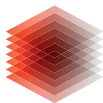


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Compatibility of DAMA/LIBRA dark matter detection with other searches

C. Savage^a, G. Gelmini^b, P. Gondolo^a and K. Freese^d

Published 14 April 2009 • *Journal of Cosmology and Astroparticle Physics*, Volume 2009, April 2009

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Abstract

The DAMA/NaI and DAMA/LIBRA annual modulation data, which may be interpreted as a signal for the existence of weakly interacting dark matter (WIMPs) in our galactic halo, are examined in light of null results from other experiments: CDMS, XENON10, CRESST I, CoGeNT, TEXONO, and Super-Kamiokande (SuperK). We use the energy spectrum of the combined DAMA modulation data given in 36 bins, and include the effect of channeling. Several statistical tools are implemented in our study: likelihood ratio with a global fit and with raster scans in the WIMP mass and goodness-of-fit (g.o.f.). These approaches allow us to differentiate between the preferred (global best fit) and allowed (g.o.f.) parameter regions. It is hard to find WIMP masses and couplings consistent with all existing data sets; the surviving regions of parameter space are found here. For spin-independent (SI) interactions, the best fit DAMA regions are ruled out to the 3 σ C.L., even with channeling taken into account. However, for WIMP masses of ~ 8 GeV some parameters outside these regions still yield a moderately reasonable fit to the DAMA data and are compatible with all 90% C.L. upper limits from negative searches, when channeling is included. For spin-dependent

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compatible. For SD neutron-only couplings we find no parameters compatible with all the experiments. Mixed SD couplings are examined: e.g. ~ 8 GeV mass WIMPs with $a_0 = \pm a_p$ are found to be consistent with all experiments. In short, there are surviving regions at low mass for both SI and SD interactions; if indirect detection limits are relaxed, some SD proton-only couplings at high masses also survive.

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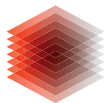
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Autor: Savage, Christopher [Link], Treves, Katherine [Link], Gondola, Paolo [Link], Spruyt, Douglas [Link]
Inhalt: The DAMANAI and DAMALIBRA astiral modulation data, which may be interpreted as a signal for the existence of weakly interacting dark matter (WIMP) in our galactic halo, are re-examined in light of new measurements of the local velocity relative to the galactic halo. In the vicinity of the Sun, the velocity of the Galactic disk has been estimat...
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Inhalt: The DAMANAI and DAMALIBRA astiral modulation data, which may be interpreted as a signal for the existence of weakly interacting dark matter (WIMP) in our galactic halo, are examined in light of null results from other experiments. We use the energy spectrum of the combined DAMA modulation data given in 36 bins, and include the effect of channe...
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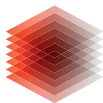
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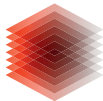
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C. Savage^a, G. Gelmini^b, P. Gondolo^c and K. Freese^d

Published 14 April 2009 • *Journal of Cosmology and Astroparticle Physics*, Volume 2009, April 2009



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Fred C. Adams, J. Richard Bond, Katherine Freese, Joshua A. Frieman, and Angela V. Olinto
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We discuss the particle physics basis for models of natural inflation with pseudo Nambu-Goldstone bosons and study the consequences for large-scale structure of the nonscale-invariant density fluctuation spectra that arise in natural inflation and other models. A pseudo Nambu-Goldstone boson, with a potential of the form $V(\varphi) = A^4 \left[1 \pm \cos\left(\frac{\varphi}{f}\right) \right]$, can naturally give rise to an epoch of inflation in the early Universe, if $f \sim M_{\text{Pl}}$ and $A \sim M_{\text{GUT}}$. Such mass scales arise in particle physics models with a gauge group that becomes strongly interacting at the grand unified theory scale. We work out a specific particle physics example based on the multiple gaugino condensation scenario in superstring theory.

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Katherine Freese, Joshua Frieman, and Andrew Gould
Phys. Rev. D **37**, 3388 – Published 15 June 1988

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If weakly interacting massive particles (WIMP's) are the dark matter in the galactic halo, they may be detected in low-background ionization detectors now operating or with low-temperature devices under development. In detecting WIMP's of low mass or WIMP's with spin-dependent nuclear interactions (e.g., photinos), a principal technical difficulty appears to be achieving very low thresholds (\lesssim keV) in large (\sim km³) detectors with low background noise. We present an analytic treatment of WIMP detection

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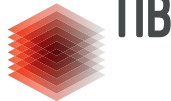
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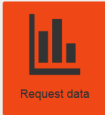



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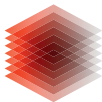
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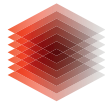
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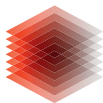


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- Lizenzinformationen wichtig

7: Aufgaben



- grundsätzlich: Vielzahl an Instrumenten, kann weiter zunehmen
- auch eigene, reproduzierbare Auswertungen möglich
- aber: Datenquellen nicht entsprechend vielfältig, häufig:
 - BASE als Aggregator
 - eigene Aggregation geharvesteter Quellen (rudimentäre Schnittstellen und Metadaten)
- Herausforderungen
 - Skills
 - Metadaten

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