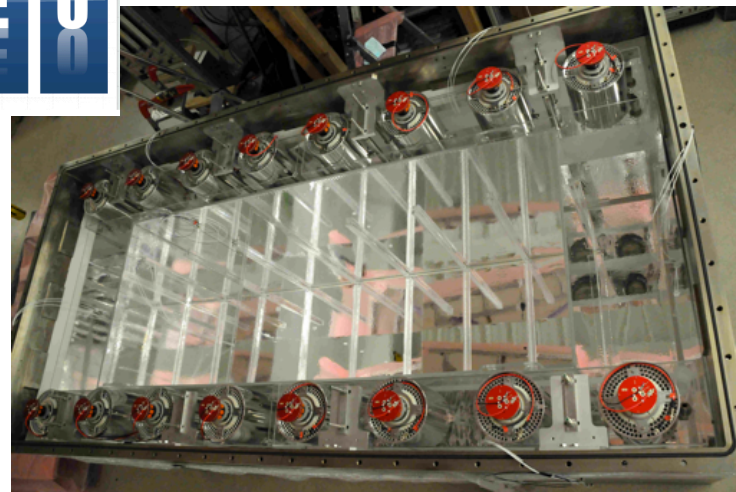
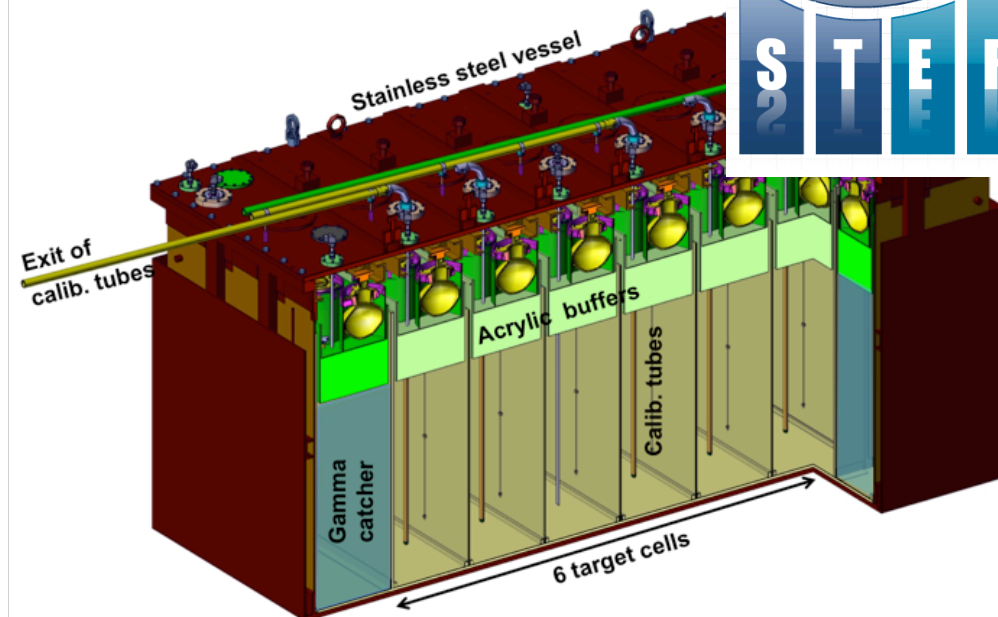




Calibration and energy scale in



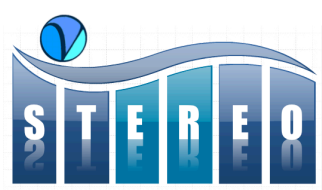
Pablo del Amo Sánchez

ESCAPE 2018

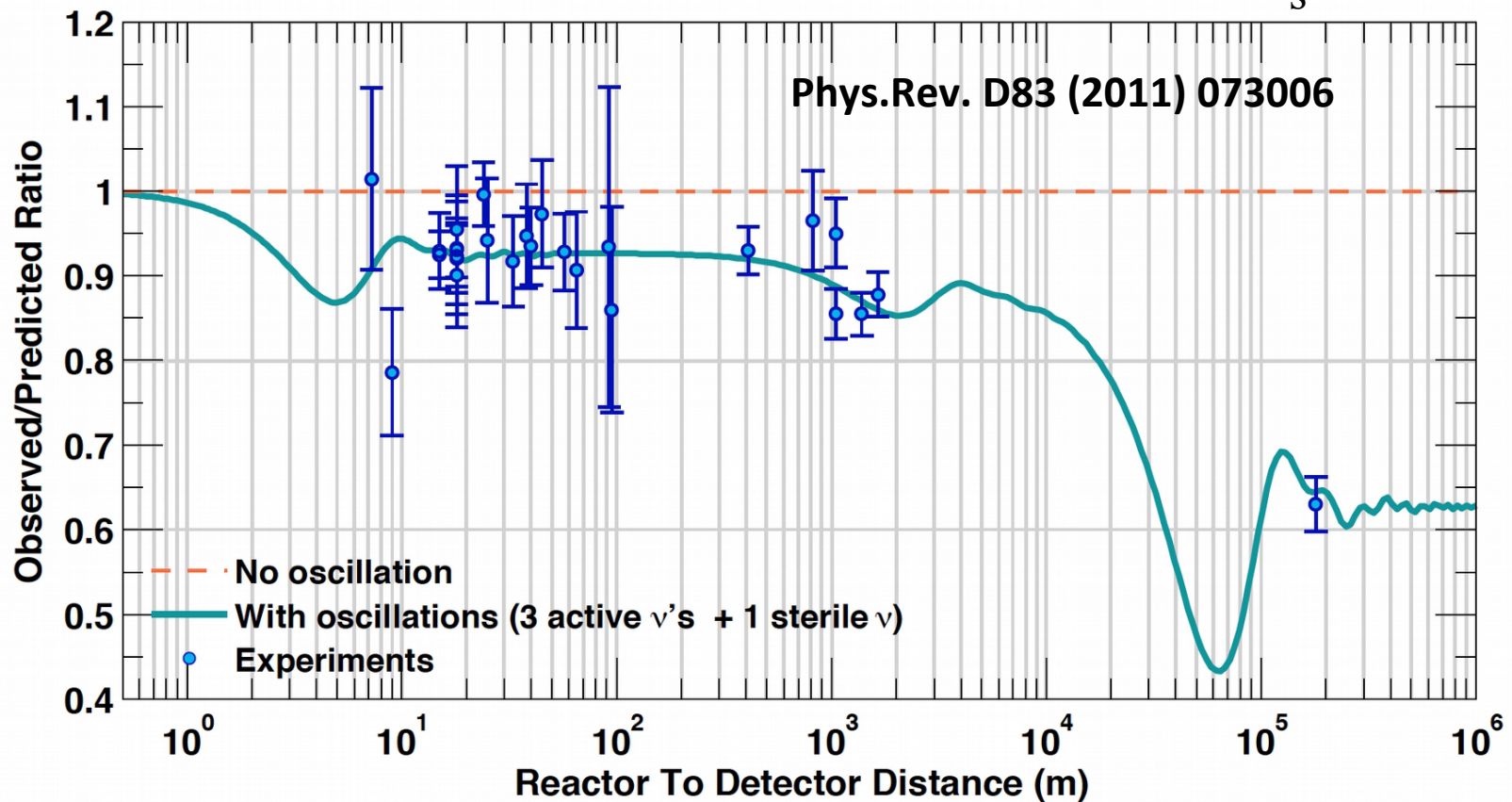
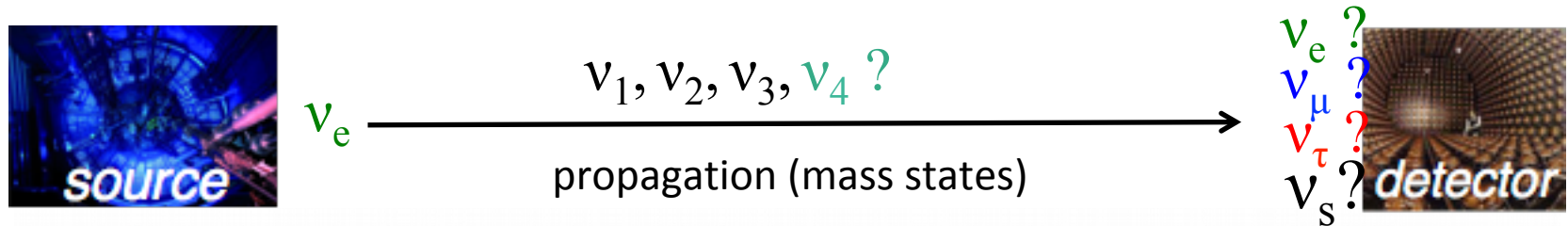
01/06/18





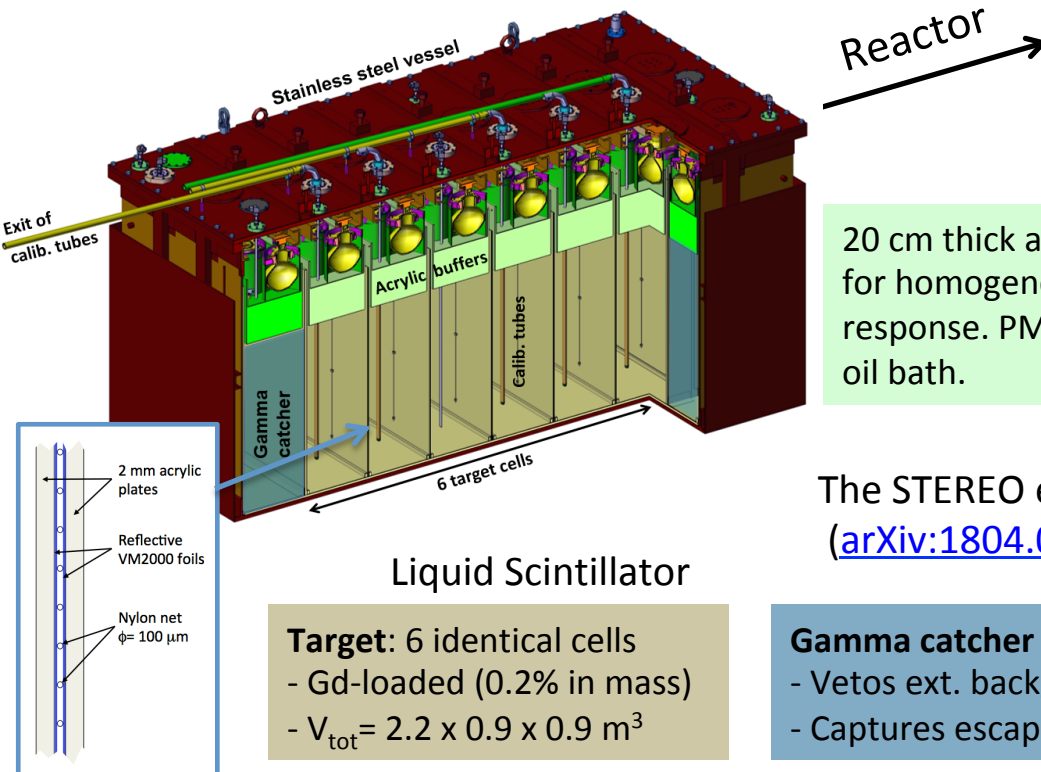


Reactor Anti- ν Anomaly

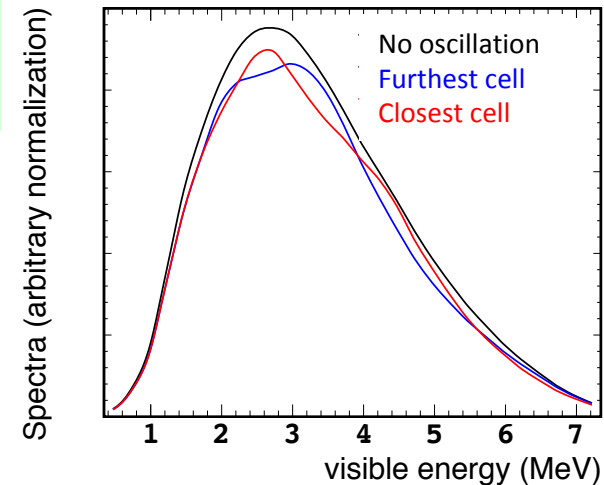
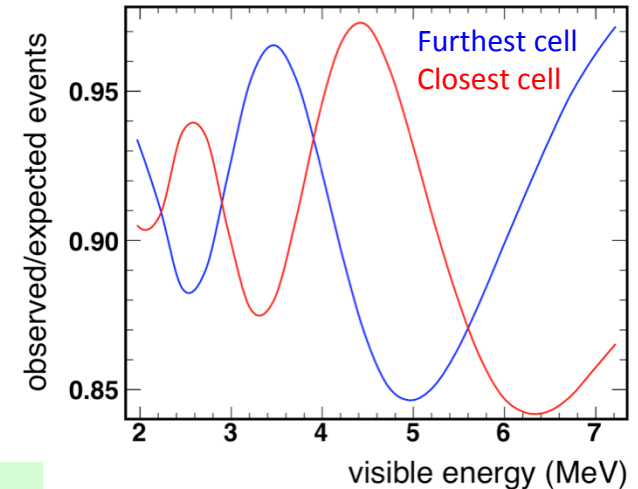


STEREO detector

- Compare 6 target cells to measure oscillation-driven distortions in the $E_{\bar{\nu}_e}$ spectrum.
- Mitigate sensitivity to predicted spectrum.



$$\sin^2(2\theta) = 0.14, \Delta m^2 = 2.4 \text{ eV}^2$$



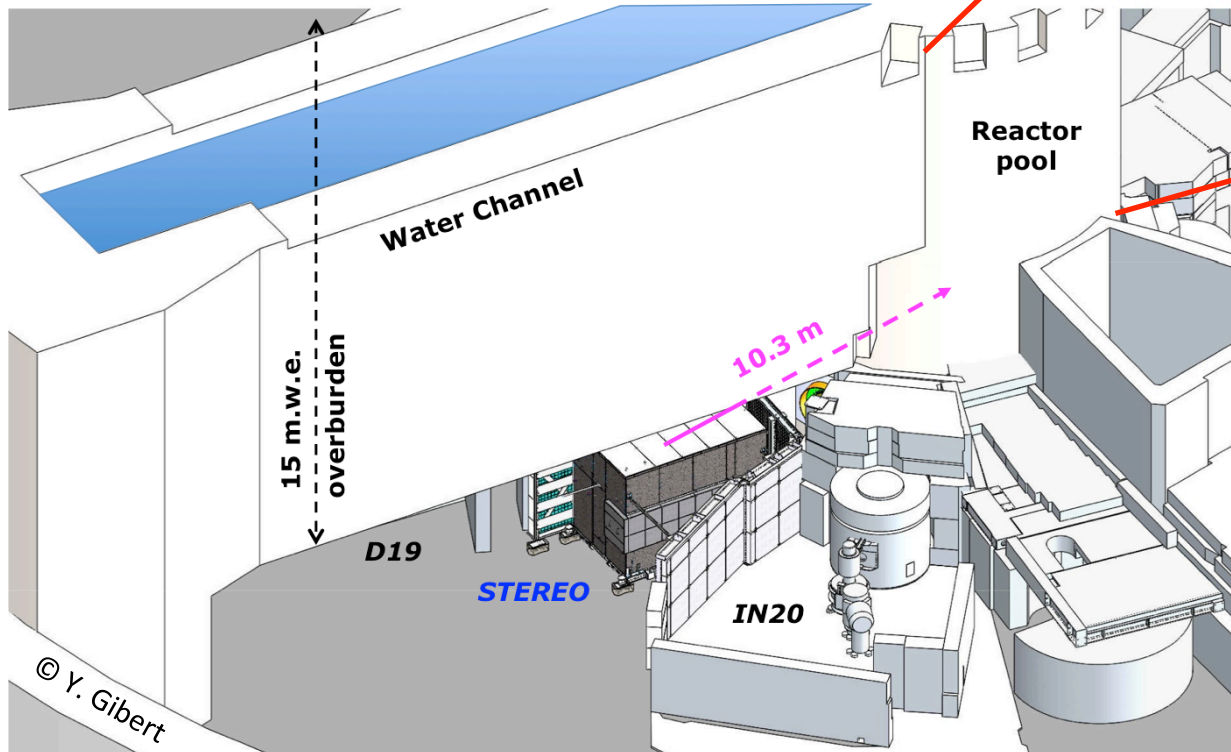
The STEREO experiment, ([arXiv:1804.09052](https://arxiv.org/abs/1804.09052)).



ILL Site

Compact core

- 58.3 MW_{thermal}
- Ø40 cm × 80 cm
- Highly enriched: 93% ²³⁵U
- 3-4 cycles/year each of 50 days
- 10^{19} s^{-1} pure $\bar{\nu}_e$ flux

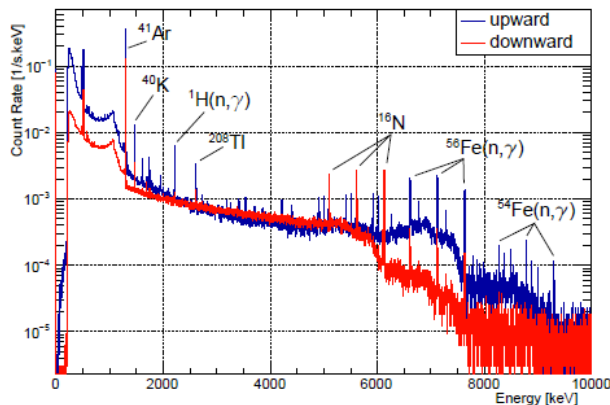


Challenging mitigation of the background generated by:

- Neighboring experiments.
- Cosmic-rays.

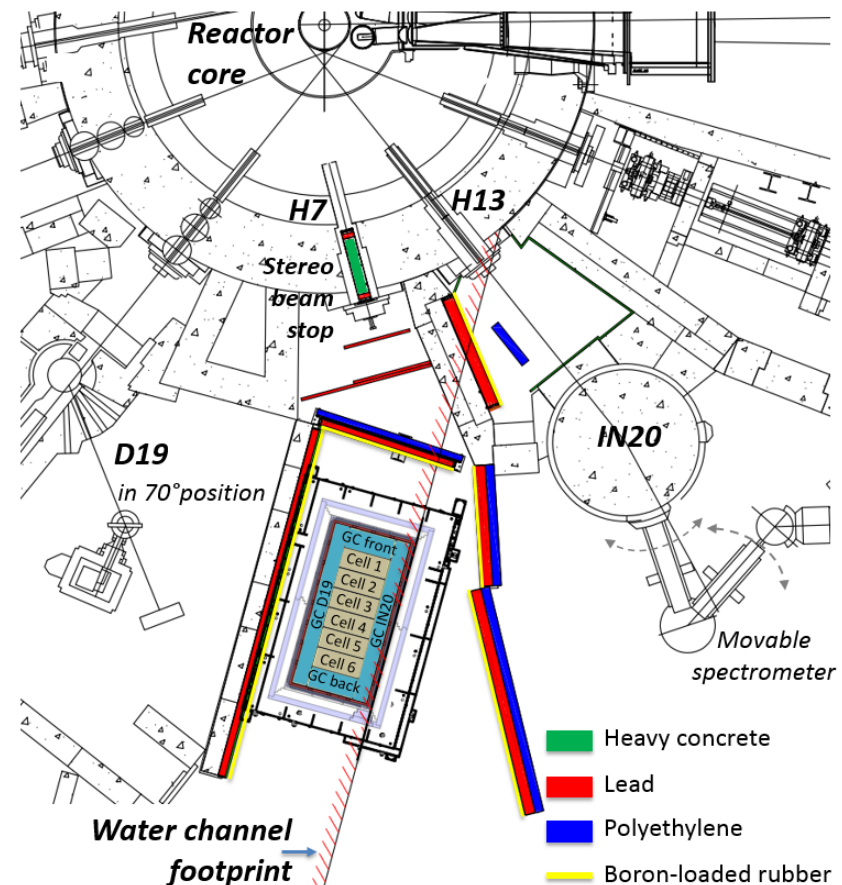
Reactor sources of background

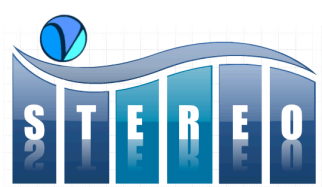
- Extraction of neutron beams for neighboring experiments.
- Extensive campaigns of characterization of n and γ sources before shielding design.



- High E γ 's from n-capture on metals, ^{41}Ar in and ^{16}O in primary water circuit
- Stray magnetic fields from IN20 magnets.

Heavy passive shielding added on front and side walls

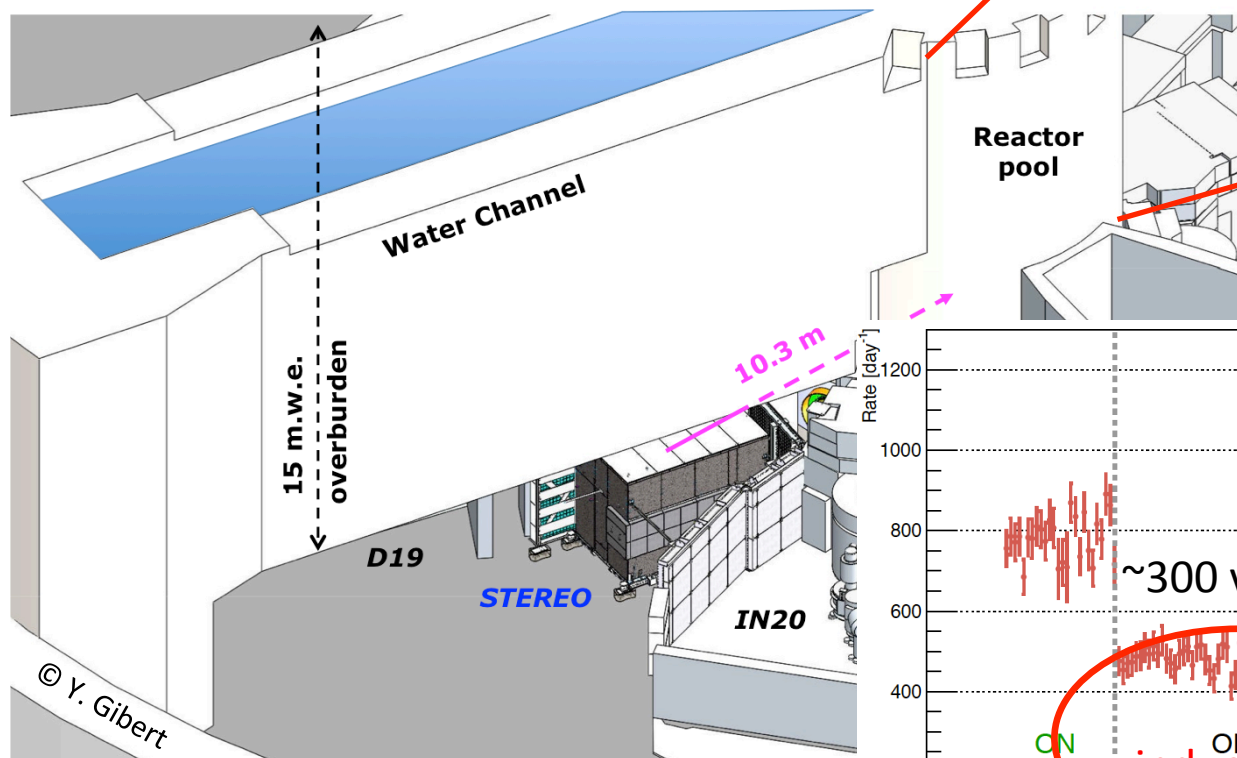
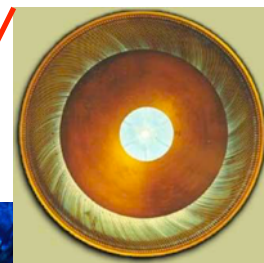




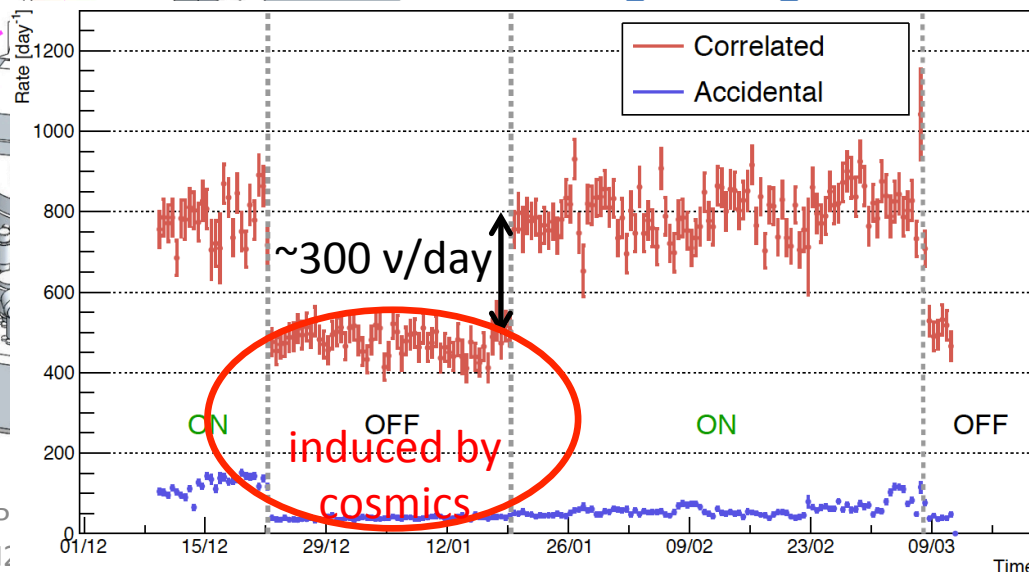
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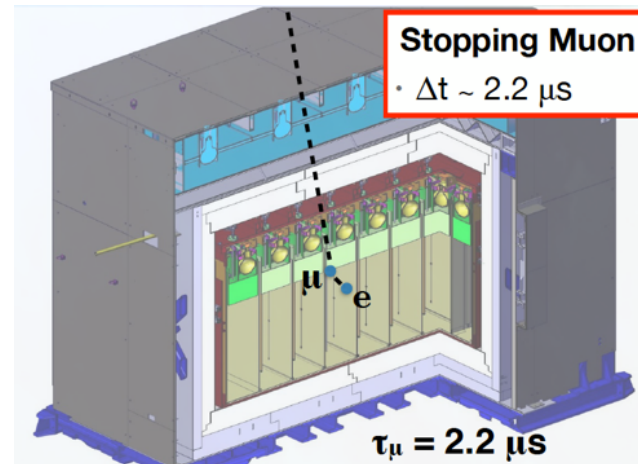
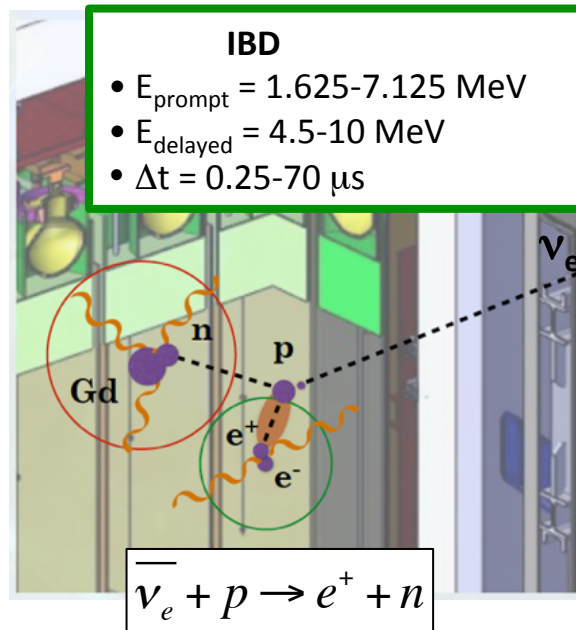


Challenging mitigation of the background generated

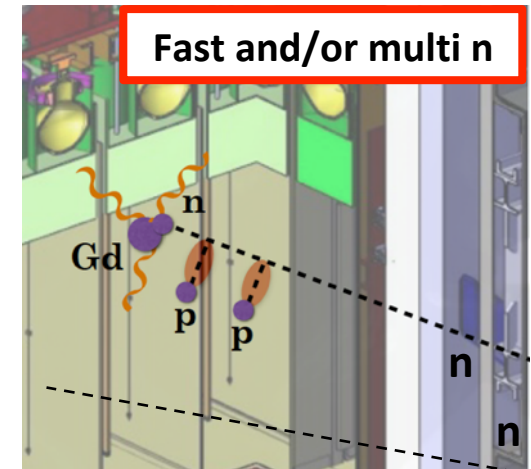


Selection cuts

Neutrino selection



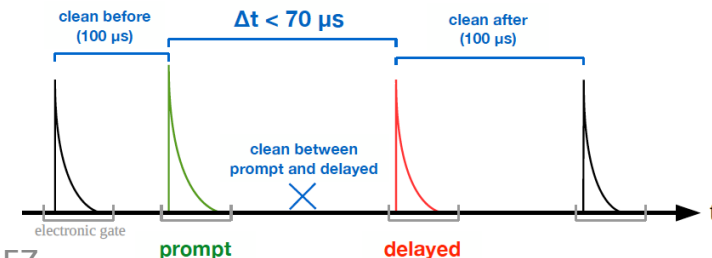
- Charge asymmetry per cell: $Q_{\text{max}}/Q_{\text{tot}} < 0.50$



- $\Delta t_{\text{last-}\mu} > 100 \mu\text{s}$
- Isolated prompt-delayed pair

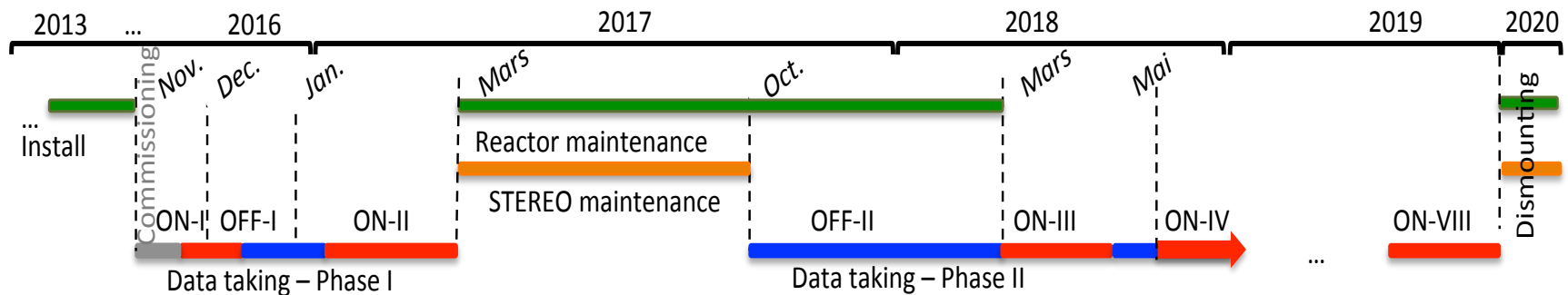
Topology cuts:

- E_{prompt} in γ -catcher $< 1.1 \text{ MeV}$
- E_{prompt} in neighboring cell $< 0.8 \text{ MeV}$
- E_{delayed} in target $> 1 \text{ MeV}$
- $D_{\text{prompt-delayed}} < 60 \text{ cm}$





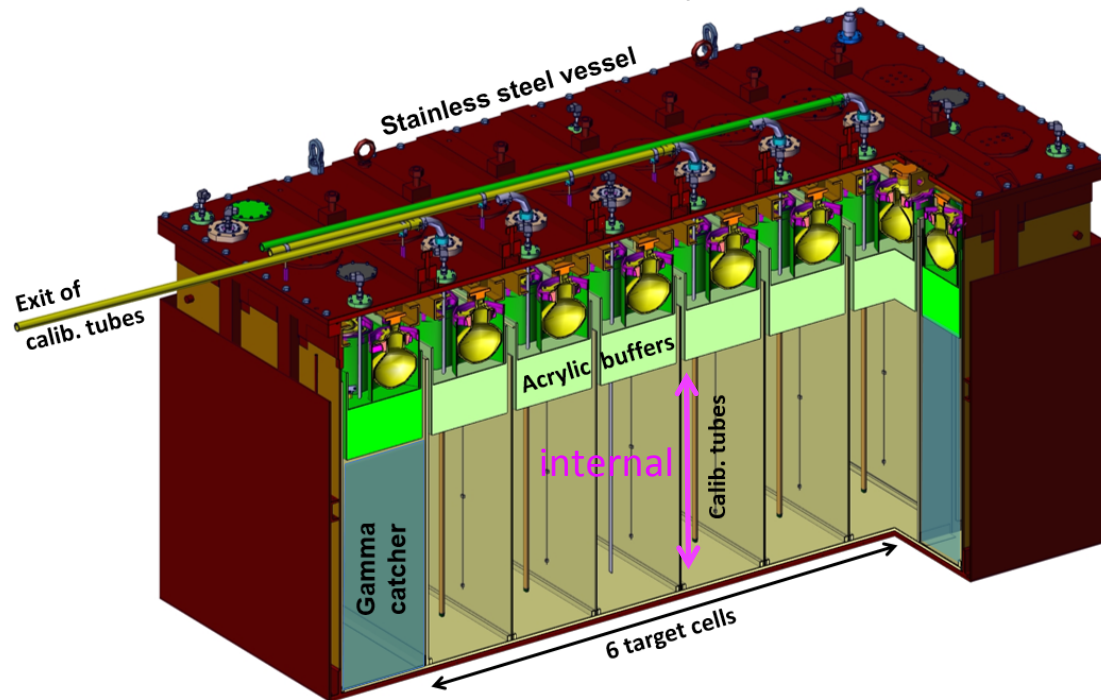
Timeline



- Data taking started in Nov 2016
- Registered 66 days reactor ON + 28 react OFF before reactor shutdown for maintenance (Phase I)
- Back to data-taking in Oct 2017 – plenty of reactor OFF data to study cosmic-induced backgrounds

STEREO calib systems

Need to calibrate every (small) active volume independently



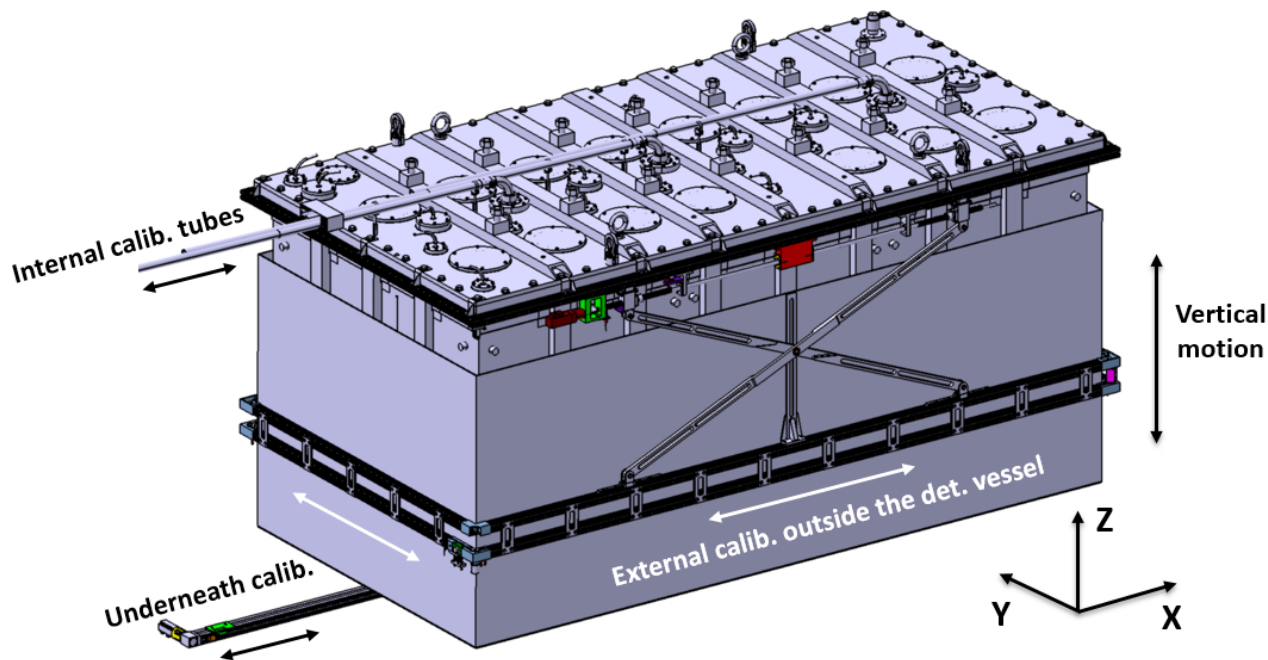
3 different ways to expose active volumes to radioactive sources:

- internal calib tubes (Target only, 5 cells)
- outside steel vessel, around GC
- outside steel vessel, below central axis

Source	^{68}Ge	^{124}Sb	^{137}Cs	^{54}Mn	^{65}Zn	^{60}Co	^{24}Na	AmBe
γ -ray energies (MeV)	0.511	0.603	0.662	0.835	1.11	1.17	1.37	2.22 (H(n, γ))
	0.511	1.69	-	-	-	1.33	2.75	4.43
Initial Activity (kBq)	90	2.4	37	90	3.3	50	5.9	$250 \cdot 10^3$ (^{241}Am)

STEREO calib systems

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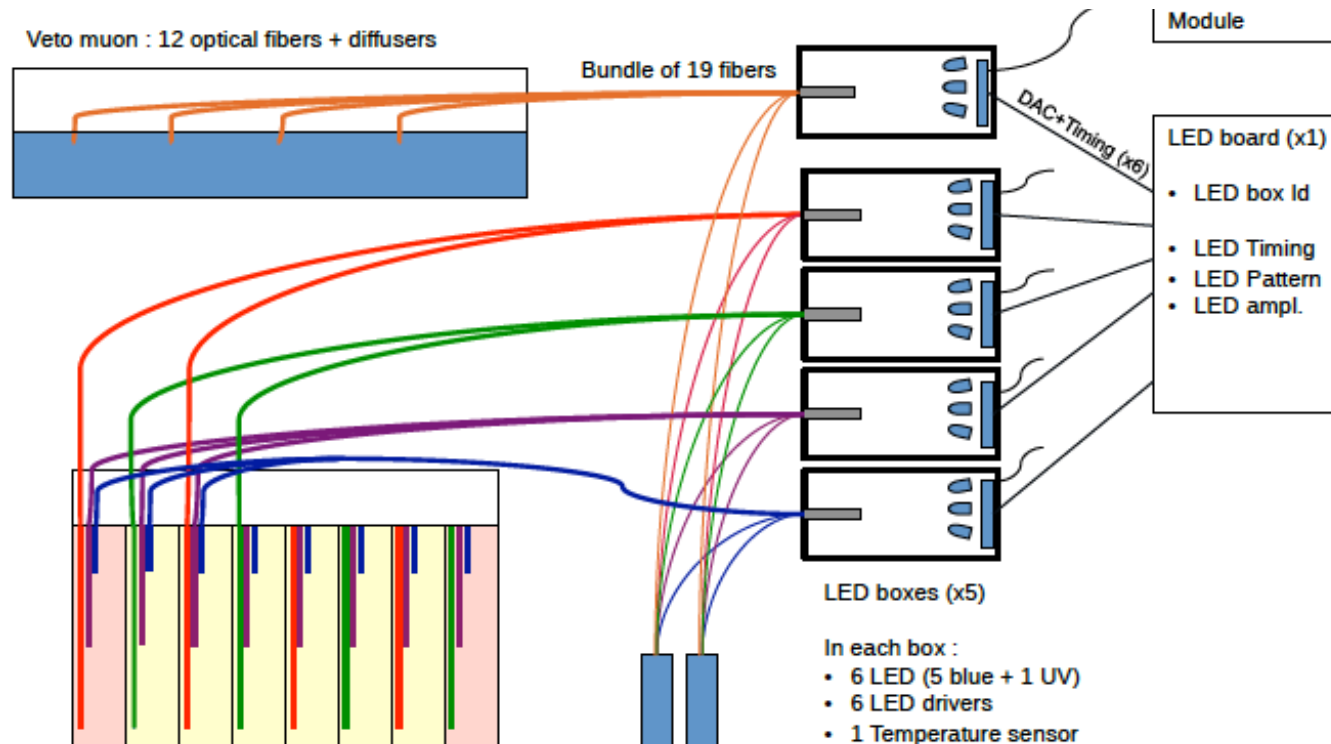
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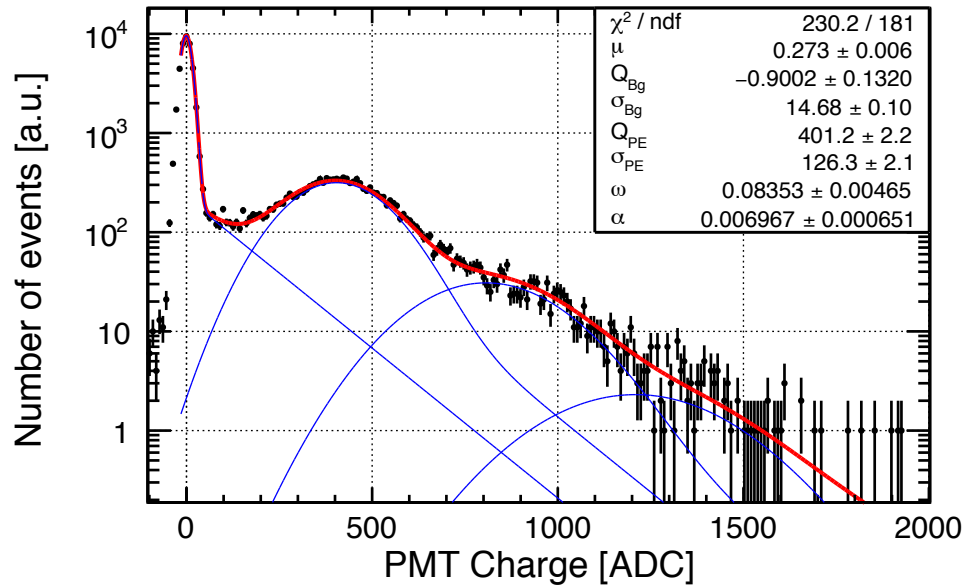
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Initial Activity (kBq)	90	2.4	37	90	3.3	50	5.9	$250 \cdot 10^3$ (^{241}Am)

STEREO calib systems

- 465nm wavelength LED coupled to 3 optic fibers/cell ending in diffusive teflon balls
- Monitoring of light attenuation, collection; PMT gain calibration; electronics non-linearities

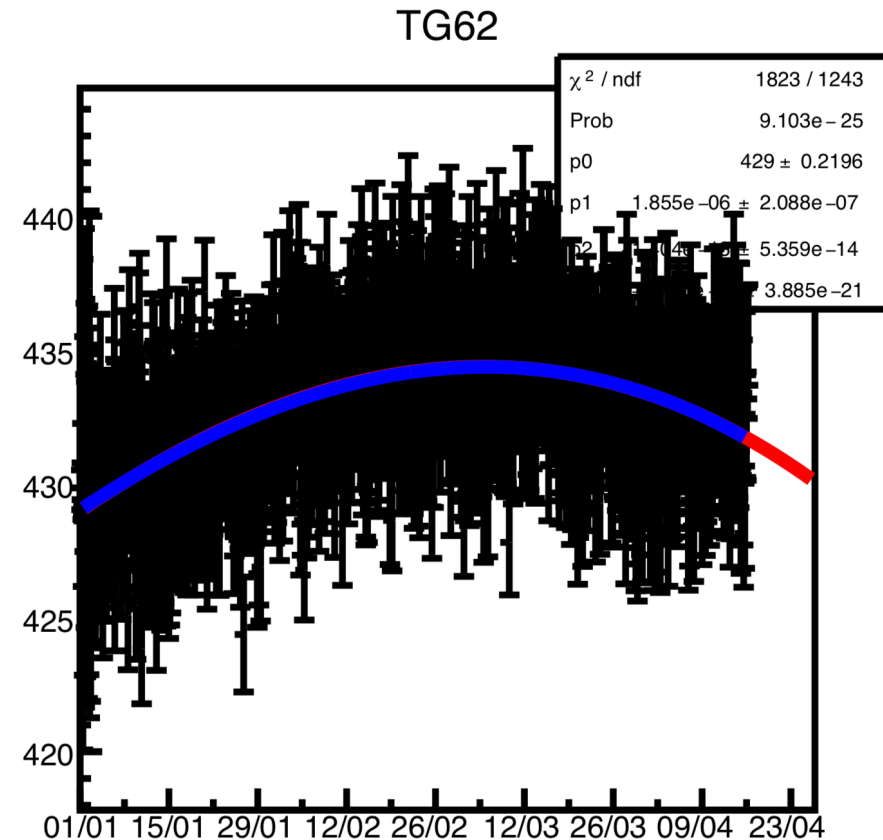


PE gain calib

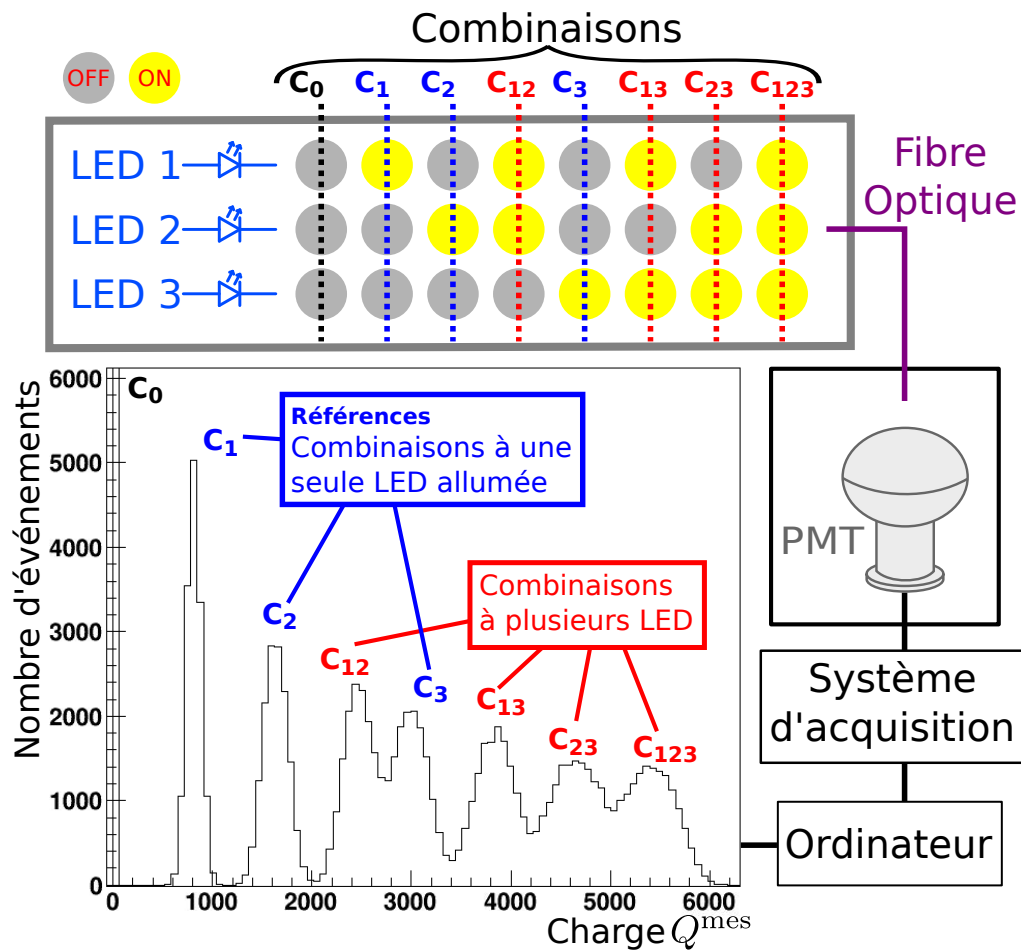


- Gain interpolated to smooth out fluctuations
- Target & GC PMTs $\sim 1\%$ variation

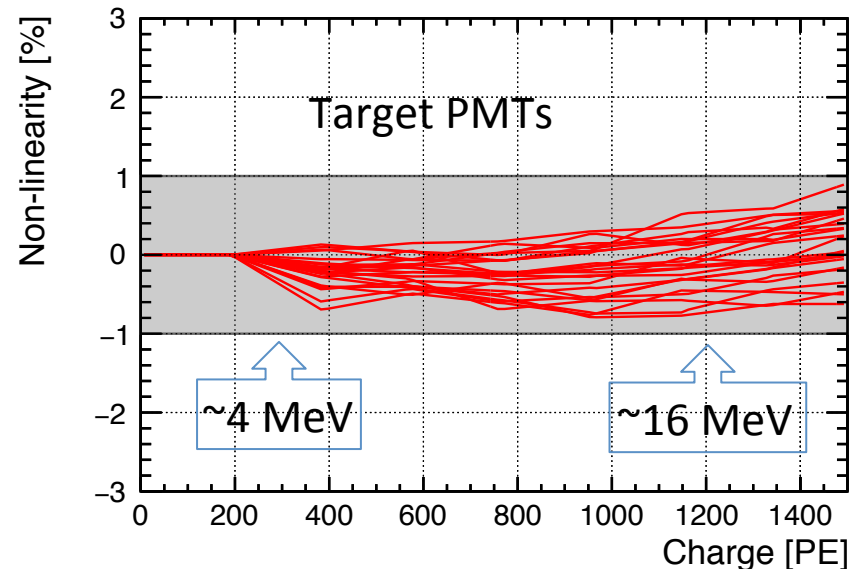
- Single PhotoElectron gain calibration several times a day

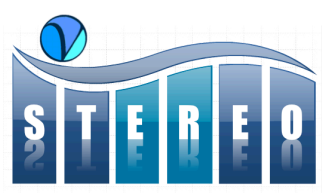


Electronics non-linearities



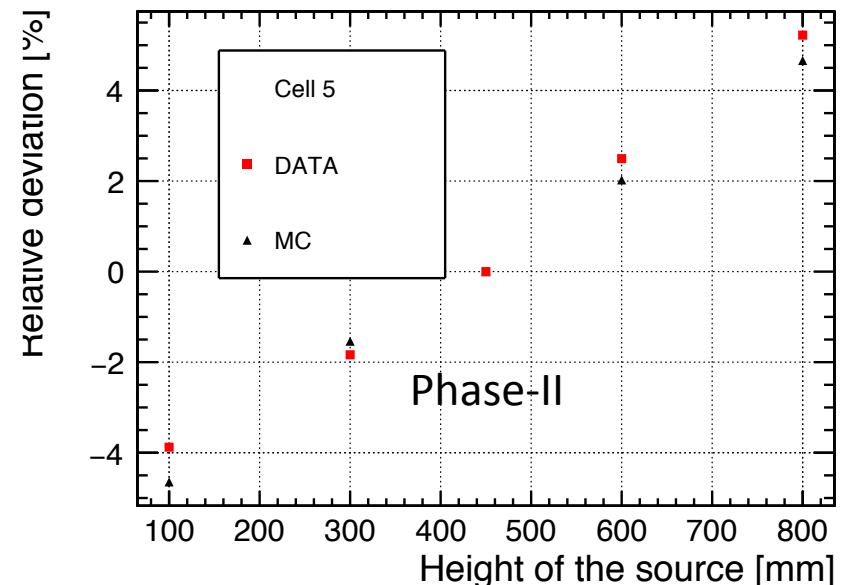
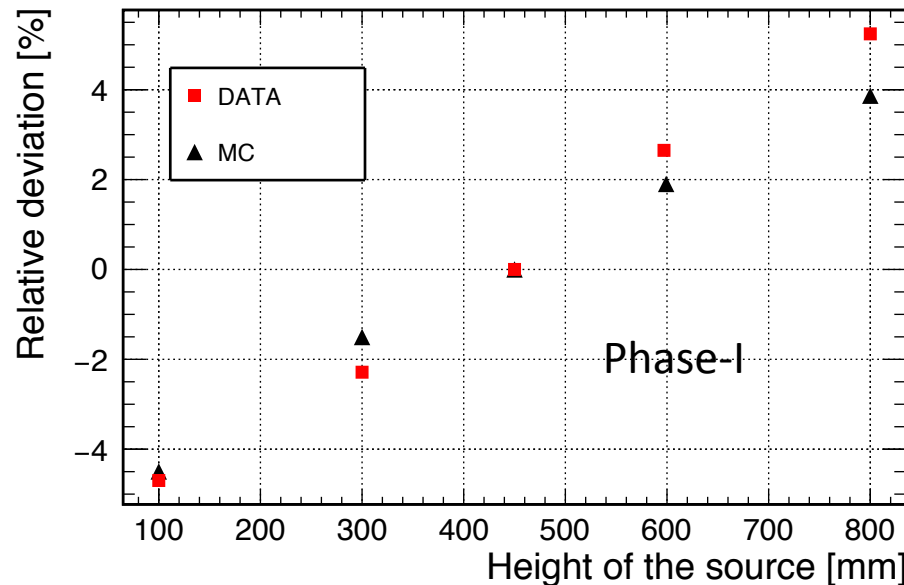
- Switch LEDs on simultaneously and compare with LEDs one at a time
- Target PMTs non-linearity well below 1%

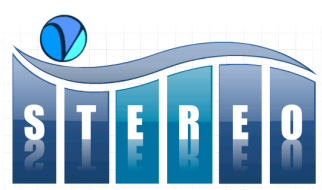




Light collection in MC

- Inhomogeneous light collection within a cell
- Little chance of correcting for it (reflection scrambles photon trajectories, localisation within a cell impossible)
- But: model accurately height dependence in MC to match rad sources data (attenuation length, wall reflectivity, refractive indices)

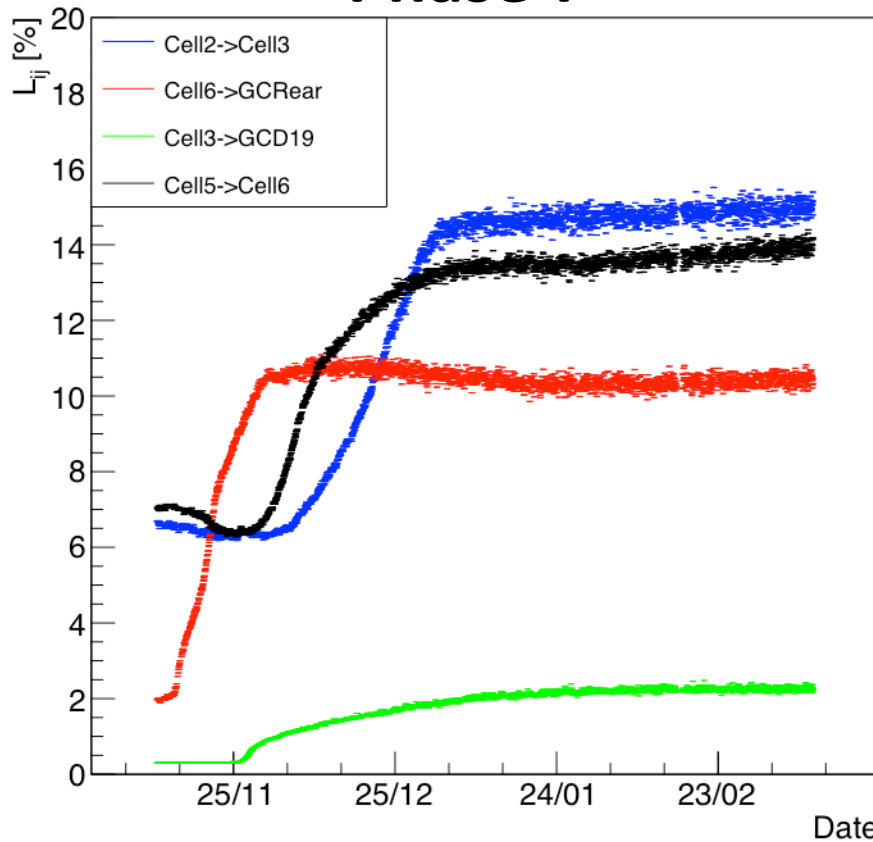




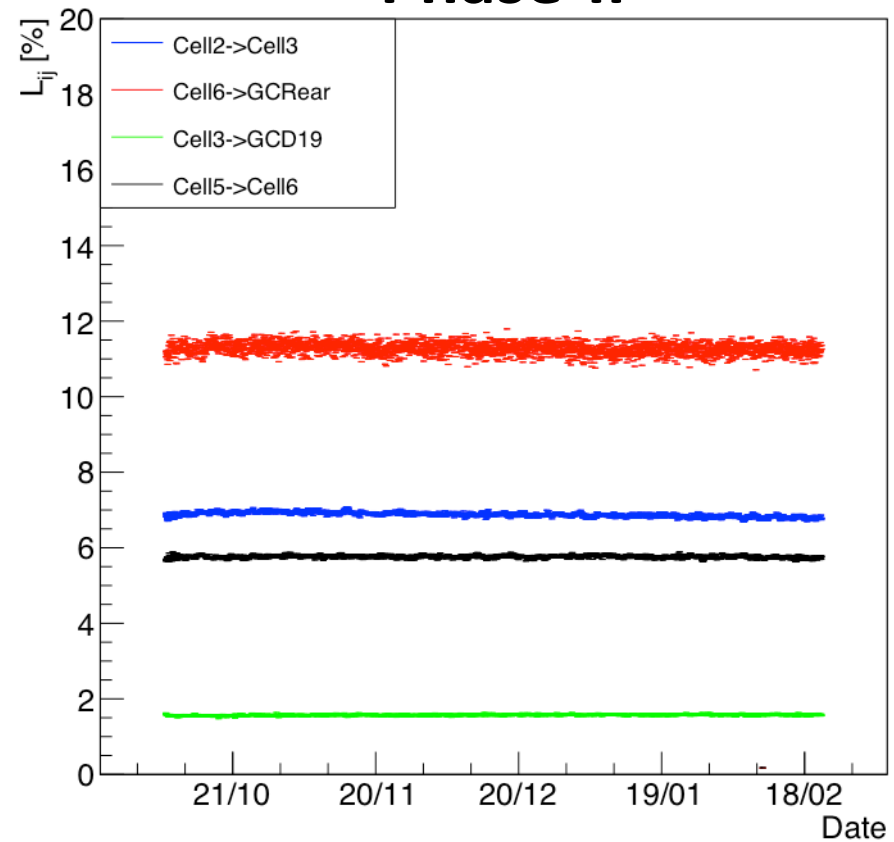
Light Crosstalk between cells

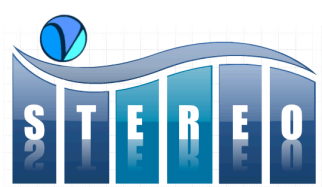
- “Bumpy” first physics run from a light-collection point of view:

Phase-I



Phase-II

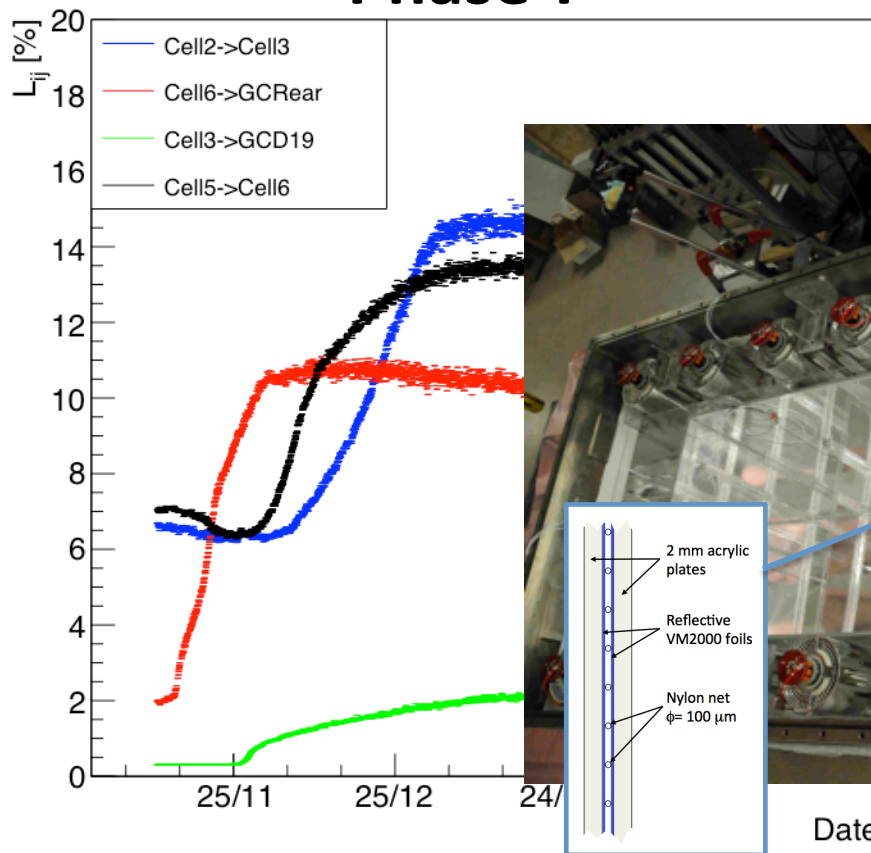




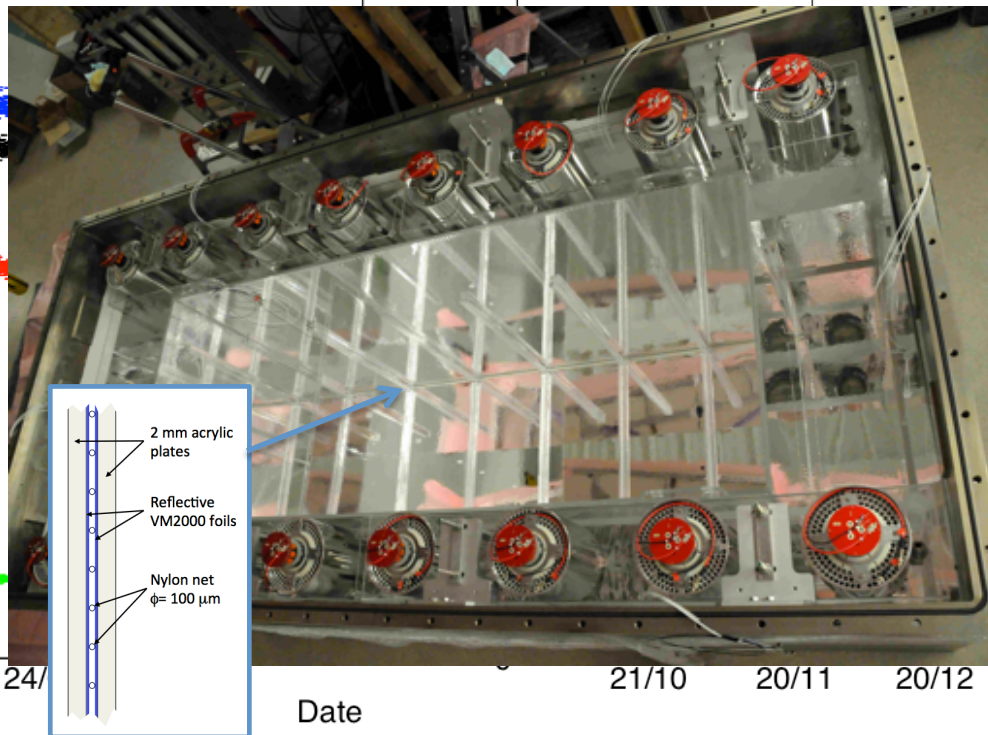
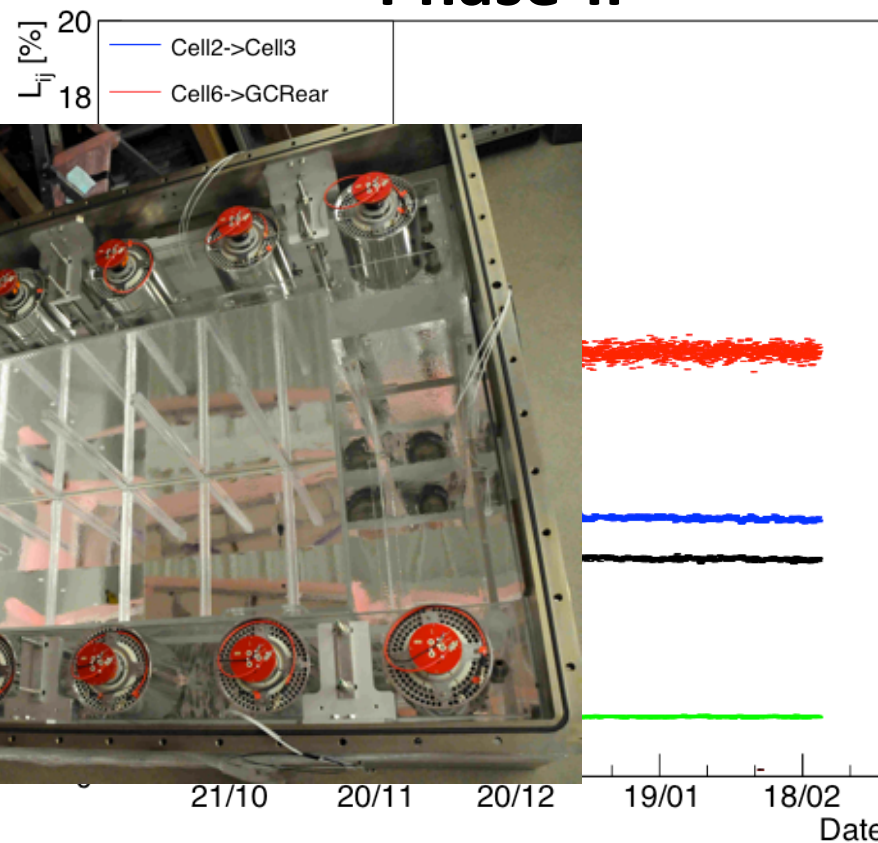
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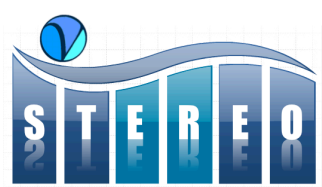
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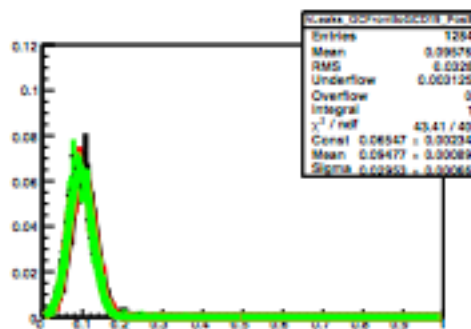
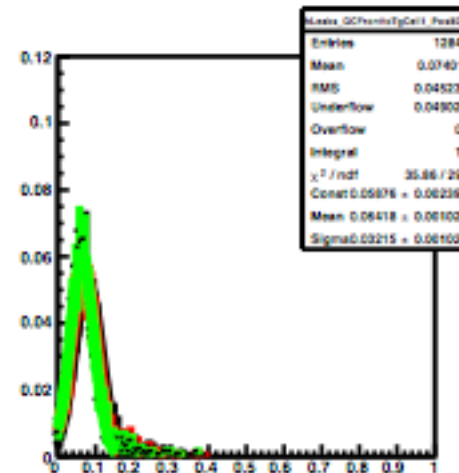
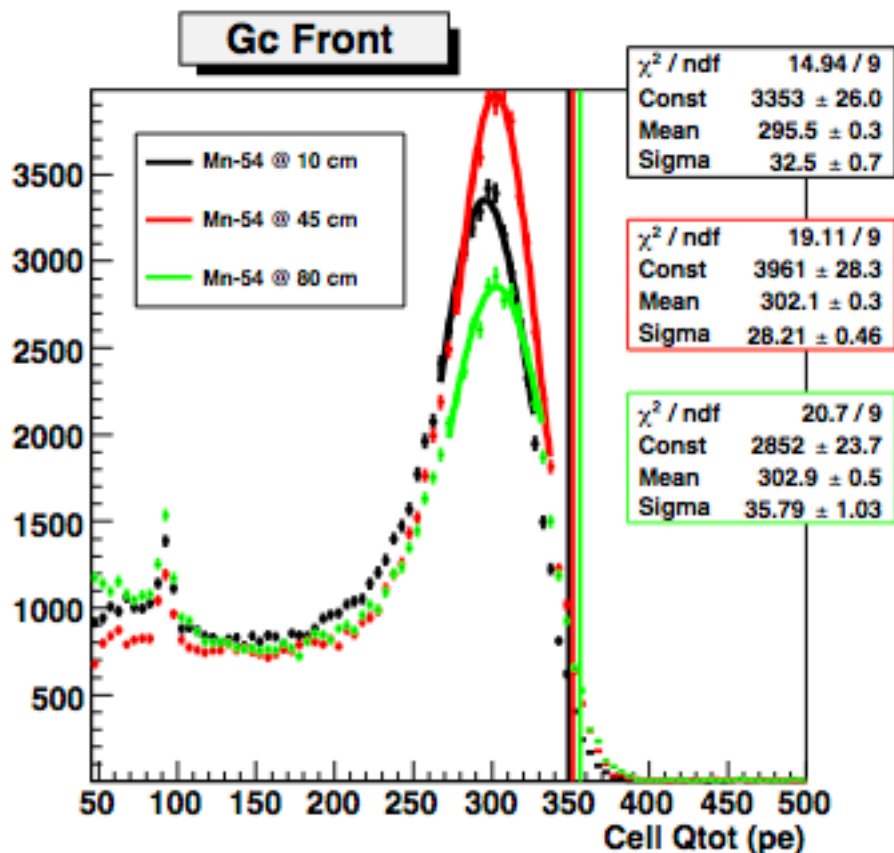
Phase-II



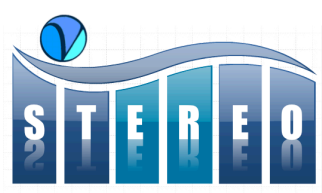


Crosstalk – Rad sources

Precisely track light crosstalk with frequent internal, low energy source (^{54}Mn) calibs

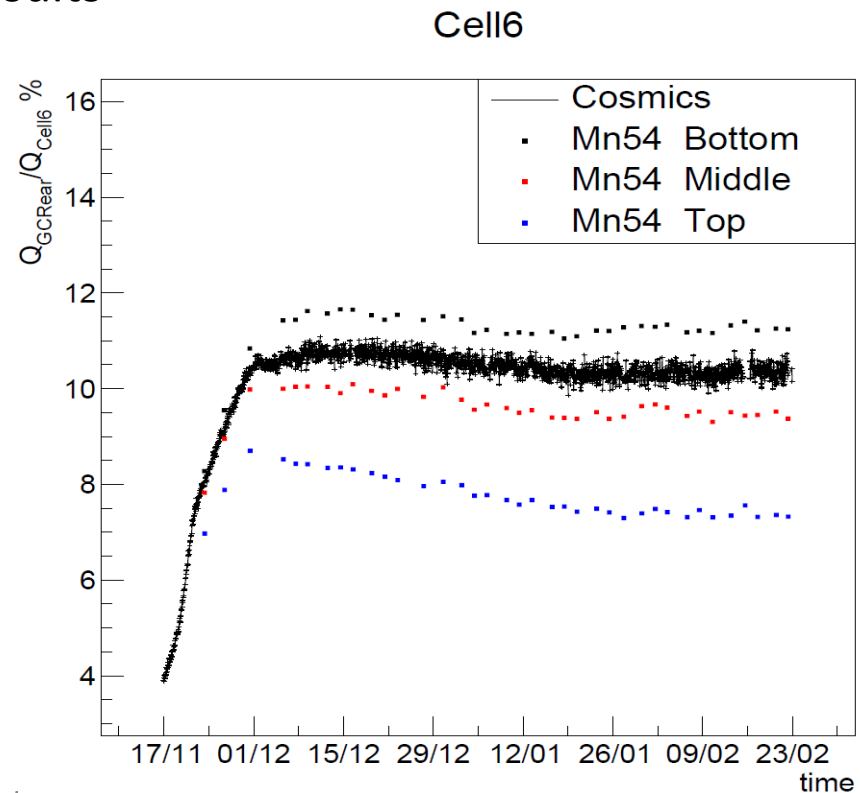
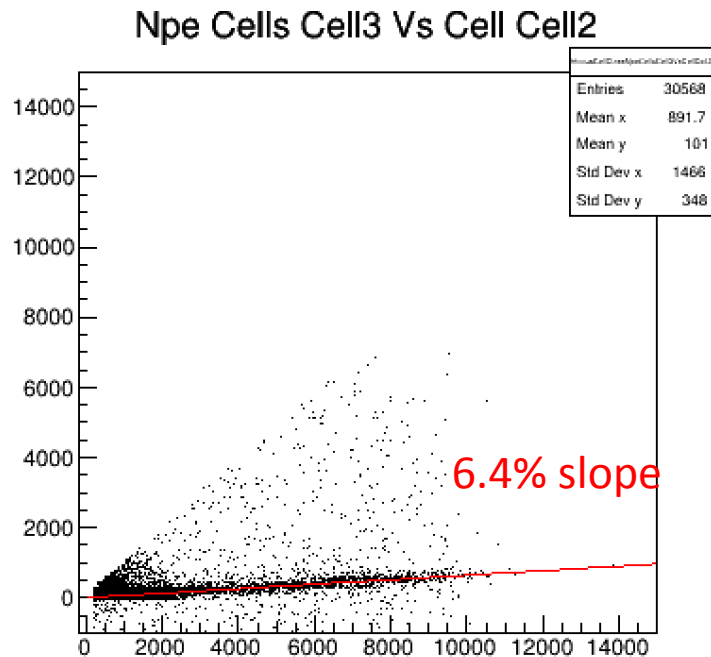


- Avoid energy deposition in nearby volumes: select only events in right-hand tail
- Height dependence of crosstalk observed



Crosstalk – Cosmics

- Independent method: look at correlation of charges in different cells in single triggers (mostly cosmics)
- “Online” measurement
- In good agreement with source results



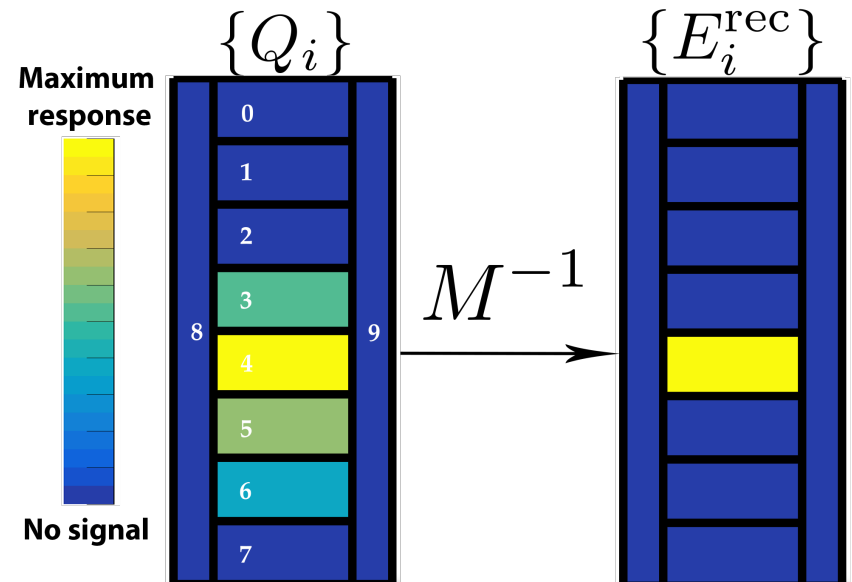
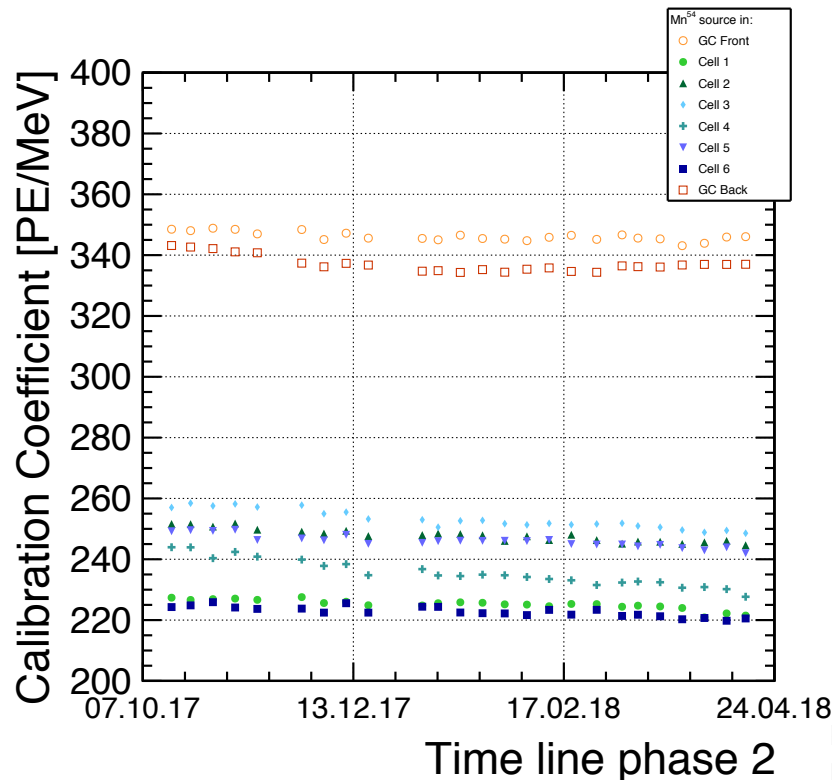


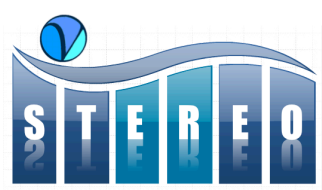
Crosstalk-correcting E reco

- Correct for cross-talk in energy reco algorithm – greatly simplifies Data/MC comparison

$$Q_i = \sum_j E_j^{dep} \times \underset{\text{Calib coeffs}}{C_j} \times \underset{\text{Light crosstalk}}{L_{ji}} = \sum_j E_j^{dep} M_{ji}$$

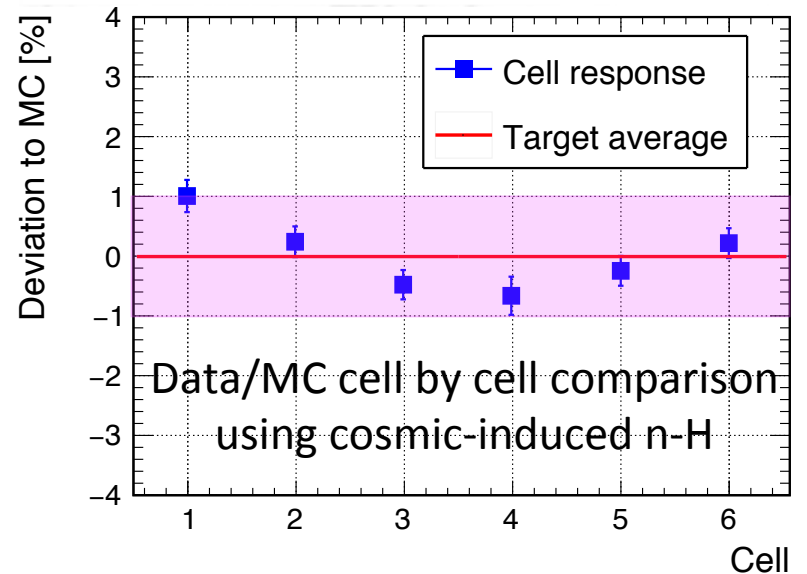
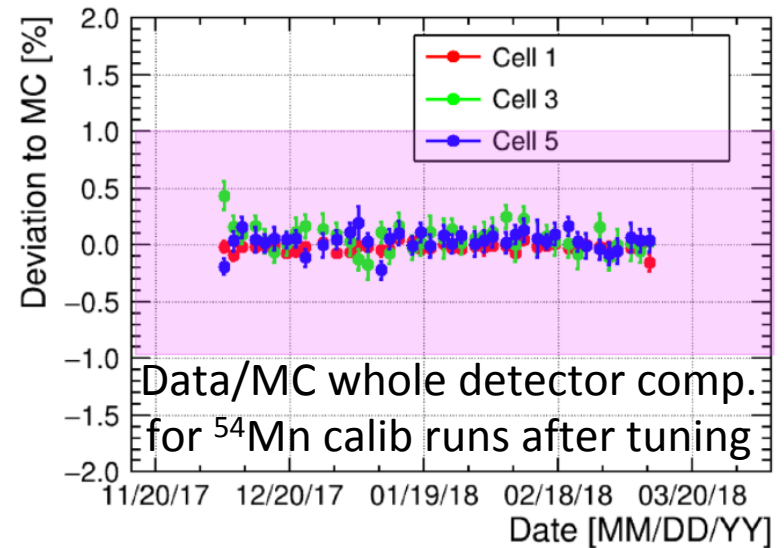
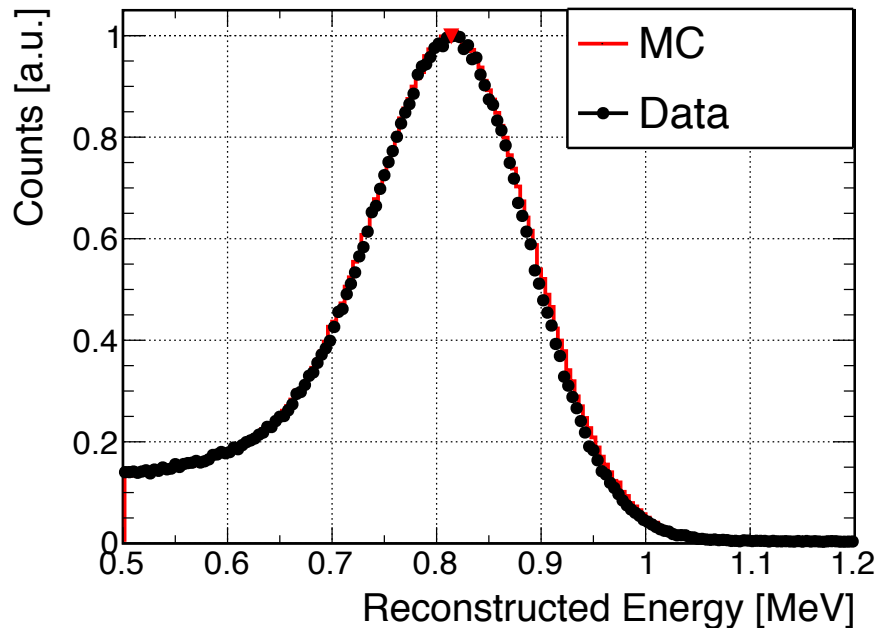
$$M^{-1} \begin{pmatrix} Q_0 \\ Q_1 \\ \vdots \\ Q_9 \end{pmatrix} = \begin{pmatrix} E_1^{rec} \\ E_2^{rec} \\ \vdots \\ E_9^{rec} \end{pmatrix}$$

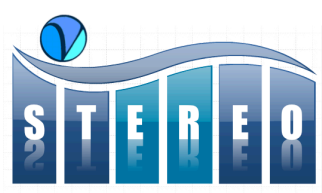




Calibration coeffs tuning

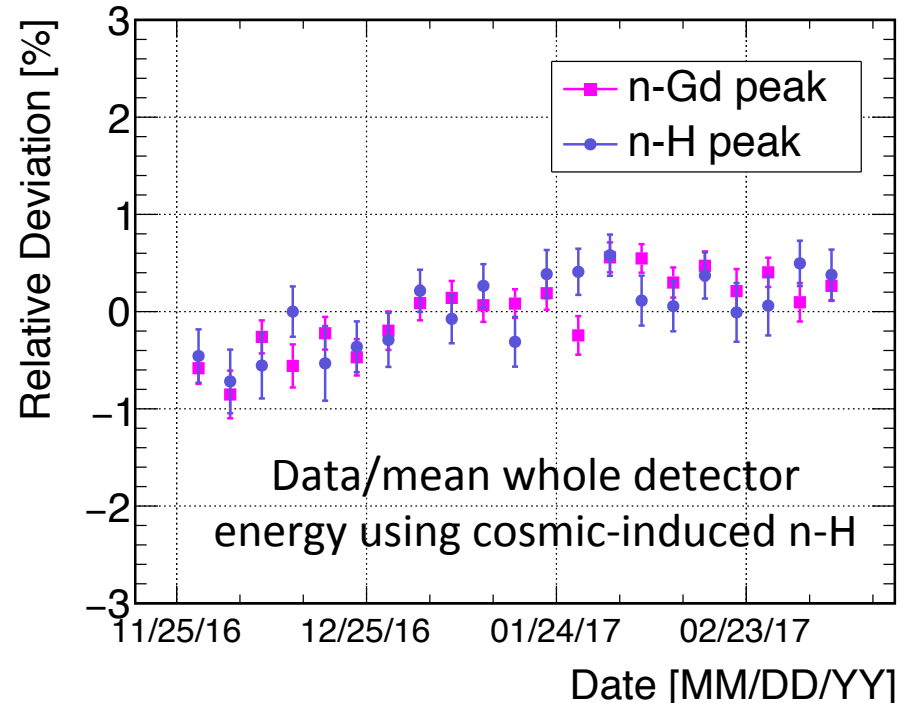
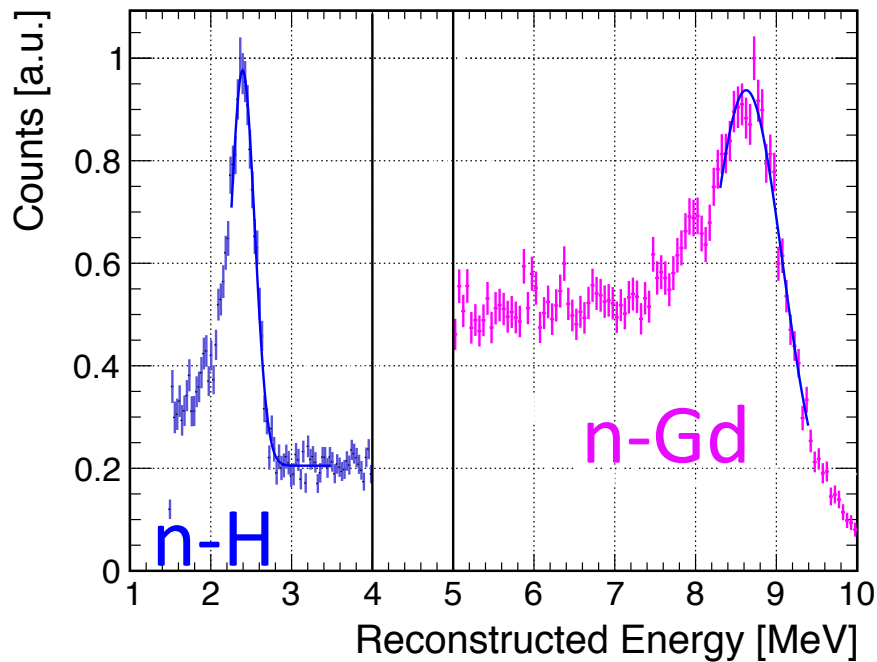
- Energy scale anchor 0.835 MeV (^{54}Mn)
- Iteratively tune C_i and L_{ij} so that energies reco'ed in Data and MC match
- 9.4% energy resolution at 0.835 MeV
- 0.2% systematics from true E deposit
- 1% cell-to-cell uncorrelated systematics

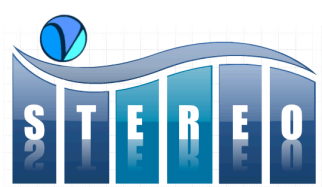




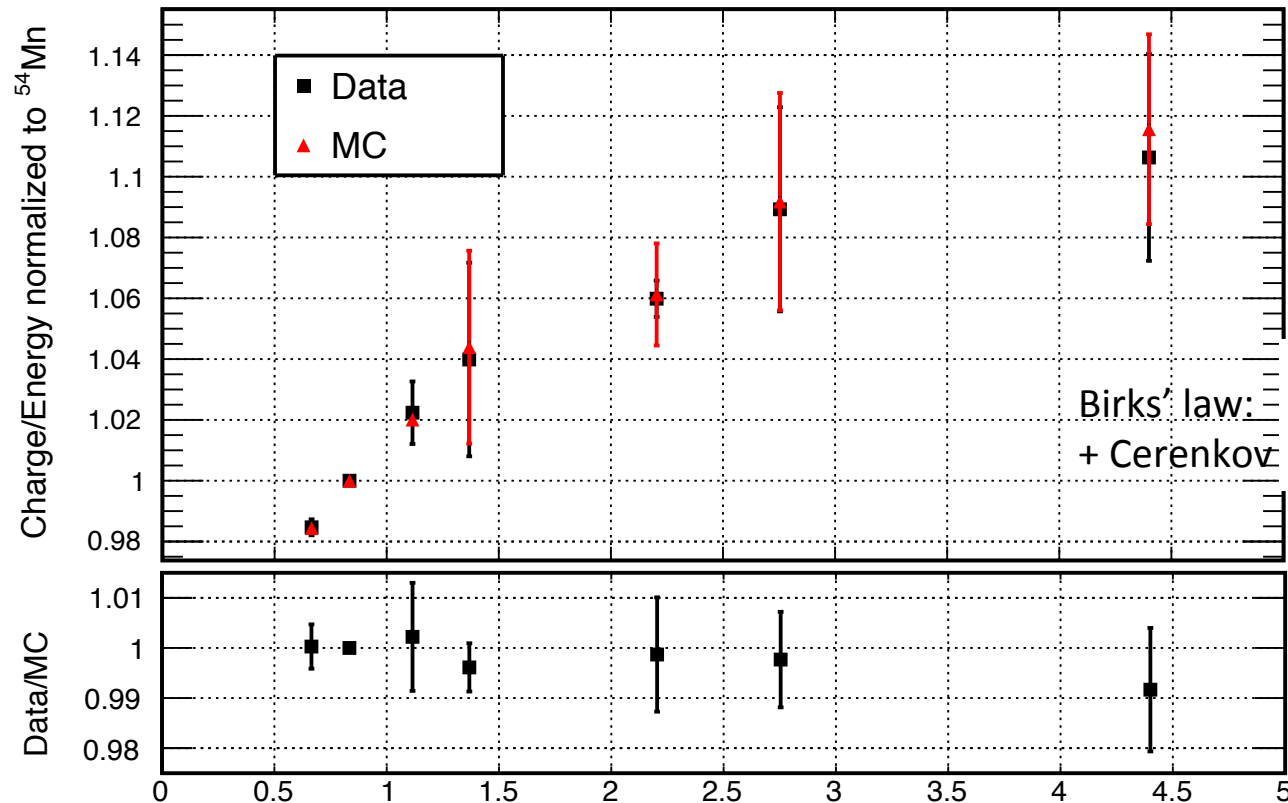
E scale time stability

- Use n captures from cosmic spallation to monitor energy scale time stability
- 0.35% variation in Phase I





Quenching

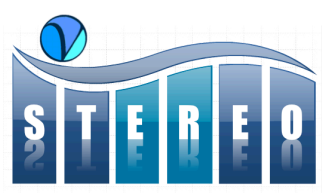


Precisely determined LS
non-linearly (1% level)

$$\frac{dL}{dx} = S \frac{\frac{dE}{dx}}{1 + K_{\text{Birks}} \frac{dE}{dx}}$$

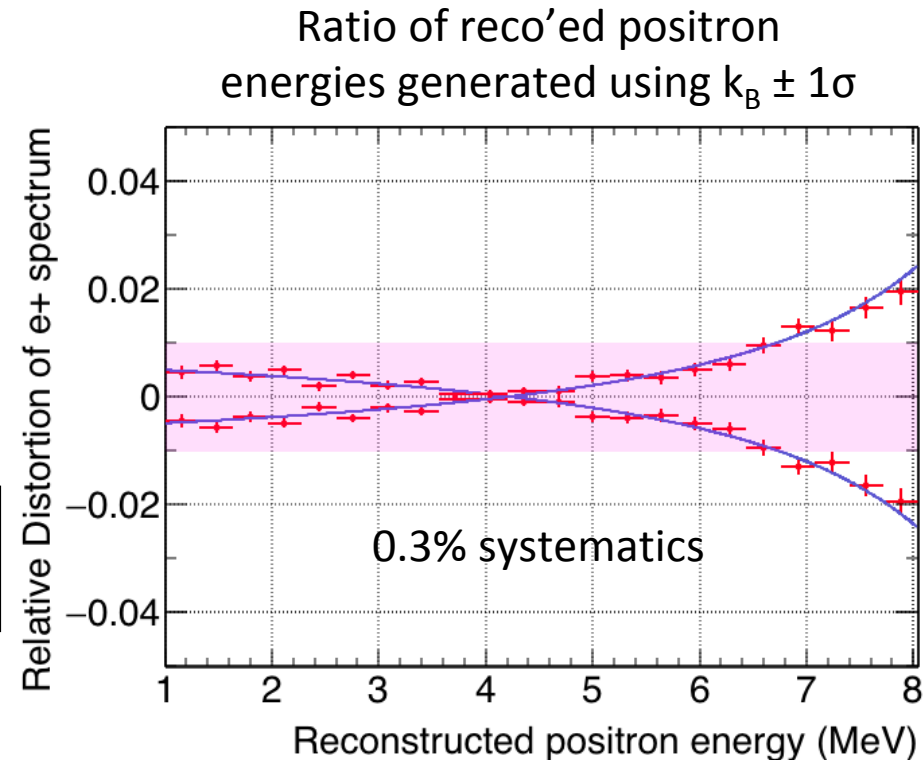
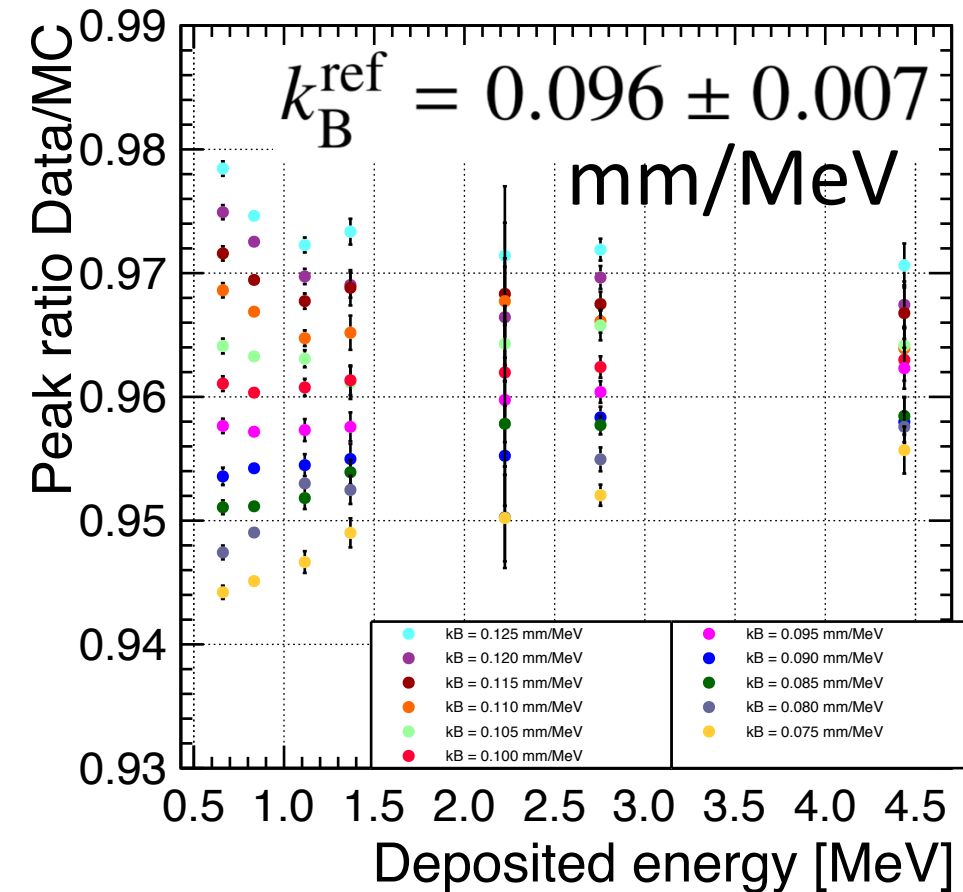
Very good Data/MC
agreement after k_B tuning

Source	^{68}Ge	^{124}Sb	^{137}Cs	^{54}Mn	^{65}Zn	^{60}Co	^{24}Na	AmBe
γ -ray energies (MeV)	0.511 0.511	0.603 1.69	0.662 -	0.835 -	1.11 -	1.17 1.33	1.37 2.75	2.22 (H(n, γ)) 4.43
Initial Activity (kBq)	90	2.4	37	90	3.3	50	5.9	$250 \cdot 10^3$ (^{241}Am)



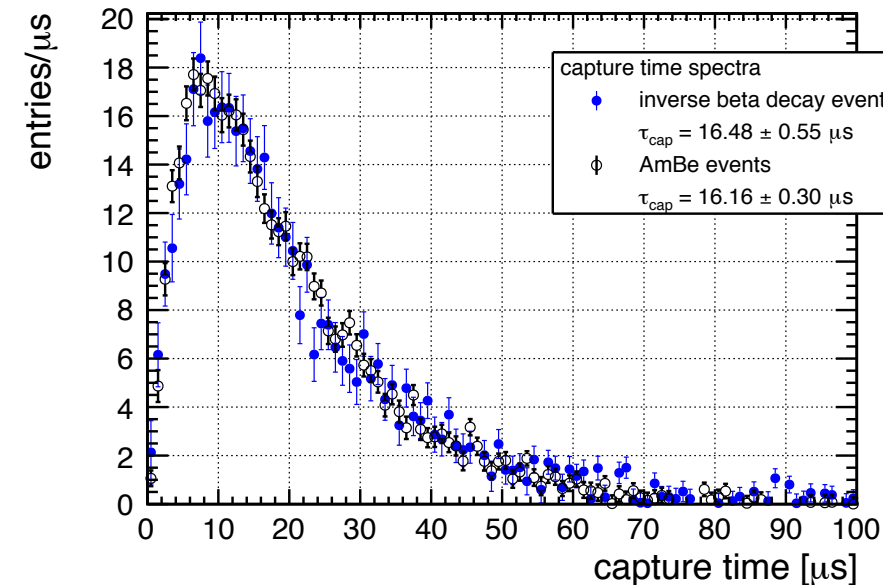
Quenching tuning in MC

- Fit to Data/MC ratio for source peak means for different MC k_B values yields

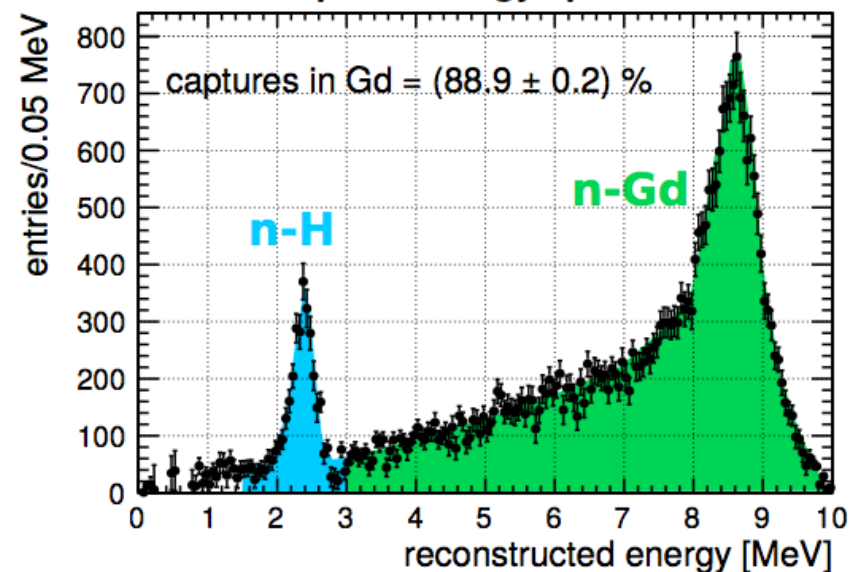


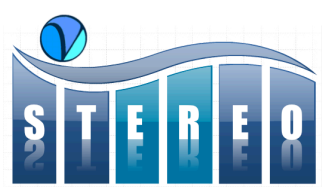
Neutron capture efficiency

- Thoroughly studied with AmBe internal calibrations



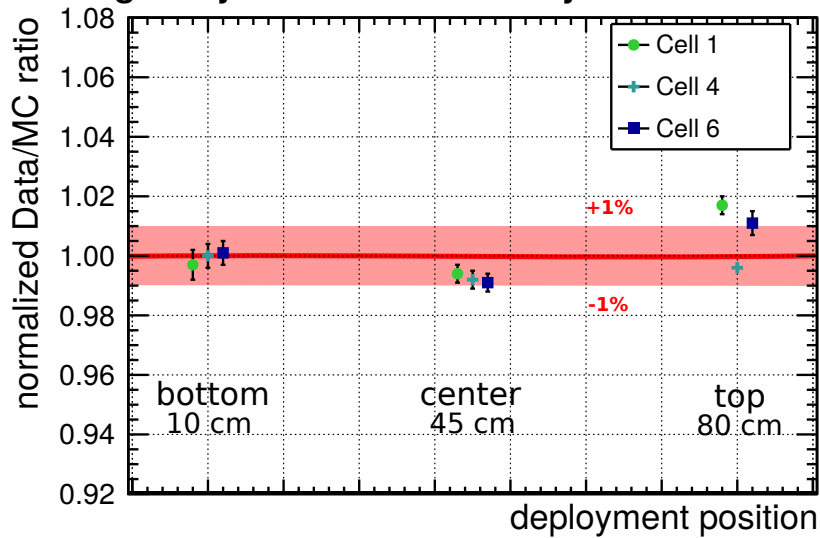
AmBe neutron capture energy spectra



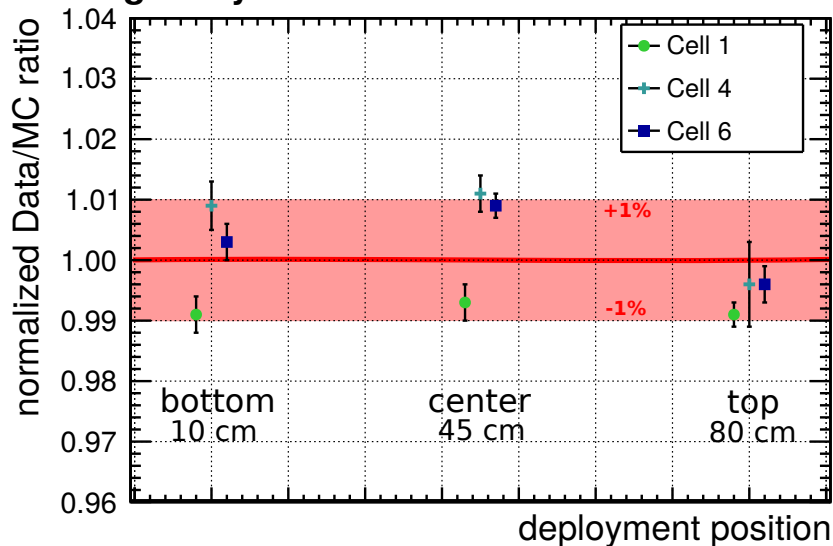


Neutron capture efficiency

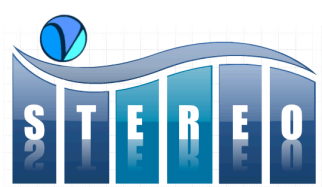
Homogeneity of the Cut Efficiency



Homogeneity of the Gd-fraction

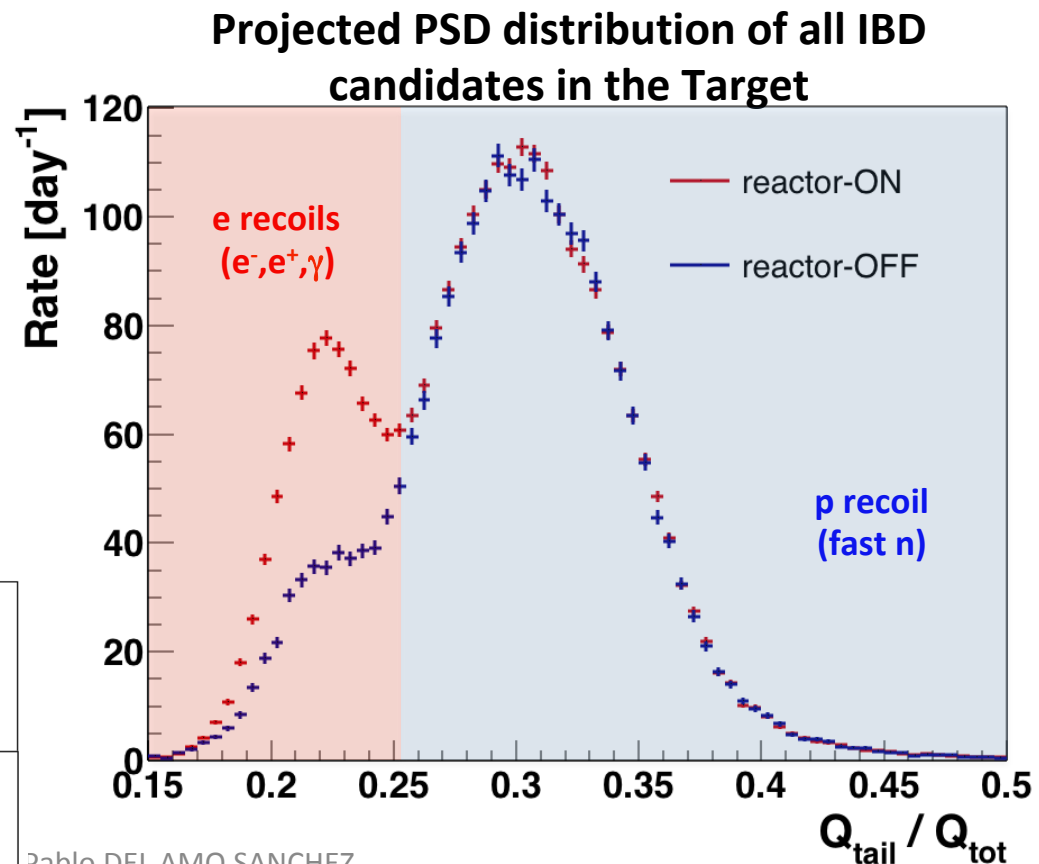
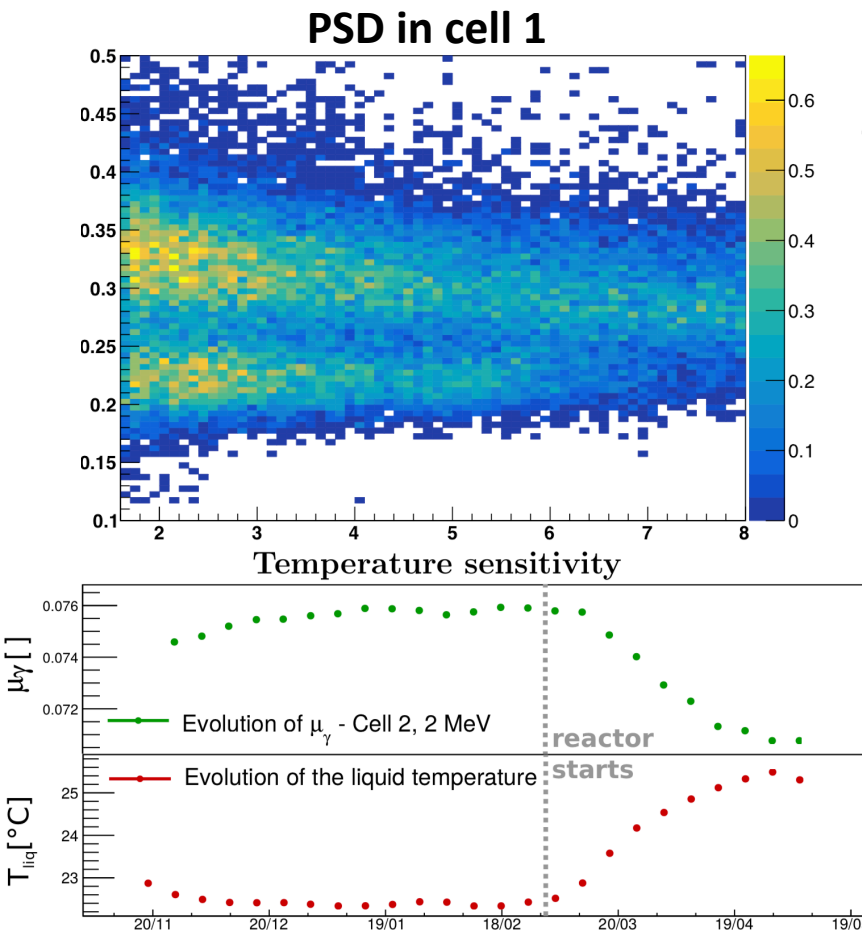


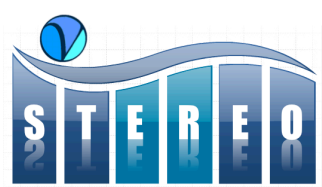
Source		Contrib $\sigma^{\text{NormUncor}}$
n-capture eff.	0.961 ± 0.008 (stat+syst)	0.81%
IBD cut efficiency (energy + DeltaT	0.981 ± 0.009 (stat+syst)	0.88%
TOTAL neutron eff.	0.943 ± 0.012 (stat+syst)	1.2%



Pulse Shape Discrimination

- Pulse Shape Discrimination response sensitive to temperature. Cosmics rate depends on atmospheric pressure → not just an ON-OFF subtraction
- In-situ PSD calibration**, PSD parameters fitted and allowed to vary with time

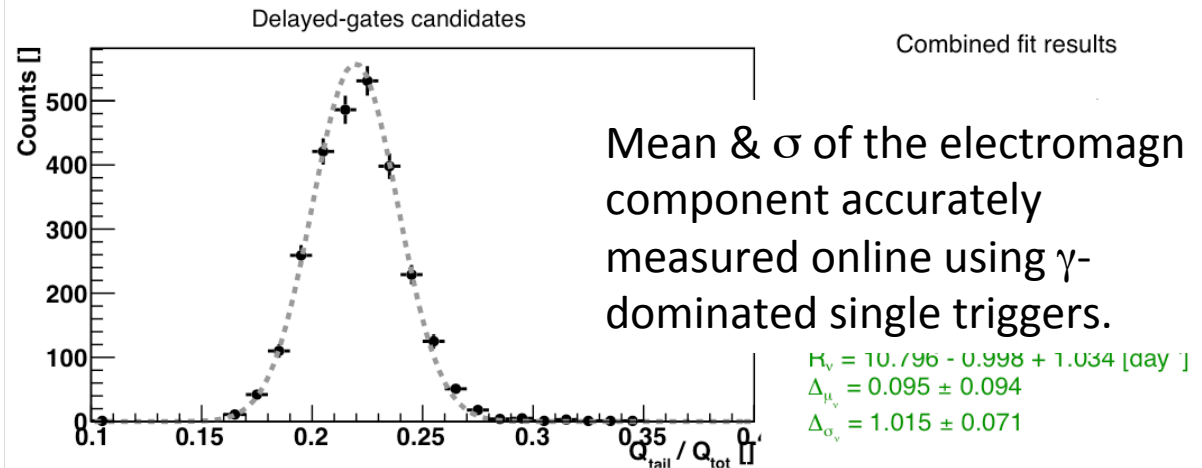
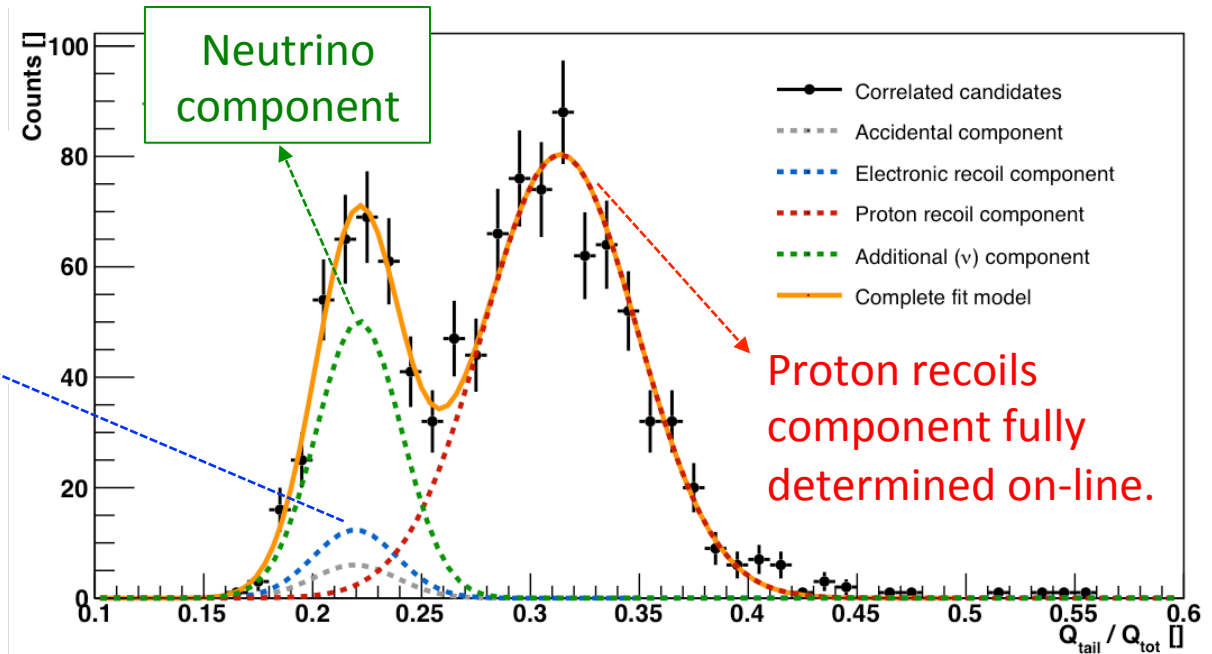


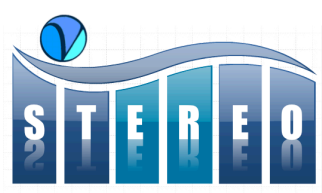


Pulse Shape Discrimination

Amplitude of the electromagn. cosmic background fixed by the A_γ / A_p ratio measured during OFF periods.

Amplitude of the accidental background determined by off-time windows.





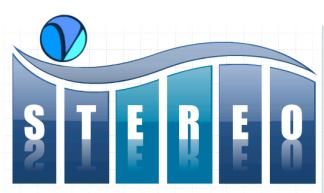
Putting all together

During Phase I:

- Energy scale systematics already at the 1% level
- Detection efficiency systematics at the ~2% level

Energy scale at 1% level	Source	Cell-to-cell correlated	Cell-to-cell uncorrelated
	E^{ref}	-	0.20%
	^{54}Mn anchor	-	0.30%
	Time stability (n-H)	0.35%	-
	Cell-wise Data-MC comparison (from calib. sources + n-H)	-	1.00%
Total		0.35%	1.06%

Source	Detection efficiency syst	Contrib to $\sigma_{\text{Cell}}^{\text{NormUncor}}$	
Cell volume		0.85 %	
n-capture efficiency		1.20 %	
Other selection cuts efficiency		0.90 %	(3.1 % cell4)
TOTAL		1.7 %	(3.4 % cell4)



See related posters # 195, 196 & 197
at Neutrino2018!

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