

The Wavelength-shifting



Optical Module (WOM)

Introduction

- One of the most important properties for Neutrino-detectors: signal-tonoise ratio.
- Bigger PMT → bigger photosensitive area BUT also bigger dark noise → SNR mostly unchanged
- The WOM employs a passive, wavelength-shifter (WLS) coated surface, total internal reflection and small PMTs \rightarrow large effective area with low dark noise



Testing setup



Task:

Characterize WLS coating and ensure uniformity of lightguiding efficiency for all tubes.

Solution:

• Orbital stage on a linear stage to guide an optical fiber along and around the tube. • Fiber is connected monochromator and a Xe-lamp. • Tube efficiency can be measured for all wavelengths and all illumination positions.

Schematic drawing of the WOM

- UV-transparent housing (quartz) + transparent tube coated with WLS. • UV light is absorbed by the WLS and re-emitted isotropically.
- 74.5% of the re-emitted photons are trapped by total internal reflection.
- Light is detected by small PMTs at tube ends
- Optionally: concentrated light on even smaller PMTs with adiabatic light guides (ALG)
- Ideal for Cherenkov-detectors due to UV-sensitivity

0.30 Emission Efficiency 0.25 PMT QE (%) **5** 0.20 [A.U.] ^{0.4} 0.1 ₩ 0.10 0.2 0.05 500 450 300 Wavelength [nm] Absorption and emission spectrum of the WLS and quantum-efficiency of the PMTs

Definition:

Efficiency = (#detected / #injected) photons, depends on WLS efficiency, trapping ratio, light guiding efficiency for the emitted light.



A photo of the Mainz WOM tester Allows semi-automatic scans of coating uniformity. Second system using 3 small PMTs is at DESY Zeuthen



spot

Sample result from a 2D scan. Relative PMT signal variation relative to signal at 0° over angular position.Incident wavelength λ =375 nm. Variation is smaller than ±5%

Building the WOM







Performance



WOM assembly: 1 = PMT, 2 = ALG, Adiabatic light guide (ALG) 3 = springs, 4 and 5 = feedthroughs

depends on speed



Performance vs. Coating speed on quartztubes. High reproducability and high total efficiency (up to 24.3% per side)

> WOM prototype inside quartz vessel

A dark noise challenge



Dark noise measured inside a Faraday cage in a climate chamber (-50° to +20°C)

Dark noise of PMT: 30-50 Hz ...with ALG and WLS tube added: 70 Hz ...with quartz pressure vessel: 630 Hz!

BUT: this quartz contains ~150 ppm ²³⁸U

 \rightarrow cleaner quartz has been found (from Heraeus and Raesch)

 \rightarrow produce vessel from clean, low activity quartz (HSQ300, SUP310 or RQ200)

- \rightarrow expected new vessel contribution <10 Hz
- expected total dark noise <80 Hz



Wavelength-scan, one-side-efficiency vs. injected wavelength, illumination 150 away from tube end. Above λ =400 nm the UV-light is not absorbed.



Timing: detection time of single-pe events using a picosecond UV-pulser, depending on illumination point distance from tube end.

Fit results: WLS-decay time = 1,35 ns Photon-attenuation length: λ_{eff} = 320 cm

Distance-scan: total efficiency vs. Distance of illumination point from the tube end. Comparison of the two test setups. Average efficiency 38-43% over 90 cm length.



Measured light propagation fitted with a simplified 2D model (flattened tube mantle with periodic boundary conditions)

SHiP decay vessel

+background tagger

using WOMs

IceCube

Summary

•41 ±1.7% avg. efficiency

compared to PMT alone

Improve SNR by factor >10

• even higher with adiabatic

Prototype built

light guide

temperature



Bedrock



- The WOM is ideal for experiments that aim to detect UV light with a high SNR
- Initially developed for IceCube Gen2 >2.5 km
 - Focus: low noise and high effective area
 - Main cost driver: quartz pressure vessel for 500 bar
- \rightarrow only necessary in ice! DeepCore

• A smaller, much cheaper WOM has been developed for SHiP, others may follow. \rightarrow Hyper-K? & Pingu



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