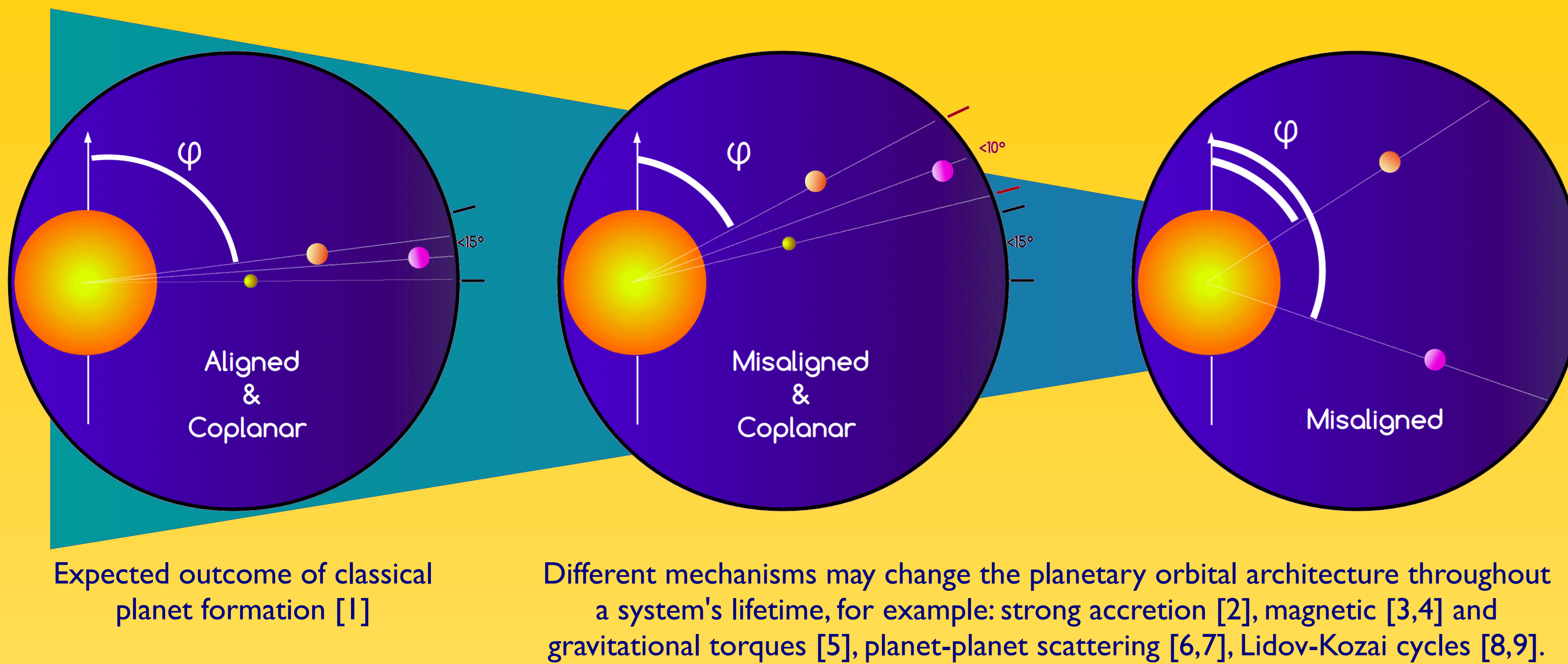


Spin-Orbit Alignment of Planetary Systems (SOAPS): The case of 64 *Kepler* planets and planet candidates

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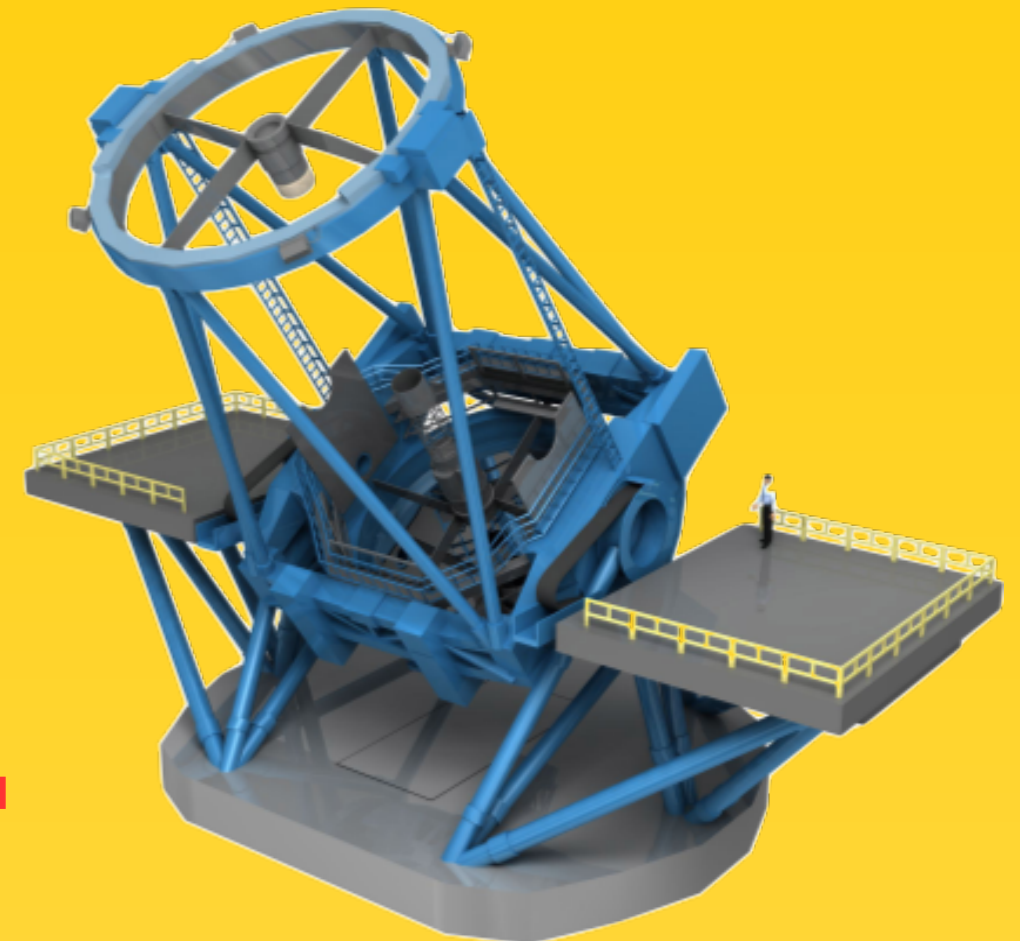
The **obliquity of a planet** is the physical property that may probe the history of formation and evolution of planetary systems describing their current orbital configuration.
SOAPS measured the obliquity in the line of sight (φ) for 34 *Kepler* planetary systems.



$$\varphi = |i_* - i_p|$$

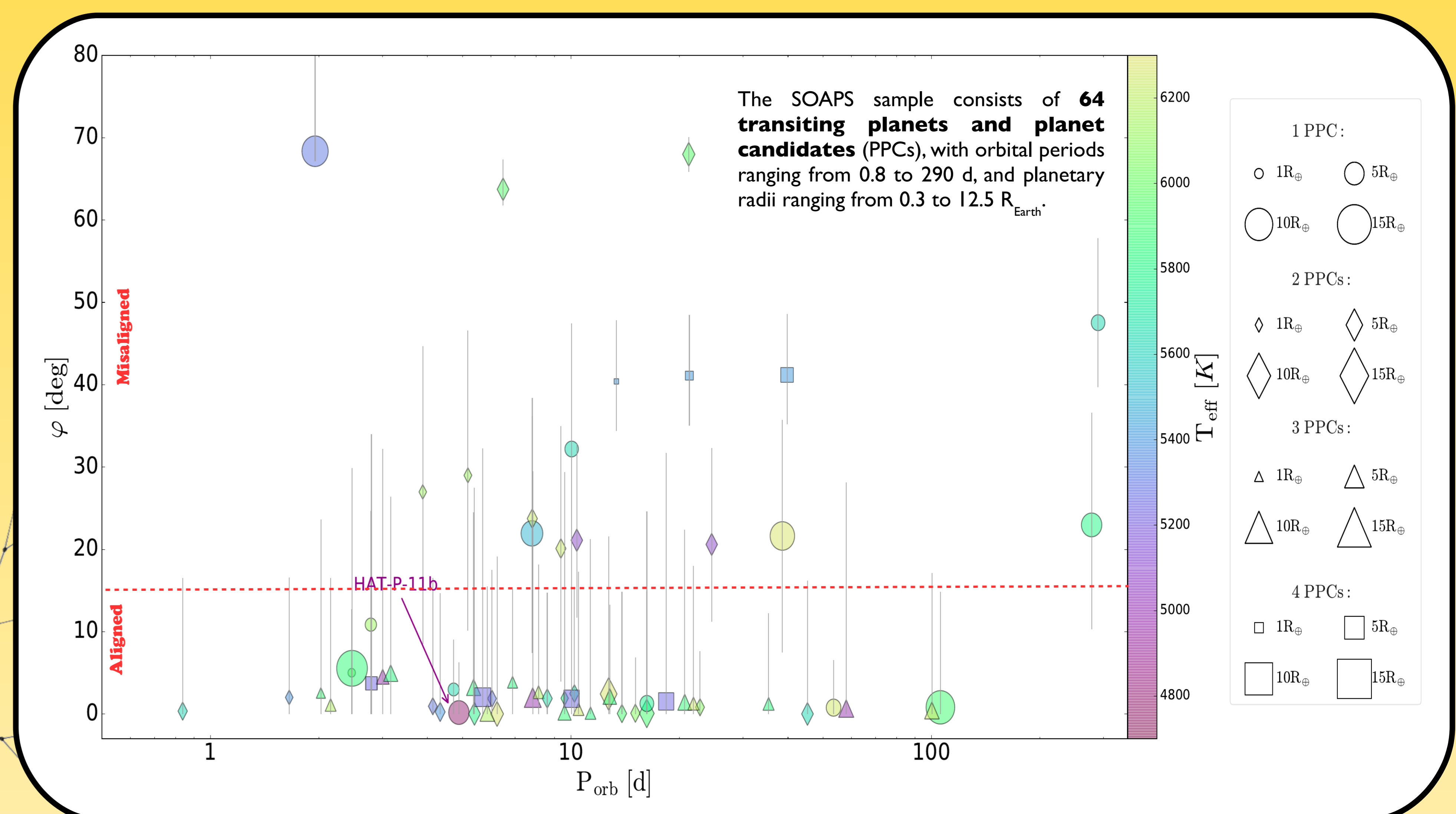
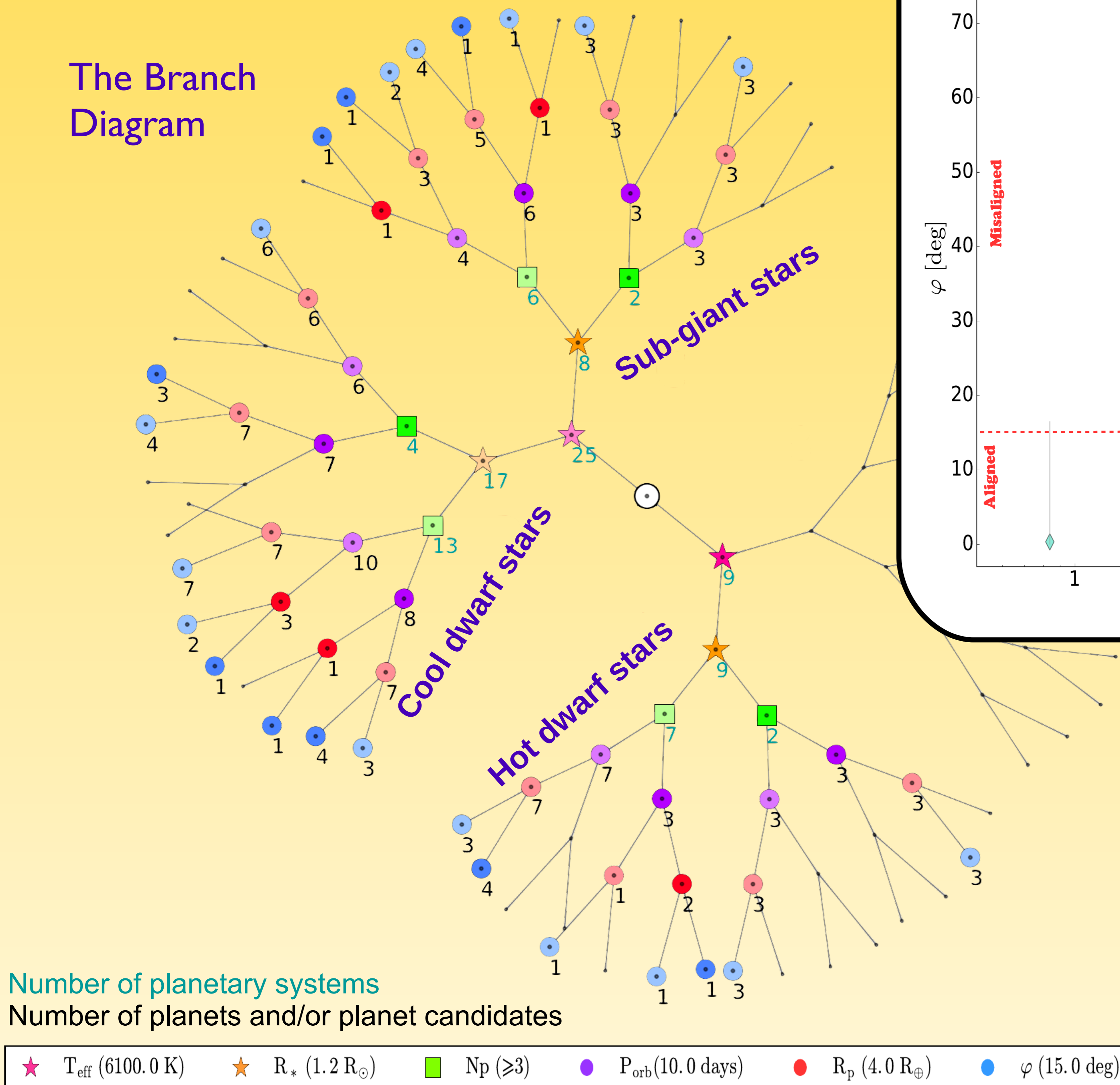
$$\sin i_* = \frac{v \sin i_* \cdot P_{rot}}{2\pi R_*}$$

Line of sight



The SOAPS measurements are based on ultra-high resolution échelle spectra acquired with Subaru+HDS ($R \sim 160,000$) for precise $v \sin i_*$ measurements, combined with robust stellar rotational periods (P_{rot}) and stellar radii (R_*) from the literature [10].

The Branch Diagram



Conclusions from SOAPS

Planet Multiplicity

- Multi-planet systems are coplanar.
- Systems with one (○) or two (◇) planets exhibit a large range of obliquities.
- Systems with three (△) or four (□) planets tend to be in aligned orbits.

Host Temperature

- Our results do not support that alignment occurs preferably for systems around relative cool stars over hotter stars.

Star's Evolution

- Hot and cool dwarf hosts: ~70% aligned
- Sub-giant hosts: 80% aligned

Stellar Companions

- We found no correlation with obliquity for stars with detected potential stellar companions and those without.

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