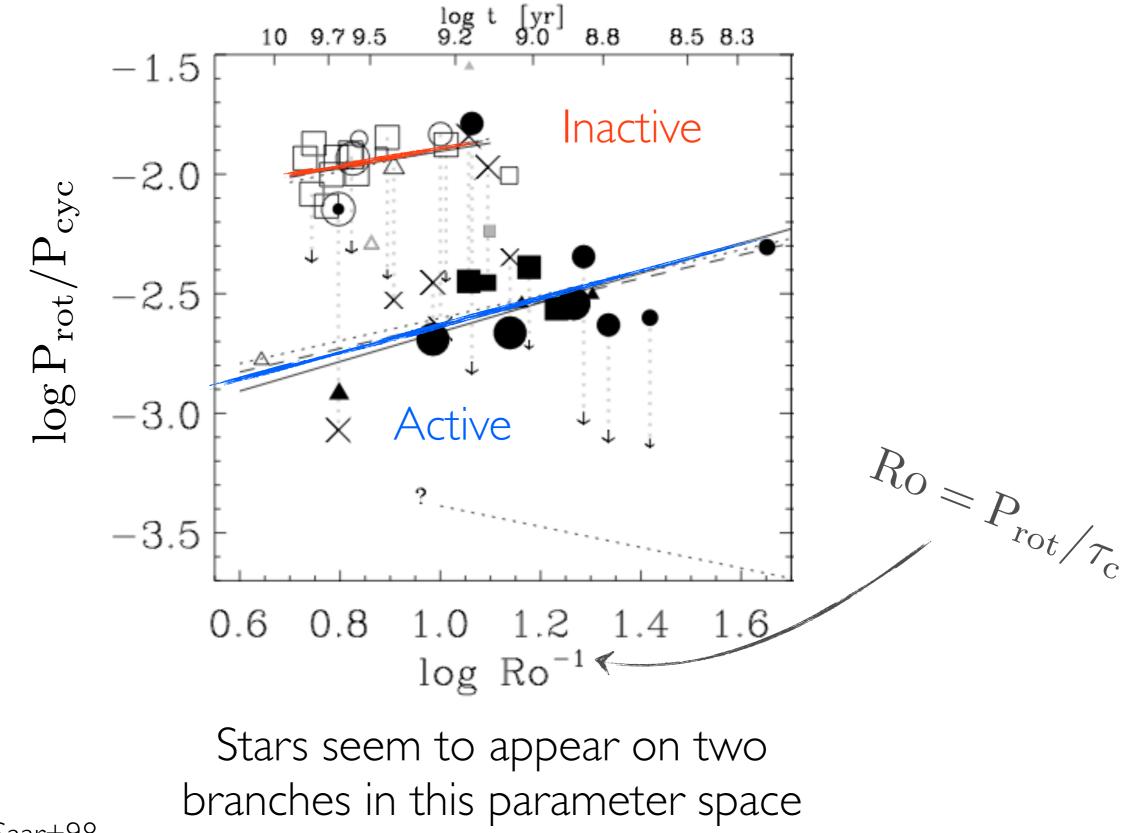
Victor See

Cool Stars 19, Uppsala

M. Jardine, A. Vidotto, J.-F. Donati, S. Boro Saikia, J. Bouvier, R. Fares, C. Folsom, S. Gregory, G. Hussain, S. Jeffers, S. Marsden, J. Morin, C. Moutou, J.D. do Nascimento, P. Petit, I. Waite + BCool

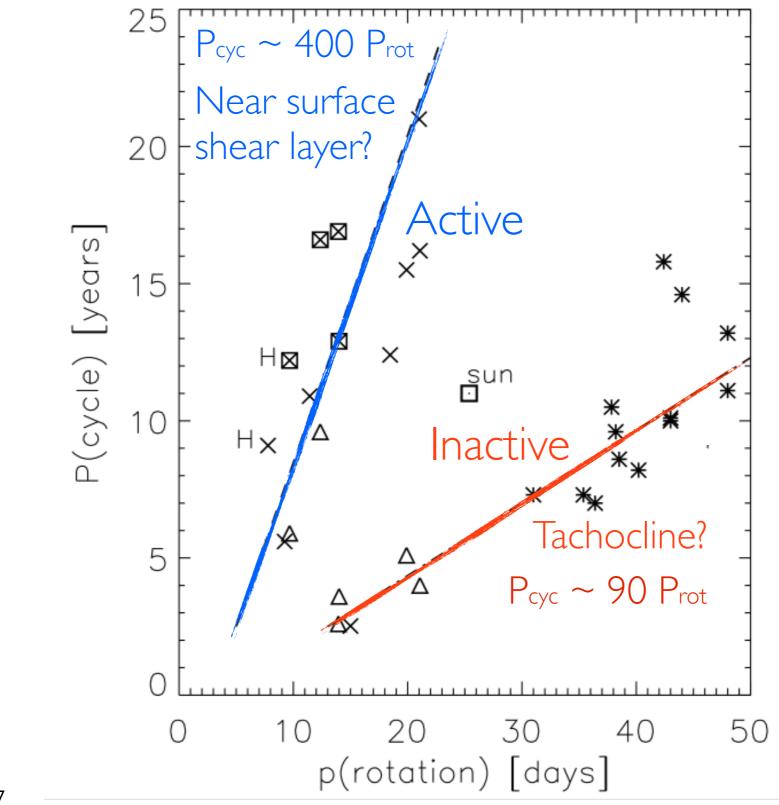


Activity branches



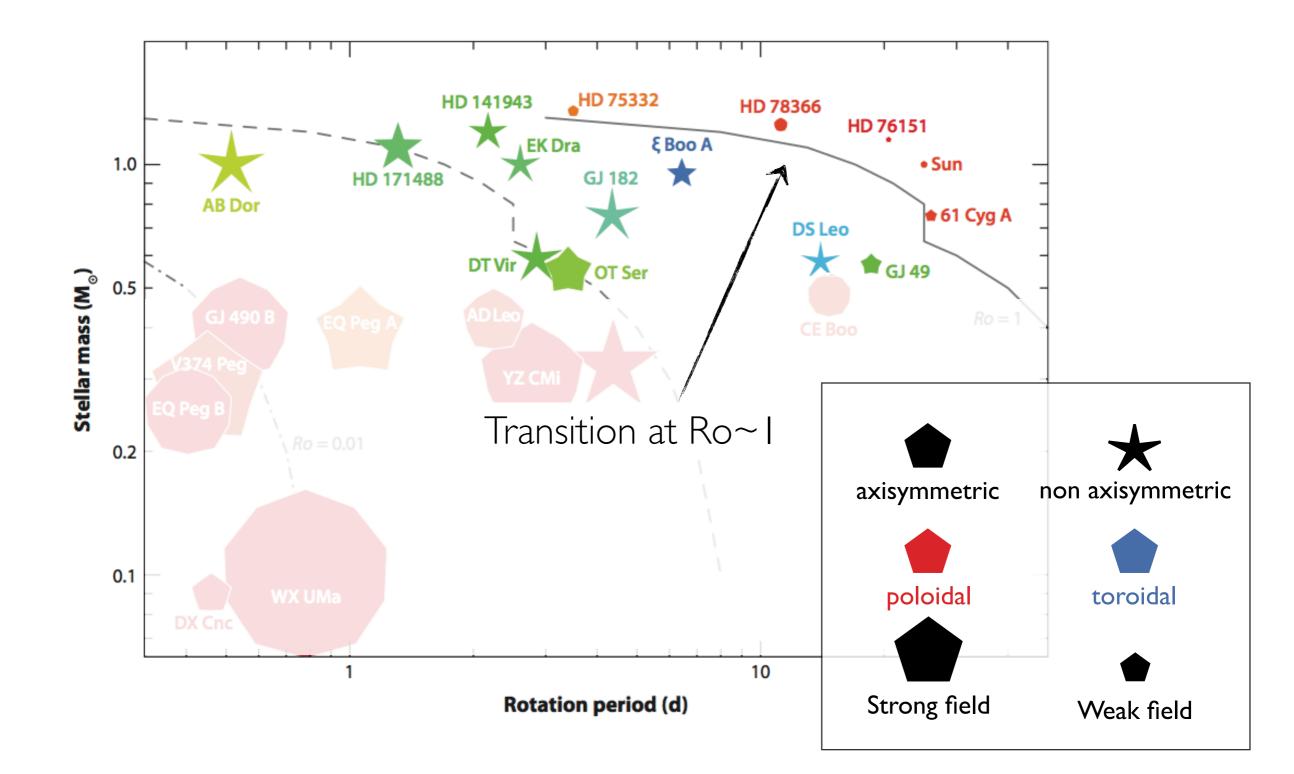
Baliunas+95, Saar+98

Activity branches



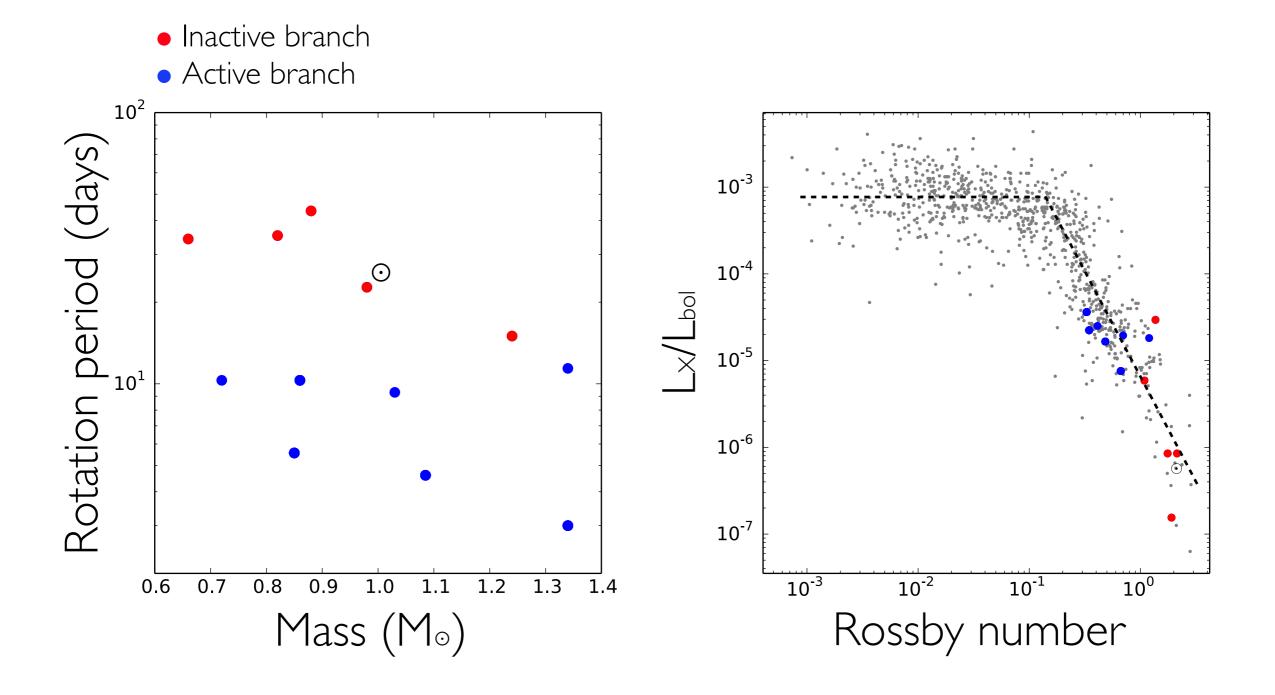
Bohm-Vitense 07

Zeeman-Doppler imaging



Donati and Landstreet 09

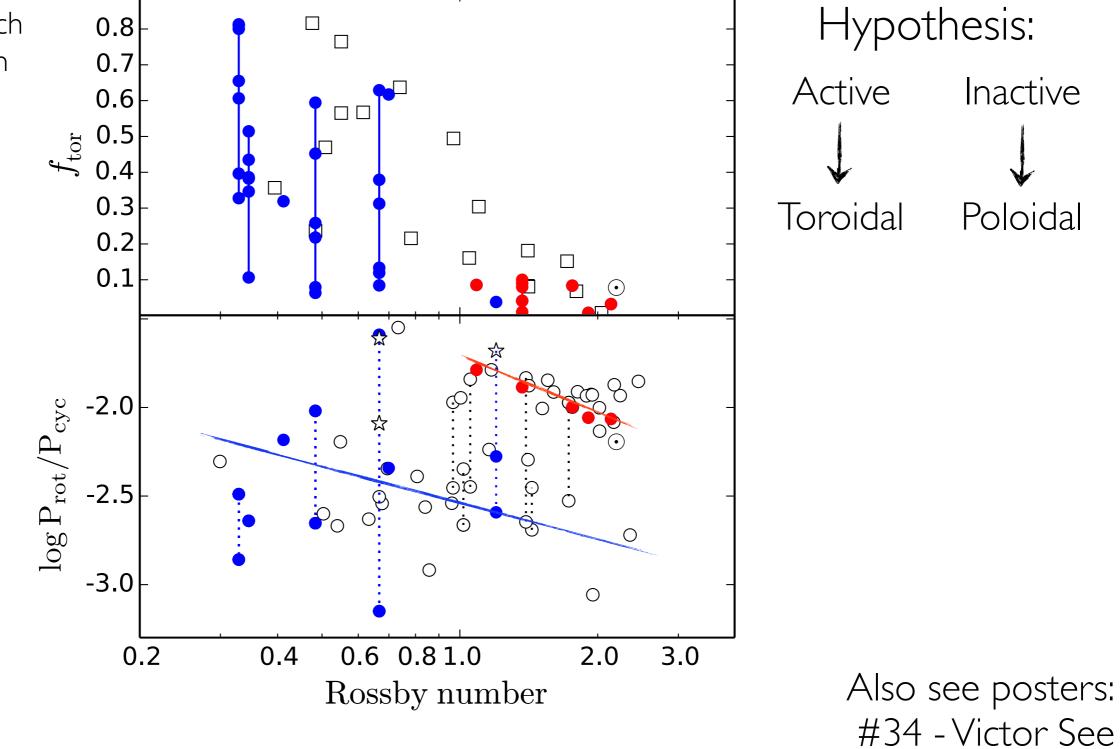
Sample + BCool collaboration



Wright+II, See+(submitted)

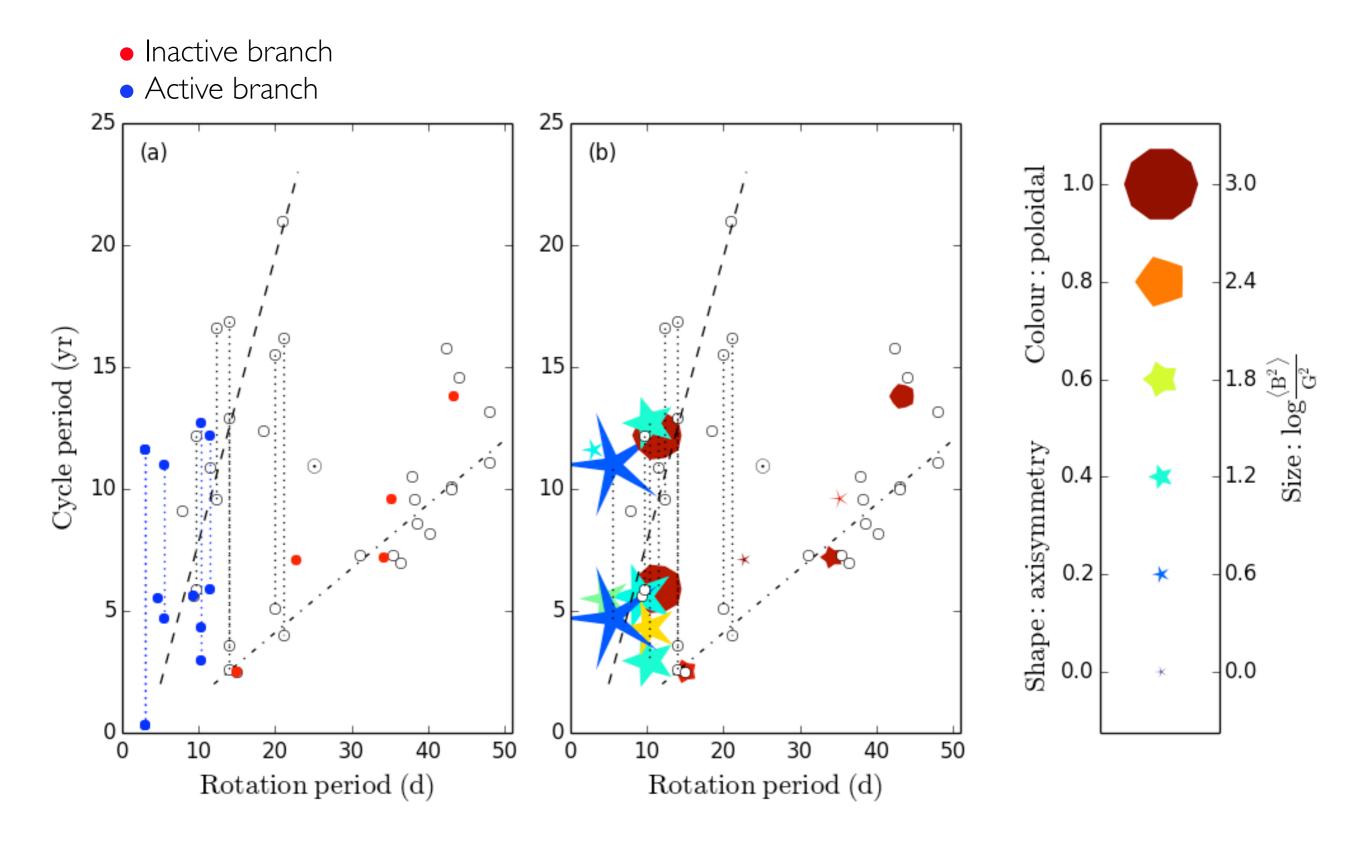
Field geometry on branches

- Inactive branch
- Active branch



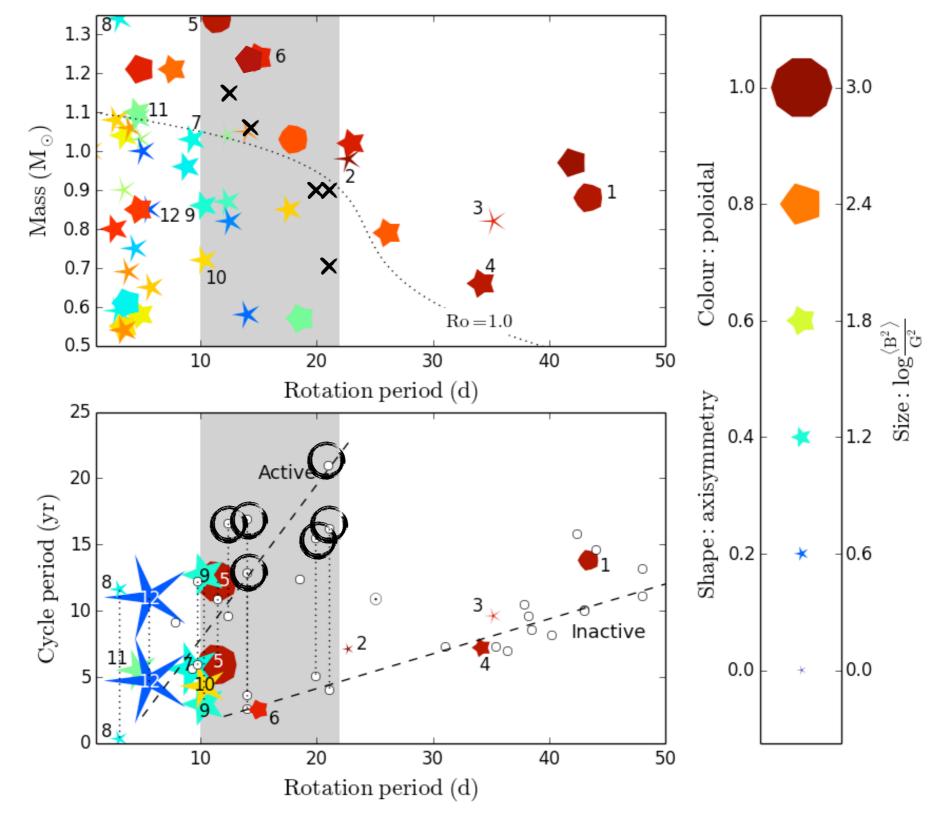
#233 - Lisa Lehmann

Field geometry on branches



Bohm-Vitense 07, See+(submitted)

Field geometry on branches



Bohm-Vitense 07, See+15, See+(submitted)

Summary

Field topologies appear to be different along different activity branches

This hypothesis is currently uncertain

Confirmation/rejection will require more ZDI mapping of stars where branches overlap

