Physics of v-A Interactions

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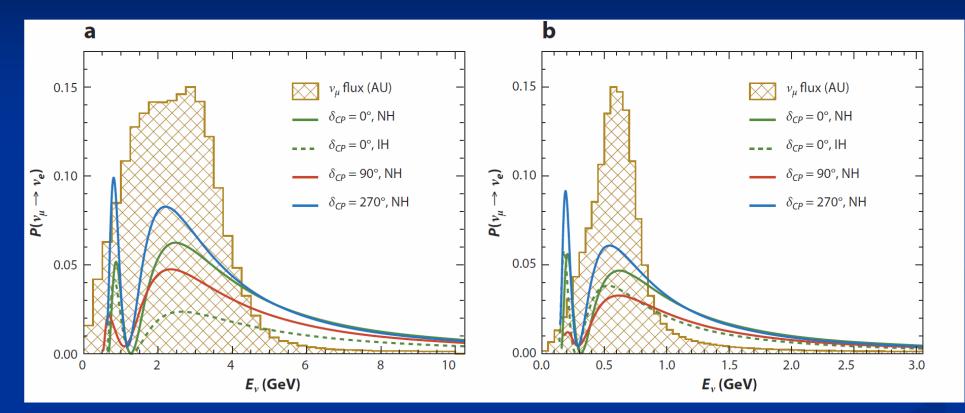
What is (not) measured in a LBL exp?

LBL experiments measure only flux-averaged cross sections
 The neutrino energy is not measured
 Oscillation Patterns as function of neutrino energy must be reconstructed
 needs nuclear theory and modeling
 Experiments require few % accuracy





Oscillation Signals as F(E_v)



From: Diwan et al, Ann. Rev. Nucl. Part. Sci 66 (2016)

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DUNE, 1300 kmHyperK (T2K) 295 kmEnergies have to be known within 100 MeV (DUNE) or 50 MeV (T2K)Ratios of event rates to about 10%Neutrino 2018

Problem: Neutrino Energy

- The incoming neutrino energy on the abscissa of all such plots is not known, but must be reconstructed; very different from Nuclear Physics and High Energy Physics where the beam energy is accurately known.
- The reconstruction has to start from an only partially observed final state (detector limitations!) and proceeds from there ,backwards' to the initial state.





Generators

- Generators are needed for this ,backwards calculation'
- The accuracy of the energy reconstruction and thus the precision of any neutrino mixing parameters depends crucially on the precision of these generators
- Generators must be an integral part of any experiment
- Generators must be able to handle:
 - the extended target size complications
 - the primary neutrino-nucleus interaction
 - the final state interactions





Neutrino Cross Sections: Nucleus

- All targets in long-baseline experiments are nuclei: C, O, Ar, Fe
- Cross sections on the nucleus:
 - QE + final state interactions (fsi)
 - Resonance-Pion Production + fsi
 - Deep Inelastic Scattering \rightarrow Pions + fsi
- Additional cross section on the nucleus:
 - Many-body effects, e.g., 2p-2h excitations
 - Coherent neutrino scattering and coh. pion production



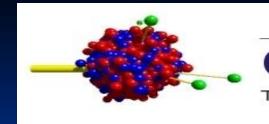


Generators for energy reconstruction

- Present day's neutrino generators combine different physics processes and models into one patchwork, creating artificial degrees of freedom
- Present day's neutrino generators often use outdated physics (RS)
- Present day's generators do not bind the nucleus
- To make up for these deficiencies, present day's neutrino generators rely on tuning, i.e. fitting to data
- Time to build new, consistent generators, based on present day's nuclear physics, both for initial interaction and the final state interactions (transport)







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Gibuu

The Giessen Boltzmann-Uehling-Uhlenbeck Project

GiBUU : Quantum-Kinetic Theory and Event Generator based on a BM solution of Kadanoff-Baym equations GiBUU propagates phase-space distributions, not particles Physics content and details of implementation in: Buss et al, Phys. Rept. 512 (2012) 1-124 Code from gibuu.hepforge.org, new version GiBUU 2017 Details in Gallmeister et al, Phys.Rev. C94 (2016) no.3, 035502





Theoretical Basis of GiBUU

- Kadanoff-Baym equation (1960s)
 - full equation not (yet) feasible for real world problems
 - Boltzmann-Uehling-Uhlenbeck (BUU) models: GiBUU
 - Boltzmann equation as gradient expansion of Kadanoff-Baym equations, in Botermans-Malfliet representation (1990s)
 - Cascade models

Simplicity

(typical event generators, GENIE, NEUT, NuWro, ...)

 no mean-fields, primary interactions and FSI not consistent, reweighting of different interaction types,



Correctness

INT 12/2013



• **GIBUU** describes: (within the same unified theory and code)

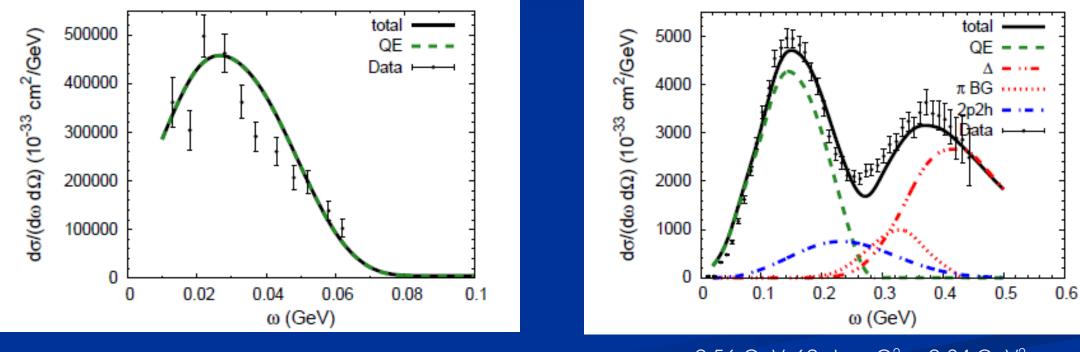
- heavy ion reactions, particle production and flow
- pion and proton induced reactions on nuclei
- photon and electron induced reactions on nuclei
- neutrino induced reactions on nuclei using the same physics input! And the same code!
 NO TUNING!





Inclusive QE Electron Scattering

a necessary check for any generator development



0.24 GeV, 36 deg, $Q^2 = 0.02 \text{ GeV}^2$

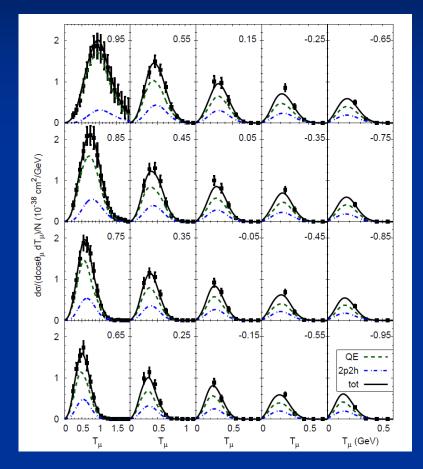
Target: C 0.56 GeV, 60 deg, $Q^2 = 0.24$ GeV²



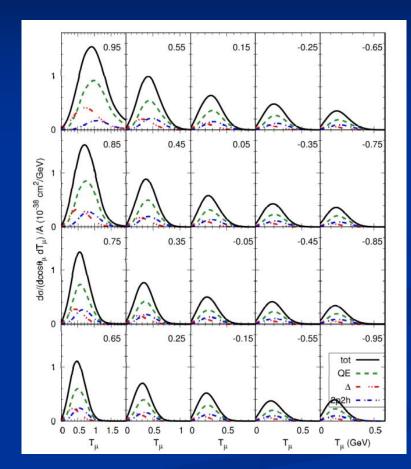
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Inclusive Lepton Kinematics

Gallmeister et al. Phys.Rev. C94 (2016) no.3, 035502



MiniBooNE C12 (QE + 2p2h)



 $2p2h \sim \Delta$

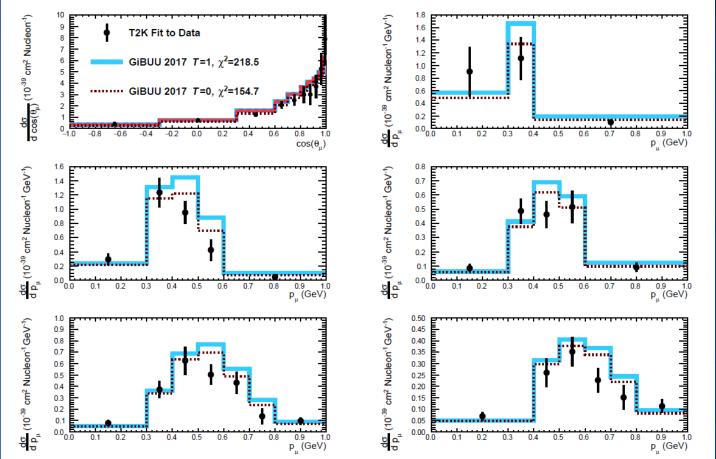
MicroBooNE Ar40 (fully inclusive)





T2K Inclusive Cross Section





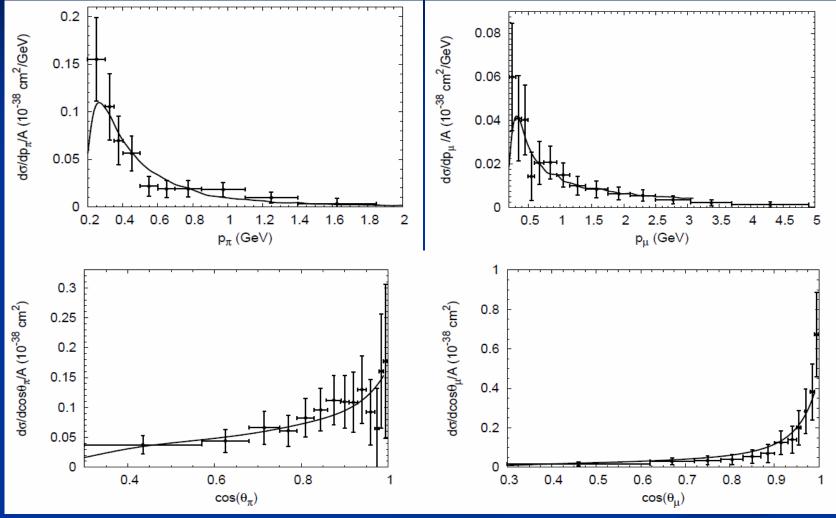
Dolan et al, arXiv:1804.09488

Poster: S. Dolan et al Wednesday, #104





T2K ND280 Pi+



Mosel, Gallmeister, Phys.Rev. C96 (2017) no.1, 015503

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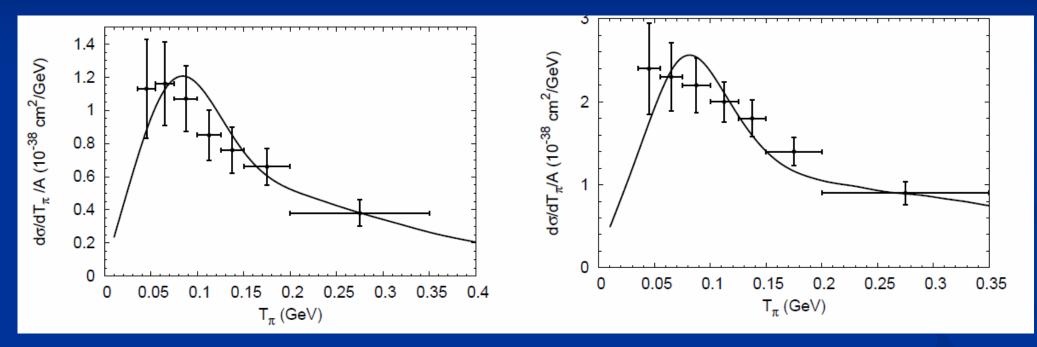




H₂O target

MINERvA Pions

CC charged pions on CH target



 $W < 1.4 \, GeV$

W < 1.8 GeV, multiple pions





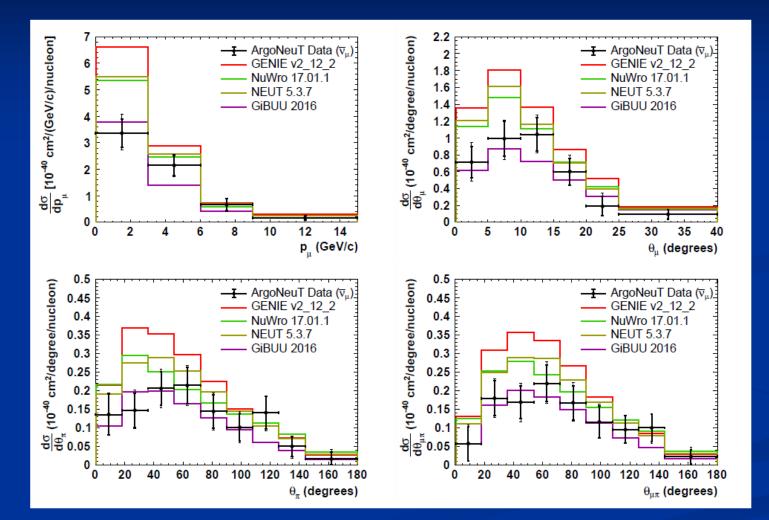
Challenge: ⁴⁰Ar

- ⁴⁰Ar not isospin symmetric, N > Z.
- QE and pion production are one-body processes
- and easy to handle, but what about 2p2h Isospindependence?
- Relation of 2p2h to electron scattering process?
 CHECK: how well does GiBUU do on ArgoNeut data? Now: pion production, fully inclusive to come





Pion Production on LAr



Neutrino 2018

ArgoNeut arXiv:1804.10294

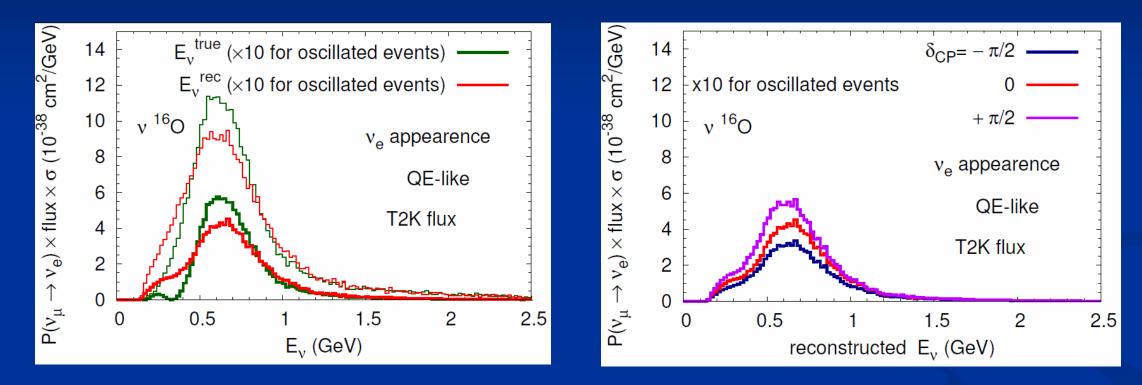
Antineutrinos

Excellent agreement of GiBUU with Ar data NO Tune



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Oscillation signal in T2K δ_{CP} sensitivity of appearance exps



O. Lalakulich et al, Phys.Rev. C86 (2012) 054606

Reconstruction error as large as δ_{CP} dependence

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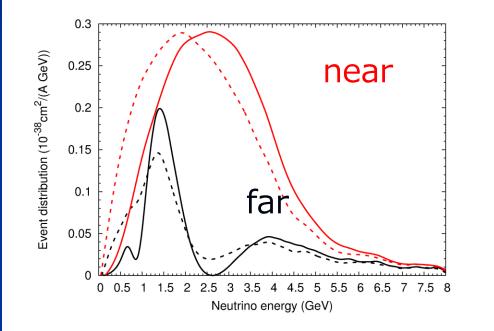


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QE Energy Reconstruction for DUNE

Muon survival in 0 pion sample



Dashed: reconstructed, solid: true energy

Reconstruction by kinematic method

Mosel et al, Phys.Rev.Lett. 112 (2014) 151802

Nearly 500 MeV difference between true and reconstructed event distributions → not a useful method

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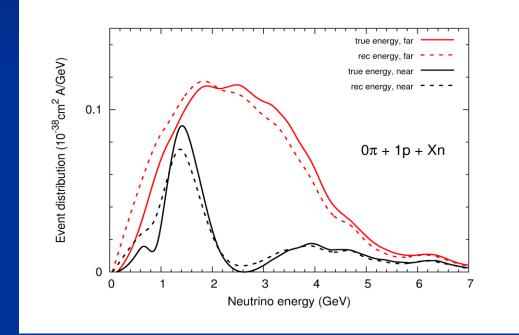


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QE Energy Reconstruction for DUNE

Muon survival in $0\pi + 1p + Xn$ sample



Dashed: reconstructed, solid: true energy

Mosel et al., Phys.Rev.Lett. 112 (2014) 151802

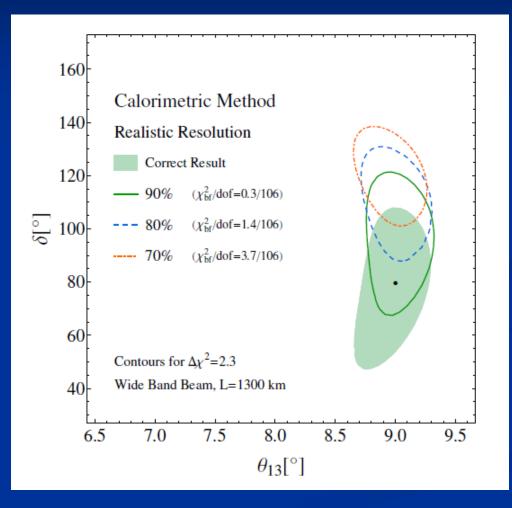
Dramatic improvement in 0 pi, 1p, Xn sample, down by only factor 3
 → Kinematic reconstruction useful also at DUNE energies

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Extraction of Oscillation Parameters



Oscill parameters in dependence % of true missing energy for DUNE

Ankowski et al, Phys.Rev. D92 (2015) 091301





Summary I

- Energy reconstruction is essential for precision determination of neutrino oscillation parameters (and nu-hadron cross sections)
- Neutrino energy must be known within about 50 (T2K) or 100 (DUNE) MeV
- Nuclear effects complicate the energy reconstruction
- Need state-of-the-art generators for reconstruction, with predictive power and no artificial degrees of freedom



Summary II

- Precision era of neutrino physics requires more sophisticated generators and a dedicated joint effort in nuclear theory and generator development
- This joint effort has to be funded as integral part of experiments

Transport Theory has to find its way into neutrino generators!





Gibuu

Essential References:

- I. Buss et al, Phys. Rept. 512 (2012) I contains both the theory and the practical implementation of transport theory
- 2. Gallmeister et al., Phys.Rev. C94 (2016), 035502 contains the latest changes in GiBUU2016
- 3. Mosel, Ann. Rev. Nucl. Part. Sci. 66 (2016) 171 short review, contains some discussion of generators
- The work reported here was done in collaboration with Kai Gallmeister and Olga Lalakulich





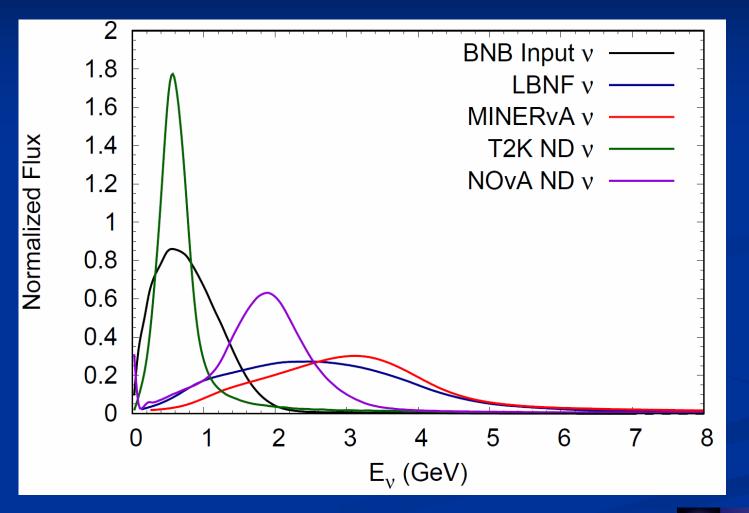




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Energy-Distributions of Neutrino Beams



Energy must be reconstructed event by event, within these distributions





Generators describe vA interactions?

- Take your favorite neutrino generator (GENIE, NEUT, ...): "a good generator does not have to be right, provided it fits the data"
- All of these generators neglect from the outset:
 - Nuclear binding
 - Final state interactions in nuclear potential
 - Same ground states for different processes
- Generators can possibly be tuned, but can they then have any predictive power? With their patchwork of recipes?







GiBUU Ingredients

- Mean field potential with local Fermigas momentum distribution, nucleons are bound (not so in generators!)
- Initial interactions are calculated by summing over interactions with all bound, Fermi-moving nucleons

2p2h interactions taken from electron scattering

- Final state interactions are handled by quantum-kinetic transport theory (off-shell transport possible)
- Calculations give the final state phase space distribution of all particles





Quantum-kinetic Transport Theory

On-shell drift term

Off-shell transport term

Collision term

$$\mathcal{D}F(x,p) - \operatorname{tr}\left\{\Gamma f, \operatorname{Re}S^{\operatorname{ret}}(x,p)\right\}_{\operatorname{PB}} = C(x,p) \ .$$

$$\mathcal{D}F(x,p) = \{p_0 - H, F\}_{\rm PB} = \frac{\partial(p_0 - H)}{\partial x} \frac{\partial F}{\partial p} - \frac{\partial(p_0 - H)}{\partial p} \frac{\partial F}{\partial x}$$

H contains mean-field potentials

Describes time-evolution of F(x,p)

$$F(x,p) = 2\pi g f(x,p) \mathcal{P}(x,p)$$

Phase space distribution

KB equations with BM offshell term

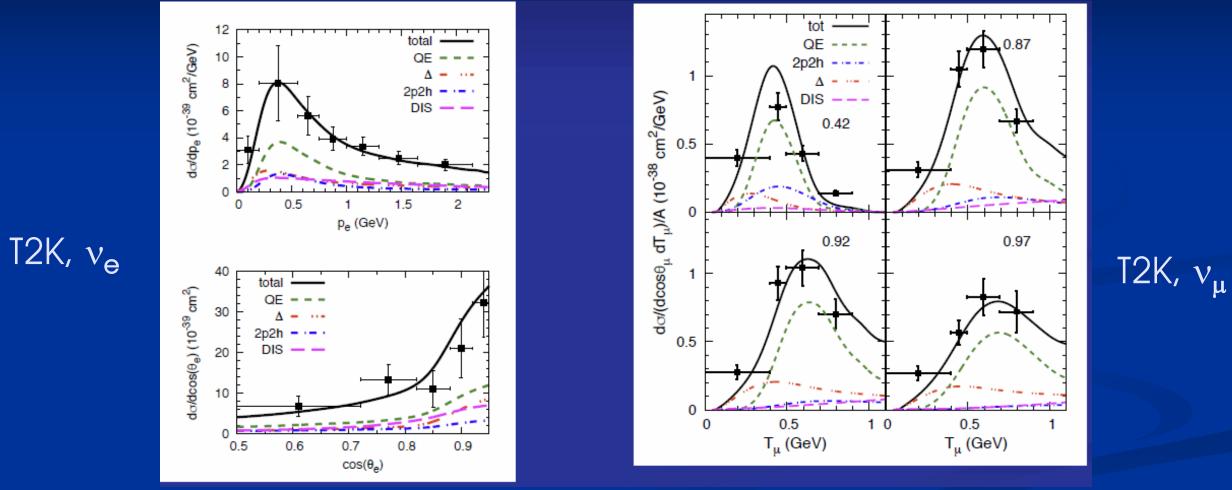
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Comparison with T2K incl. Data

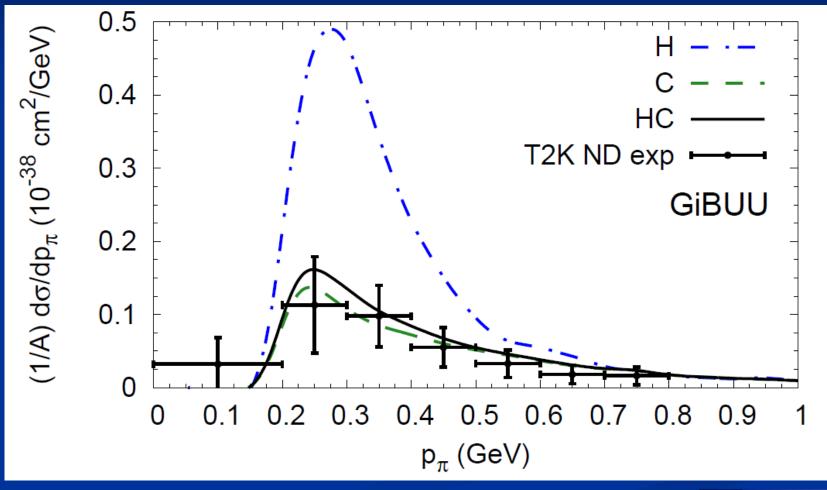


Agreement for different neutrino flavors





Neutrino-induced Reactions: Pion Production





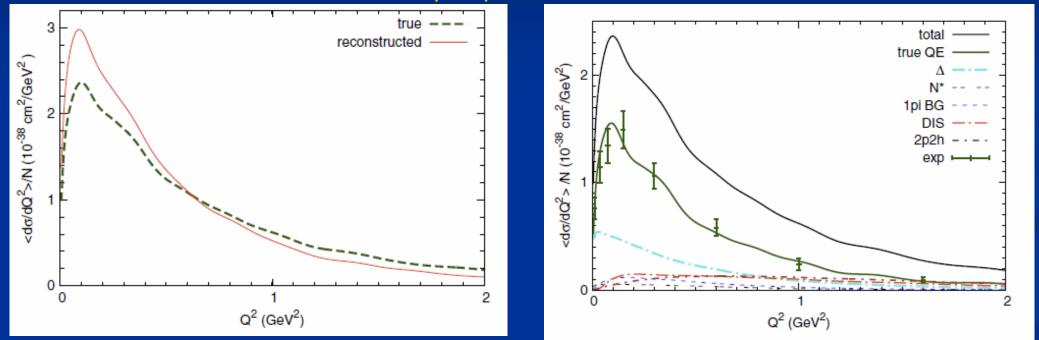


T2K

CH target

MINERvA Q² Reconstruction

Only 0-pion events



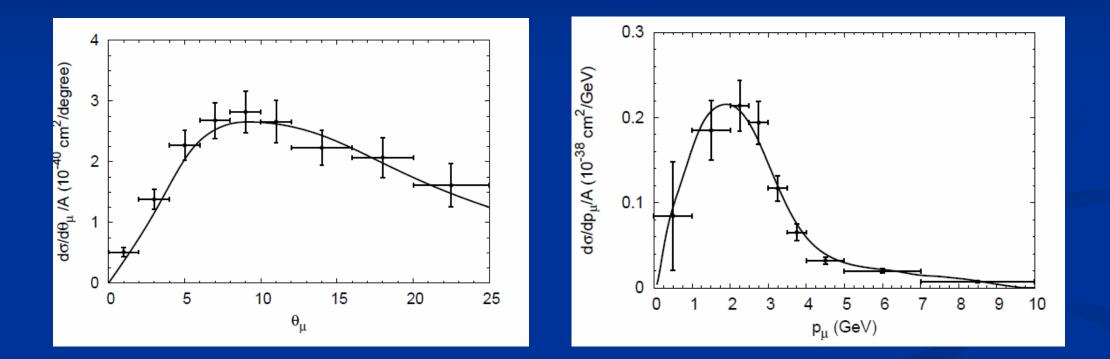
Dramatic sensitivity to reconstruction in peak area: accuracy of ,data'??

Mosel et al., PR D89 (2014) 093003





MINERvA Pions



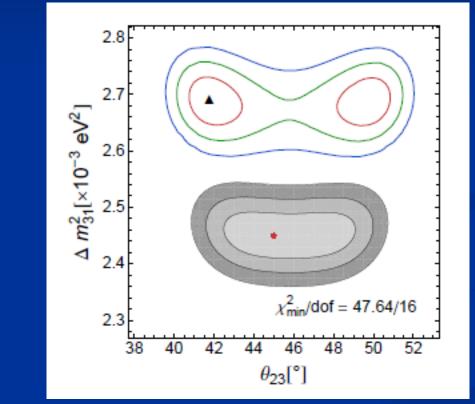
W < 1.8 GeV

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Generator Dependence of Oscillation Parameters



GIBUU-GENIE

Gibuu-Gibuu

From: P. Coloma et al, Phys.Rev. D89 (2014) 073015

Nature: GiBUU Generator: GENIE



