



ESO Public Surveys

Status and forward look

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Motivation & Policies

- Since 2010, ESO operates dedicated telescopes for surveys: VST and VISTA. Also FLAMES@UT2, UVES@UT2, VIMOS@UT3, EFOSC/SOFI@NTT spectrographs were supporting survey projects
 - ESO public surveys: such scheme implements a partnership between ESO and its community:
 - ESO organize calls, support telescope operations, deliver raw data to teams, acts as data standard authority, facilitate the publication of products through the archive, and organize peer reviews activities coordinated by the ESO Survey Team
 - Community define the science projects, the observing strategy & observations, final scientific QC and delivery of science data products
- The ESO Science Archive Facility (SAF; <u>http://archive.eso.org/cms.html</u>) is the primary point of publication/availability of the survey products (as per ESO Council Meeting 104, 17–18 December 2004).



Outline

Motivation & Policies – Public Surveys @ ESO

On-going ESO Public Surveys

- VST surveys
- VISTA first & second cycle
- > Spectroscopic Surveys
- Scientific impact & returns for the community
- Forward look
- Conclusions
- Acknowledgments





On-going Public Surveys http://www.eso.org/sci/observing/PublicSurveys/sciencePublicSurveys.html

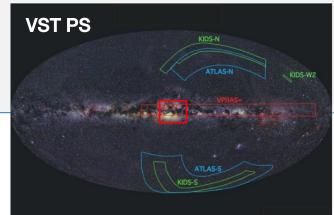
Public Surveys in a nut-shell

- Legacy value for astronomical community at large
- Very large programmes (>2 years); very diversified observing strategies
- Synergy between optical and NIR imaging, from 0.33 to 2.15 micron
- Spectroscopic follow-up of sources; multiplexing capabilities (FoV ~ 10 arcmin) to single slit; Spec_res 10⁴ to ~1000.
- All raw observations are immediately public
- Survey teams commit to deliver reduced images/spectra and catalogues within ~yearly releases



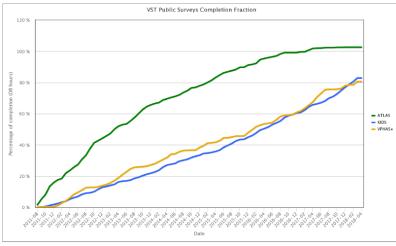


VST PS

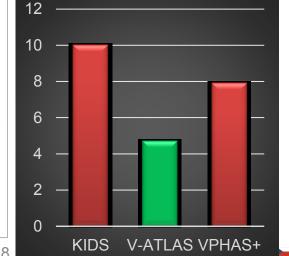


The VST surveys started operations on October 2011

Survey ID and home page	Science topic	Area deg ²	Filters	Magnitude limits	Observing time completed (hrs) to Oct 1 st 2017
KIDS http://kids.strw.leidenuniv.nl/	Extragalactic	1500	u'g'r'I'	24.1 24.6 24.4 23.4	2297
ATLAS http://astro.dur.ac.uk/Cosmolog y/vstatlas/	Wide area/BAO	4000	u'g'r'I'z	22.0 22.2 22.2 21.3 20.5	1422
VPHAS+ http://www.vphas.eu	Stellar astrophysics	2000	U' g' Ha r' I'	21.8 22.5 21.6 22.5 21.8	985





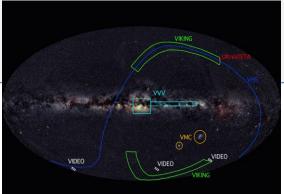


- ATLAS extended 700 sqdeg in u'
- VPHAS+ reprioritization
- KIDS may be completed in 2019

+ES+ 0 +

VISTA PS

VISTA PS



- First cycle of VISTA surveys started operations
- on April 2010.
- Current assessment is that they will all be completed in P101!

Survey ID and home page	Science topic	Area (deg ²)	Filters	Magnitude limits	Observingtimecompleted(hrs)to Oct. 1st 2017
Ultra-VISTA http://home.strw.leidenuniv.nl/~ ultravista/	Complete	d ^{' deep} '3 ultra deep	Y J H Ks NB118	25.7 25.5 25.1 24.5 26.7 26.6 26.1 25.6 26.0	1809
VHS http://www.ast.cam.ac.uk/~rgm/ vhs/	Whole sky	17800	Y J H Ks	21.2 21.1 20.6 20.0	4519
VIDEO http://www- astro.physics.ox.ac.uk/~video	Deep high-z	12	Z Y J H Ks	25.7 24.6 24.5 24.0 23.5	1876
VVV http://vvvsurvey.org/	Complete	d 0	Z Y J H Ks	21.9 21.1 20.2 18.2 18.1	2157
VIKING http://www.astro- wise.org/projects/VIKING/	Complete	d 00	Z Y J H Ks	23.1 22.3 22.1 21.5 21.2	2410
VMC http://star.herts.ac.uk/~mcioni/v mc/	Resolved SFH	180	Y J Ks	21.9 21.4 20.3	1779





VISTA PS

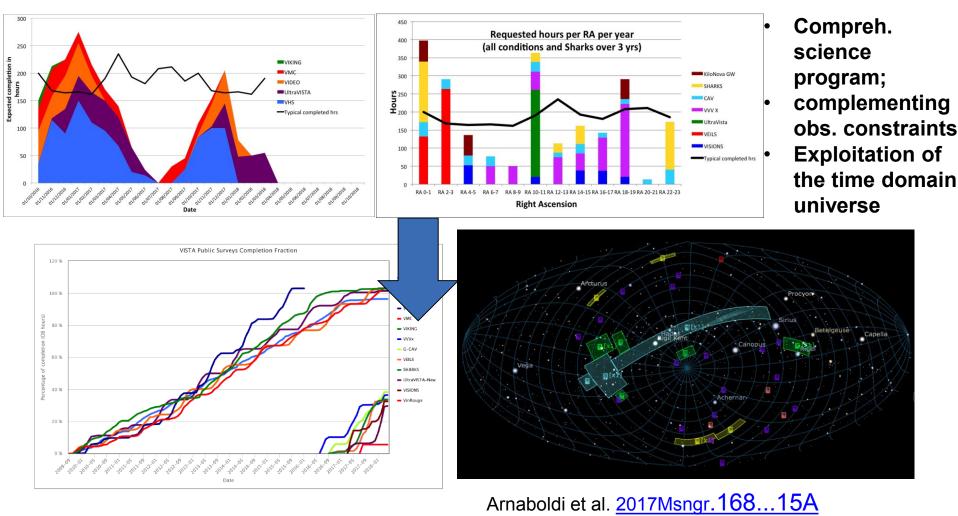
- Second cycle of VISTA surveys, ESO call in 2015
- 13 Loi submitted by the community, oversubscription > 2
- 7 proposals selected; started operations in April 2017



Name P.I.	Short Title	Filters	Tot. Time (hrs)	Area (deg²)
GW; N. Tanvir	Kilonova counterparts to Gravitational wave sources	Y J Ks	420	300
UltraVISTA; J. Dunlop	Completing the legacy of UltraVISTA	J H Ks	756	0.75
VVVX; D. Minniti	Extending VVV to higher Gal lat.	J H Ks	1900	1700
VEILS; M. Banerji	VISTA Extragalactic Infrared Survey	J Ks	1180	9
CAV; M. Nonino	Clusters at VIRCAM	Y J Ks	560	30
VISIONS; J. Alves	VISTA star formation atlas	J H Ks	553	550
SHARKS;I. Oteo	Southern Herschel-Atlas Regions K-band survey	Ks	1200	300

Lessons learnt and time domain

VISTA 1st cycle surveys reaching completion in P101!





Spectroscopic Surveys

- the Milky Way (MW) and in 100 open cluster satellite survey. It provides the physical statellite survey. It provides the physical statellite survey. It provides the physical statellite survey of operations. The MW state st
- PESSTO: spectroscopic follow-up of above physics in an unbiased sample of nearby galaxies to understand of nearby galaxies to unders
- VANDELS: study of the star forming galaxies in the redshift range 2.5-z<7.0 and

South and UDS. Goal is to measure metallicities and OOOOY fields, CDF south and UDS. Goal is to measure metallicities and OOOY onized gas in these systems. 914 hours allocated on VING OOSMOS field in the redshift range 0.6 < z < 1.0. Unde OOOV with the redshift range O.6 < z < 1.0. Unde OOV with the redshift range O.6 < z < 1.0. Unde OOV with the redshift range O.6 < z < 1.0. Unde OOV with the redshift range O.6 < z < 1.0. Unde OOV with the redshift range O.6 < z < 1.0. Unde OOV with the redshift range O.6 < z < 1.0. Unde OOV with the redshift range OOV with the redshift range of the redshift r VIMOS@UT3. (http://www.mpia.de/home/legac/index.html)



Spectroscopic Surveys

- Gaia ESO: this survey targets 10⁵ stars distributed among the major components of the Milky Way (MW) and in 100 open cluster a strong synergy with the Gaia satellite survey. It provides the phone of operations. Target selection comes from the imaging surveys VHS and VVV, among others. (<u>http://www.gaia-eso.eu/</u>)
- of nearby galaxies to understanc 01/2012 on EFOSC/SOFU@NTT
- **VANDELS:** study of the star forming galaxies in the redshift range 2.5<z<7.0 and passive galaxies in the redshift range 1.5 < z < 2.5, in the two 2018 LS fields, CDF South and UDS. Goal is to measure metallicities and 2018 is ionized gas in these systems. 914 hours allocated on VIII March 2018 is ionized gas in LEGA-C: study of 3000 field in the redshift range 0.6 < z < 1.0. Understic Completions grow in mass through measurements of their
- dynamical masses, ste ages and metallicities. 1010 hours allocated on VIMOS@UT3. (http://www.mpia.de/home/legac/index.html)



SCIENTIFIC IMPACT & RETURNS FOR THE COMMUNITY





Impact of ESO PS



The Messenger

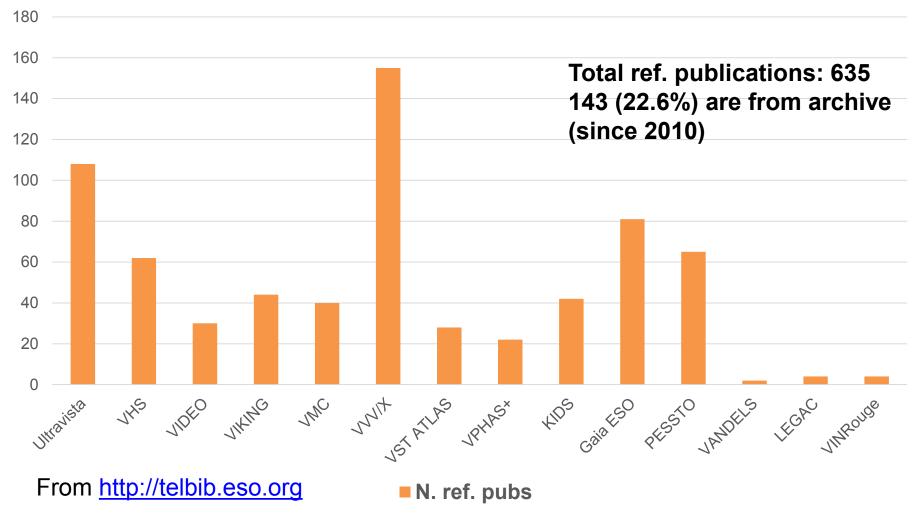
Ref. Publications: 635 total, 22% from archive papers

http://telbib.eso.org

VST in the era of the large surveys, OAC, June 2018

Impact of ESO PS

Cumulative number refereed publications for ESO PS





The Messenger

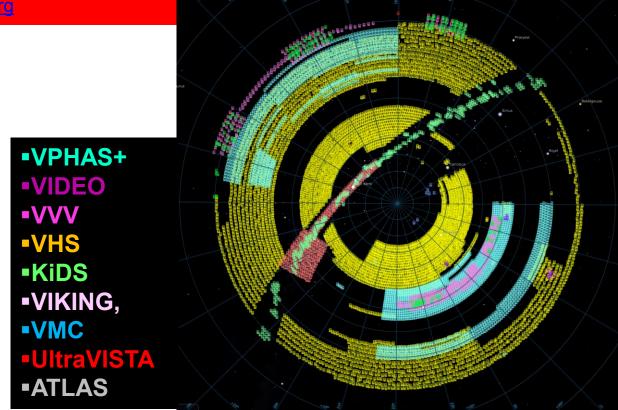
CREWS-INTICOURSE N

Impact of ESO PS

Ref. Publications: (22% from archive p

All PS have published data product releases through ESO SAF : >40 TB, 270k+ files, > 30k spectra; Opt./NIR: 4336/9445 deg²

http://telbib.eso.org





Impact of ESO PS

The Messenger

Ref. Publications: 63 22% from archive

Catalogues with aperture matched magnitudes in all relevant bands, lights curves, spectral time series

http://www.eso.org/qi

Building a community

>2600 unique users of science ready data...and counting!
> 10 requests per user;~30% of these users are new to ESO, having not applied for time



Two ESO workshops dedicated to Public surveys in 2012 & 2015

VST in the era of the large surveys, OAC, June 2018



VHS DR4.1

catalogue



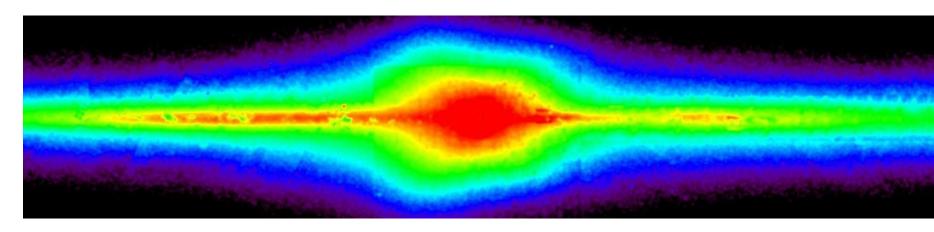
From VISTA 1st Cycle

MILKY WAY AND LOCAL UNIVERSE



Impact of ESO PS

The structure of the Milky Way outside the Bulge(*)



This image shows the surface density of stars in the Milky Way as seen from the Sun, taken from four different surveys (UKIDSS, VVV, 2MASS, and GLIMPSE) and corrected for extinction. The bulge is the thicker region near the center; it is asymmetric because it is barred. The asymmetry in the disk towards the left of the image is due to the thinner long bar outside the bulge. Wegg, Gerhard & Portail 2015, MNRAS, 450, 4050

* MPE PR http://www.mpe.mpg.de/6333402/News_20150521





From the VISTA 2nd cycle & Public Spectroscopic Surveys

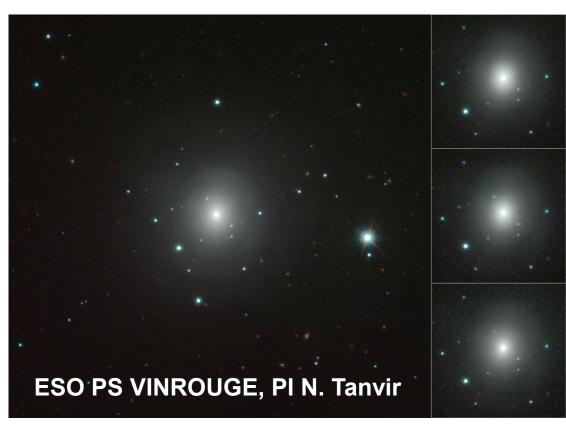
GRAVITATIONAL WAVE ALERTS AND EM COUNTERPARTS



IMPACT of ESO PS

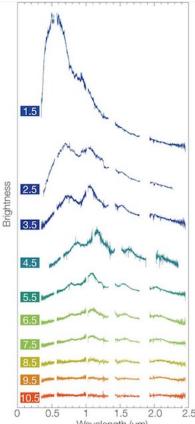
ESO PR 1733 - ESO Telescopes Observe First Light from Gravitational Wave Source

Mosaic of VISTA images of NGC 4993 showing changing kilonova



Montage of X-shooter spectra showing changes in the kilonova in NGC 4993 over 12 days.

ESO/Pian et al./Smartt & ePESSTO



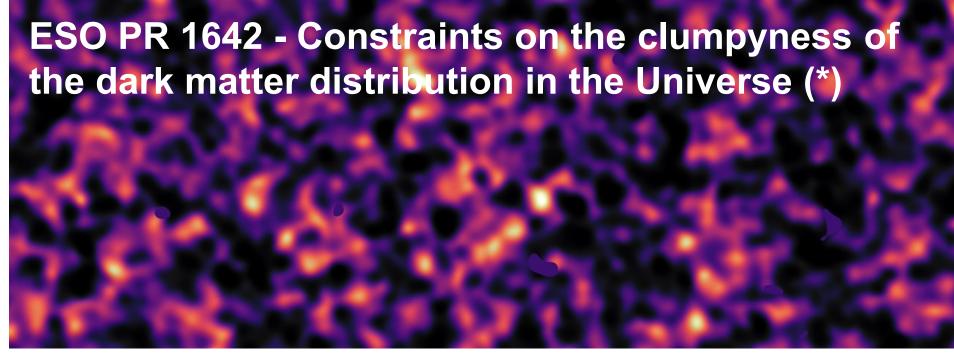


From the VST Public Surveys

COSMOLOGY



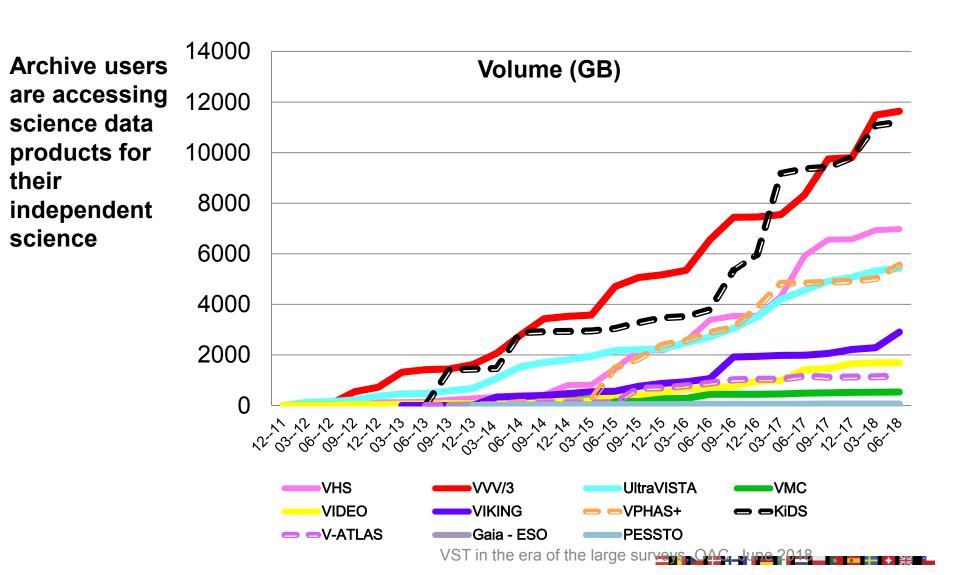
Impact of ESO PS



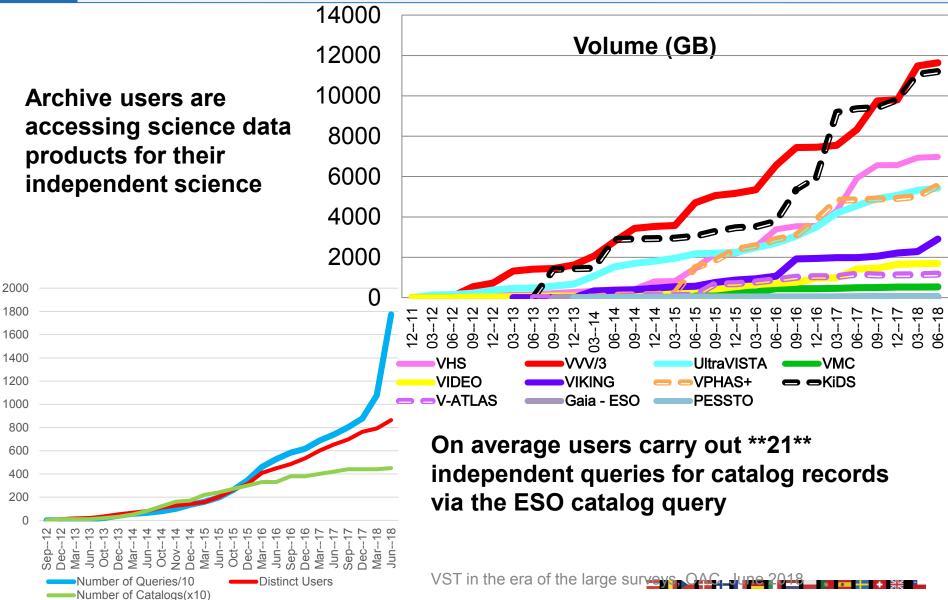
Dark matter may be less dense and more smoothly distributed throughout space than previously thought. An international team used data from the Kilo Degree Survey (KiDS, PI Kujken) to study how the light from about 15 million distant galaxies was affected by the gravitational influence of matter on the largest scales in the Universe. The results appear to be in disagreement with earlier results from the Planck satellite. (*) Hildebrandt et al. 2017, MNRAS,465,1454



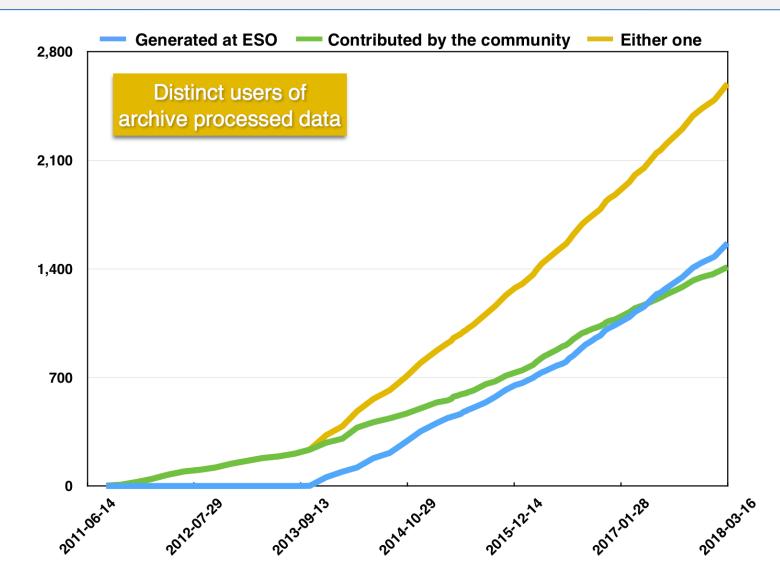
Returns for the community http://archive.eso.org/scienceportal



Returns for the community http://archive.eso.org/scienceportal



Science data products







FORWARD LOOK

VST in the era of the large surveys, OAC, June 2018

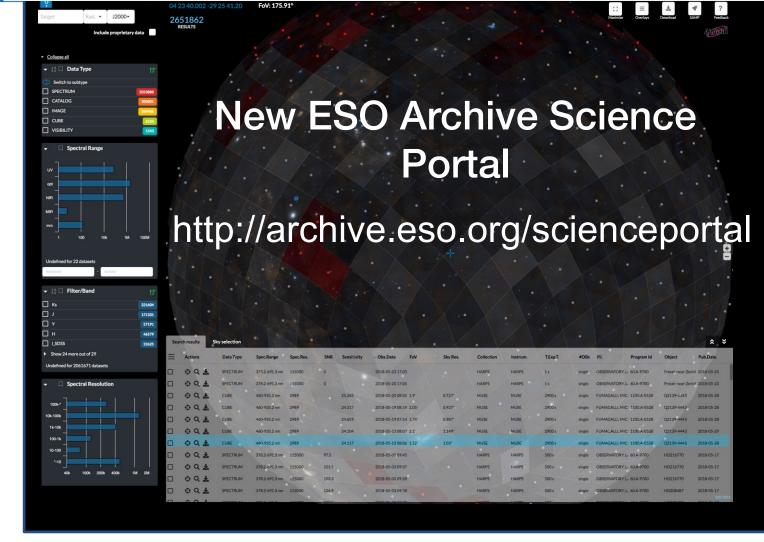


Forward look (I)

- Next Generation Transit Surveys (NGTS) started operations in 01.04.2016 & data products delivery ongoing (Pub. 07.2018)
- Construction and deployment in operations of two wide field spectrographs – <u>the Multi Object Optical and Near Infrared</u> <u>Spectrograph</u> (MOONS) – and the <u>4 meter Multi Object</u> <u>Spectroscopic Telescope</u> (4MOST) in 2020- 2022
- These spectrographs on VLT/VISTA have large multiplexing (1000, ~2400 fibers), wide-field (500 sq. arcmin, 4.1 sq. deg), extended wavelength coverage (0.6-1.8 micron; 0.37-0.95 micron)
- Effective follow-up of the interesting candidates identified via their colors or morphological properties provided by the public imaging surveys and/or space missions, like Gaia and eROSITA among others



Forward look (II)





Archive Science Portal



Signal-to-Noise Ratio

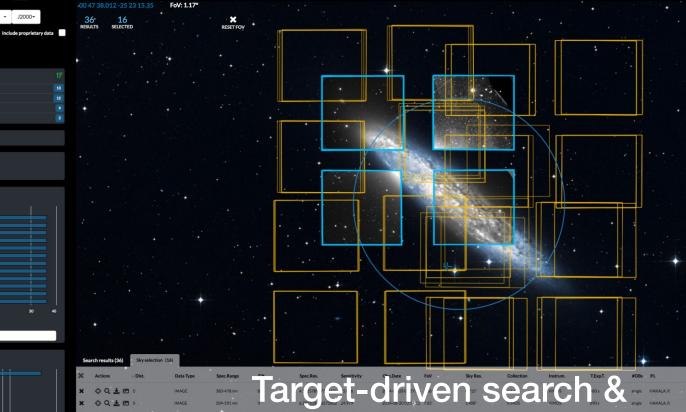
Date of Obs

2009

2008 2005

No data to displ





014-08-20 03:33:23 7.6

Iti-resolution preview of imaging data VIMOS

IMAGE 0Q 10 0 × 0Q 10 0 014-08-21 04:40:18 7.65 ADP.2018-03-26T12:59:35.239 OQLE

10

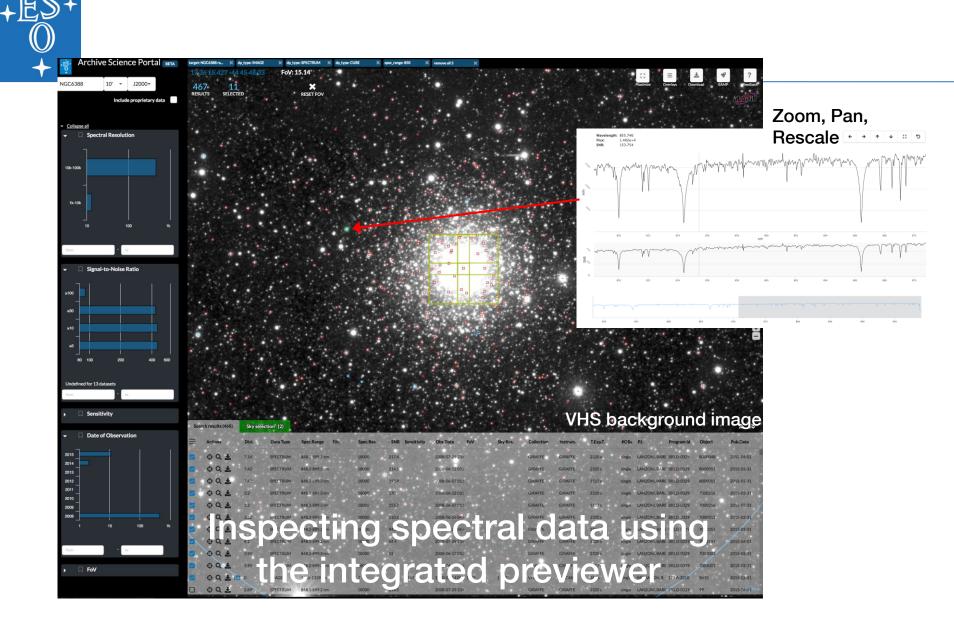
OQ 1

383-478 nm

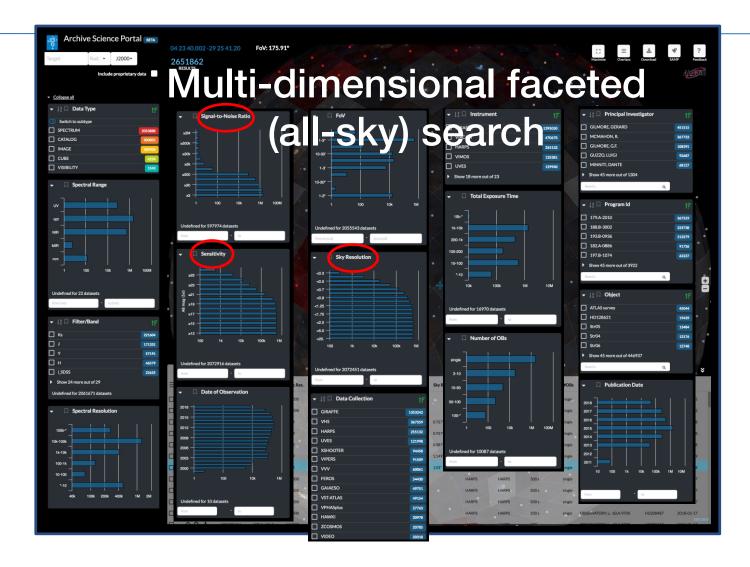
VST in the era of the large surveys, OAC, June 2018

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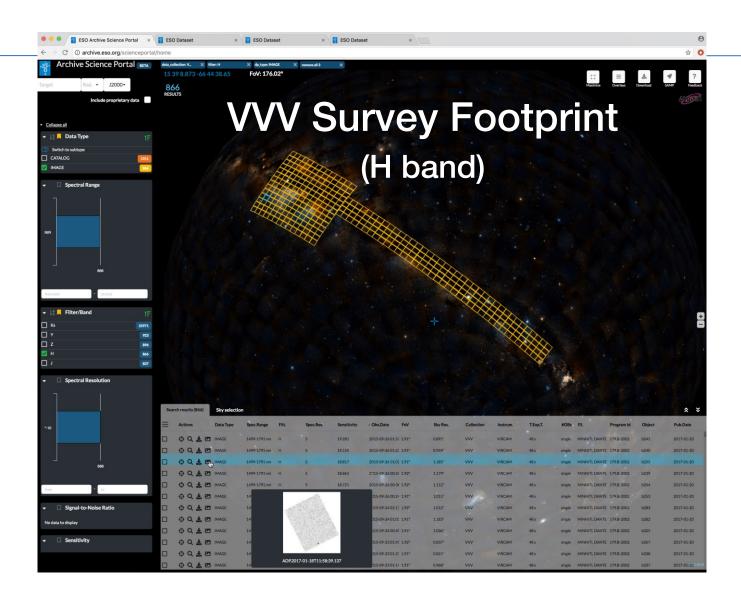
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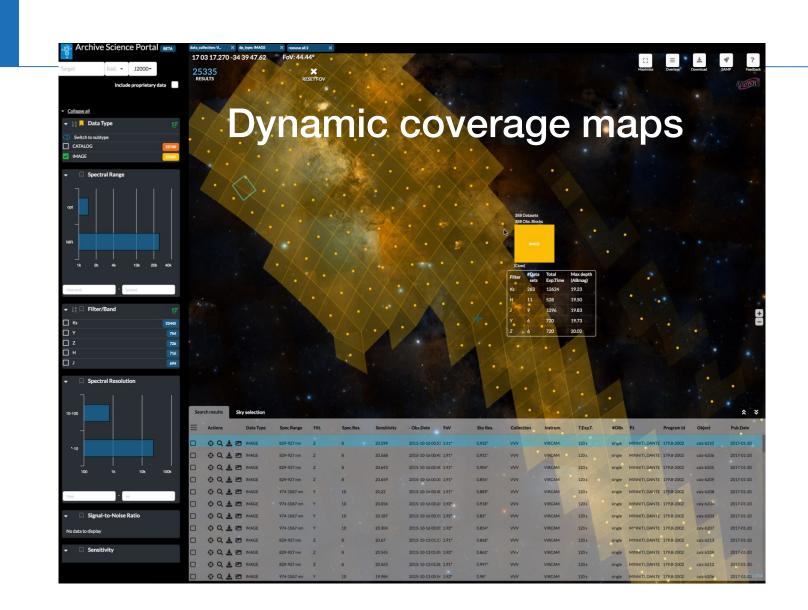








VST in the era of the large surveys, OAC, June 2018



VST in the era of the large surveys, OAC, June 2018



Summary of ASP Features

- Multi-dimensional faceted search
- Zooming and panning the sky view (powered by Aladin light, CDS)
- Aggregations expose the entire archive content
- Iterative queries (adding and dropping constraints)
- Footprints & dynamic coverage maps
- Highlighting of datasets across views
- Previews for all products
- Spectral data viewer
- Color-coding according to data type

- HiPS multi-resolution previews of imaging datasets
- Search by target name or around a position
- Background Imagery from HiPS network
- Autocompletion of text-based fields
- Interval constraints
- Server-side sorting
- One-click download
- Dataset detail view
- Interoperability via SAMP & programmatic access



Conclusions

Public surveys are leading the way in observational astronomy

- Projects such as these are characterized by large investments in ``survey systems" that include <u>dedicated telescopes and</u> <u>instruments</u>, a large community of astronomers involved in the science projects and <u>large networks for the data distribution</u>.
- The scientific success of survey projects includes the legacy values of the science products that become available through the archives for further scientific analysis by the community
- ESO Public Surveys are examples of effective implementations of such systems, and empower the community at large to discover the Universe, in the spirit of the IAU statement "The Universe: yours to discover"



Acknowledgements

- We would like to thank our La Silla Paranal colleagues, for their work and support to the science operations of the ESO Public Surveys.
- We wish to acknowledge our colleagues from the ESO Department of Engineering for their support in developing the tools required for carrying out Phase1, Phase 2 and Phase 3 for the ESO Public Surveys, and the ESO library team for the careful monitoring of the refereed publications.
- We wish to thank the PIs of the Public Surveys and their collaborators, including the data centers at CASU, WFAU, ASTROWISE and TERAPIX, for their hard work and support to the ESO mission.