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UNIVERSITY OF CAPE TOWN
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Thanks to Mattia Negrello (Cardiff) Stephen Serjeant (OU),
Steve Crawford (SAAO), Mattia Vaccari (UWC)...



Strong Gravitational Lensing Studies with Herschel
and Multi-Wavelength Follow-up Observations:
synergies with KiDS

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 HERMES

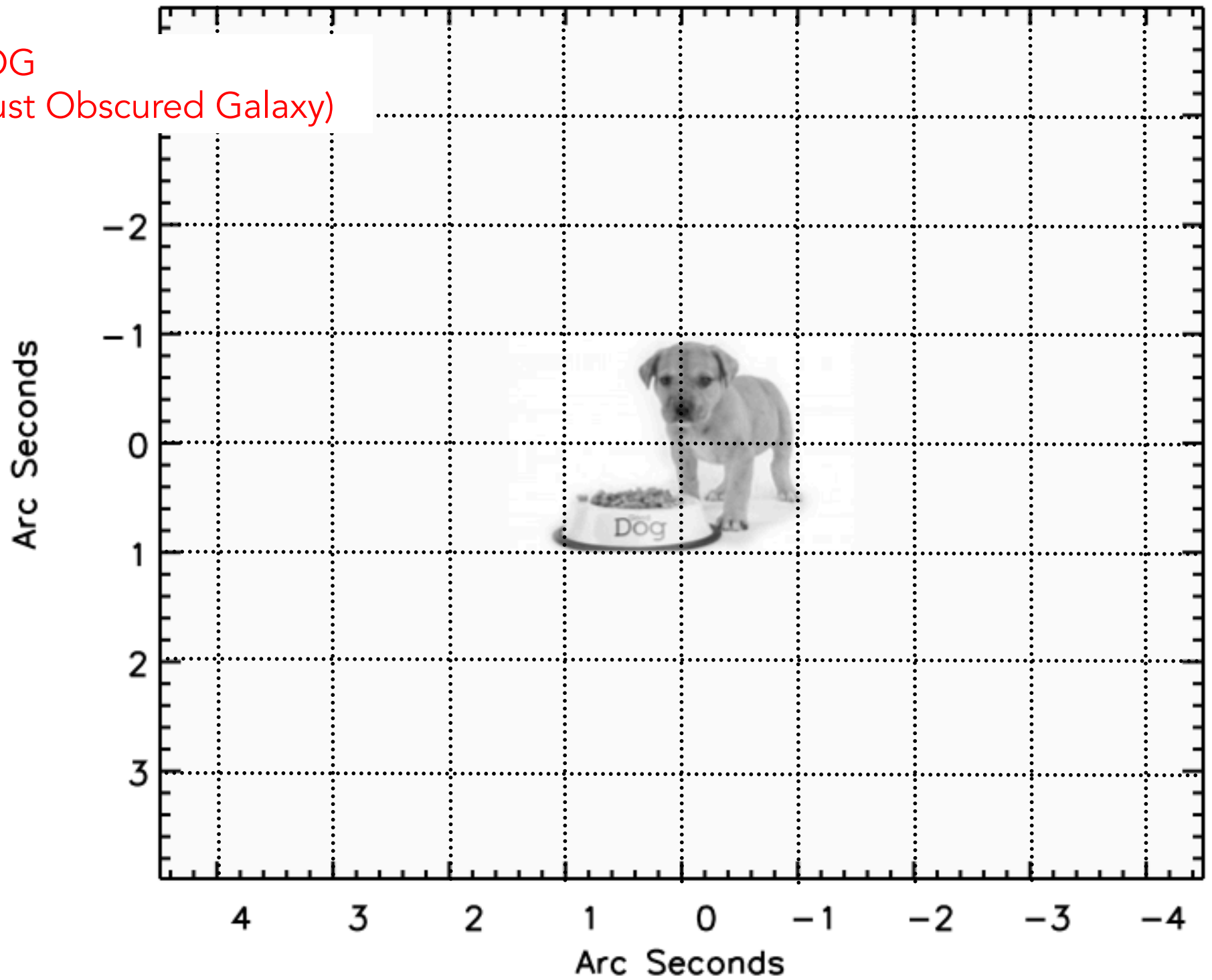


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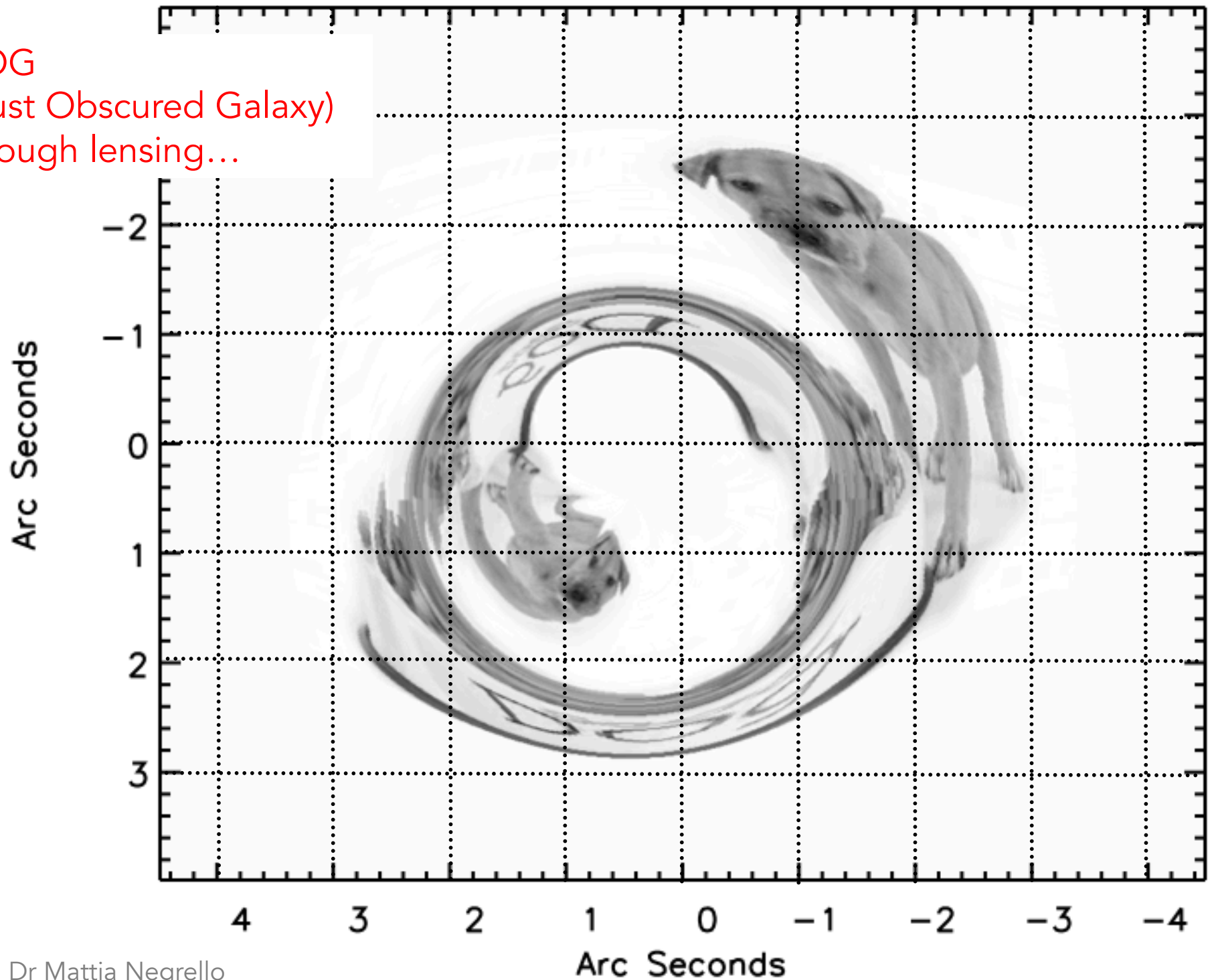


Strong Gravitational Lensing Studies with Herschel
and Multi-Wavelength Follow-up Observations:
synergies with KiDS

DOG
(Dust Obscured Galaxy)



DOG
(Dust Obscured Galaxy)
through lensing...



Credits: Dr Mattia Negrello

What is gravitational lensing?



- Gravitational lensing **conserves surface brightness**
- Angular magnification means **flux magnification**
- Multiple images and “Einstein ring” are possible

How to find lenses (before Herschel)



Gravitational lensing is a rare phenomenon.

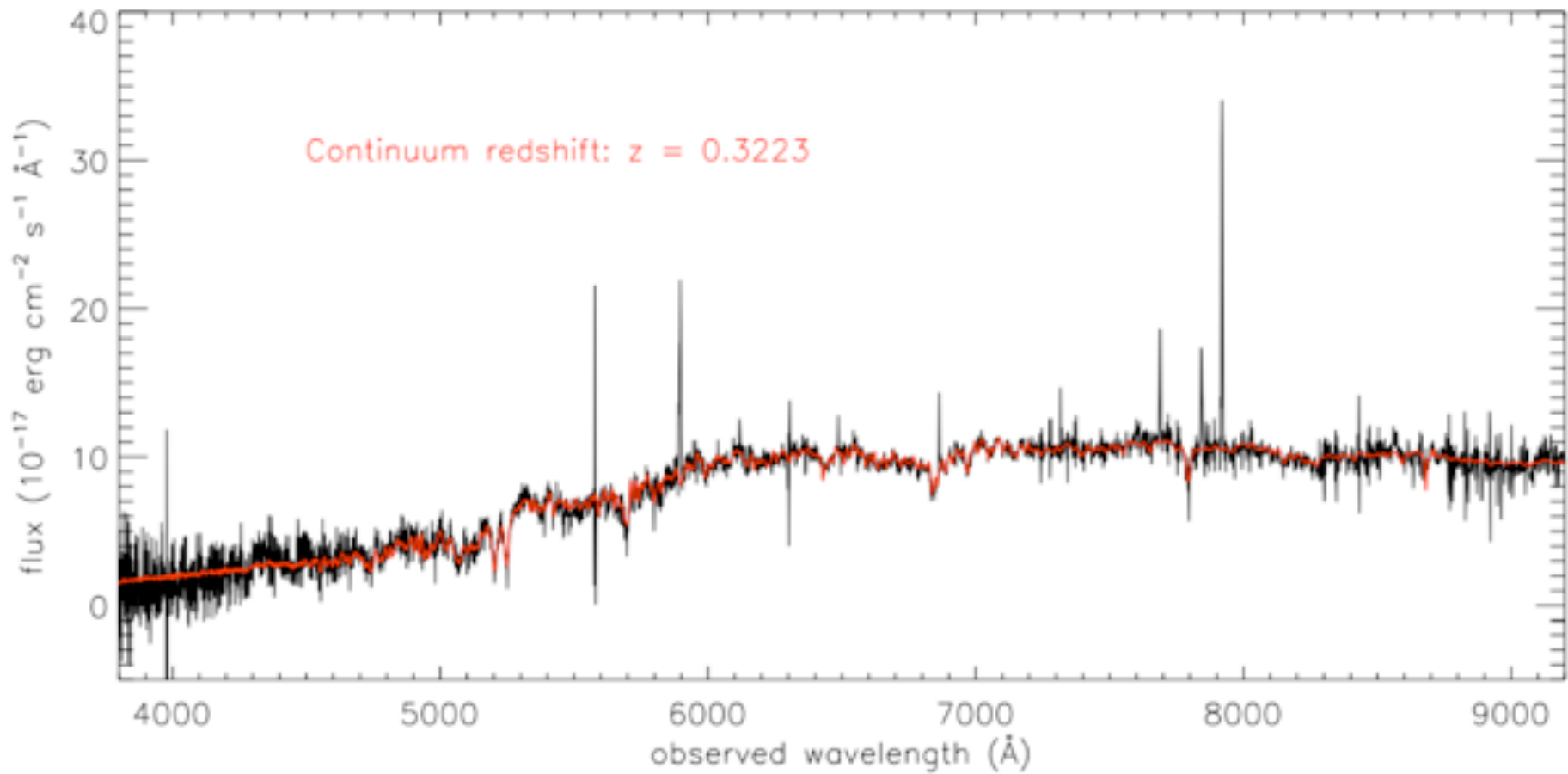
A number of different strategies:

- **Look for** potential lenses or **potential deflectors** and use additional information to identify the (small) subset of strong gravitational lensing events.
- **Searching for** gravitational **lensing morphologies** in high-resolution data (e.g. HST, Marshall et al. 2009)

How to find lenses - Optical



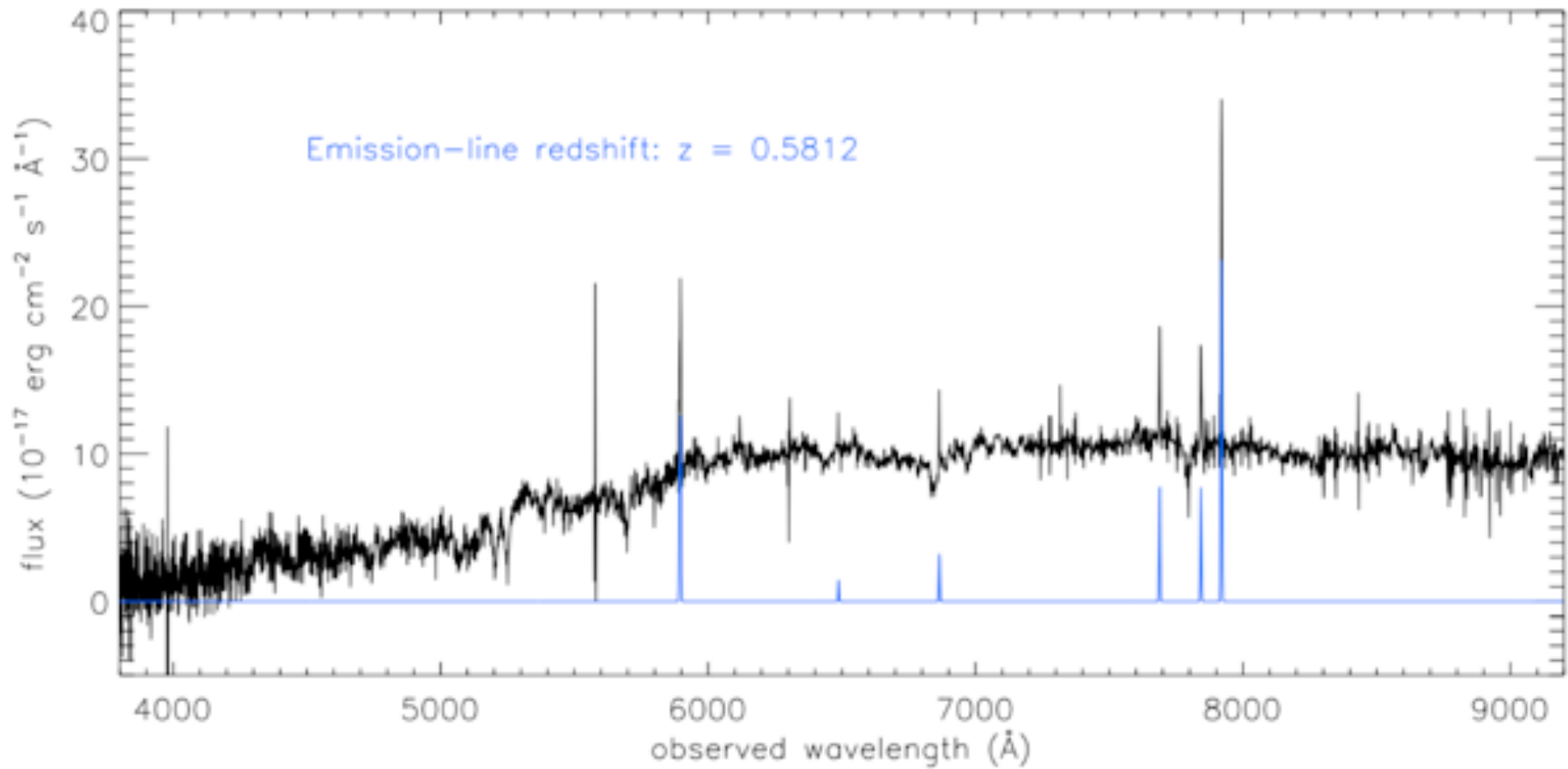
SLACS



How to find lenses - Optical



SLACS

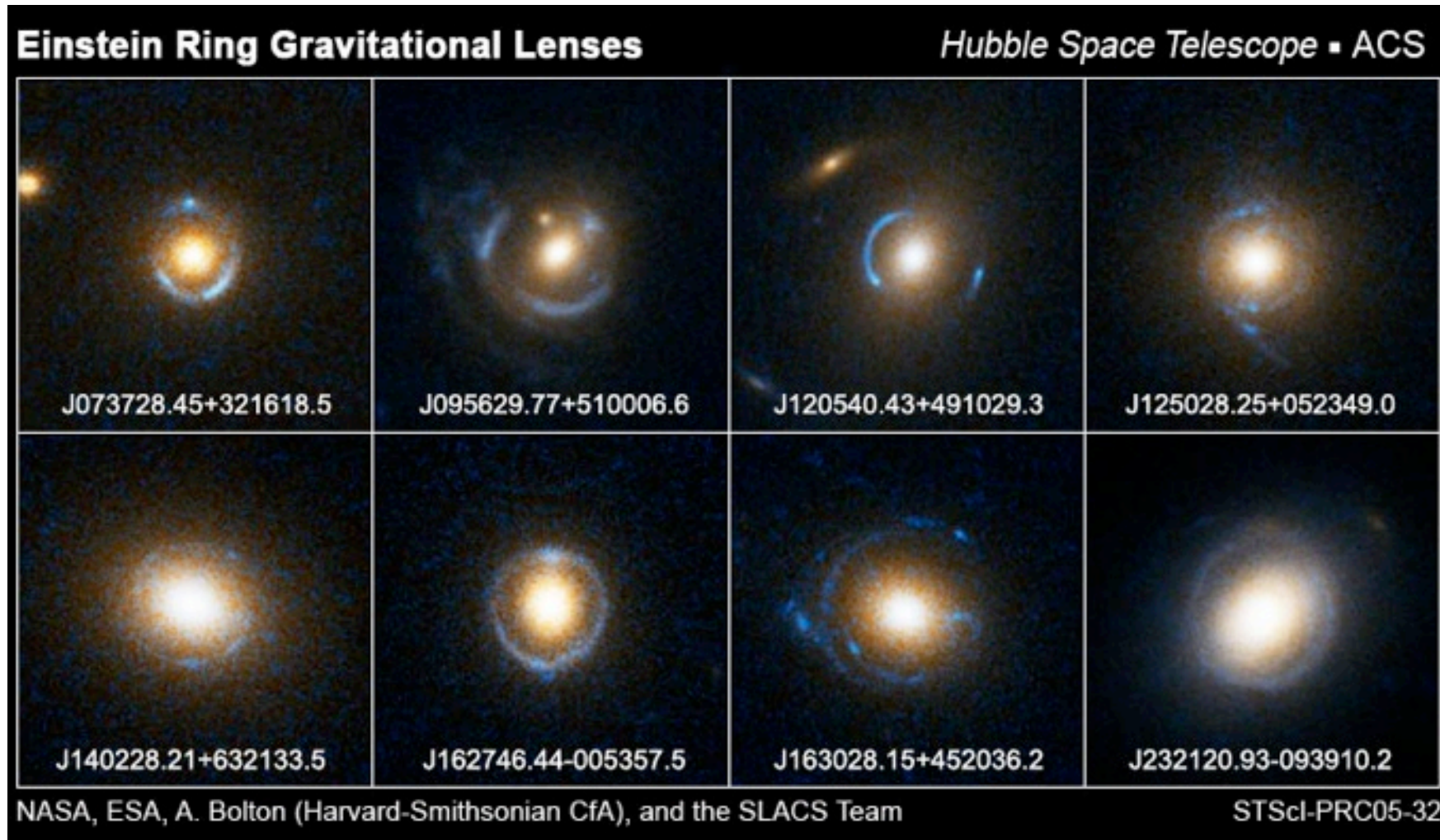


How to find lenses - Optical

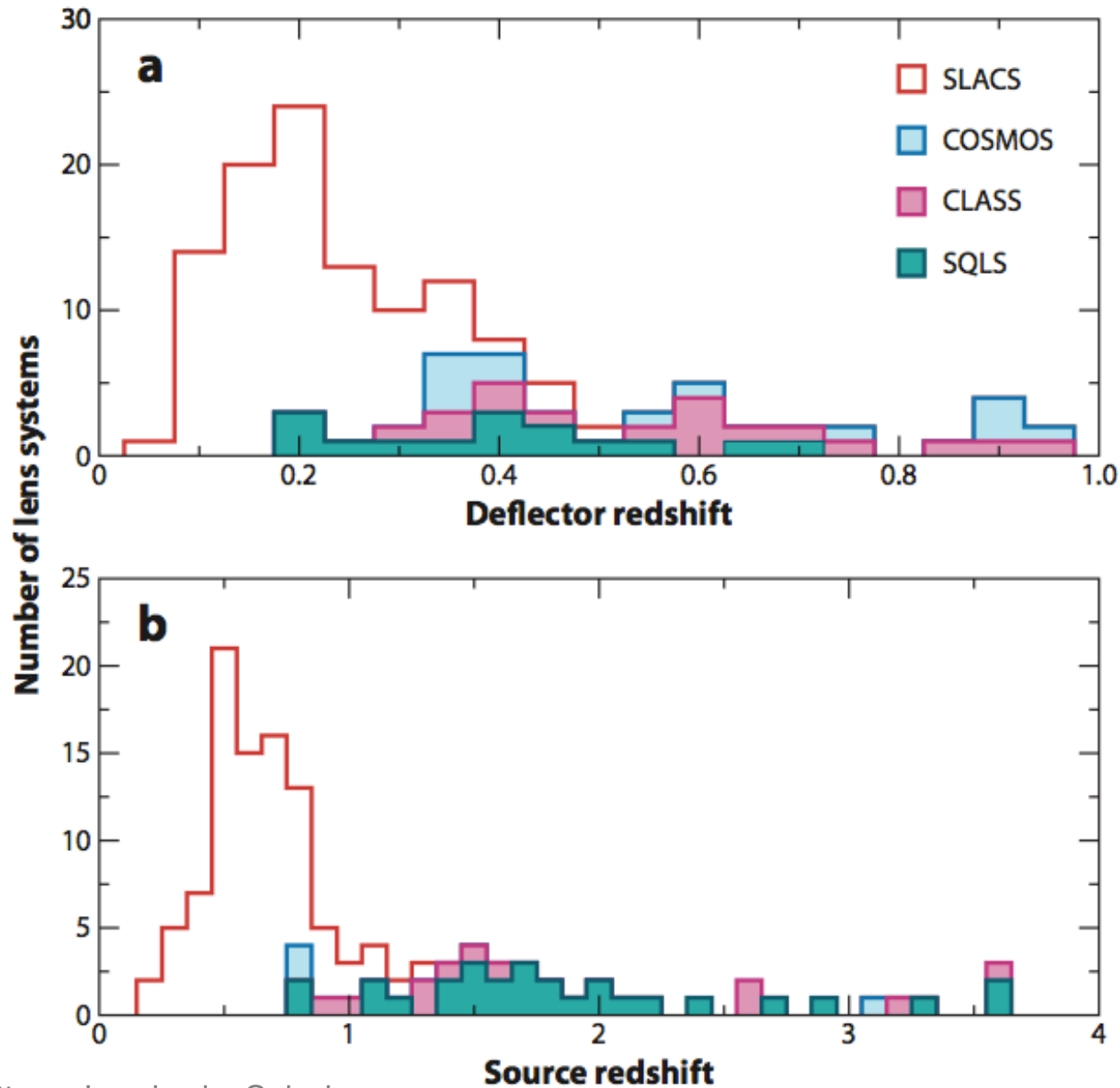


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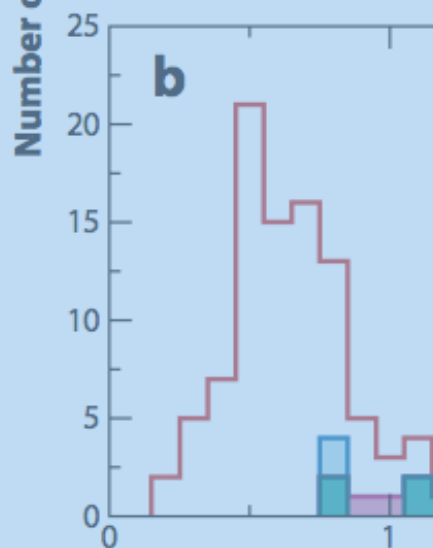
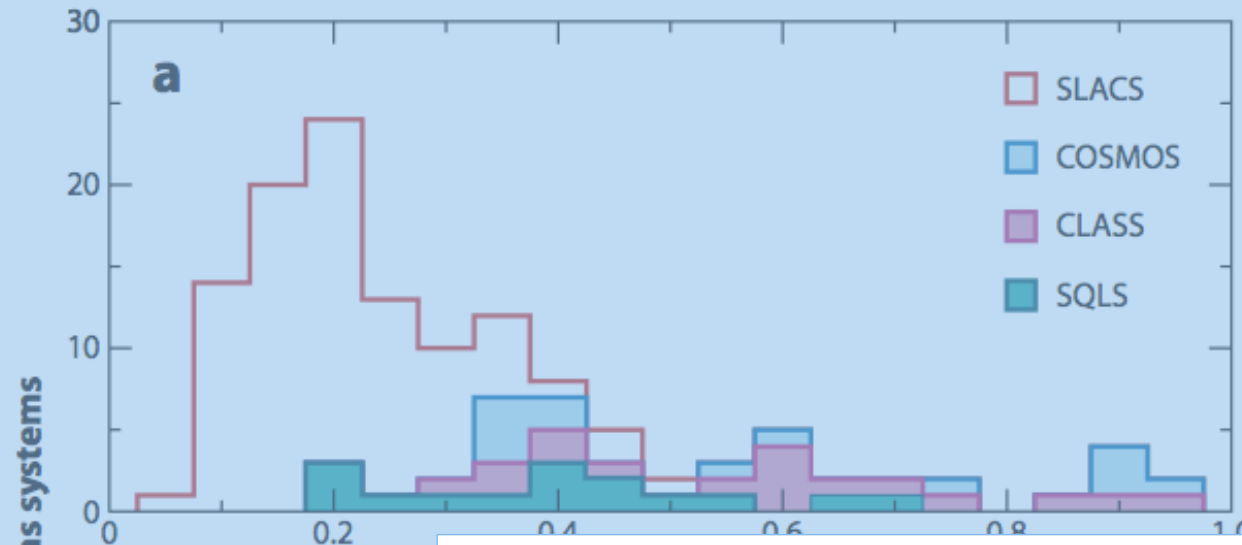
SLACS



How to find lenses - Optical



How to find lenses (before Herschel)



TWO MAJOR LIMITS:

1- most lenses have been found at $z < 0.4$, which is

→ very favorable regime for detailed follow-up

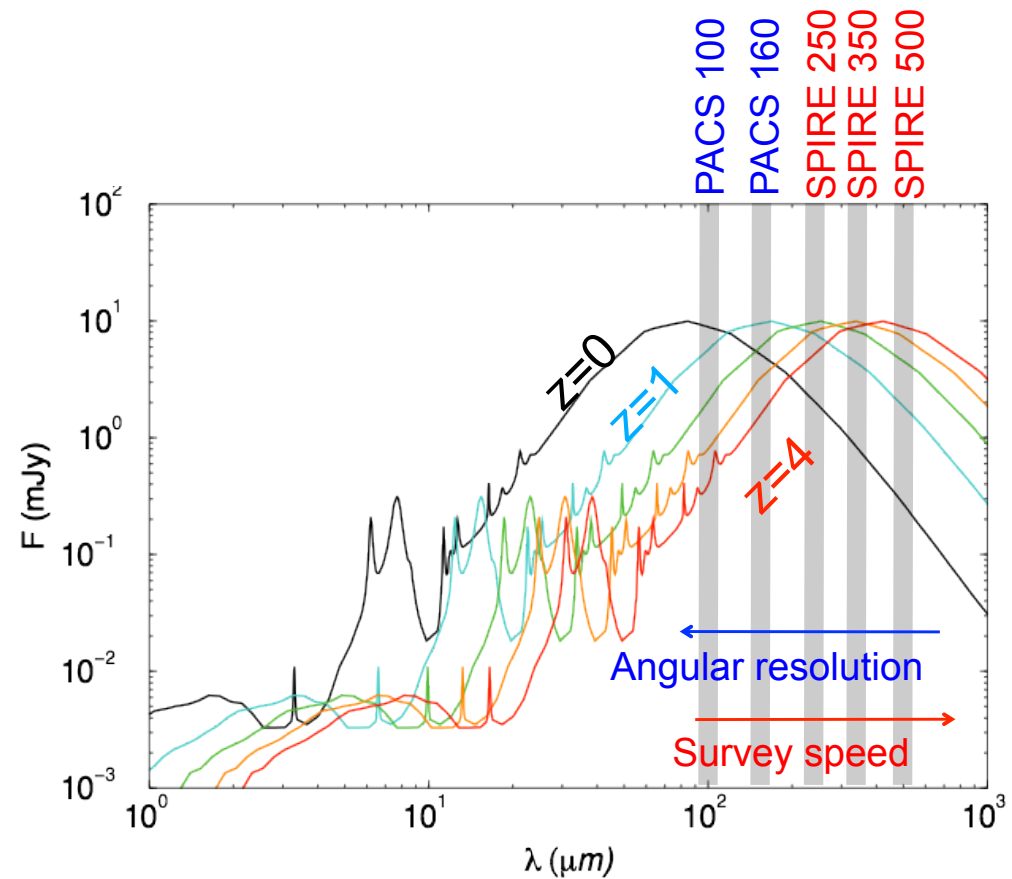
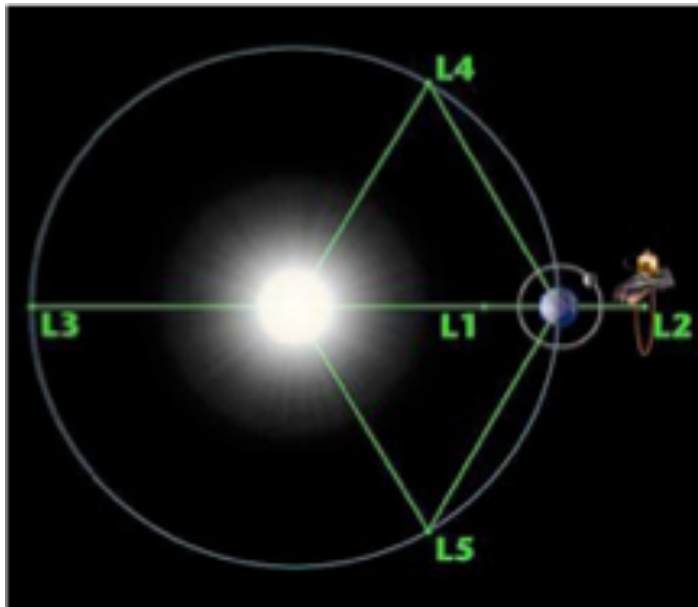
→ but limits the look-back time baseline for evolutionary studies and the spatial scales probed by lensing.

2. Many gravitational lens systems still lack source or deflector redshifts.

A breakthrough: Herschel

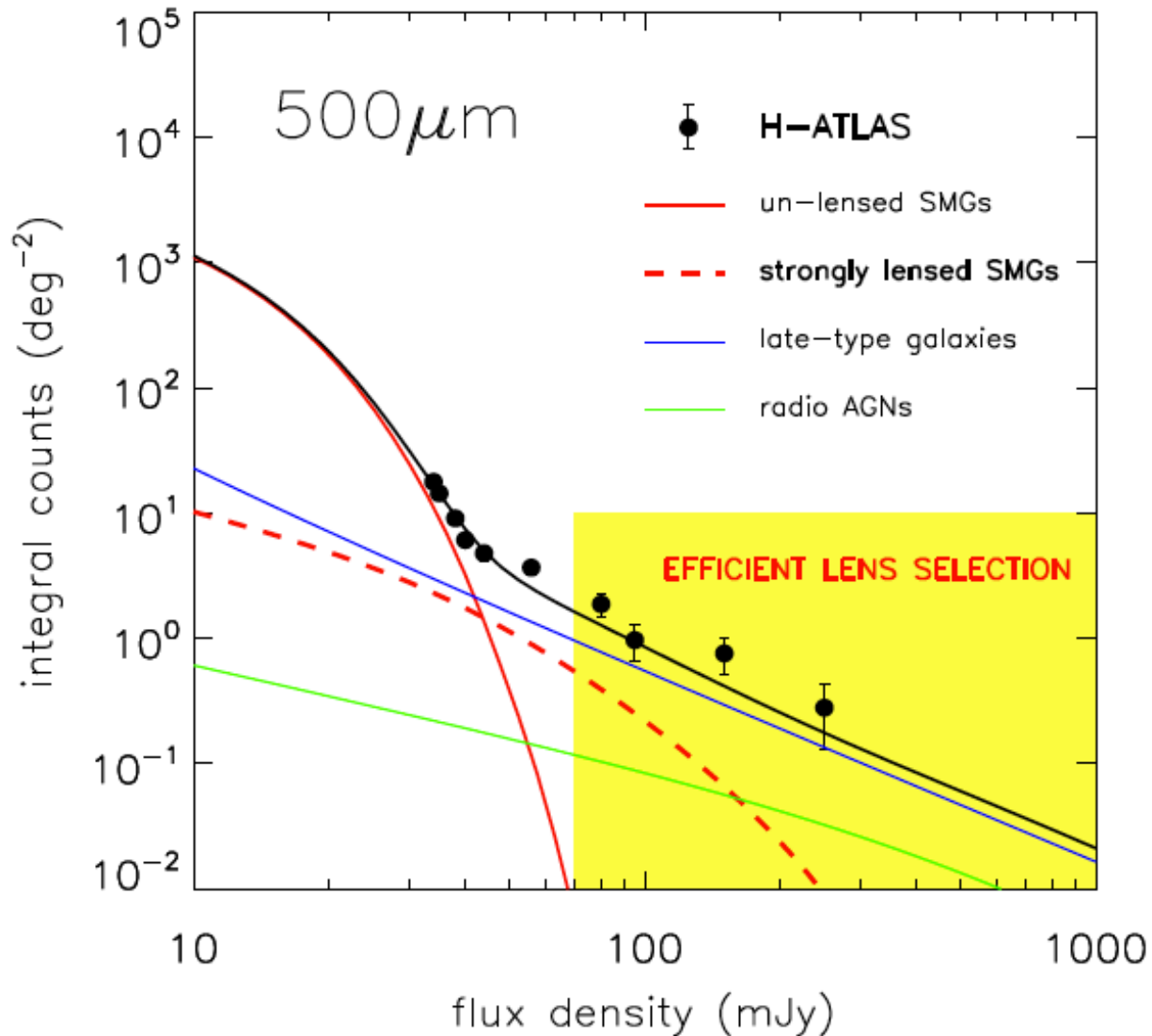
Herschel is an **ESA cornerstone mission (2009-2013)**

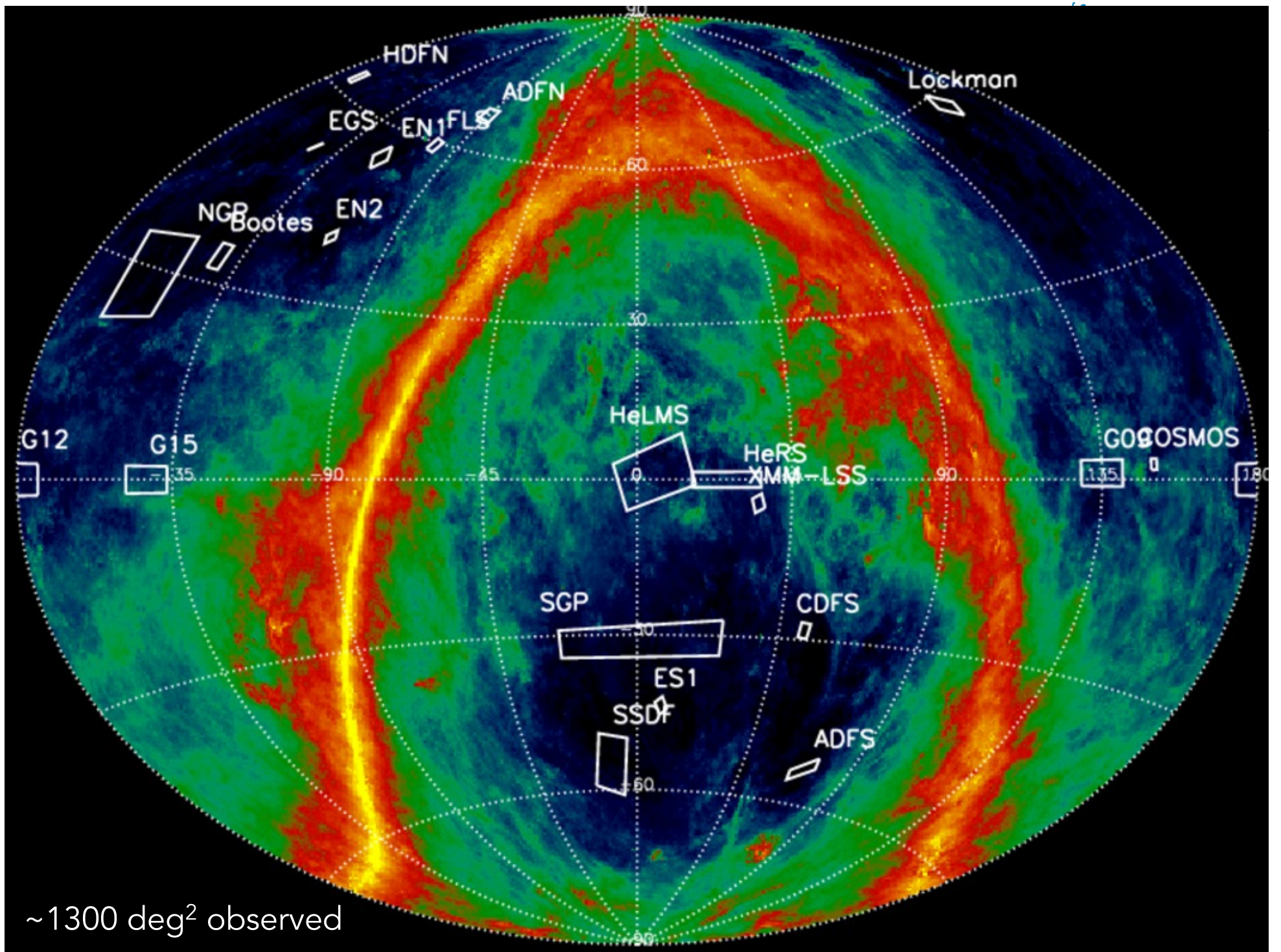
- large (3.5 m) aperture, low emissivity ($\sim 5\%$), passively cooled (70-90 K)
- cryogenically cooled focal plane science instruments



A breakthrough: Herschel

Negrello et al. 2010 Science 330, 800

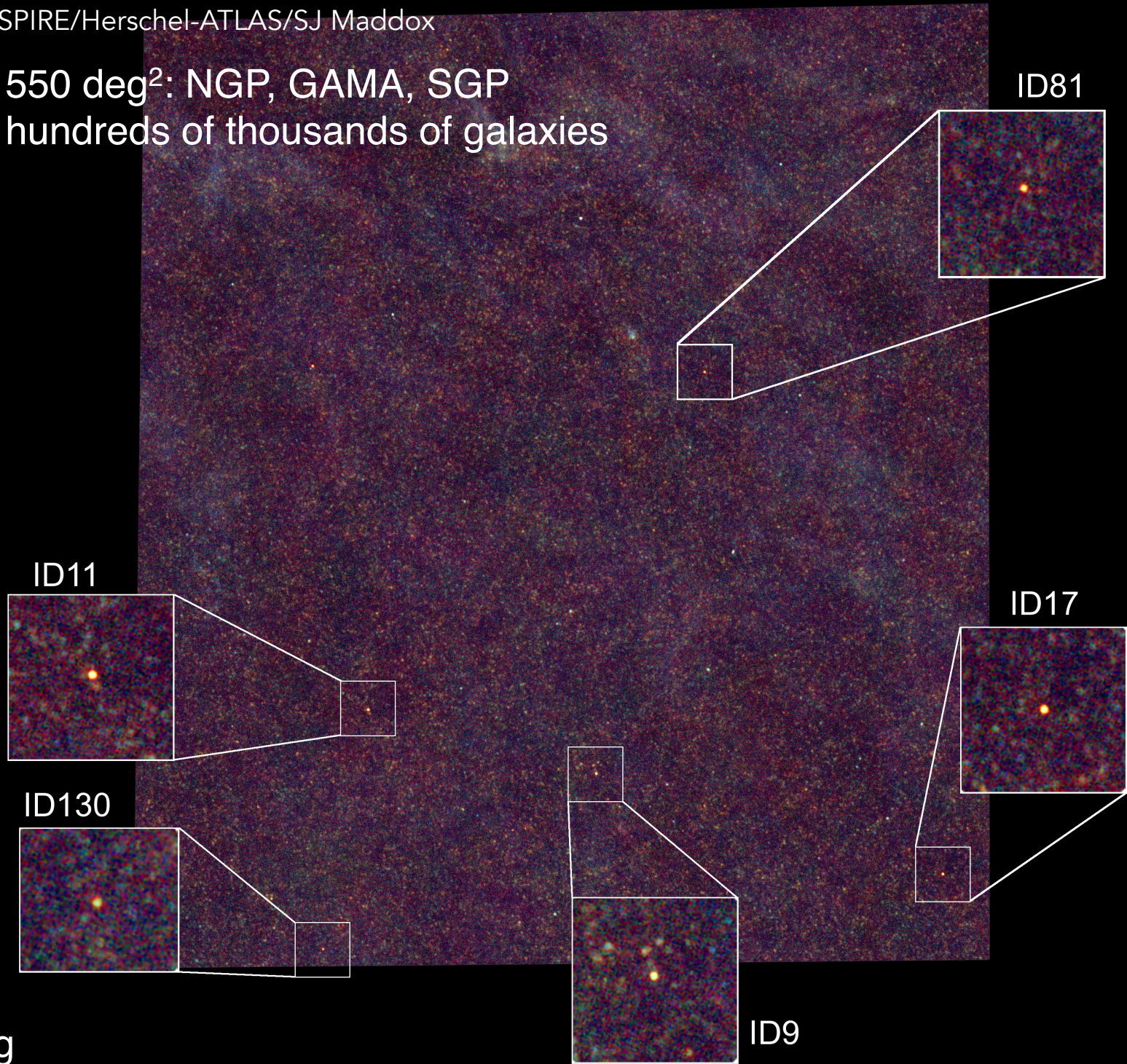




Credit: ESA/SPIRE/Herschel-ATLAS/SJ Maddox

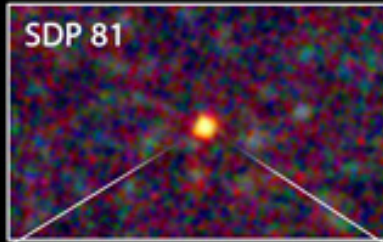


550 deg²: NGP, GAMA, SGP
hundreds of thousands of galaxies

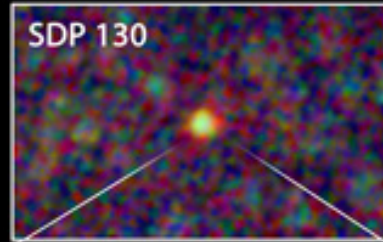


Herschel:

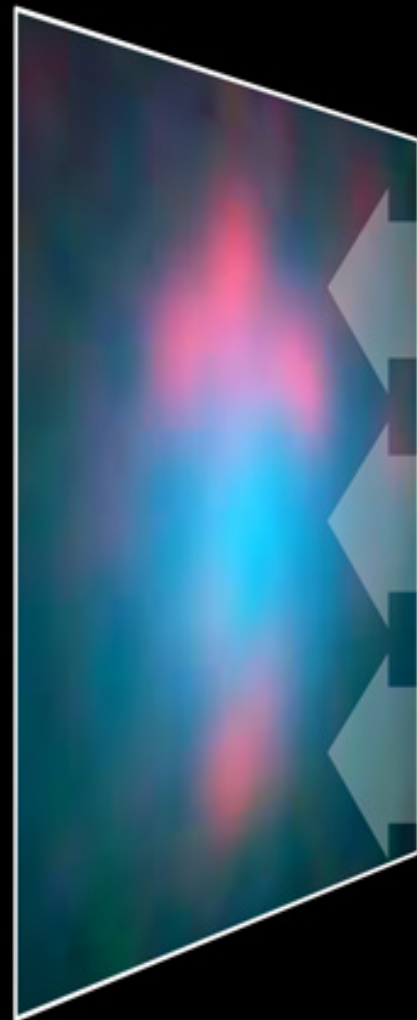
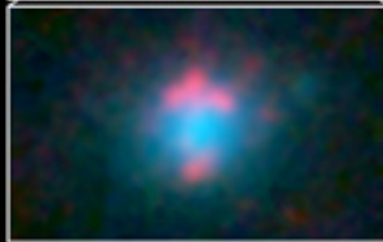
SDP 81



SDP 130



Keck & SMA:



DISTANT GALAXY



FOREGROUND GALAXY



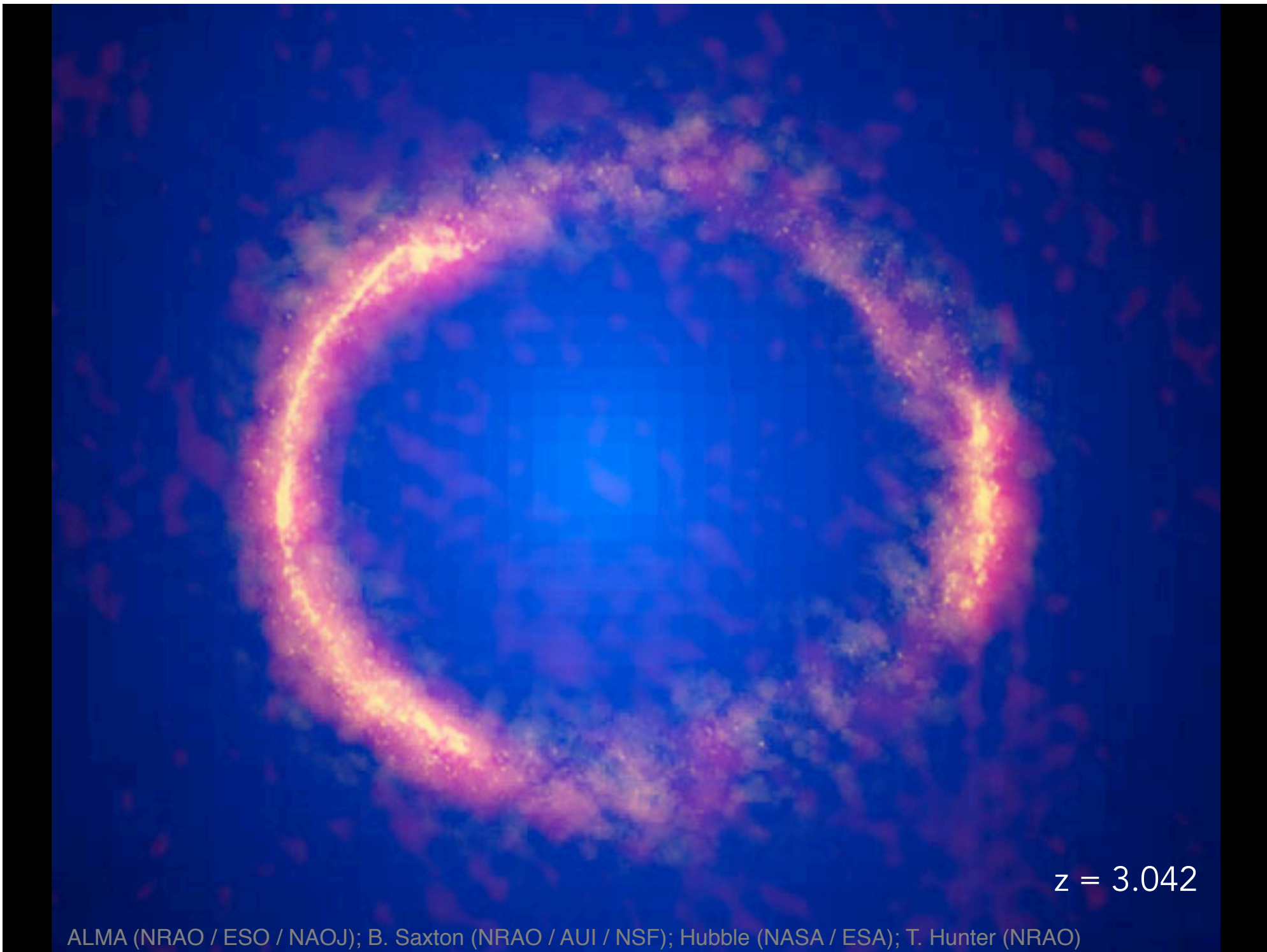
EARTH

LENSED IMAGES OF DISTANT GALAXY

3 BILLION YEARS

11 BILLION YEARS

Credit: ESA/NASA/JPL-Caltech/Keck/SMA



$z = 3.042$

ALMA (NRAO / ESO / NAOJ); B. Saxton (NRAO / AUI / NSF); Hubble (NASA / ESA); T. Hunter (NRAO)

A breakthrough: Herschel



Negrello et al. 2010 Science 330, 800

- ~100% lens selection efficiency
- ~1/deg² strong gravitational lenses in Herschel ATLAS
- Can find lenses at much higher redshifts

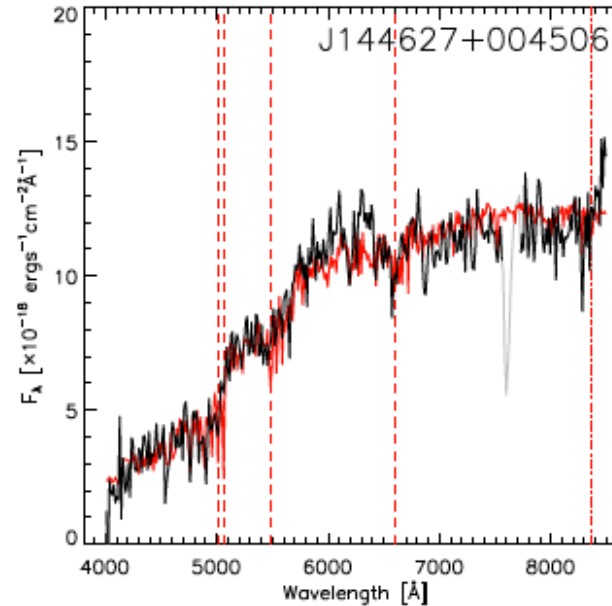
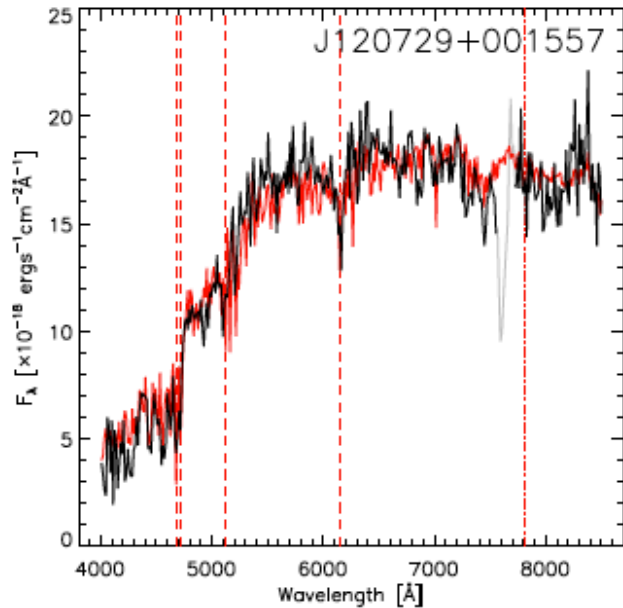
A breakthrough: Herschel



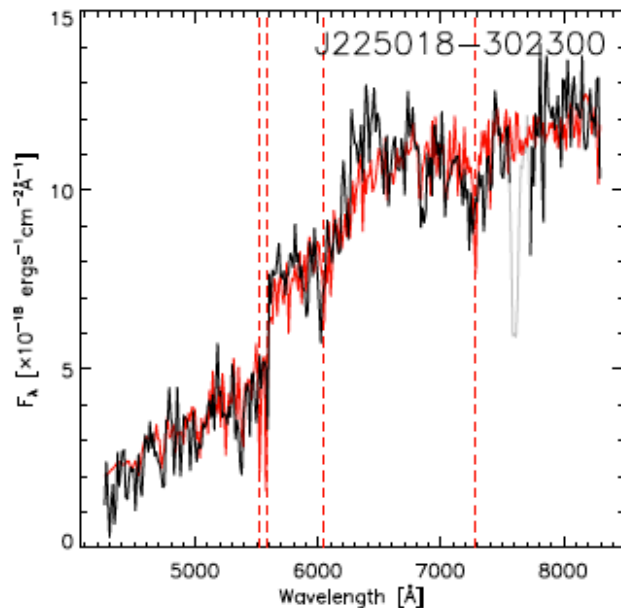
Negrello et al. 2010 Science 330, 800

- ~100% lens selection efficiency
- $\sim 1/\text{deg}^2$ strong gravitational lenses in Herschel ATLAS
- Can find lenses at much higher redshifts than the optical selection
- Purely magnification-based selection
- Need to follow up: optical spectroscopy for foreground redshifts, CO for background redshifts

Multi- λ follow-ups: NTT spec-z



3.58 m telescope

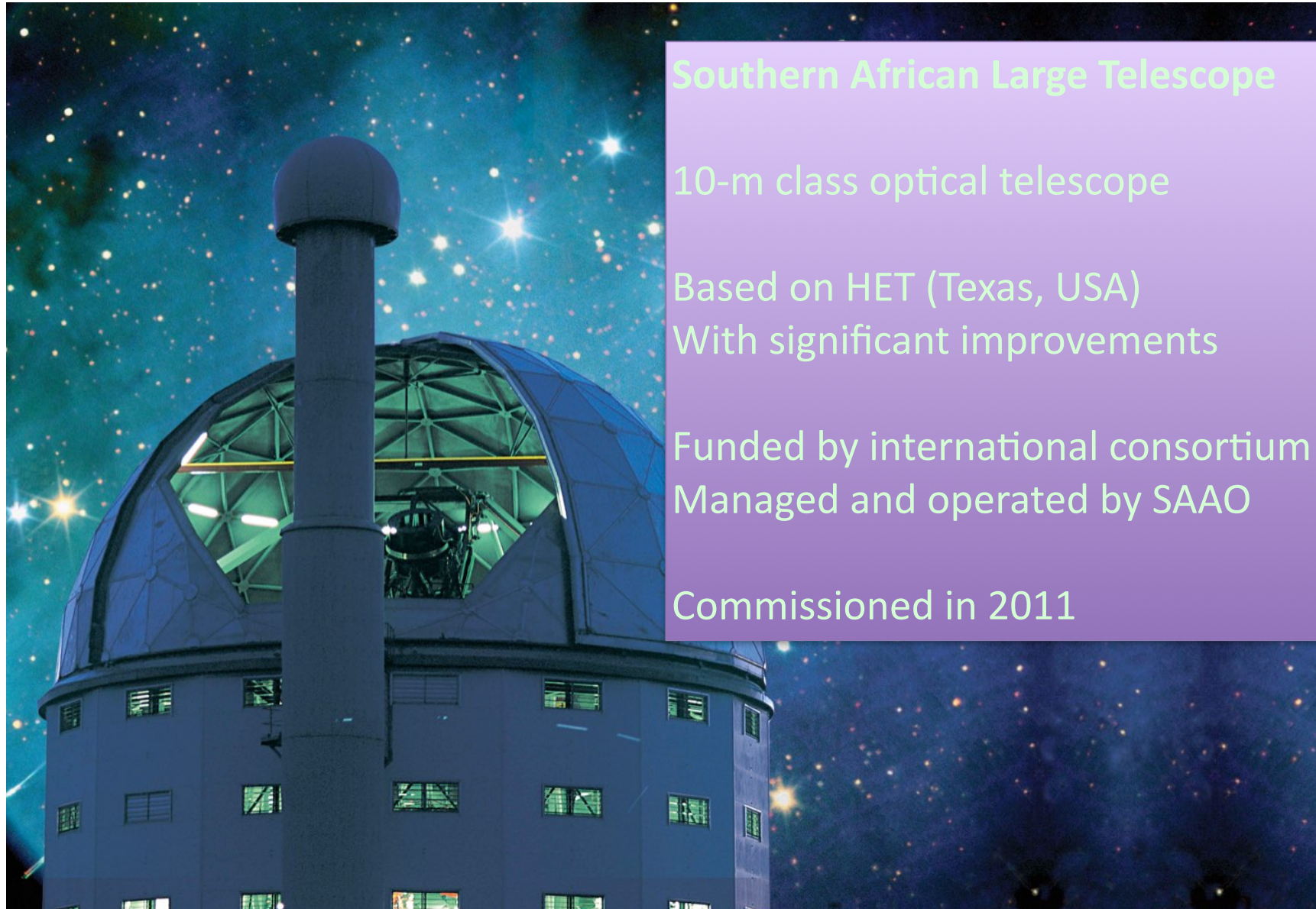


BUT not great success:

- Bad weather conditions
- Need more time per target

Amber, 2015, PhD thesis

Multi- λ follow-ups: The SALT Strong Gravitational Lensing Legacy Project



Southern African Large Telescope

10-m class optical telescope

Based on HET (Texas, USA)

With significant improvements

Funded by international consortium

Managed and operated by SAAO

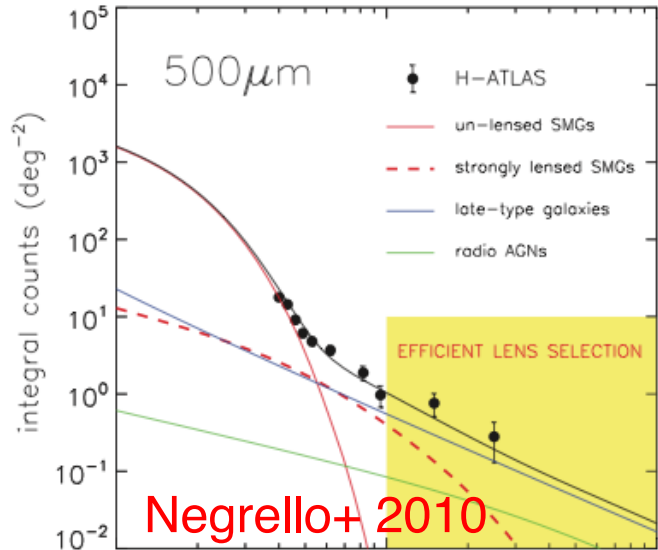
Commissioned in 2011

Multi- λ follow-ups: The SALT Strong Gravitational Lensing Legacy Project

Grating Name	Wavelength Coverage (nm)	Usable Angles (deg)	Bandpass per tilt (nm)	Resolving Power (1.25" slit)
PG0300	370–900		390/440	250–600
PG0900	320–900	12–20	~300	600–2000
PG1300	390–900	19–32	~200	1000–3200
PG1800	450–900	28.5–50	150–100	2000–5500
PG2300	380–700	30.5–50	100–80	2200–5500
PG3000	320–540	32–50	80–60	2200–5500



Multi- λ follow-ups: The SALT Strong Gravitational Lensing Legacy Project

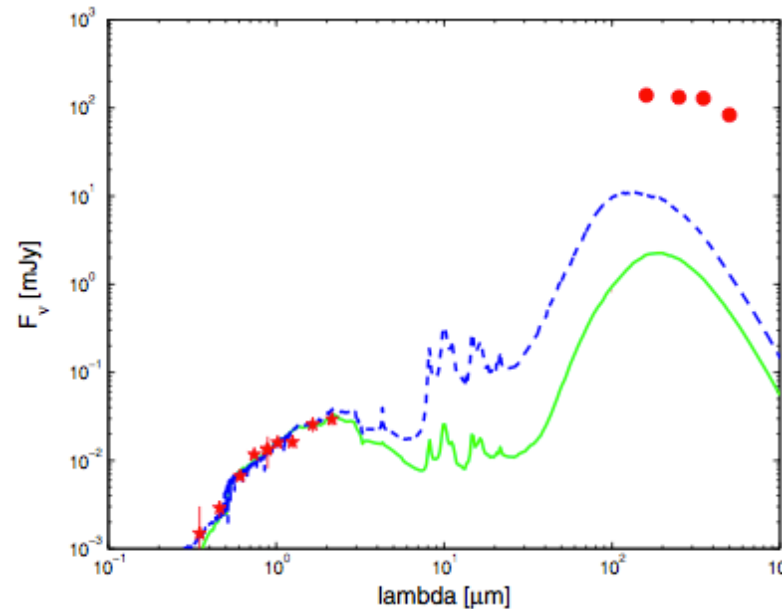
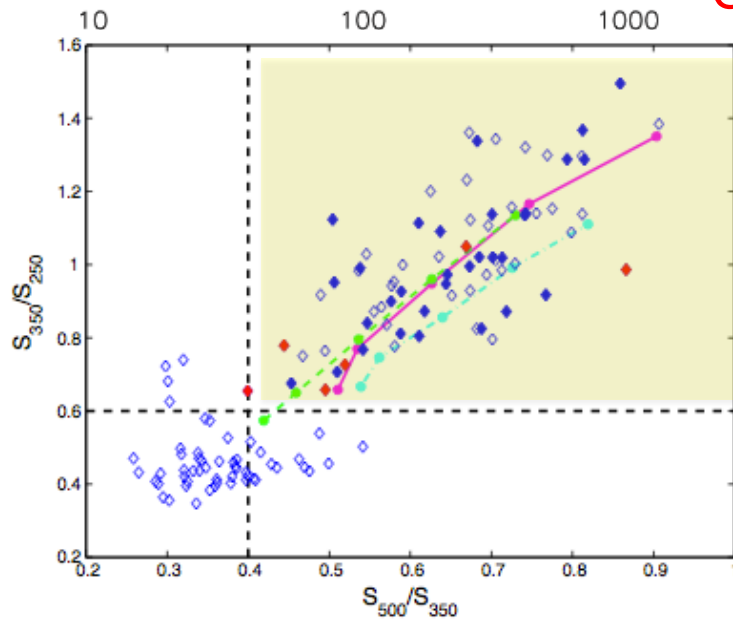


~100% lens selection efficiency
S500 > 100mJy

+

~80% lens selection efficiency
Combining Submm + NIR (+ Optical)
S500 > 80mJy

Gonzalez-Nuevo + 2012



Multi- λ follow-ups: The SALT Strong Gravitational Lensing Legacy Project



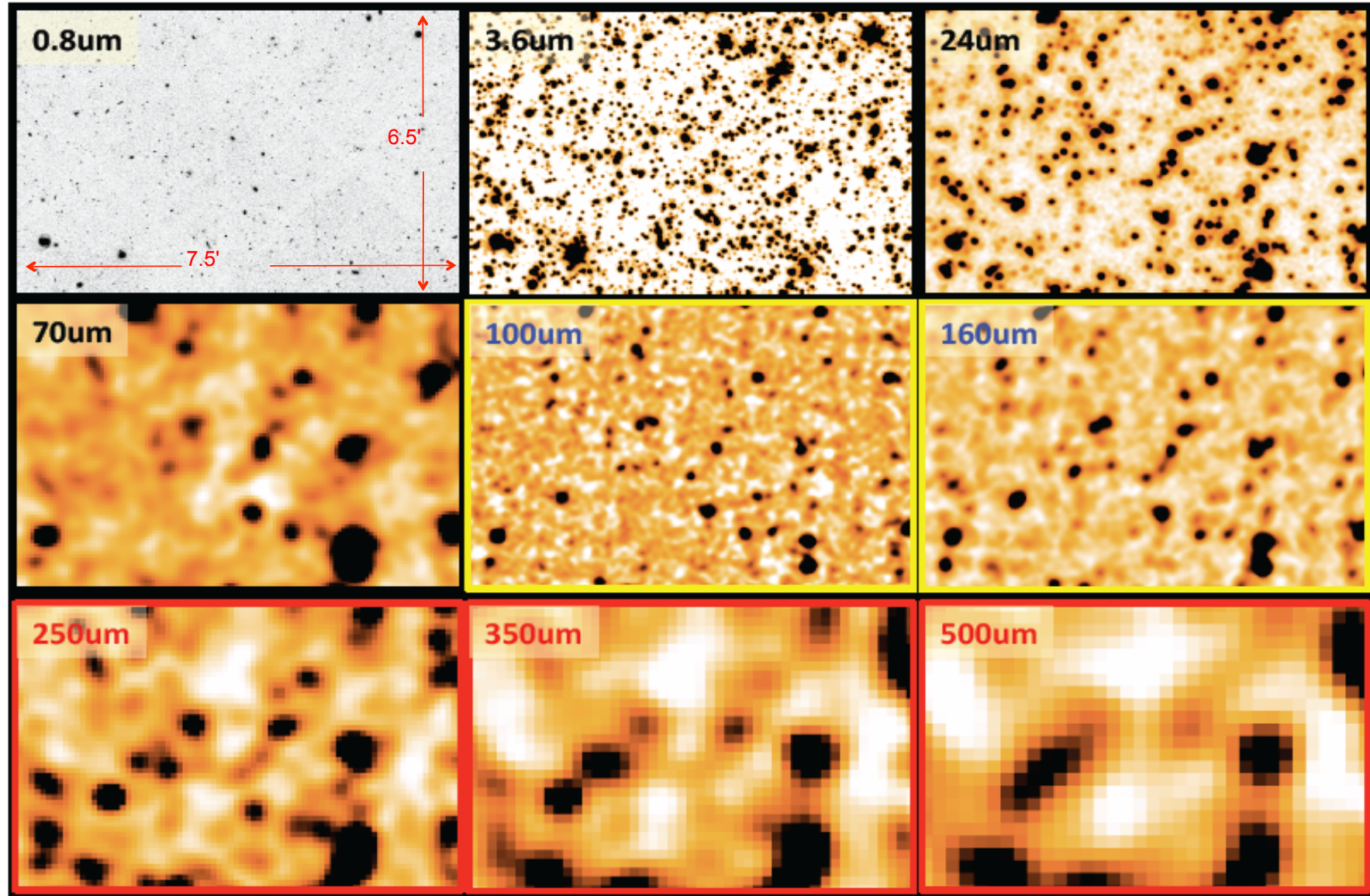
- ~1000 deg² Herschel surveys at all RAs
- ~500 Southern strong lens candidates

Multi- λ follow-ups: The SALT Strong Gravitational Lensing Legacy Project



- ~1000 deg² Herschel surveys at all RAs
- ~500 Southern strong lens candidates
- ~2 lenses per hour / long-slit RSS spec.
- 300hrs grey nights, 3-yr long term program
- ~1.5'' seeing, poor transparency

How to identify the right counterpart?



From Herschel to the SALT Strong Gravitational Lensing Legacy Project



<https://herschel.sussex.ac.uk>

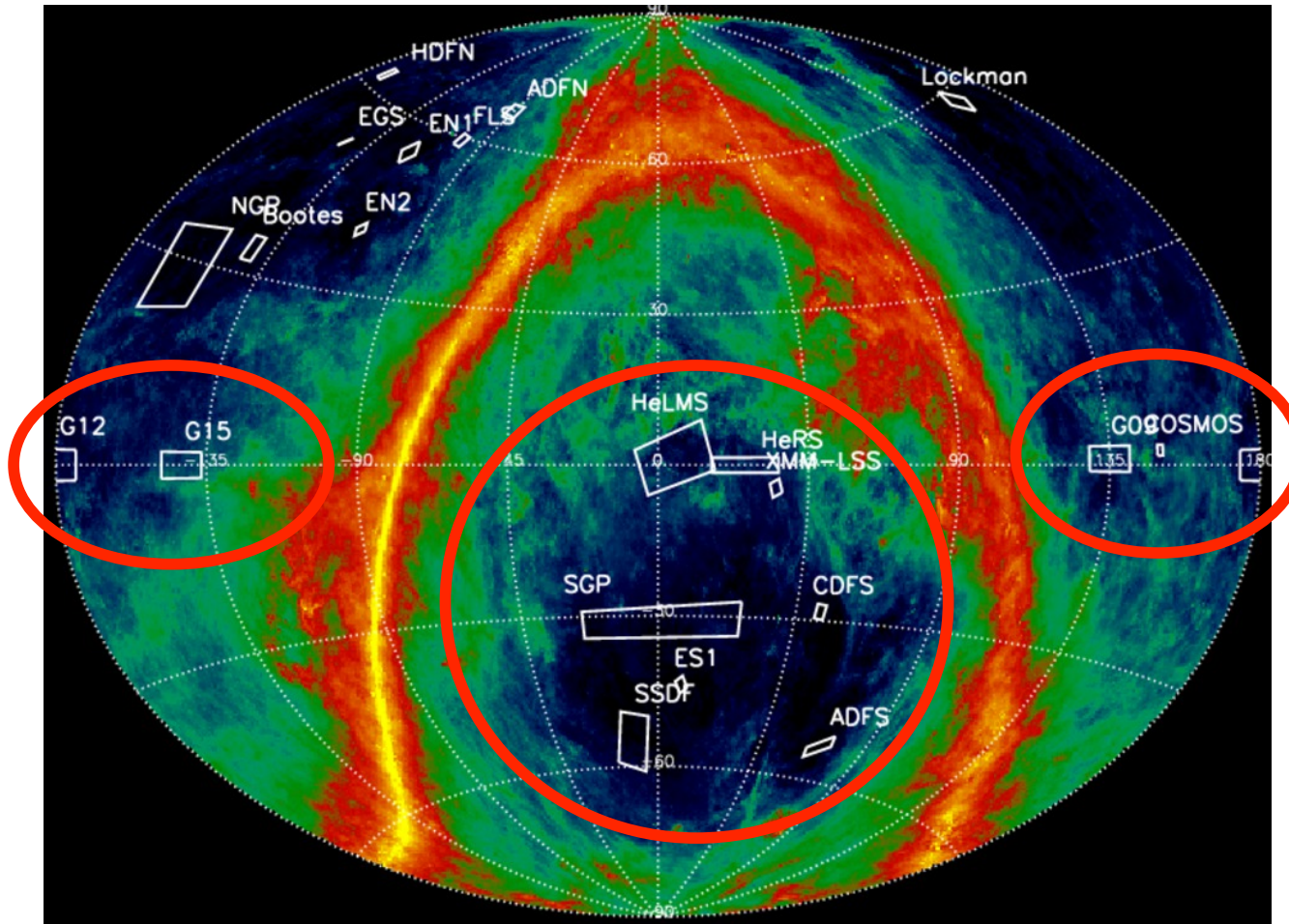


HERSCHEL EXTRAGALACTIC LEGACY PROJECT

HOME ABOUT CONTACT US MEETINGS SURVEYS TEAM MEMBERS VISUAL MEDIA

- 2.5 MEuros EC-REA FP7-SPACE Project co-funded by South African DST
- Bringing together Multi-Wavelength Data Over Herschel Extragalactic Fields
- Releasing Data Products and Software Tools to the Astronomical Community
- PI : Seb Oliver (Sussex)
- Project Scientists : Peter Hurley (Sussex) & Mattia Vaccari (UWC)

Multi- λ follow-ups: The SALT Strong Gravitational Lensing Legacy Project



- GALEX (UV)
- VST (ugriz)
- DES (grizy)
- VISTA (ZYJHK)
- IRAC & MIPS
- PACS & SPIRE
- GMRT/JVLA/LOFAR/WSRT
- ASKAP/MeerKAT

See also Spitzer Data Fusion @ <http://mattiavaccari.net/df/>
Vaccari 2015, Marchetti PhD thesis & Marchetti et al. 2016

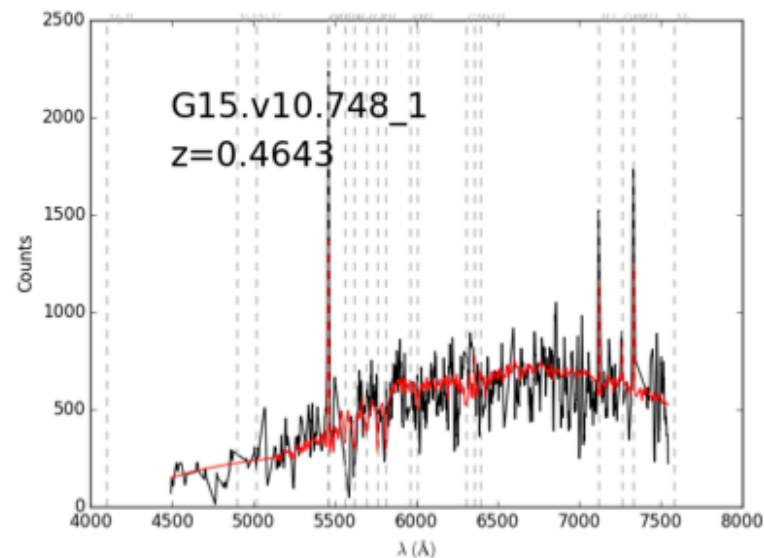
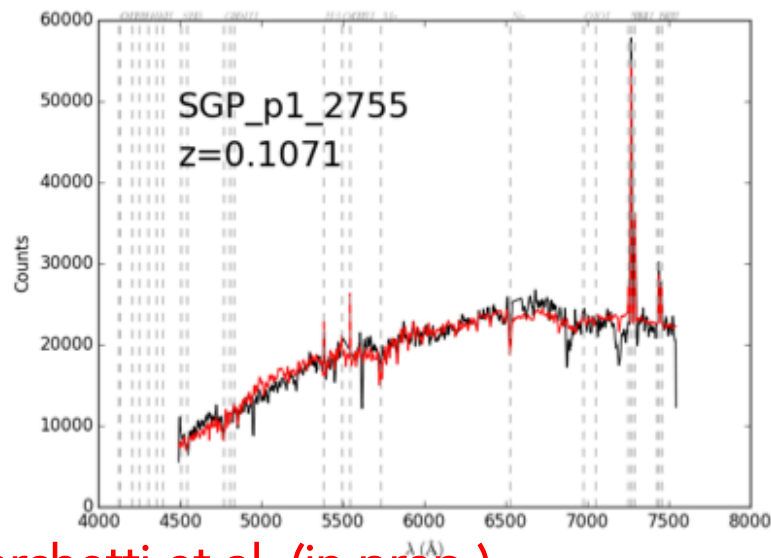
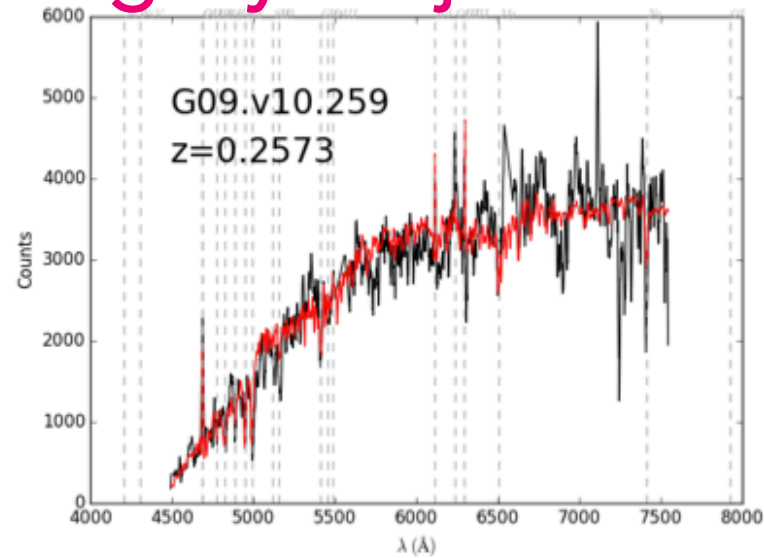
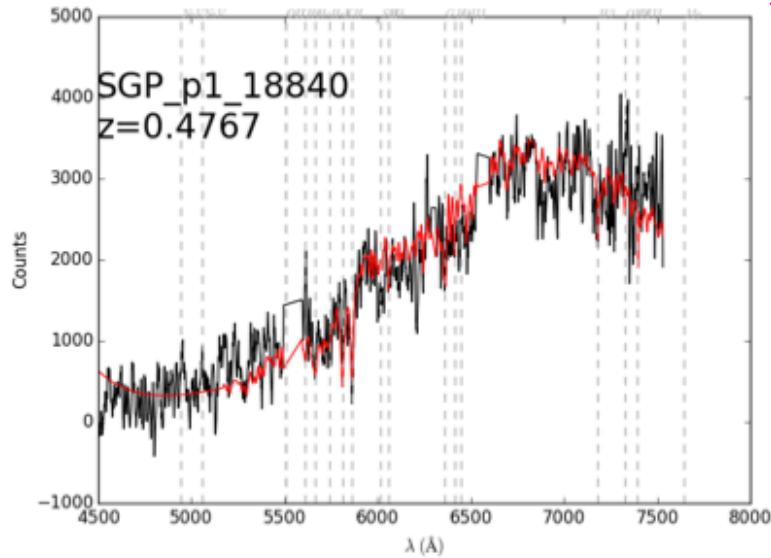
Multi- λ follow-ups: The SALT Strong Gravitational Lensing Legacy Project



- PI : Stephen Serjeant (OU)
- Project Scientist : Lucia Marchetti (UCT/UWC)
- SALT Data Manager: Steve Crawford (SAAO)
- Data Fusion Manager : Mattia Vaccari (UWC)

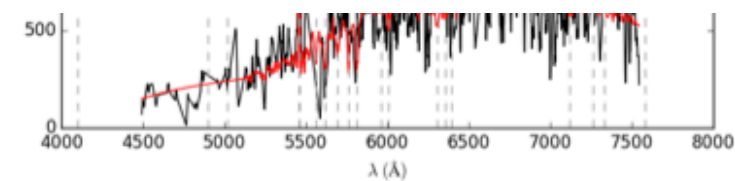
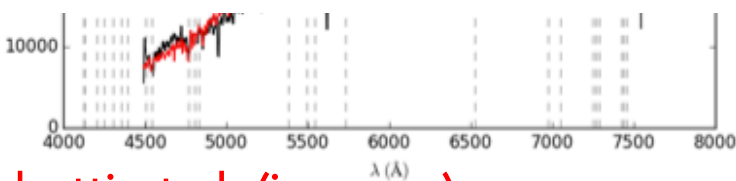
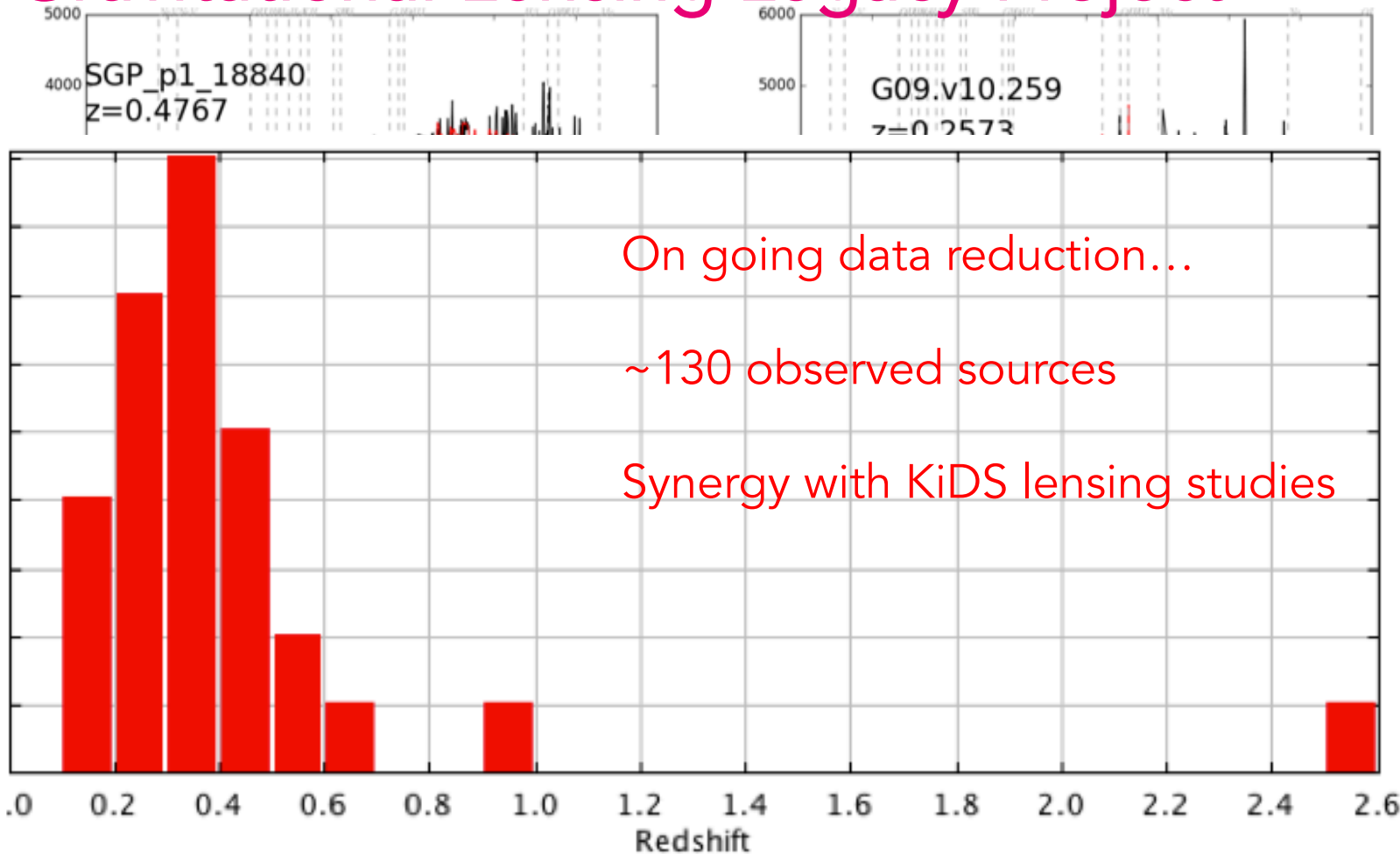
- 2015-2: 25.7hrs P1 + 9.6hrs P3
- 2016-1: 17.5hrs P2 + 48hrs P3
- 2016-2: 2.9hrs P1 + 5.5hrs P2 + 92.2hrs P3
- 2017-1: 9.1hrs P1 + 12.64hrs P2 + 34.38hrs P3
- 2017-2: 20.21hrs P2 + 18.8 hrs P3
- 2018-1 (current)

Multi- λ follow-ups: The SALT Strong Gravitational Lensing Legacy Project



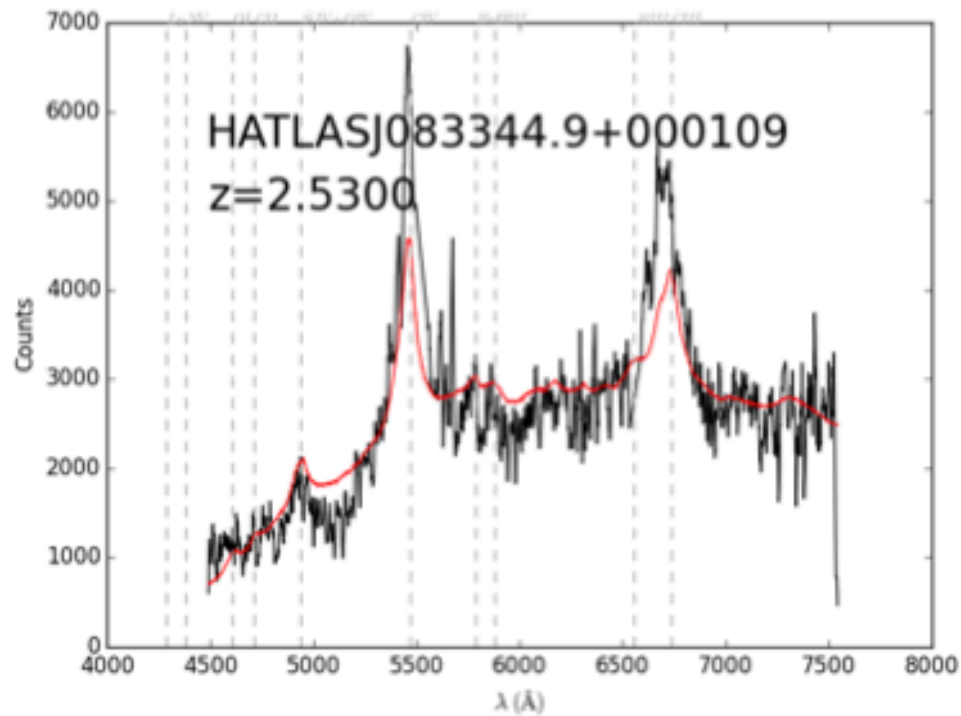
Marchetti et al. (in prep.)

Multi- λ follow-ups: The SALT Strong Gravitational Lensing Legacy Project

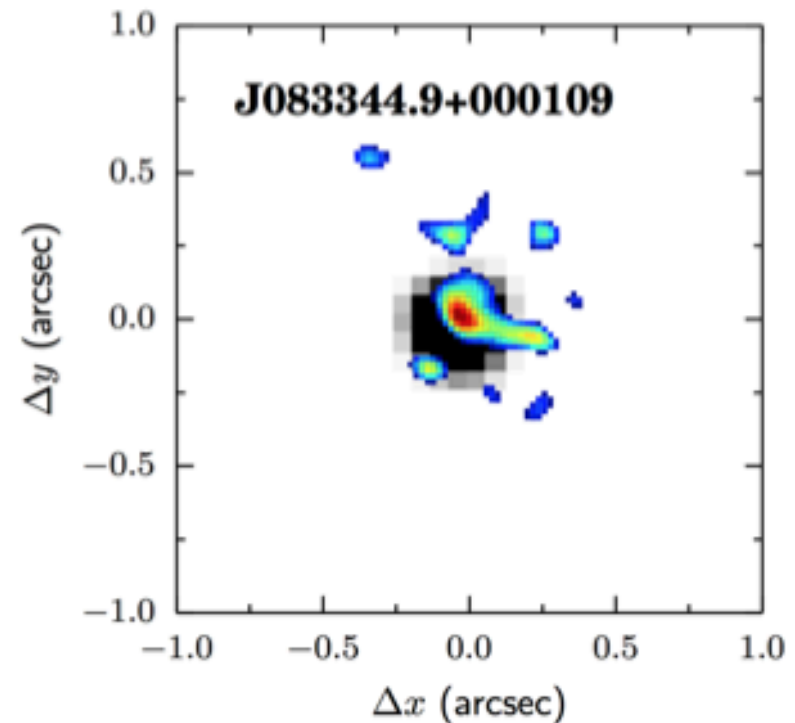


Marchetti et al. (in prep.)

Multi- λ follow-ups: The SALT Strong Gravitational Lensing Legacy Project

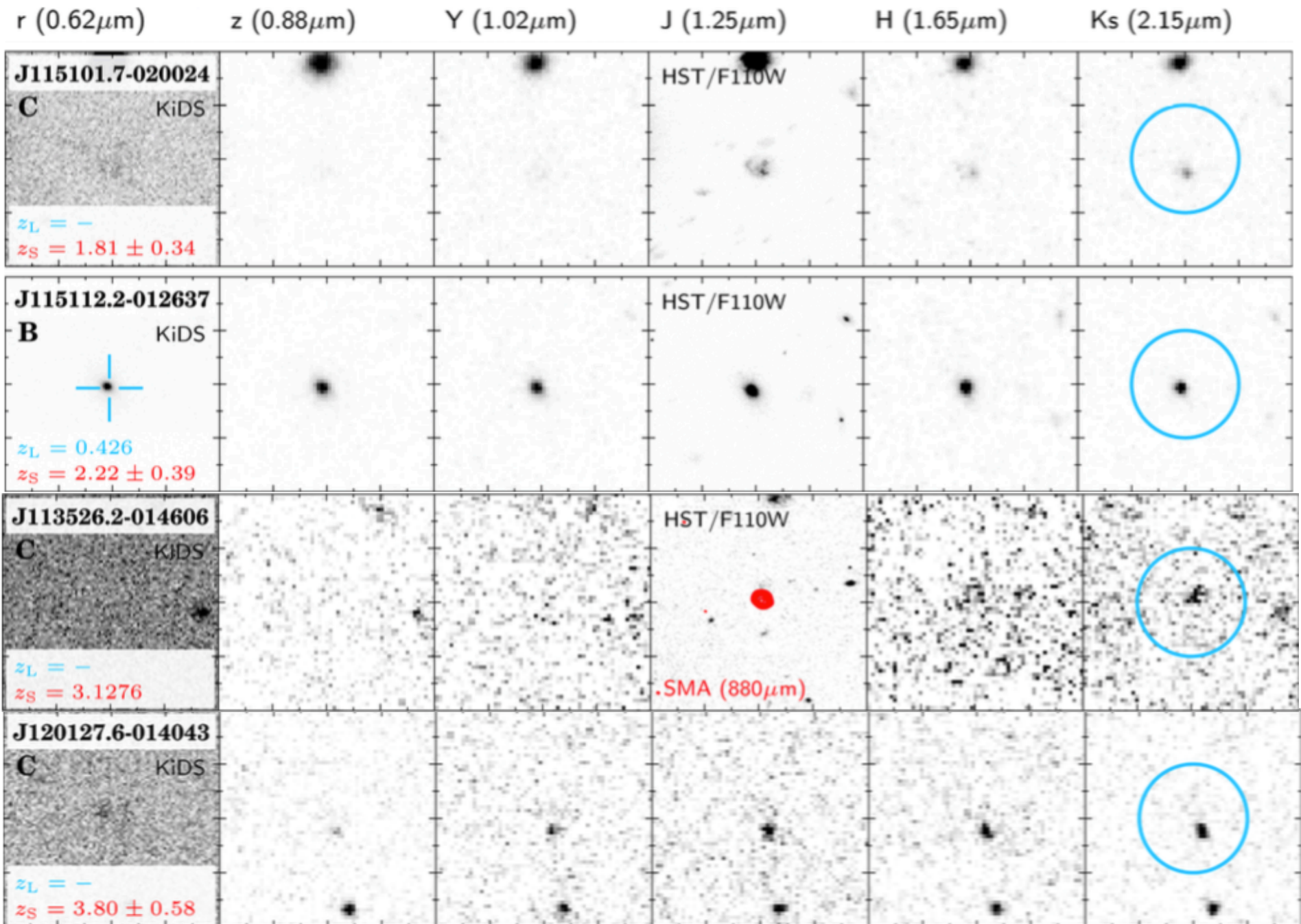


SALT



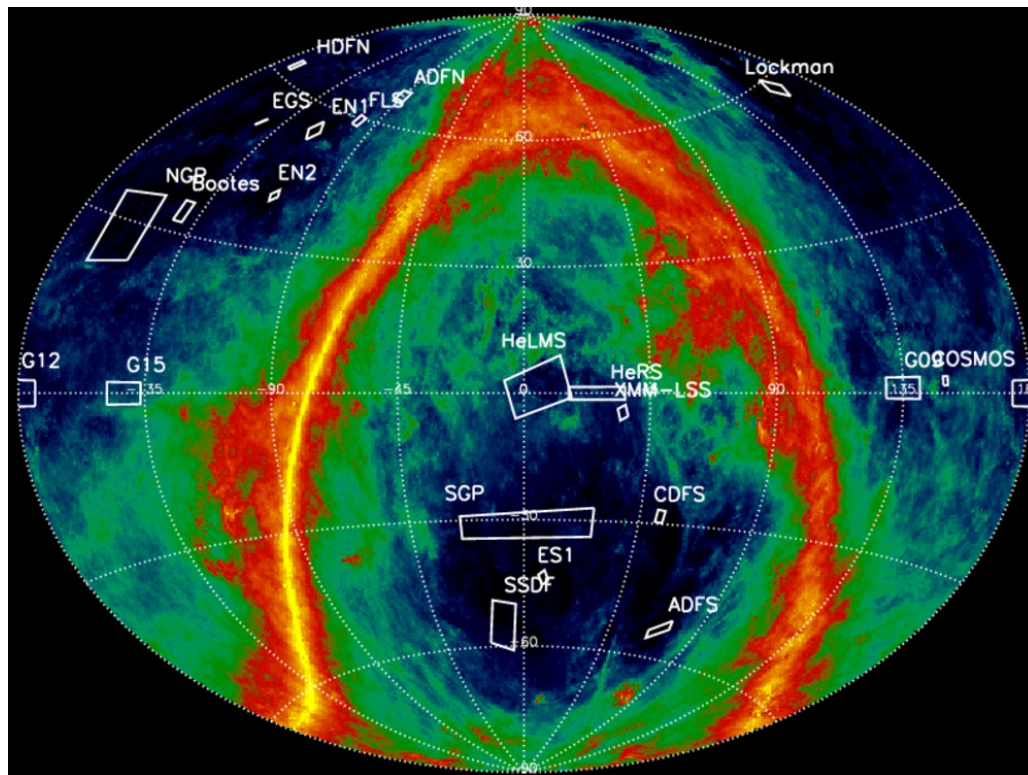
HST + ALMA

Multi- λ follow-ups (Negrello et al. 2017)



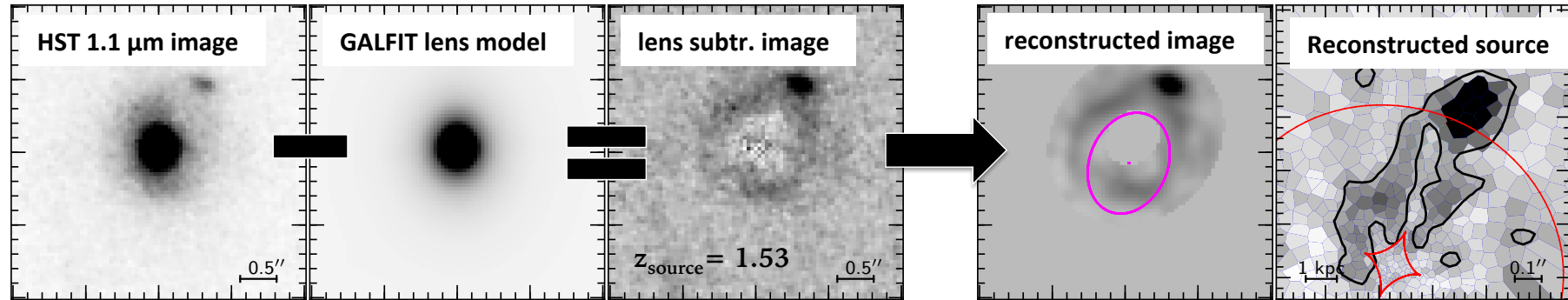
Multi- λ follow-ups: HST

- Cycle-18 – 5 HATLAS SDP sources (PI: Negrello)
- Cycle-19 – 200 F110W snapshots (PI: Negrello) – ~85% completed
Negrello et al. in prep Enia et al. in prep
[Etc..]
- Cycle-25 – 200 F110W Snapshots (PI: Marchetti) – Observing

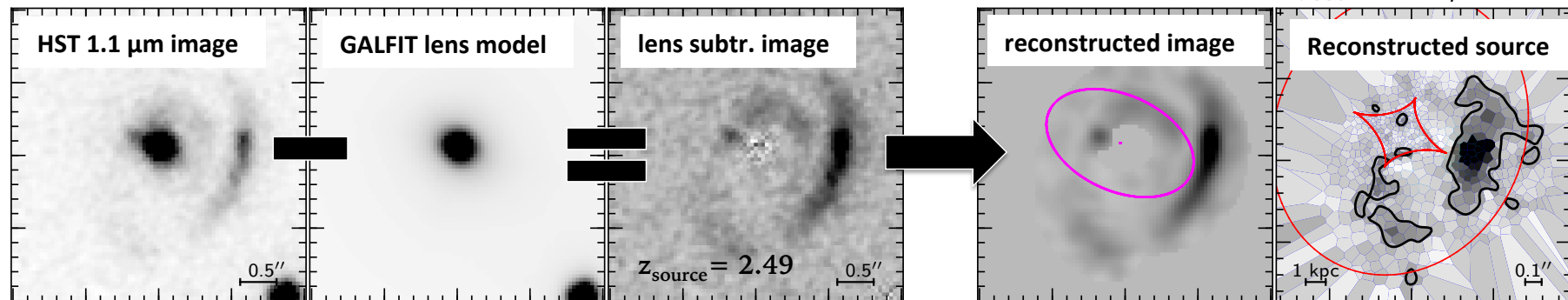


Multi- λ follow-ups & modeling

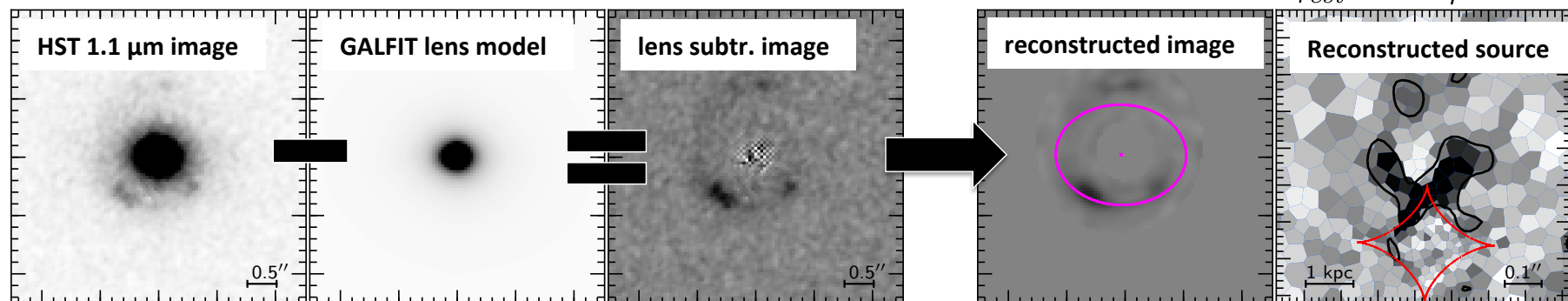
HATLASJ125759.5+224558



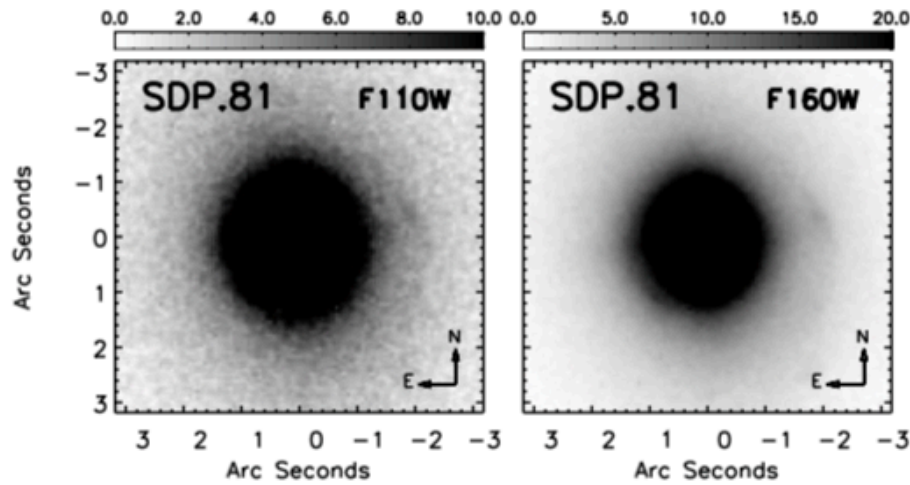
HATLASJ133846.5+255054



HATLASJ133008.5+245900

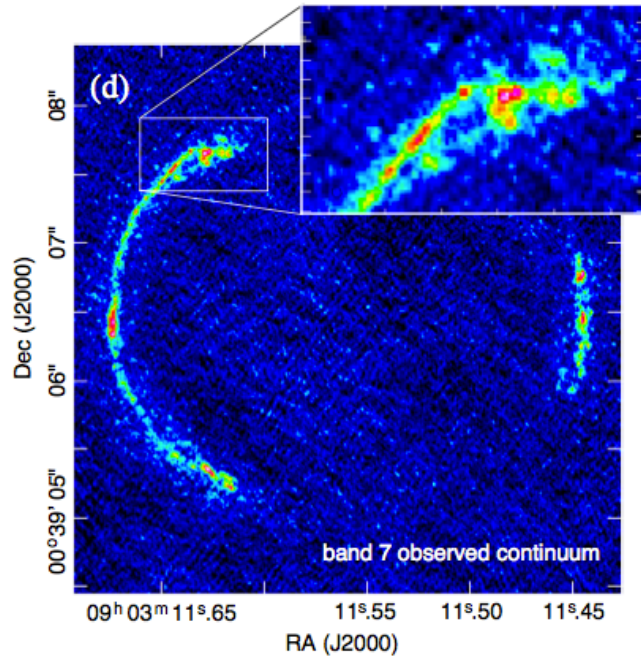


Multi- λ follow-ups: HST + ALMA

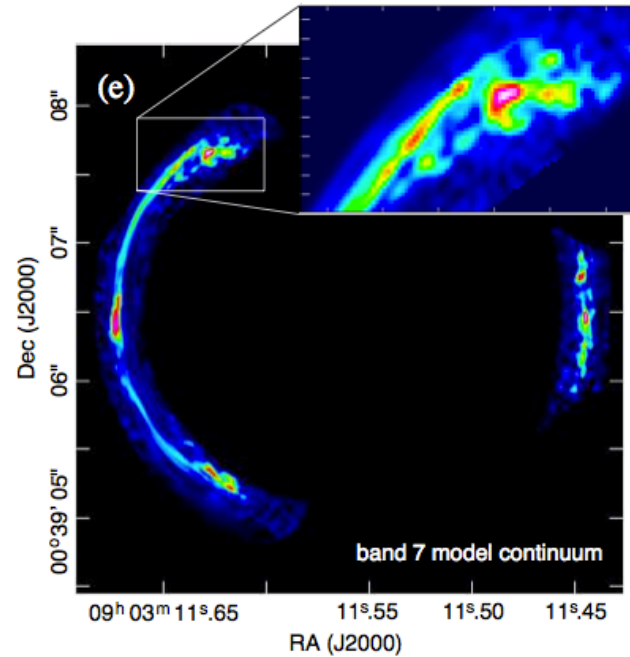


Negrello et al. 2013

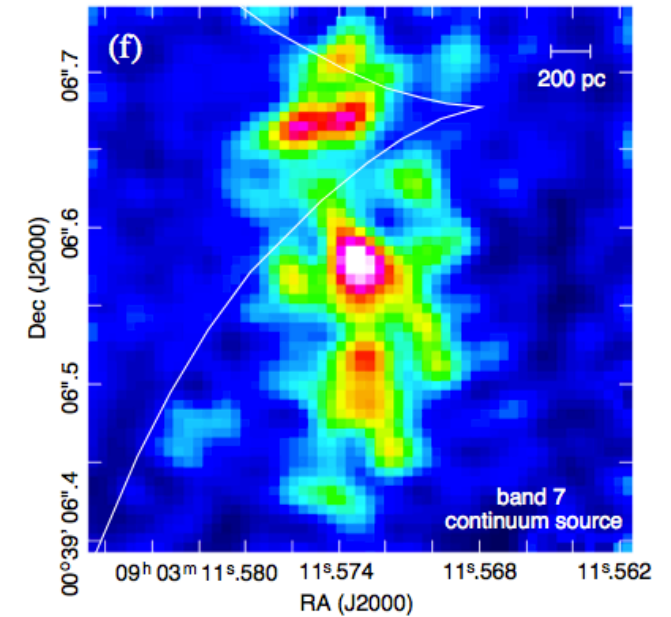
Observations (Band 7)



Best-fit model

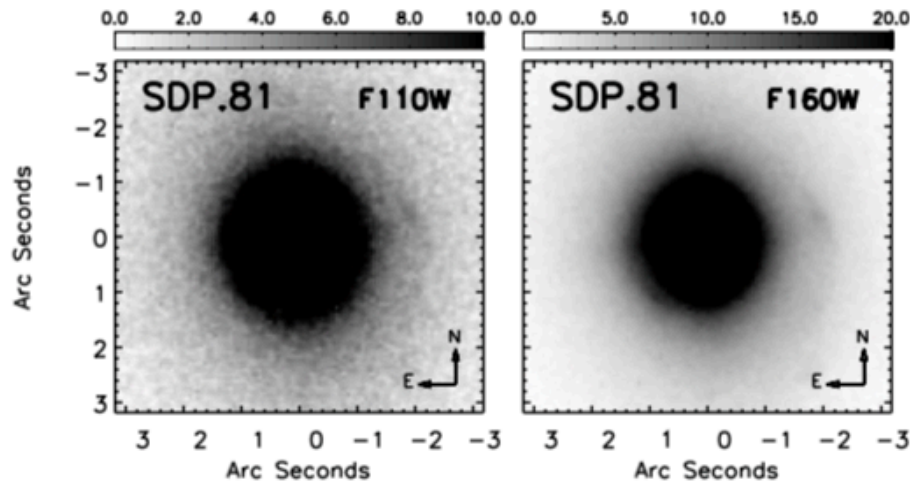


Reconstructed galaxy



Dye et al. 2015

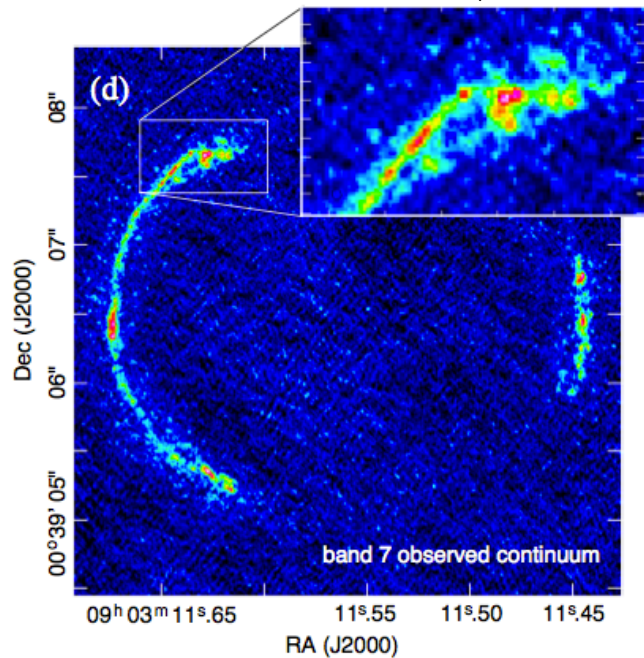
Multi- λ follow-ups: HST + ALMA



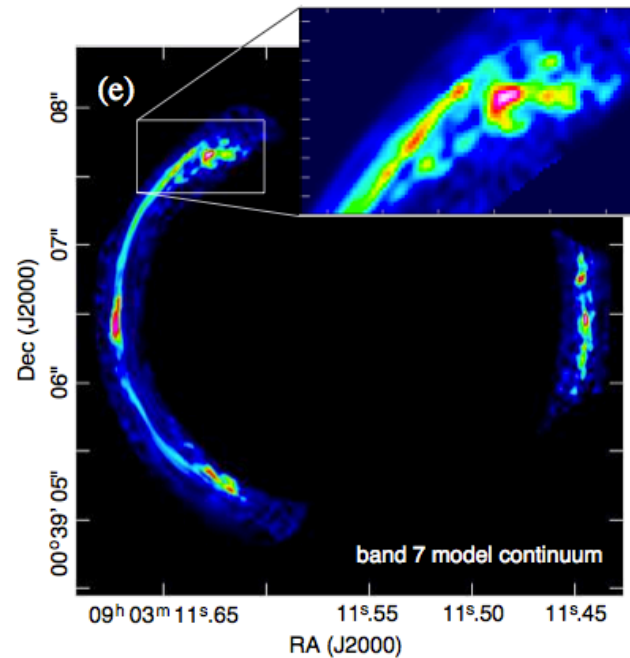
Negrello et al. 2013

Magnification due to lensing: $\sim 6-30$
 \Rightarrow resolution of $180\text{pc}/\sqrt{(6-30)} = 30-70 \text{ pc} !!$

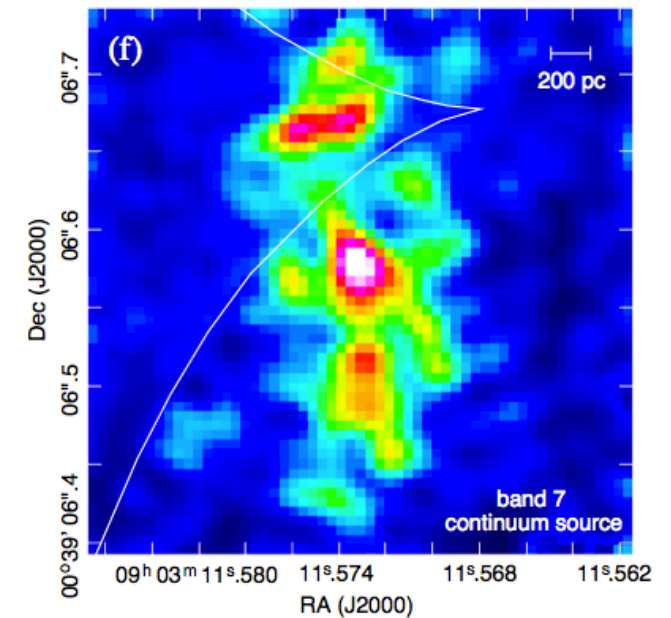
Observations (Band 7)



Best-fit model



Reconstructed galaxy

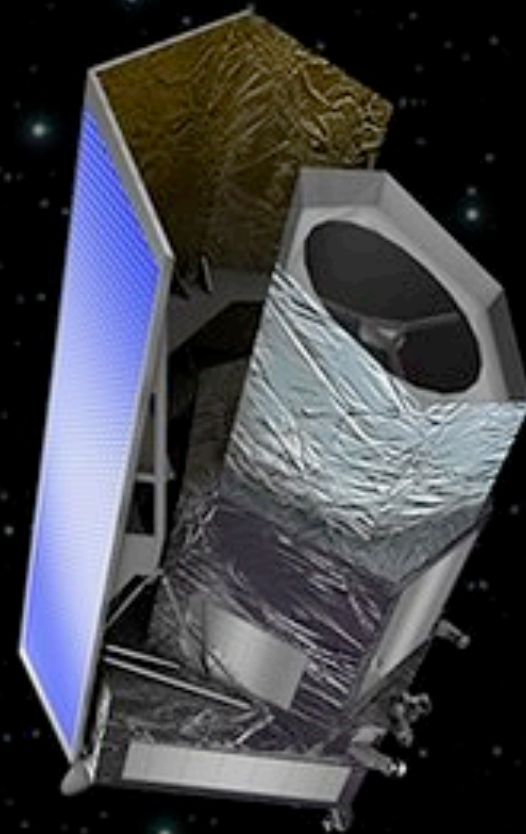


Dye et al. 2015

A golden future: Euclid



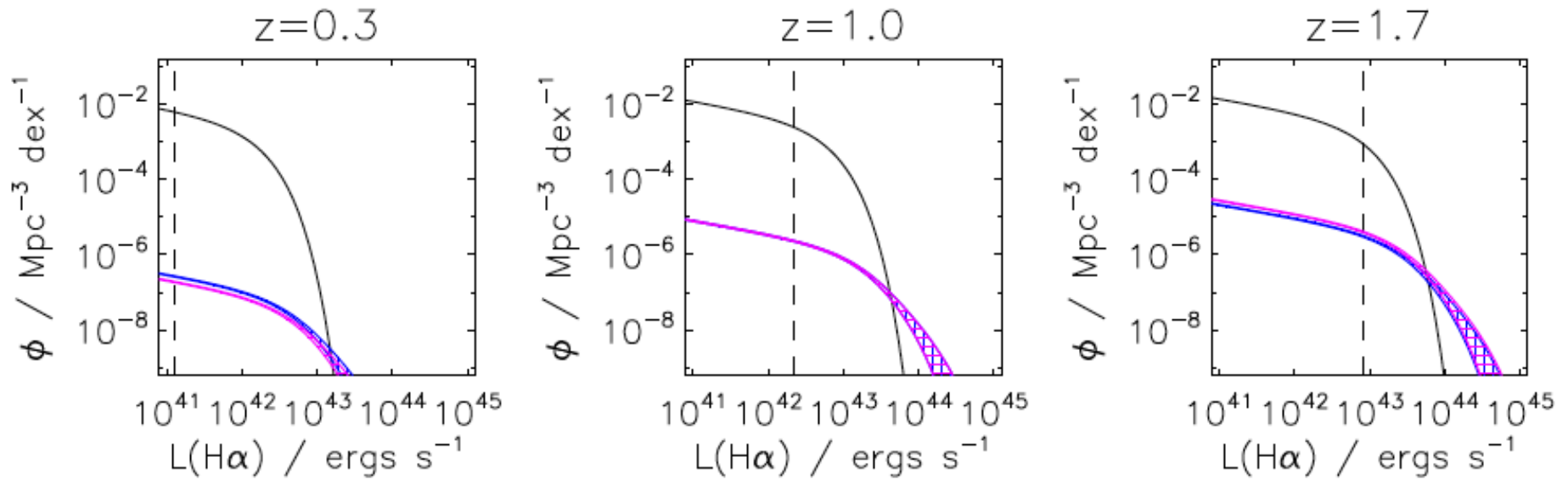
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Euclid lenses – H α LF



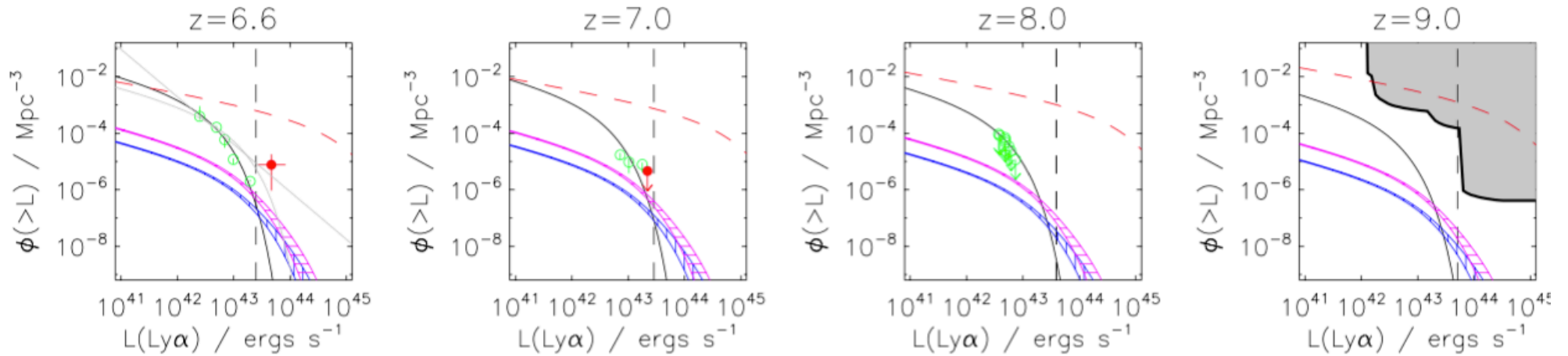
~1000 H α lenses (97% pure), and ~10⁵ arcs (needs arcfinders)



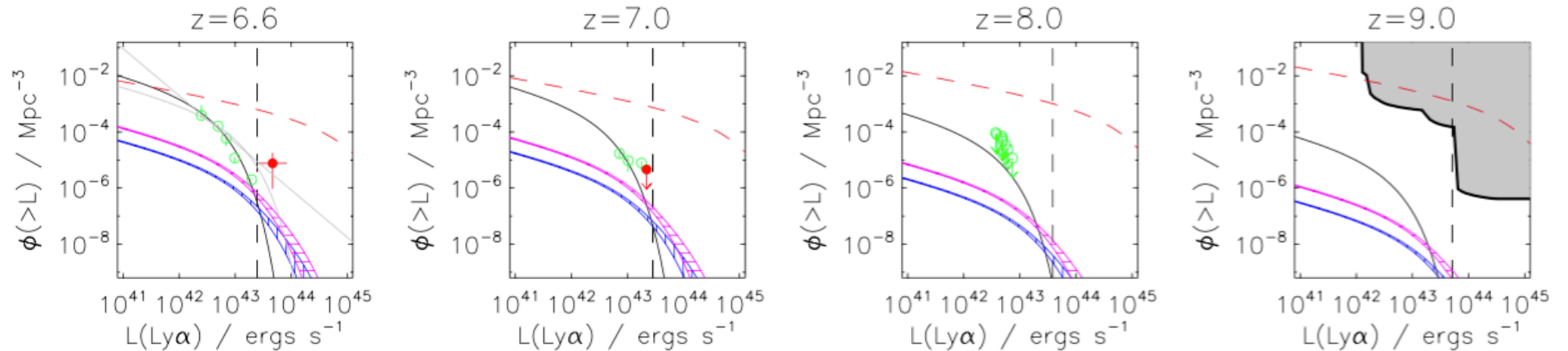
Euclid lenses – Lyman α LF

Marchetti, Serjeant, Vaccari 2017

Using LF from Ouchi et al. 2010



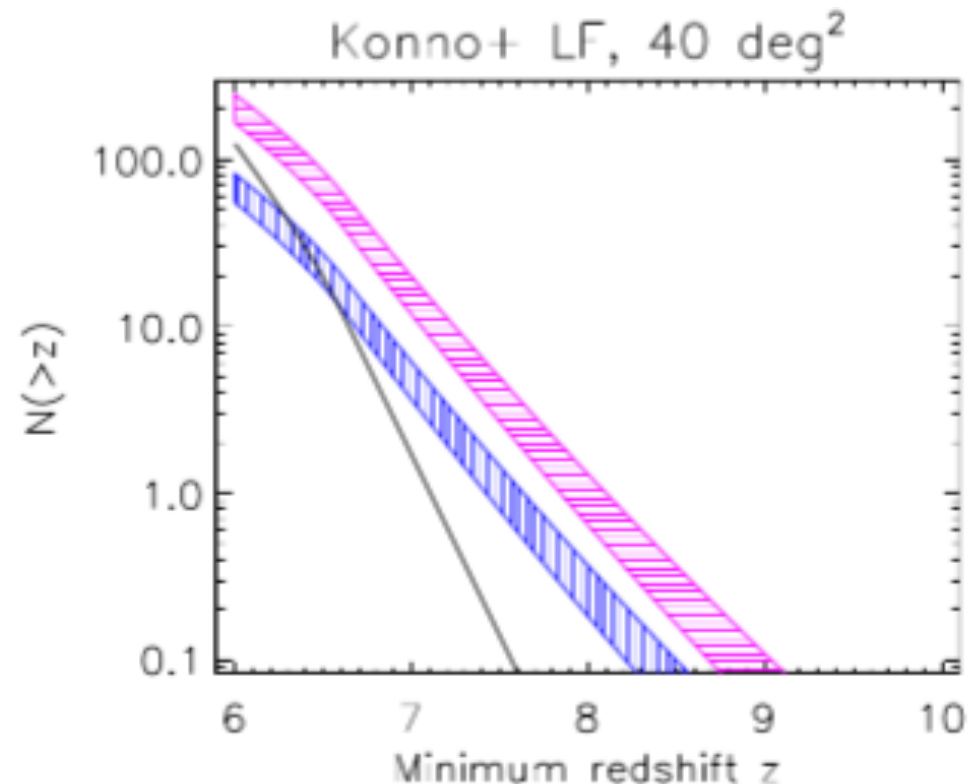
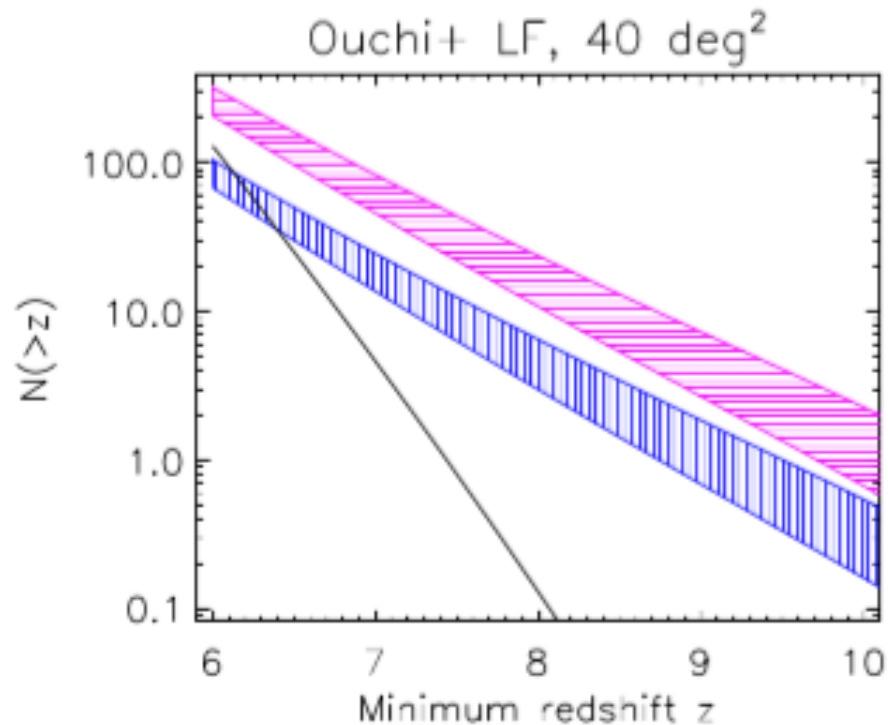
Using LF from Konno et al. 2014



Euclid lenses – Lyman α LF



Strong lensing roughly doubles the numbers of $z > 6$ galaxies with Euclid-Deep surveys!

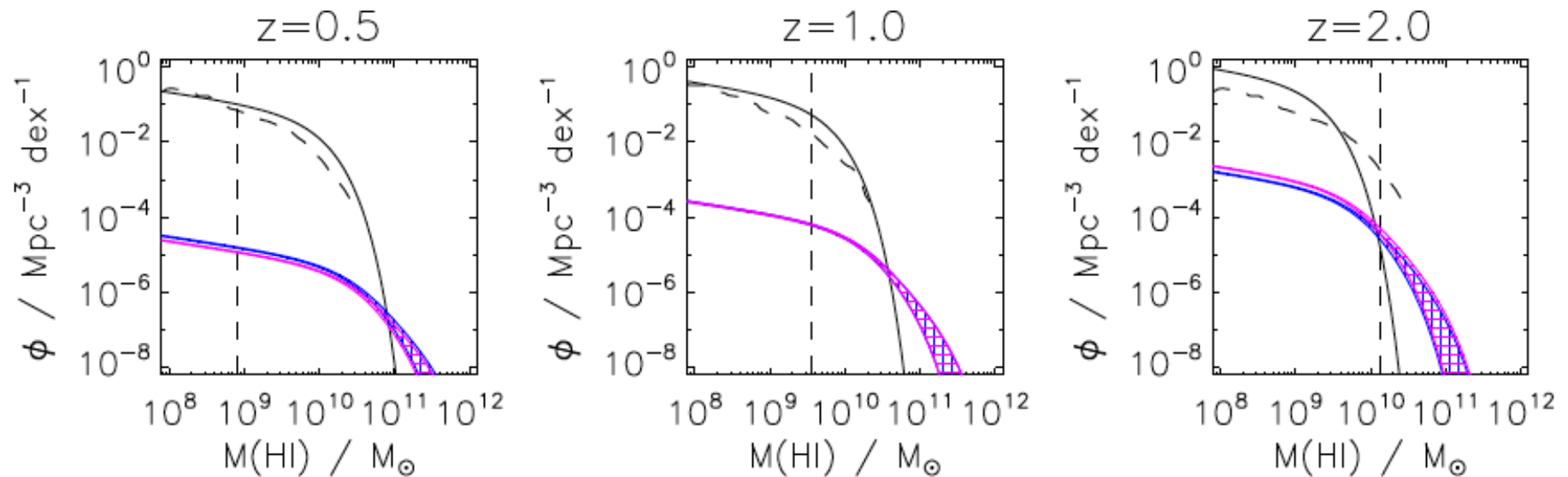


~ 0.85 and ~ 1.82 deg⁻² lensed Ly α emitters at $z > 6.6$

SKA lenses – HI Mass function



$\sim 10^4$ to $\sim 10^5$ strong lenses in SKA-2 dark energy survey!

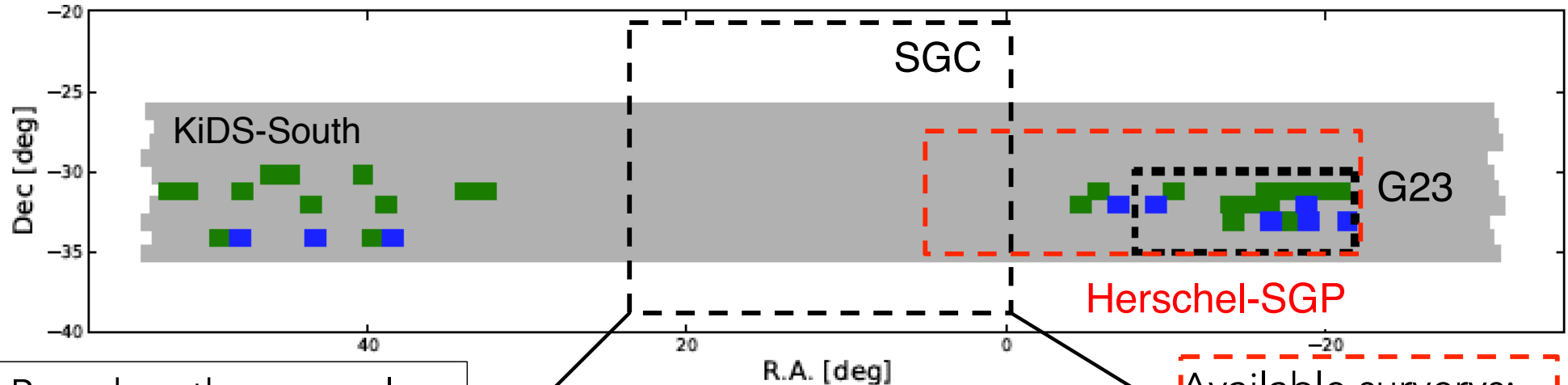


Summary

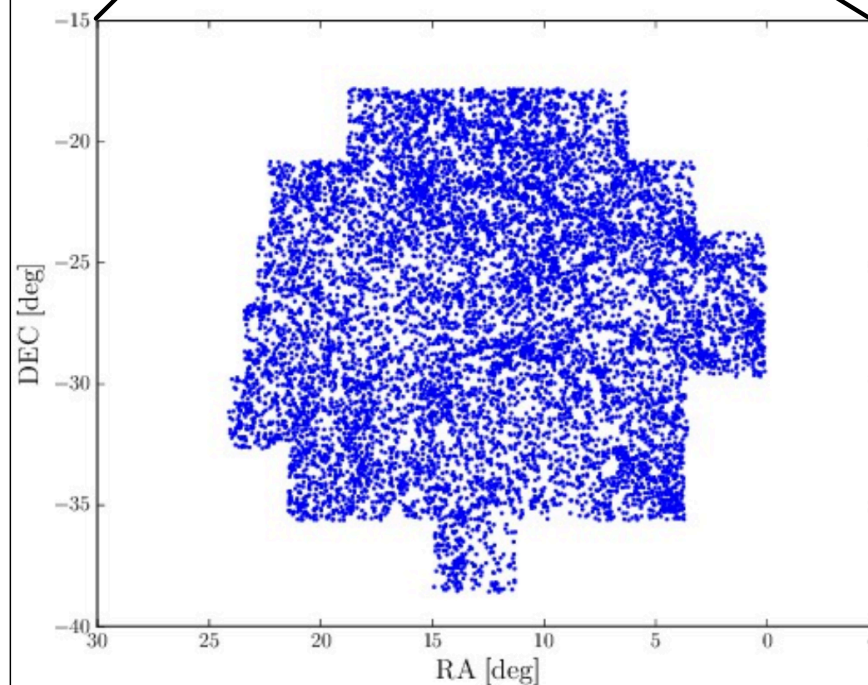


- Gravitational lensing is a powerful tool for cosmology
- Herschel has opened the door to a new strong lensing selection criterion – mostly complementary to Optical/ KiDS selection
- Redshifts and a multi-wavelengths observations are crucial to identify and characterise the lensing phenomena
- Euclid, SKA... will provide many more lensing systems enabling to put stronger constraints on cosmological parameters

EXTRAS: NEW WISE G23 & South Galactic Cap Project



Based on the example of G12 (Jarrett et al. 2017) We aim to characterise the WISE & Multi- λ properties of these sources to prepare the ground for Radio surveys. e.g. **KIDS** is crucial for **photo-z** ...



Available surveys:

- WISE
- GALEX (UV)
- **VST (ugriz)**
- DES (grizy)
- **VISTA (ZYJHK)**
- IRAC & MIPS
- PACS & SPIRE
- GMRT/JVLA/
- LOFAR/WSRT
- ASKAP/MeerKAT

Lucia Marchetti

SARChi research fellow

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THANKS!