

Automated Kinematics Analysis of Off-Limb Coronal Bright Fronts

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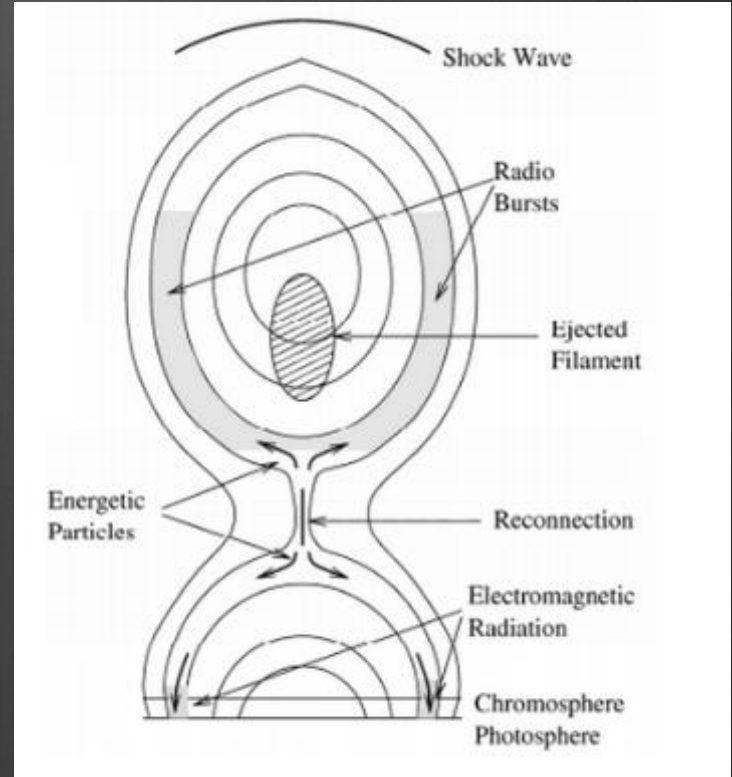
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Outline

- Introduction to Coronal Bright Fronts (CBFs)
- Automatic Processing Procedures
 - Radial Wave Kinematics
 - Tangential Wave Kinematics
- Current Kinematics Information
- Next Steps with Geometric CBF Models
- Summary

Coronal Mass Ejections

- Massive solar explosions
- Can drive shock waves
- Release broadband electromagnetic radiation, including drifting radio bursts
- Release a cloud of charged particles
- Can couple with Earth's magnetic field



From *Space Physics: An Introduction to Plasmas and Particles in the Heliosphere and Magnetospheres*, Kallenrode (p. 192)

Solar Energetic Particle (SEP) Events

- Particles (mostly protons) accelerated by the CME
- Gain their energy from shocks very close to the sun, and/or interplanetary space
- Energetic particles can damage satellites and irradiate astronauts or polar airline routes
- Understanding CME shock kinematics is key in predicting and making informed decisions regarding SEP events

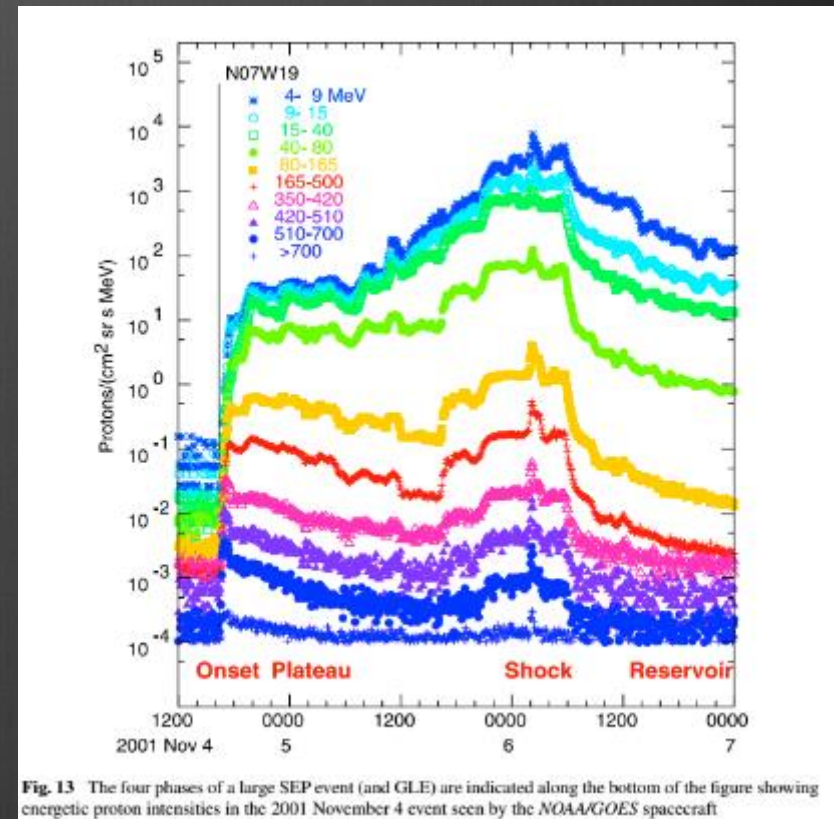
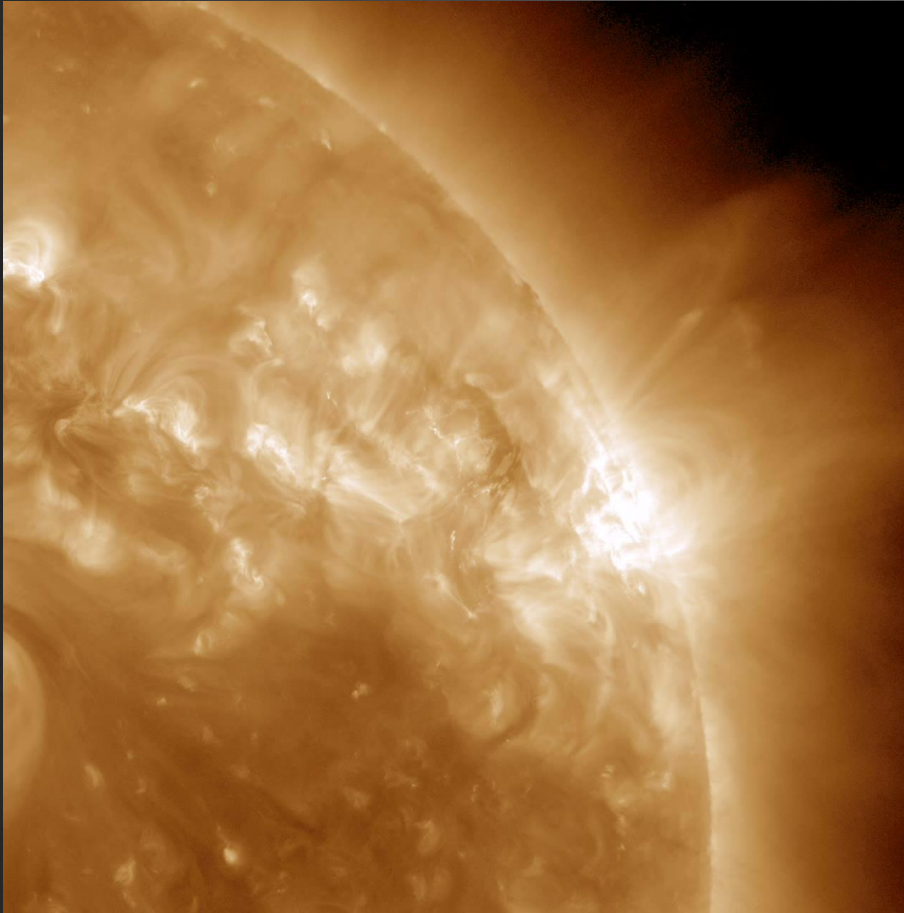


Fig. 13 The four phases of a large SEP event (and GLE) are indicated along the bottom of the figure showing energetic proton intensities in the 2001 November 4 event seen by the NOAA/GOES spacecraft

From D. Reames, Space Sci. Rev., 22 (2013)

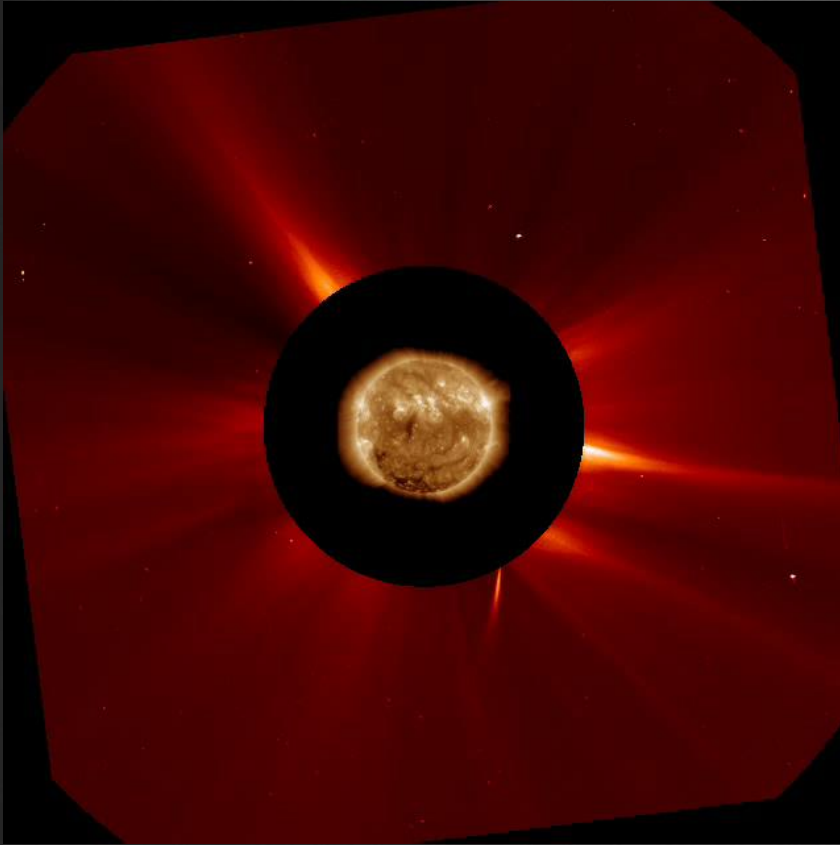
Coronal Bright Fronts (CBFs)



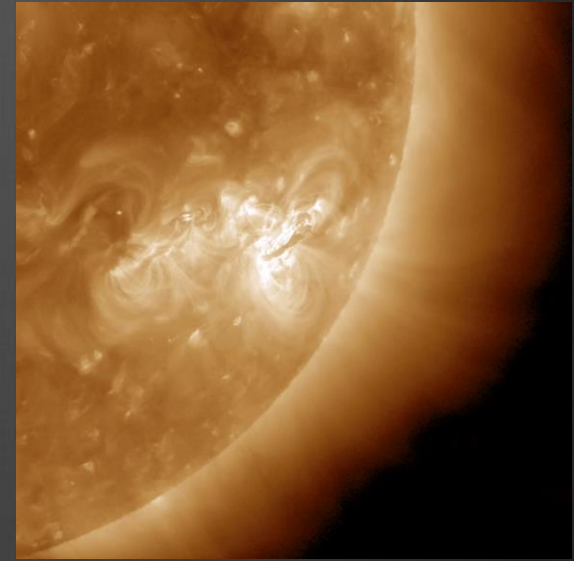
- Observed primarily in the 193 and 211 Å bands of the Atmospheric Imaging Assembly (AIA) aboard the Solar Dynamics Observatory (SDO)
- Result of a compressive or shock wave in the low corona, driven by CMEs
- How efficient are these waves at accelerating particles to SEP energies?
- Our project aims to characterize off-limb CBF kinematics using automated image processing techniques

Coronal Bright Fronts (CBFs)

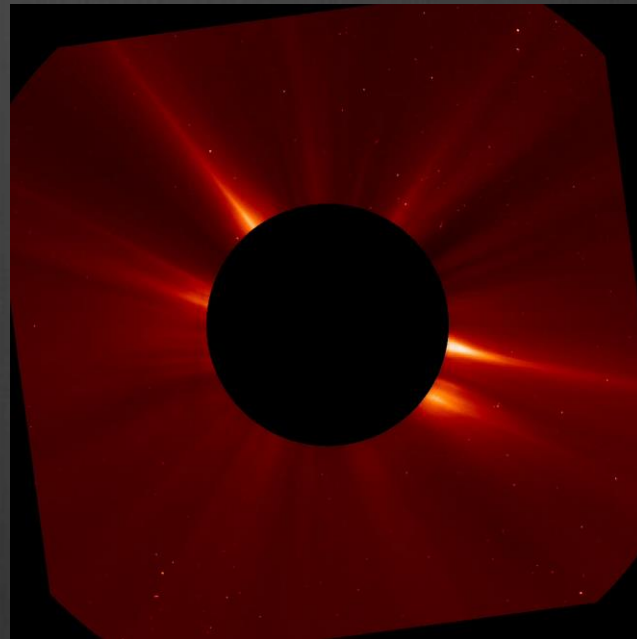
May 11, 2011 CBF Event LASCO C2



June 7, 2011
CBF Event
AIA 193 Å

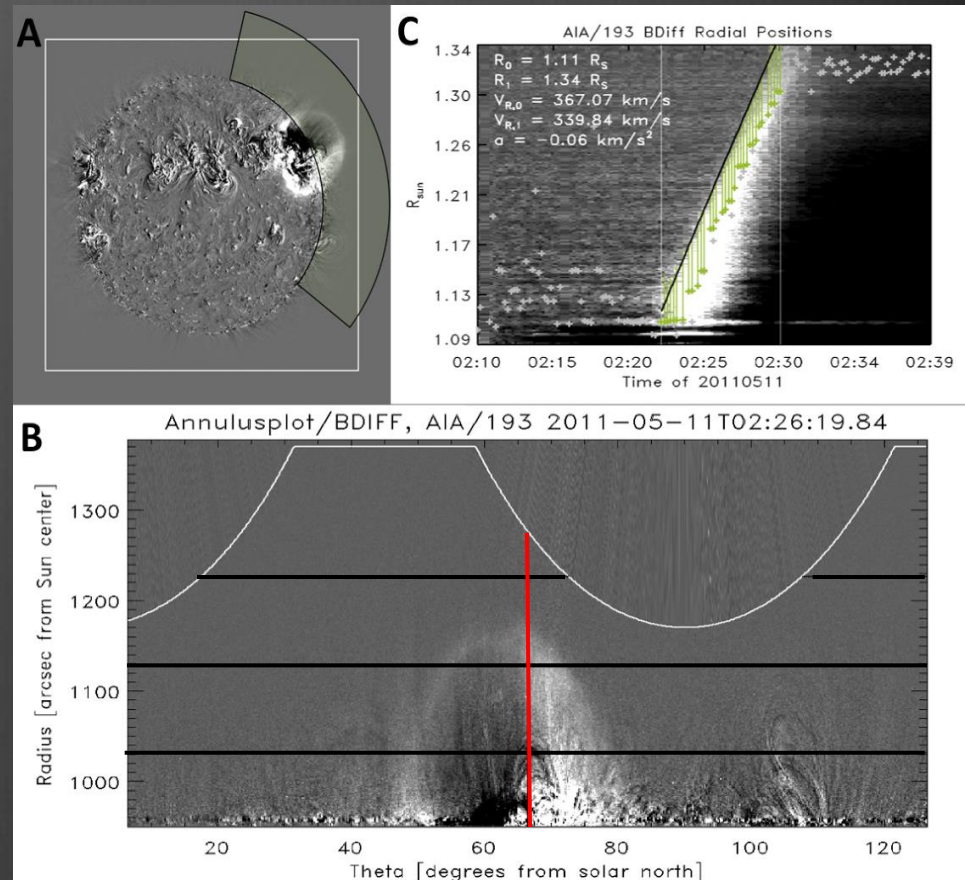


June 7, 2011
CBF Event
LASCO C2



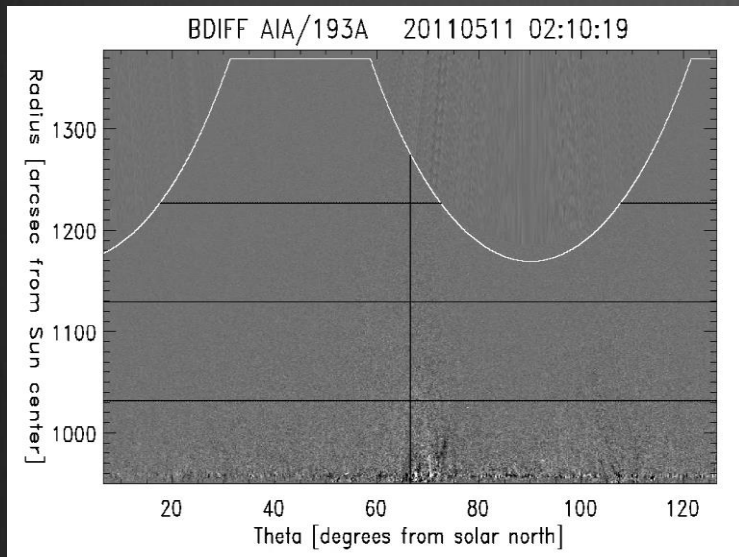
First Steps

- Before summer, two events were extensively studied 20110511 and 20100613
- Initial Goal: Process and add many new events to CBF database
- Prepared code for automation procedure

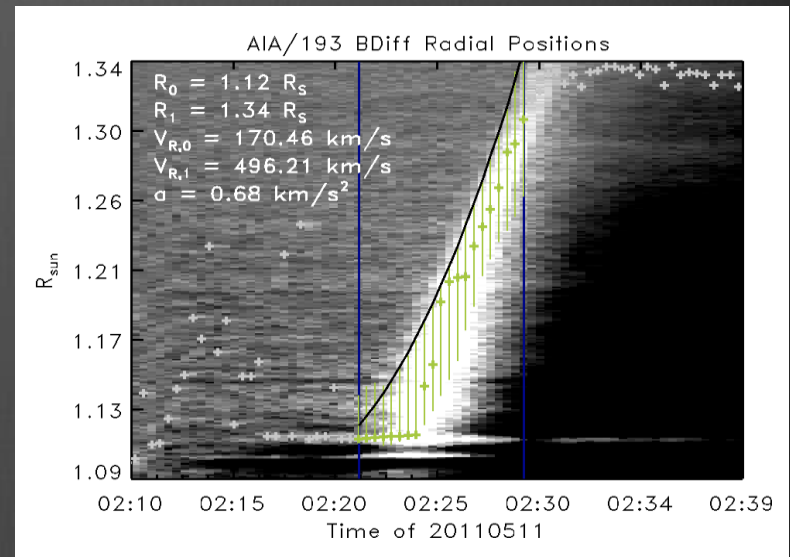


Processing pipeline from (A) Raw AIA 193 Å data to (B) Base Difference Images to (C) Pixel Intensities along red center line

Automated Kinematics Procedure

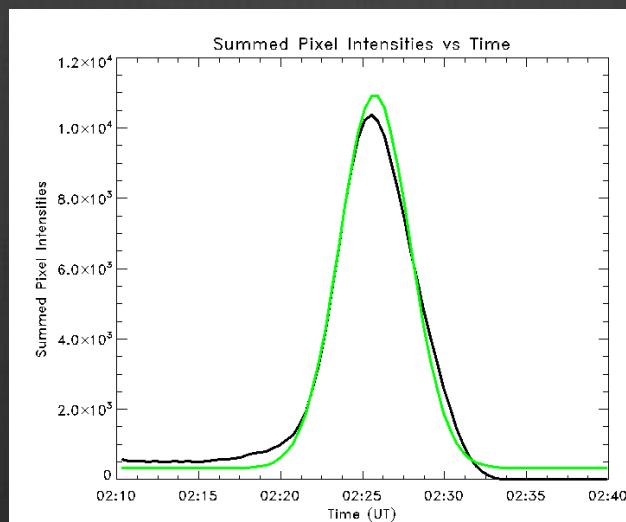


AIA Base Difference Movie of
May 11, 2011 CBF



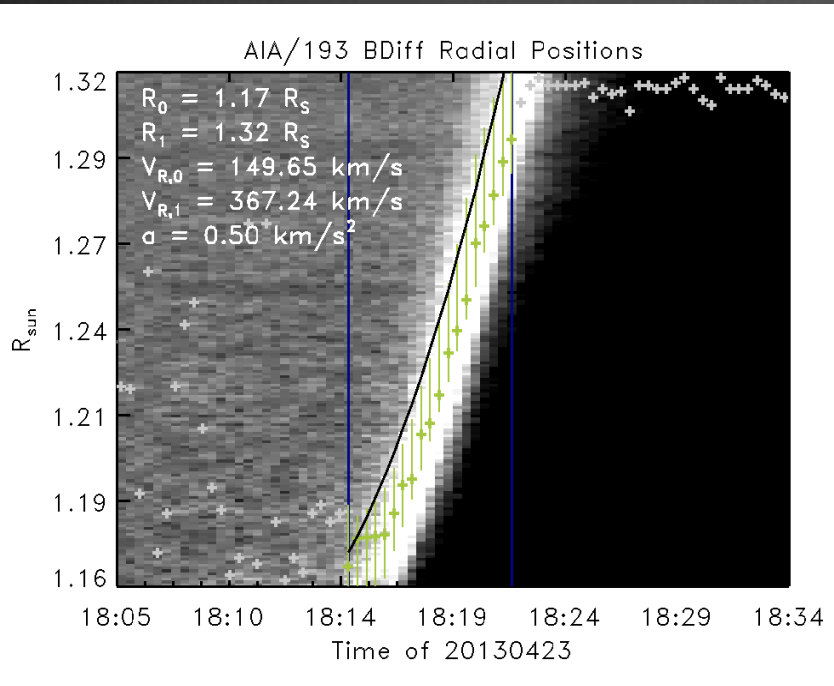
AIA Base Difference Radial
Positions of May 11, 2011 CBF

Gaussian fit of summed
pixel intensities used to
find start and end times
of CBF



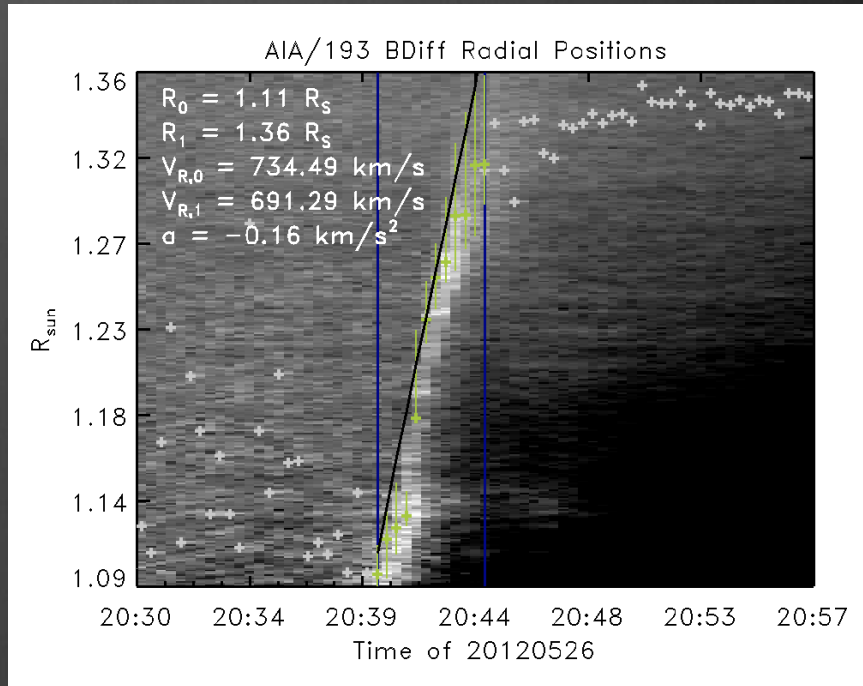
Radial Kinematics Data

20130423 CBF



- Large-scale, slower event with possible slight acceleration

20120526 CBF



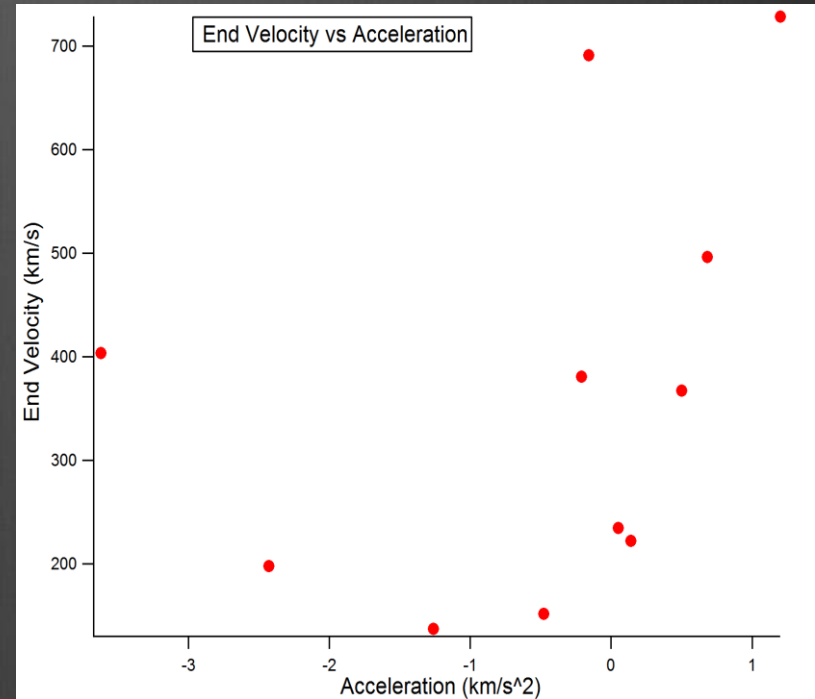
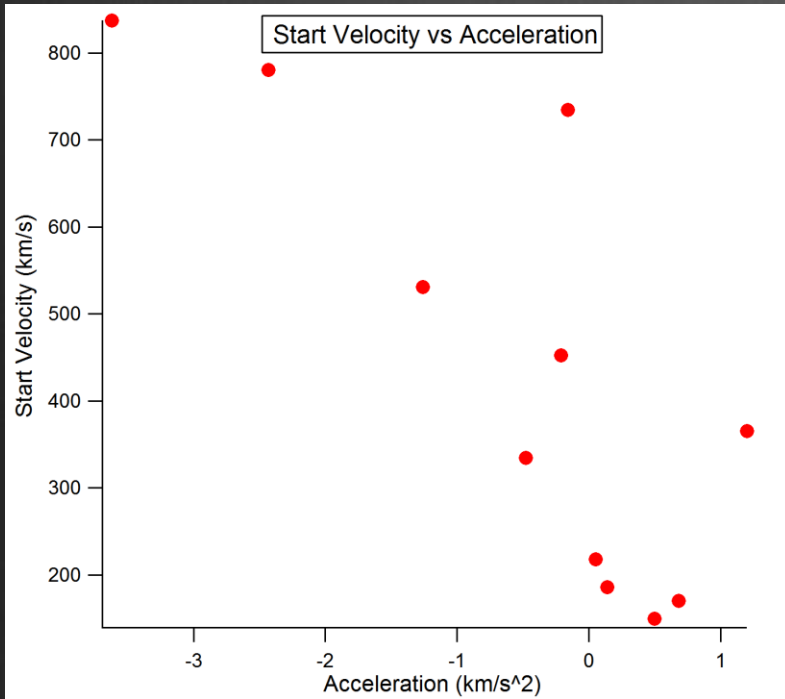
- Smaller, very fast event with possible slight deceleration

Radial Kinematics Data

Date	Start Position (R _S)	End Position (R _S)	Extent (R _S)	Mean Wave Thickness (R _S)	Start Velocity (km/s)	End Velocity (km/s)	Acceleration (km/s ²)	Relative Mean Wave Intensity
20110511	1.12	1.34	0.22	0.06	170.46	496.21	0.68	62.56
20140708	1.06	1.23	0.17	0.04	531.01	137.21	-1.26	70.49
20131212	1.11	1.32	0.21	0.06	452.68	380.73	-0.21	39.30
20130517	1.1	1.22	0.12	0.12	837.44	403.39	-3.62	41.56
20130423	1.17	1.32	0.15	0.03	149.65	367.24	0.5	84.20
20120915	1.12	1.25	0.13	0.04	334.74	151.96	-0.48	57.43
20120526	1.11	1.36	0.25	0.05	734.49	691.29	-0.16	21.55
20120424	1.1	1.26	0.16	0.04	780.68	197.62	-2.43	41.56
20110607	1.1	1.33	0.23	0.07	365.66	728.58	1.2	34.53
20110211	1.1	1.2	0.1	0.04	218.19	234.49	0.05	28.93
20110125	1.09	1.18	0.09	0.04	185.88	222.25	0.14	37.96

Mean	1.11	1.27	0.17	0.05	432.81	364.63	-0.51	47.28
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Radial Kinematics Data



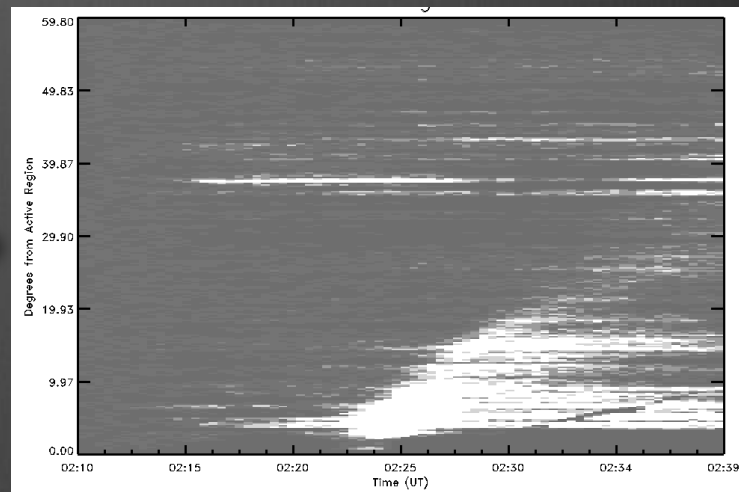
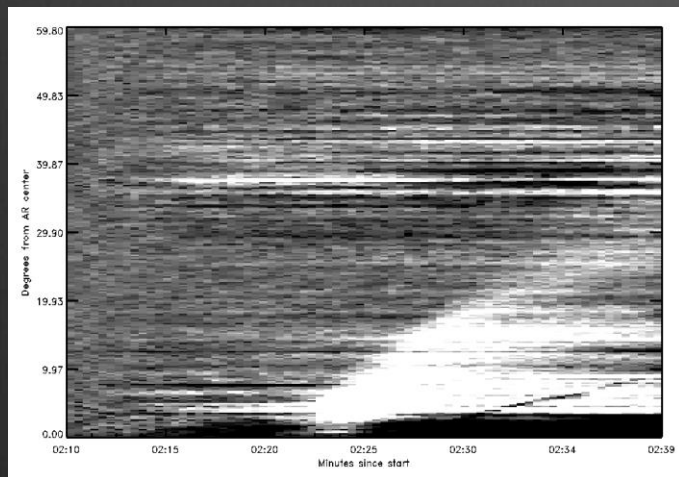
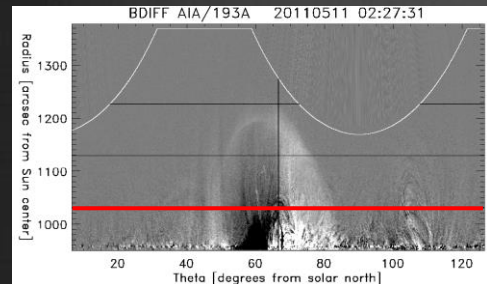
- Initially fast events tend to decelerate as they leave AIA's field of view

- Events with large end velocities tend to have positive acceleration

Tangential Kinematics Procedure

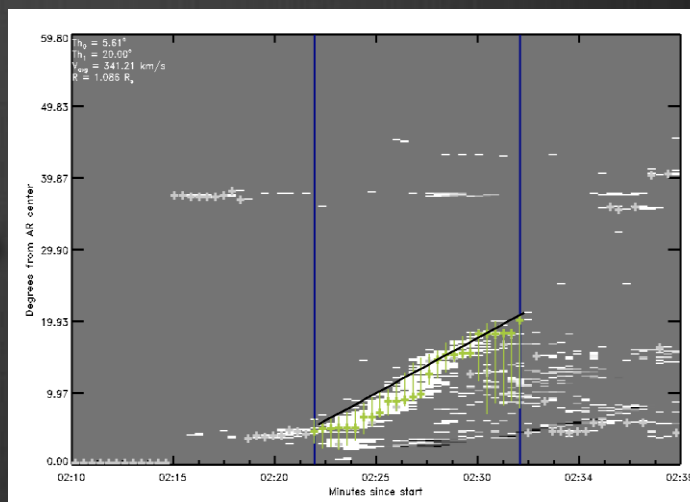
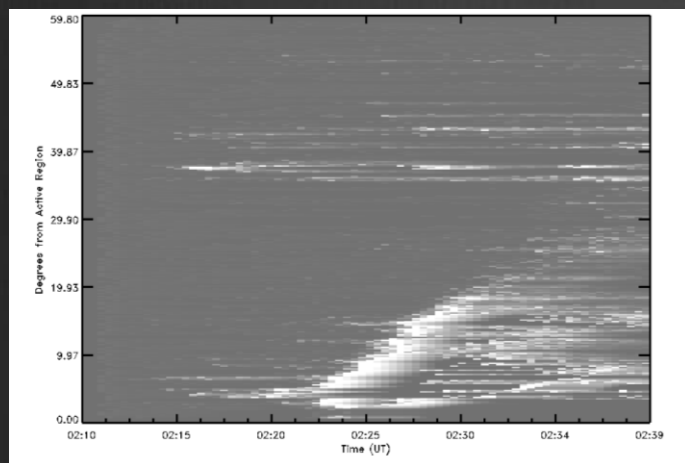
110511 Left Base
Difference
Image R = 1.086

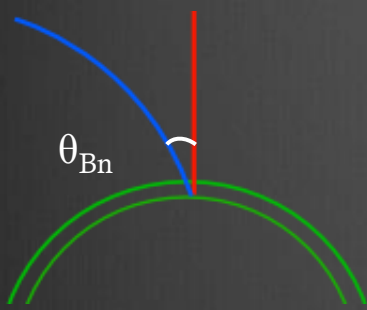
Binned 110511 Left Base
Difference Data



X-Gradient 110511 Left Base
Difference Data

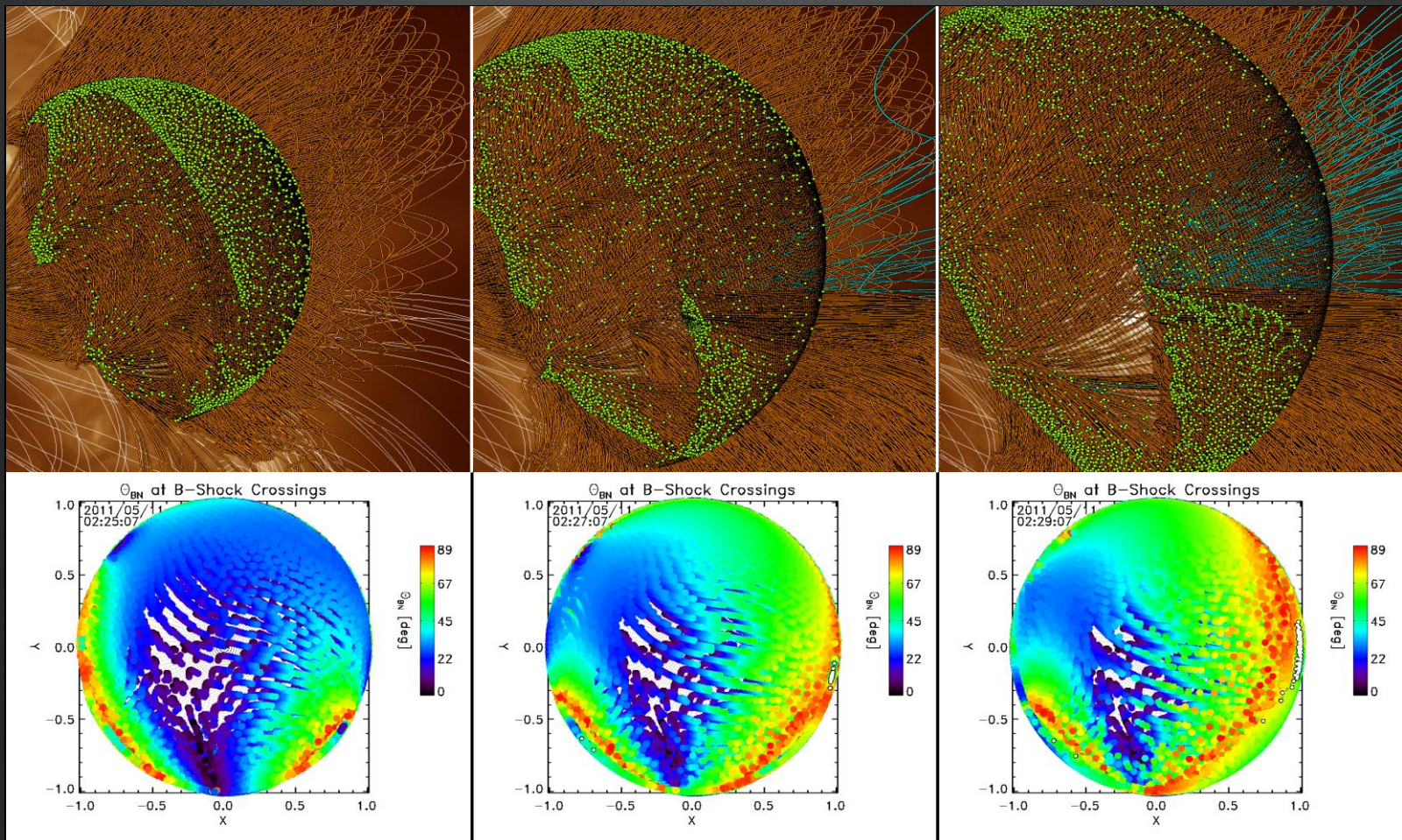
Filtered 110511 Left Base
Difference Data





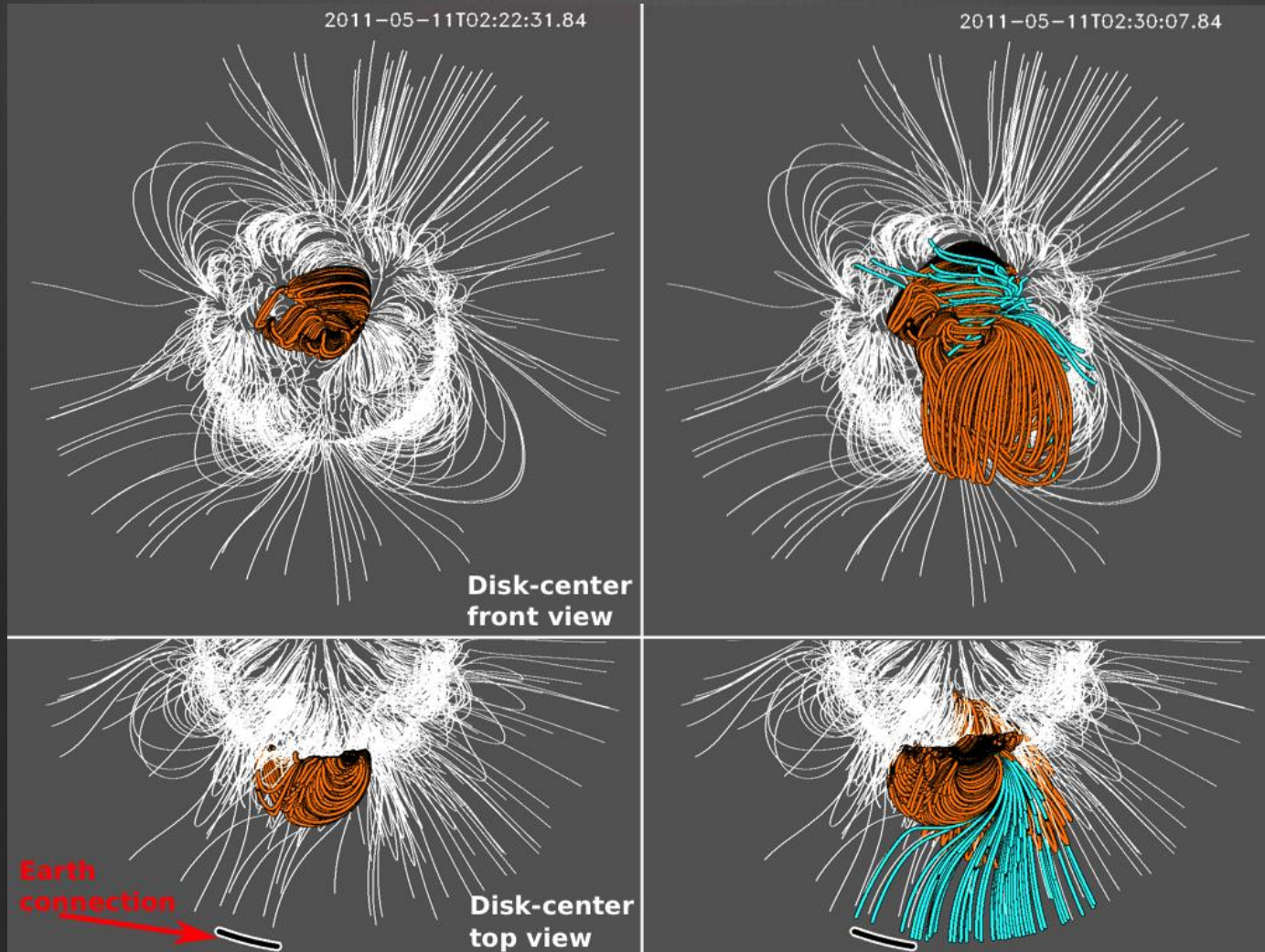
Next Steps

Kinematics Information will be used to drive time-dependent geometric models of CBFs, interacting with PFSS coronal magnetic field lines...



Next Steps

... and estimate heliospheric magnetic connectivity to Earth, etc.



Summary

- CBF kinematics information will drive geometric models providing insights to potential SEP acceleration in the lower corona
- Automated kinematics procedures has finished a pipeline of programs which can take event metadata to kinematics information
- On average, CBFs identified so far are decelerating in the lower corona with most speeds between 300-400 km/s
- Analyzed CBF events are added to a public database to be used as a research tool
- Radial kinematics could be improved by using a gradient based approach, current wave position centering location could be better chosen using an iterative technique

Acknowledgements

- A big thanks to:
 - My advisors, Kamen Kozarev and Patricia Jibben
 - The Solar REU Program, Henry “Trae” Winter and Katharine Reeves
 - Jonathan Sattelberger
- NSF-REU solar physics program at SAO, grant number AGS-1263241
- AIA:
contract SP02H1701R from Lockheed-Martin to SAO

Current CBF Database at http://helio.cfa.harvard.edu/cashew/

Questions?

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Welcome to the Solar Coronal Off-limb Wave Analysis Database and Repository.
The repository consists of data and information about solar coronal off-limb bright fronts related to CME shocks.
Feel free to contact me with any questions about the database.

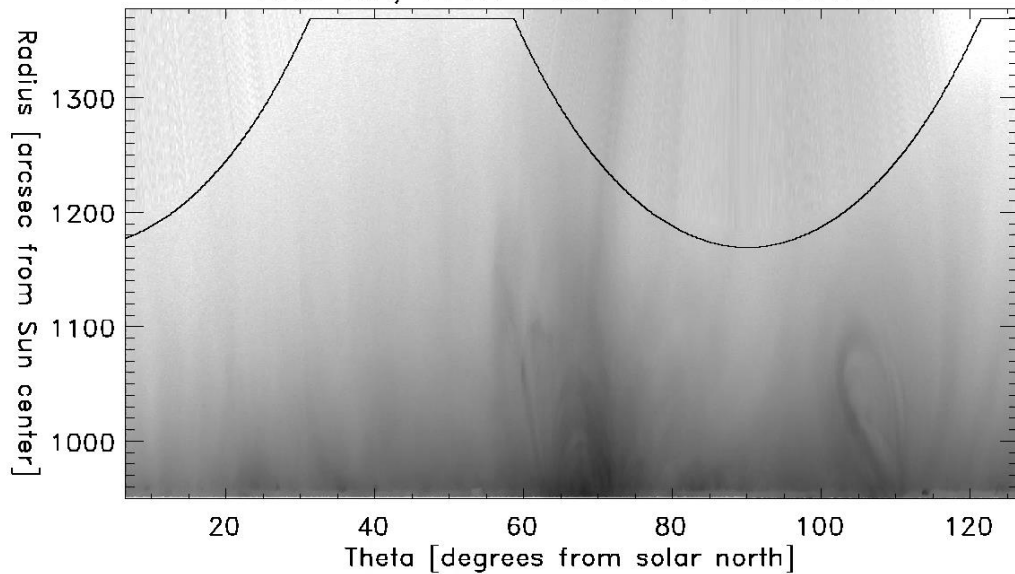
OFF-LIMB CORONAL BRIGHT FRONTS DATABASE (CURRENTLY 30 EVENTS) LAST UPDATED ON Wed Aug 13 01:03:12 2014

Date	Start (UT)	End (UT)	Flare Class	X (") East/West	Y (")	Filament Erupt	Loop Erupt	Metric Radio Spectra			EUV Raw / Run Diff. Movies		EUV Annulus plot Raw / Run Diff. Movies		Kinematics		TEAM DEM Movies	PFSS / Shock Model
								eCallisto	Nancay	RSTN	ATA/193	ATA/193	Radial	Tangential				
2014/07/08	16:08	16:28	M6.5	-810	161	Yes	No	eCallisto			RAW	RDIFF	ARAW	ARDIFF		PNG		
2014/04/07	20:50	21:23		900	-608	Yes	No				RAW	RDIFF	ARAW	ARDIFF				
2014/02/25	00:40	00:55	X4.9	-972	-214	Yes	No			RSTN	RAW	RDIFF	ARAW	ARDIFF		PNG		
2014/01/08	03:40	03:52	M3.6	950	200	No	Yes			RSTN	RAW	RDIFF	ARAW	ARDIFF		PNG		
2013/12/12	03:03	03:33	B2.2	750	-400	Yes	No	eCallisto		RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG	DEM	
2013/12/07	07:17	07:36	M1.2	690	-270	No	Yes			RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG		
2013/11/19	10:17	10:42	X1.0	900	-220	No	Yes	eCallisto	NRH	RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG		
2013/11/07	10:14	10:25		-999	100	No	Yes	eCallisto	NRH	RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG		
2013/11/05	22:10	22:40	C1.5	-563	-203	No	Yes	eCallisto		RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG		
2013/10/29	21:40	22:10	X2.3	990	38	Yes	Yes	eCallisto		RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG		
2013/05/17	08:41	09:09	M3.2	-576	192	Yes	Yes	eCallisto	NRH	RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG	DEM	
2013/05/01	02:28	02:45		-865	237	Yes	Yes	eCallisto		RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG		
2013/04/23	18:05	18:35		999	-399	No	Yes	eCallisto		RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG	DEM	
2013/04/11	06:54	07:20	M6.5	-200	200	No	Yes	eCallisto		RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG		
2012/10/07	20:15	20:45	C1.2	-932	378	No	Yes			RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG		
2012/09/15	22:50	23:20	B9.6	806	422	No	Yes			RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG	DEM	
2012/07/28	20:43	20:59	M6.1	-745	-662	No	Yes			RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG		
2012/05/26	20:30	20:58		899	416	No	No			RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG	DEM	
2012/04/24	07:39	07:55	C3.7	-1019	247	Yes	No	eCallisto		RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG	DEM	
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2011/01/25	11:56	12:26		-955	-248	No	Yes			RSTN	RAW	RDIFF	ARAW	ARDIFF	PNG	PNG		
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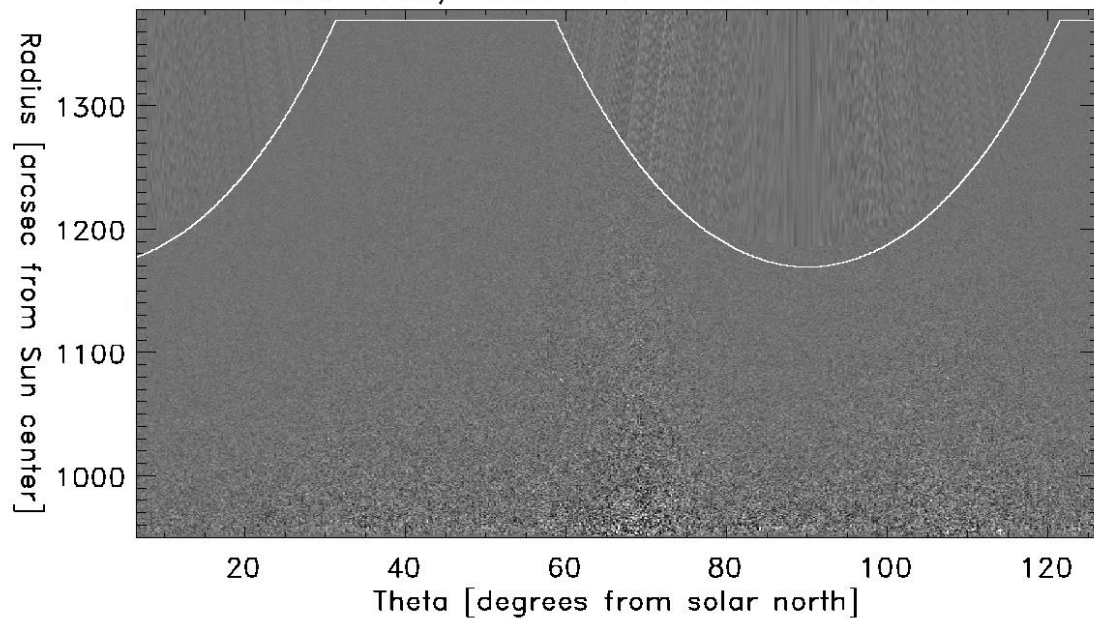


CASHew documentation Contact

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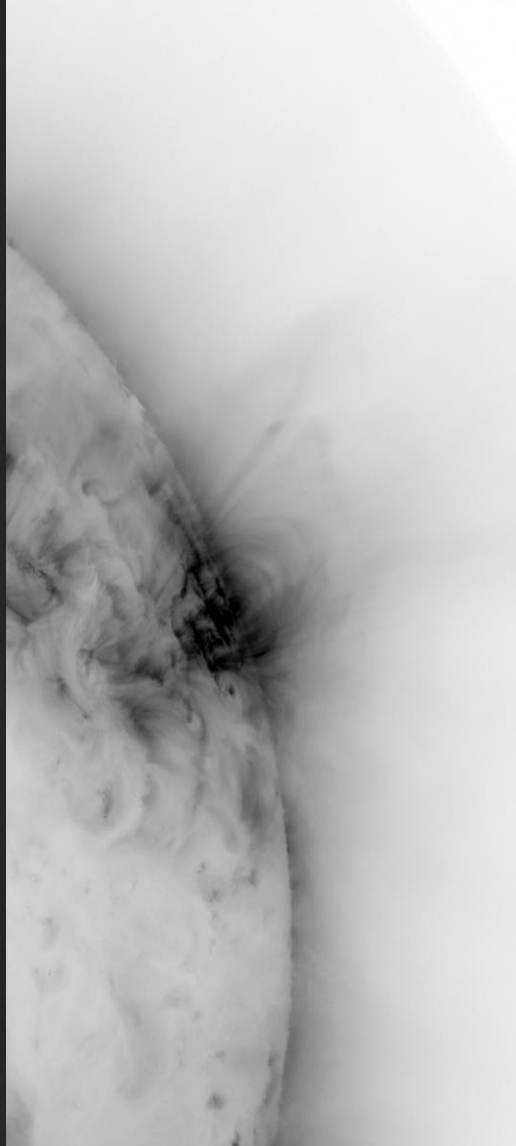


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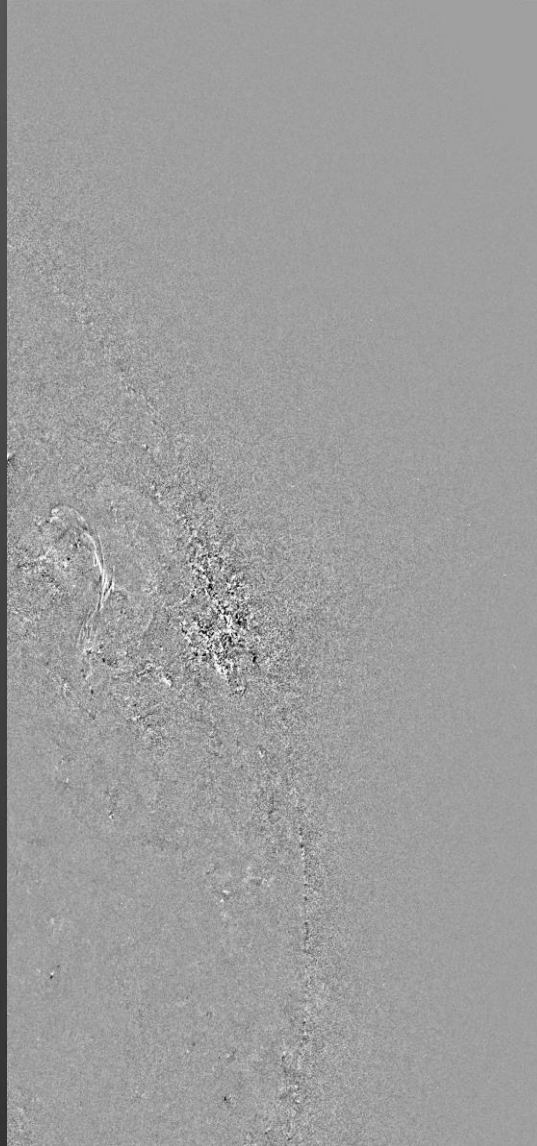
Raw

2011-05-11T02:10:43.84 / AIA:193 001



Base Difference

2011-05-11T02:10:43.84 / AIA:193 B 001



Running Difference

2011-05-11T02:10:43.84 / AIA:193 R 001

