

LOW-RADIOACTIVITY
UNDERGROUND

ARGON



Pacific Northwest
NATIONAL LABORATORY

Proudly Operated by **Battelle** Since 1965

Developing the First Supply of Underground Argon

THOMAS ALEXANDER

Pacific Northwest National Laboratory

Argon on Earth

Atmospheric isotopic abundance

J.-Y. Lee, et al., *Geochim. Cosmochim. Acta* 70 (2006) 4507-4512

^{36}Ar – 0.334%

^{38}Ar – 0.063%

^{40}Ar – 99.604%

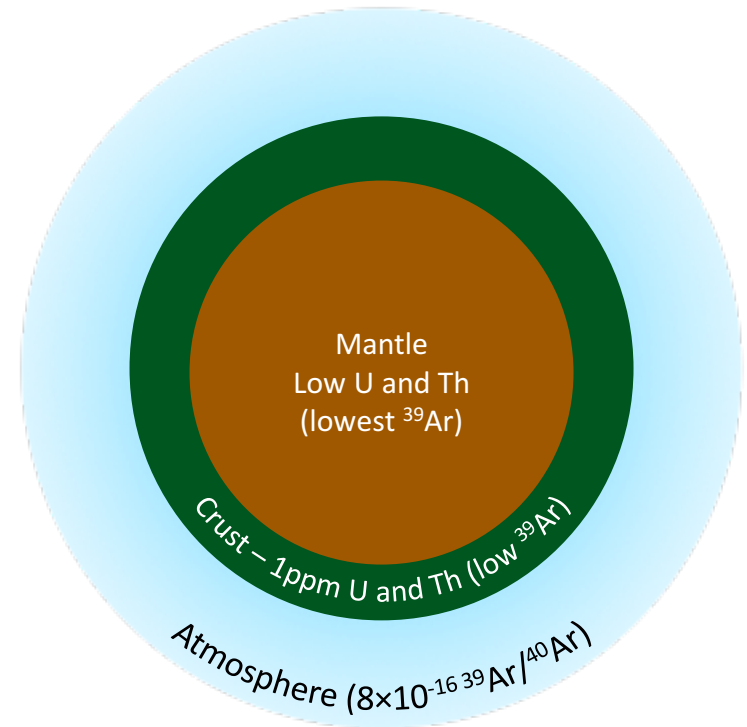
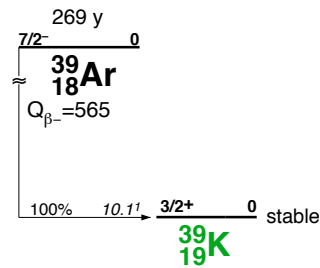
▶ ^{40}Ar comes from ^{40}K decay

▶ ^{39}Ar Origins

■ $^{40}\text{Ar}(n,2n)\rightarrow^{39}\text{Ar}$

■ $^{40}\text{Ar}(p,np)\rightarrow^{39}\text{Ar}$

▶ Underground via U + Th fission neutrons



The Hunt



2007: Helium Reserve ³⁹Ar content < 5% atmospheric levels

Discovery of underground argon with low level of radioactive ³⁹Ar and possible applications to WIMP dark matter detectors

D. Arcata-Kane^a, R. Acciari^a, O. Amaiz^a, M. Antonello^b, B. Balbusinov^d, M. Baldo Collin^d, C. J. Ballentine^e, R. Banaś^f, I. Bangali^g, A. Barakho^h, P. Benettiⁱ, J. Benziger^j, A. Burgers^k, F. Calaprice^l, E. Calligaris^m, M. Camlinighiⁿ, N. Casali^o, F. Carbonara^p, M. Cassidy^q, F. Cavanna^r, S. Centro^s, A. Chavarria^t, D. Cheng^u, A. G. Cocco^v, P. Collin^w, F. Dalnoki-Veress^x, E. de Haas^y, F. Di Pompeo^z, G. Fiorillo^{aa}, F. Fitch^{ab}, V. Gallo^c, C. Galbiati^{ac}, M. Gault^{ad}, S. Gazzana^{ae}, L. Grandi^{af}, A. Goretta^{ag}, T. Hightail^{ah}, T. Hightail^{ai}, T. Hohmann^{aj}, An. Inami^{ak}, A. LaCava^{al}, M. Lantierstein^{am}, H. Y. Lee^{an}, M. Leung^{ao}, B. Loefer^{ap}, H. H. Looi^{aq}, B. Lyons^{ar}, D. Marks^{as}, F. McCarty^{at}, G. Meng^{au}, C. Montanari^{av}, S. Mukhopadhyay^{aw}, A. Nelson^{ax}, O. Palamara^{ay}, L. Pandola^{az}, F. Pietropolo^{ba}, T. Pivonka^{bb}, A. Pocar^{bc}, R. Puntelero^{bd}, A. Rappoldi^{be}, G. Raella^{bf}, P. Rosati^{bg}, D. Robertson^{bh}, M. Roncaldelli^{bi}, M. Ronella^{bj}, C. Rubbia^{bk}, J. Ruderman^{bl}, R. Saldanha^{bm}, C. Schmitt^{bn}, R. Scott^{bo}, E. Segreto^{bp}, A. Shirley^{bq}, A. M. Szele^{br}, R. Tagaglia^{bs}, T. Tosliano^{bt}, S. Ventura^{bu}, C. Vignoli^{bv}, C. Vianje^{bw}, R. Vondracek^{bx}, A. Vianje^{by}

^aDepartment of Physics, Princeton University, Princeton, NJ 08542, USA
^bDepartment of Engineering, Princeton University, Princeton, NJ 08542, USA
^cINFN, Laboratori Nazionali del Gran Sasso, Assago (AQ), 67100, Italy
^dINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^eINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^fINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^gINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^hINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
ⁱINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^jINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^kINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^lINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^mINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
ⁿINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^oINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^pINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^qINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^rINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^sINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^tINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^uINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^vINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^wINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^xINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^yINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^zINFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{aa}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ab}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ac}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ad}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ae}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{af}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ag}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ah}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ai}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{aj}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ak}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{al}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{am}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{an}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ao}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ap}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{aq}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ar}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{as}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{at}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{au}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{av}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{aw}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ax}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ay}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{az}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{ba}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bb}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bc}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bd}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{be}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bf}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bg}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bh}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bi}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bj}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bk}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bl}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bm}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bn}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bo}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bp}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bq}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{br}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bs}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bt}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bu}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bv}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bw}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{bx}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
^{by}INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy

2007: Begin Searching CO₂ Wells

Notes on Trip to Suoyaroc - May 21-26 2007 and Proposal for Small Scale Argon Production in the 2008 Campaign
 Notes on Trip to Suoyaroc - May 23-26 2007 and Proposal for Small Scale Argon Production in the 2008 Campaign
 F. Calaprice, E. de Haas, C. Galbiati, A. Goretta, A. Inami, A. Nelson
 Physics Department of Princeton University
 February 10, 2008

	Gas from Liquid Trap	Gas from the PSA
Ar	290 ppm	2.54
CH ₄	11.54	1200 ppm
CO ₂	79.34	6500 ppm
C ₂ H ₆	0 ppb	0 ppb
N ₂	467 ppm	750 ppm
N ₂ O	2860 ppm	5000 ppm
Ne	700 ppm	22.34
N ₂	8.78	73.84
O ₂	550 ppm	120 ppm

Table 1: Composition of the stream from the liquid trap and of the stream from the PSA unit.

2008: First Production

First Large Scale Production of Argon Depleted in ³⁹Ar from Underground Wells

H. Babcock¹, C. Ballentine², R. Banaś³, I. Benziger⁴, A. Burgers⁵, F. Calaprice⁶, M. Cassidy⁷, A. Chavarria⁸, P. Collin⁹, F. Dalnoki-Veress¹⁰, E. de Haas¹¹, C. Galbiati¹², M. Gault¹³, S. Gillilan¹⁴, A. Goretta¹⁵, W. Hayes¹⁶, T. Hohmann¹⁷, An. Inami¹⁸, B. Loefer¹⁹, H. H. Looi²⁰, D. Montanari²¹, S. Mukhopadhyay²², A. Nelson²³, A. Pocar²⁴, R. Puntelero²⁵, D. Robertson²⁶, R. Saldanha²⁷, G. Schlowsky²⁸, C. Schmitt²⁹, S. Vandeberg³⁰, V. Vandeberg³¹

¹Department of Physics, Princeton University, Princeton, NJ 08542, USA
²Department of Chemical Engineering, Princeton University, Princeton, NJ 08542, USA
³Department of Physics, University of Notre Dame, Notre Dame, IN 46556, USA
⁴Department of Geosciences, University of Houston, Houston, TX 77004, USA
⁵Department of Earth, Atmospheric, and Environmental Sciences, University of Massachusetts, Haverhill, MA 01830, USA
⁶School of Earth, Atmospheric, and Environmental Sciences, University of Massachusetts, Haverhill, MA 01830, USA
⁷Department of Earth and Planetary Sciences, Harvard University, Cambridge, MA 02138, USA
⁸School of Geosciences, The University of Edinburgh, Edinburgh, EH8 9JX, United Kingdom
⁹Department of Physics, Stanford University, Stanford, CA 94305, USA
¹⁰Physics Institute, University of Bern, 3012 Bern, Switzerland
¹¹INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
¹²INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
¹³INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
¹⁴INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
¹⁵INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
¹⁶INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
¹⁷INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
¹⁸INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
¹⁹INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
²⁰INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
²¹INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
²²INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
²³INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
²⁴INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
²⁵INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
²⁶INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
²⁷INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
²⁸INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
²⁹INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
³⁰INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
³¹INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy

Abstract

We report on the first large-scale production of depleted argon from underground gas wells. We processed the exhaust stream of the CO₂ liquefier of the Holston Drift Plant in Buoyton, VA, with a special Vacuum Swing Adsorption plant. The CO₂ gas fed directly from the well into the liquefier contains argon at the concentration of 40-70 ppm, and the argon concentration in the exhaust stream of the liquefier is in the range 300-450 ppm. The Vacuum Swing Adsorption plant produces crude argon, concentrating the argon to the level of 80,000-100,000 ppm (8-10%) in a single pass. The argon production rate is 0.5-1.0 kg/day. We determined that the underground argon is depleted in ³⁹Ar by a factor 10 or more relative to atmospheric argon activated by cosmic rays, by analysis in a low-background proportional counter. Mass-spectroscopic analysis of the ³⁹Ar activity is under way. Depleted argon is of interest for the construction of large-scale WIMP dark matter searches and of detectors of reactor neutrinos for non-proliferation efforts. WIMP dark matter searches of high-sensitivity may require depleted argon targets of 10 tons or more. Underground argon offers an affordable solution for the production of depleted argon targets. Prior to this work, only a few grams of depleted argon from underground wells were separated and purified from natural gas. To the best of our knowledge, the work reported in this paper is the first production of depleted argon from underground sources of the world.

2009: Move to Colorado

First Large Scale Production of Low Radioactivity Argon From Underground Sources

H. O. Back¹, F. Calaprice², C. Condon³, E. de Haas⁴, R. Foc⁵, C. Galbiati⁶, A. Goretta⁷, T. Hohmann⁸, An. Inami⁹, B. Loefer¹⁰, D. Montanari¹¹, A. Nelson¹², A. Pocar¹³

¹Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542
²INFN and Dipartimento di Fisica, University of Trieste, Pavia 30123, Italy
³Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542
⁴Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542
⁵Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542
⁶Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542
⁷Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542
⁸Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542
⁹Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542
¹⁰Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542
¹¹Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542
¹²Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542
¹³Department of Physics, Princeton University, Jadwin Hall, Princeton, NJ 08542

arXiv:0712.0381v1 [astro-ph] 3 Dec 2007

Abstract

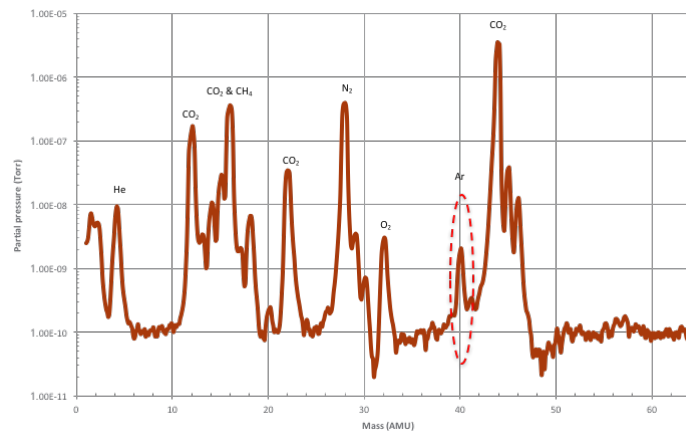
We report on the first measurement of ³⁹Ar in argon from underground natural gas reservoirs. The gas stored in the US National Helium Reserve was found to contain a low level of ³⁹Ar. The ratio of ³⁹Ar to stable argon was found to be 5.4-10⁻¹⁷ (0.4% CL), less than 1% the value in atmospheric argon (³⁹Ar/(Ar+³⁹Ar)). The total quantity of argon currently stored in the National Helium Reserve is estimated at 1000 tons. ³⁹Ar represents one of the most important backgrounds in argon detectors for WIMP

Abstract
 We report on the first large-scale production of low radioactivity argon from underground gas wells. Low radioactivity argon is of general interest, in particular for the construction of large scale WIMP dark matter searches and detectors of reactor neutrinos for non-proliferation efforts. Atmospheric argon has an activity of about 1 Bq/kg from the decay of ³⁹Ar; the concentration of ³⁹Ar in the underground argon we are collecting is at least a factor of 100 lower than this value.
 The argon is collected from a stream of gas from a CO₂ well in southwestern Colorado with a Vacuum Swing Adsorption (VPSA) plant. The gas from the well contains argon at a concentration of 400-600 ppm, and the VPSA plant produces an output stream with an argon concentration in the level of 30,000-50,000 ppm (3-5%) in a single pass. The gas is sent for further processing to Fermilab where it is purified by cryogenic distillation. The argon production rate is presently 0.5 kg/day.
 Keywords: depleted argon, dark matter, pressure swing adsorption

Exploration at Doe Canyon



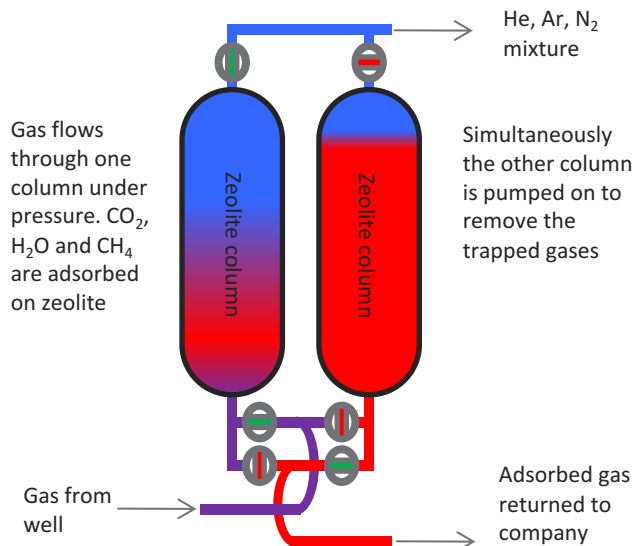
CO₂ well in SW Colorado with 400ppm Ar
(Nucl. Phys. B, 197 (2009) 70-73)
(Nucl. Instr. Meth. A 587 (2008) 46-51)



Argon Extraction-Cortez, Colorado

Approximate Input:

Gas Type	Concentration from well
Carbon Dioxide	96%
Nitrogen	2.4%
Methane	0.57%
Helium	0.43%
Other hydrocarbons	0.21%
Argon	440 ppm



Gas Type	Concentration in Output
Helium	85-95%
Argon	3-6%
Nitrogen	1-10%
Methane, Oxygen	Trace
Carbon Dioxide	Trace
Other Hydrocarbons	Trace

Average Production: 140g/day



First Measurement of Underground Argon

2011-2012

A Study of the Residual ^{39}Ar Content in Argon from Underground Sources

J. Xu^a, F. Calaprice^{a*}, C. Galbiati^a, A. Goretti^b, G. Guray^a, T. Hohman^{a,1},
D. Holtz^{a,2}, A. Ianni^b, M. Laubenstein^b, B. Loer^{a,3}, C. Love^{a,4},
C.J. Martoff^a, D. Montanari^{a,3}, S. Mukhopadhyay^d, A. Nelson^a,
S.D. Rountree^e, R.B. Vogelaar^a, A. Wright^a

^aDepartment of Physics, Princeton University, Princeton NJ 08544, USA

^bINFN Laboratori Nazionali del Gran Sasso, SS 17 bis Km 18.910, 067010 Assergi (AQ), Italy

^cPhysics Department, Temple University, Philadelphia, PA 19122, USA

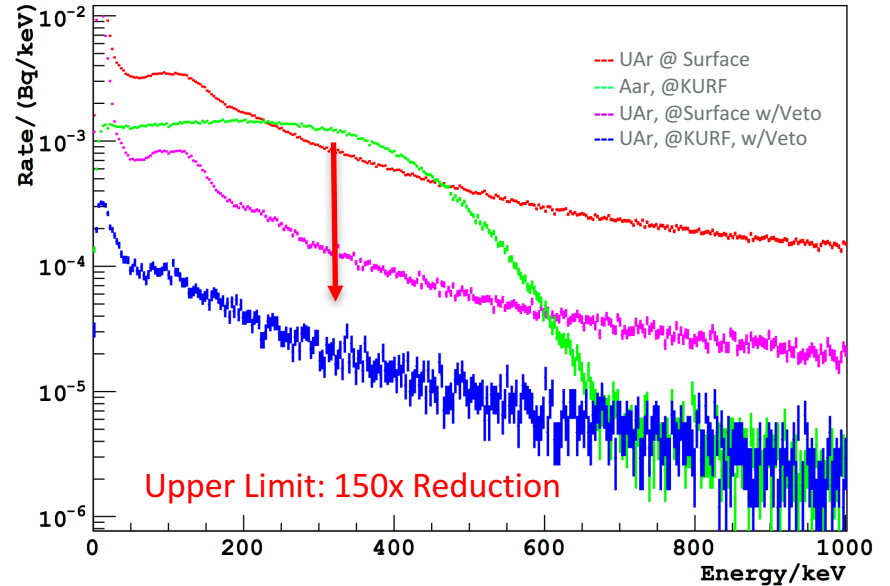
^dDepartment of Earth and Planetary Sciences, Harvard University, Cambridge, MA 02138, USA

^ePhysics Department, Virginia Polytechnic Institute and State University, Blacksburg, VA 24061, USA

Abstract

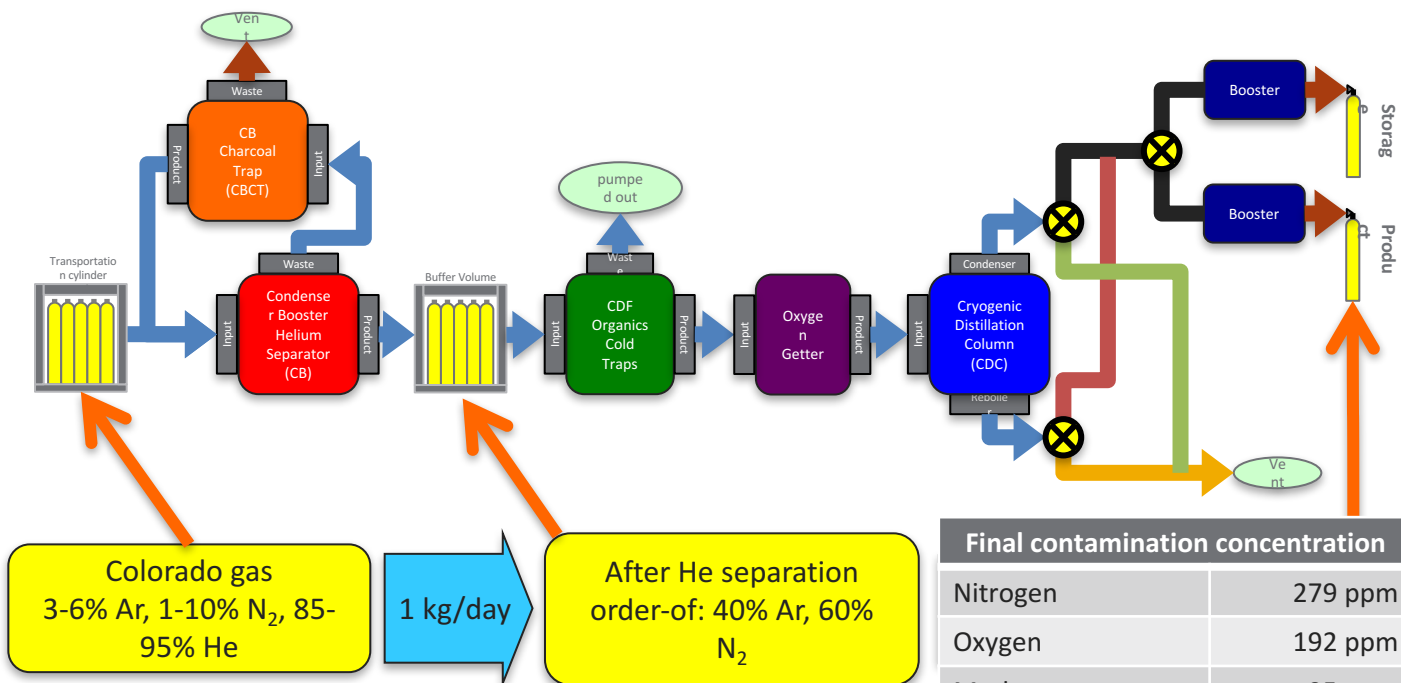
The discovery of argon from underground sources with significantly less ^{39}Ar than atmospheric argon was an important step in the development of direct-detection dark matter experiments using argon as the active target. We report on the design and operation of a low background detector with a single phase liquid argon target that was built to study the ^{39}Ar content of the underground argon. Underground argon from the Kinder Morgan CO₂ plant in Cortez, Colorado was determined to have less than 0.65% of the ^{39}Ar activity in atmospheric argon.

Keywords: underground argon, dark matter search technique,



v:1204.6011v1 [physics.ins-det] 26 Apr 2012

Purification— Fermilab, Illinois, USA

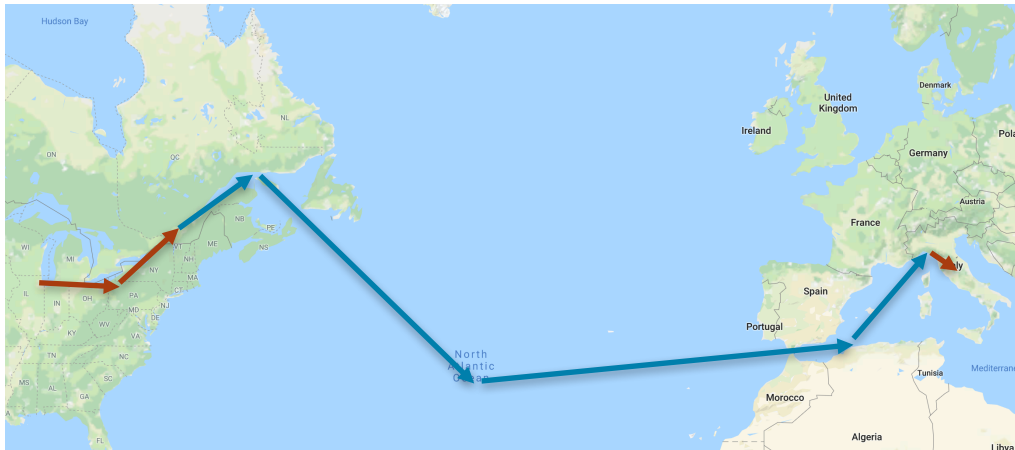


Total DS-50 production: 157 kg

Final contamination concentration	
Nitrogen	279 ppm
Oxygen	192 ppm
Methane	95 ppm
Helium	3 ppm
Carbon Dioxide	14 ppm

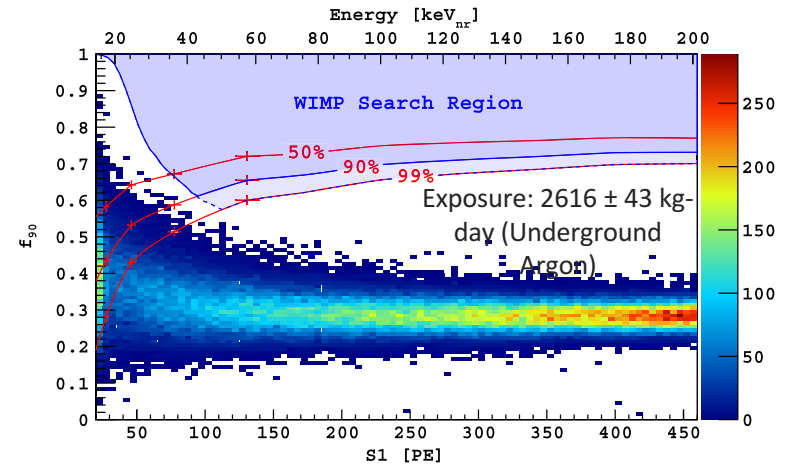
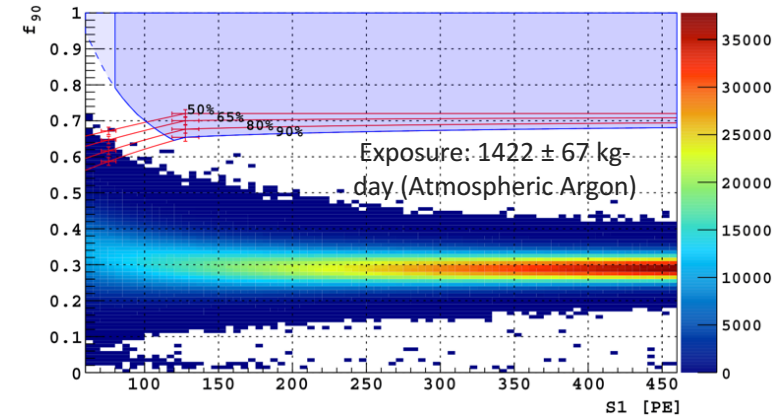
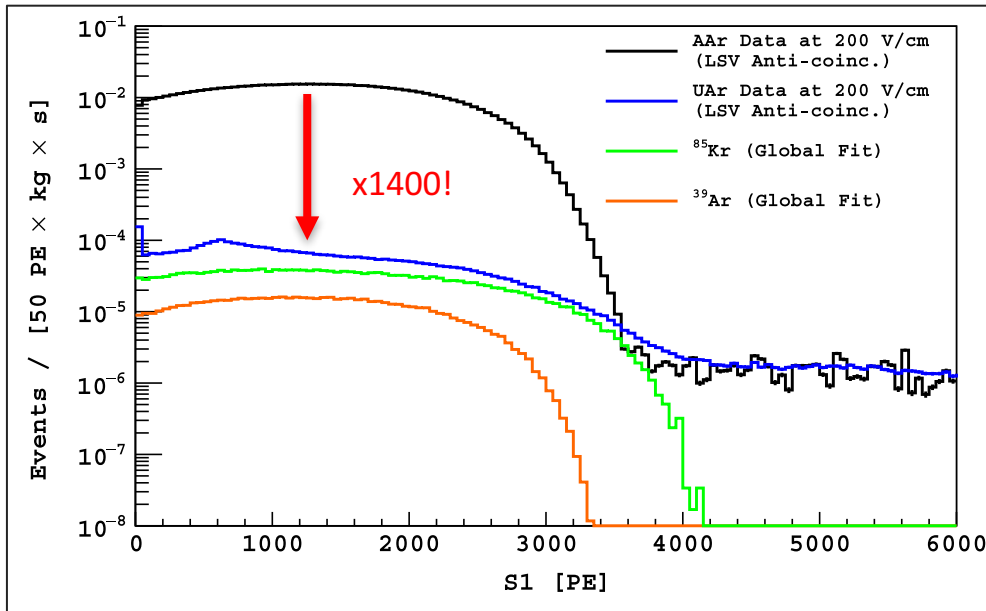
The Target Leaves FNAL for LNGS

- ▶ Underground Argon tested in SCENE cryostat
 - Test final polishing technique (Zirconium Getter)
 - Recaptured in Condenser Booster, "Lossless test"
- ▶ Moved to coast by truck, transported by sea
 - Reduce possible cosmogenic activation
- ▶ Final ~10 kg flown
- ▶ Detector filled April 2015



DarkSide-50: The First Underground Argon-based Dark Matter Detector. Physics Letters B, 743 (2015) & Phys. Rev. D 93, 081101(R) (2016)

- 157 kg produced in 5 years
- ^{39}Ar – 0.73 mBq/kg
- ^{85}Kr – 2.05 mBq/kg



Tackling Future Challenges

▶ VPSA Challenges:

- Operating R&D plant as production facility
- Lower production rate than expected
- Zeolite poisoning
 - Reduced adsorption efficacy over time
 - 900 lbs of zeolite was replaced 3 times

▶ Distillation/Purification Challenges

- Unexpected helium and oxygen levels
- Minor contaminations in VPSA output clogged distillation column
- Low production rate from VPSA (recapture of UAr from waste streams)



Contaminants found in Zeolite and Purification Systems

C ₃ H ₈	C ₅ H ₁₀ O	C ₇ H ₁₄	C ₆ H ₁₂ O	C ₇ H ₁₆	C ₈ H ₁₈
C ₅ H ₁₀ O	C ₅ H ₁₀ O	C ₆ H ₁₃ l	C ₆ H ₁₂ O	C ₆ H ₁₂ O	C ₈ H ₁₈
C ₅ H ₁₂	C ₆ H ₁₄	C ₆ H ₁₃ l	C ₇ H ₁₆	C ₅ H ₈ O ₂	C ₆ H ₁₀ O ₂
C ₆ H ₁₄	C ₆ H ₁₂ O	C ₇ H ₁₆	C ₆ H ₆	C ₈ H ₁₆	C ₈ H ₁₈
C ₅ H ₁₀	C ₆ H ₁₂	C ₇ H ₁₆	C ₆ H ₆	C ₈ H ₁₆	C ₉ H ₂₀

Why Did We See Oxygen? One Possibility.

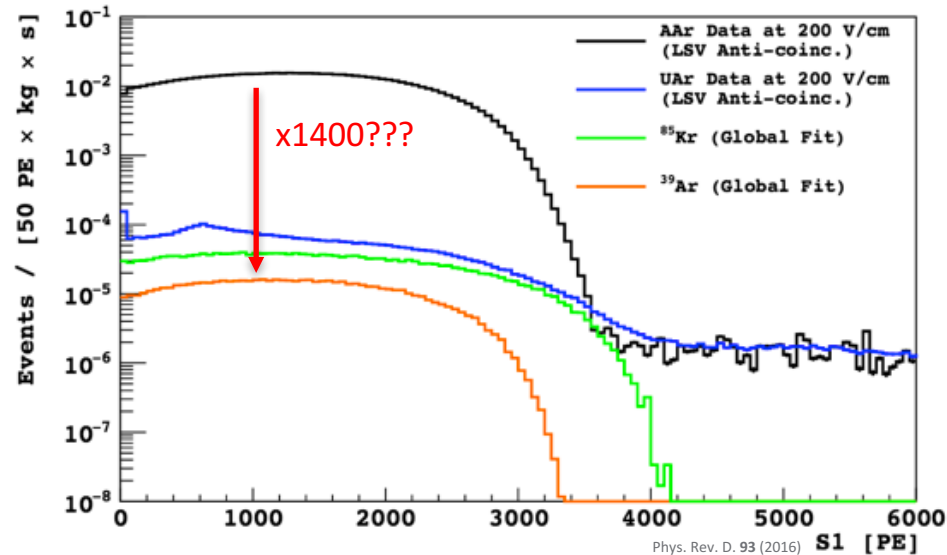
- ▶ Could all the ^{39}Ar have come from an air infiltration?

- ^{39}Ar concentration air equivalent:

$$1.95 \text{ mols}_{\text{Air}}/\text{kg}_{\text{UAr}}$$

- ▶ Air equivalent for O_2 contamination in CO_2

- Assume all O_2 is from air
 - Precision gas analysis of CO_2 [one snapshot]:
 - $\text{O}_2 = 6.7 \text{ ppm}$
 - $\text{Ar} = 427 \text{ ppm}$
 - DS-50 target production history:
 - Always measured $\text{O}_2/\text{Ar} \sim 1.5\text{-}2\%$



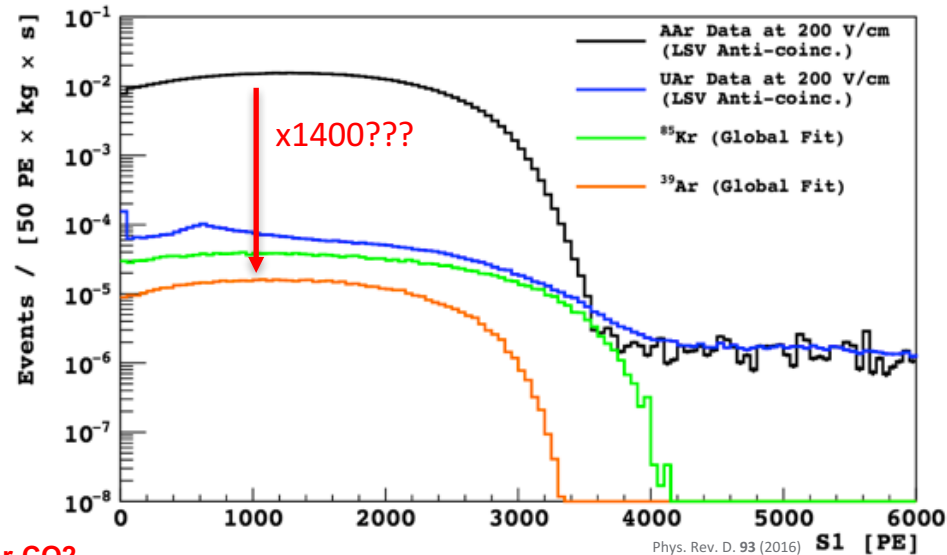
$$6.7 \text{ ppm}_{\text{O}_2} / 427 \text{ ppm}_{\text{UAr}} \rightarrow 0.392 \text{ mols}_{\text{O}_2}/\text{kg}_{\text{UAr}} \rightarrow 1.87 \text{ mols}_{\text{Air}}/\text{kg}_{\text{UAr}}$$

Why Did We See Oxygen? One Possibility.

- ^{39}Ar concentration air equivalent: **$1.95 \text{ mols}_{\text{Air}}/\text{kg}_{\text{UAr}}$**
- O2 Contamination in Argon: **$1.87 \text{ mols}_{\text{Air}}/\text{kg}_{\text{UAr}}$**

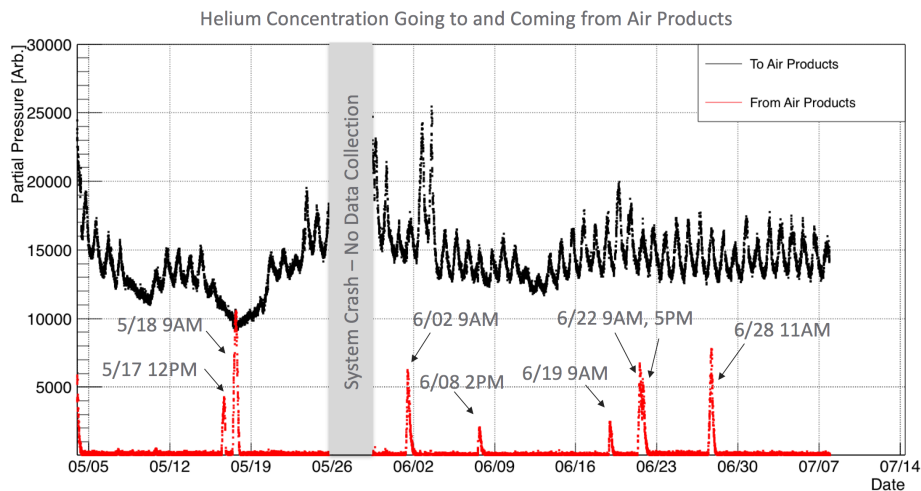
Air equivalent for ^{85}Kr contamination

- ▶ Assume all ^{85}Kr is from air
 - ^{85}Kr concentration in air = $1.3 \text{ Bq}/\text{m}^3_{\text{air}}$
 - $2.05 \text{ mBq}/\text{kg}_{\text{UAr}} \rightarrow 0.070 \text{ mols}_{\text{Air}}/\text{kg}_{\text{UAr}}$
- ▶ Assume ^{85}Kr can absorb on zeolite ~ like N_2
 - N_2 reduction factor through VPSA = 35
- ▶ $0.070 \text{ mols}_{\text{Air}}/\text{kg}_{\text{UAr-DS50}} \times 35 = 2.45 \text{ mols}_{\text{Air}}/\text{kg}_{\text{UAr-CO2}}$
 - Very rough, NOT proof of contamination



LOGAN: The LOnG-term Gas ANalyzer

- ▶ Based on an SRS UGA
 - Great for qualitative, or if well calibrated, quantitative, measurements
- ▶ Ruggedized
- ▶ Custom control software
- ▶ Autonomous



UAr production – Onwards and Upwards

