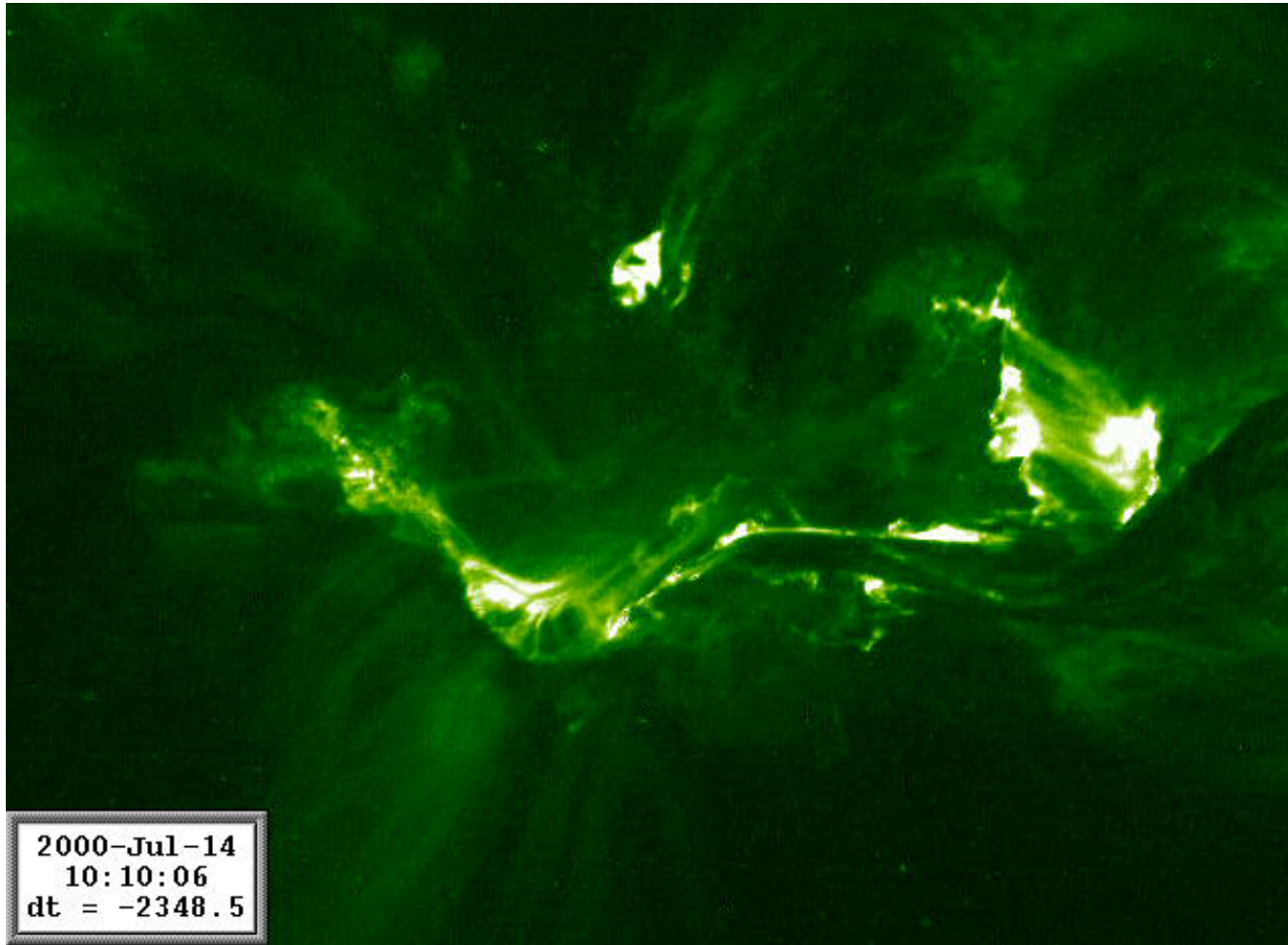


# Searching for Signs of Termination Shocks in Solar Flares

**Giselle Galan** - Harold Washington Community College

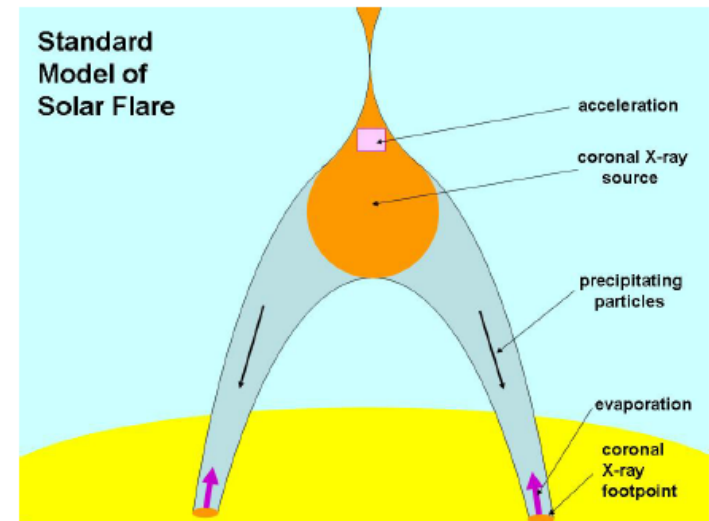
**Vanessa Polito** - Harvard-Smithsonian Center for Astrophysics

**Katharine Reeves** - Harvard-Smithsonian Center for Astrophysics



# Solar Flares

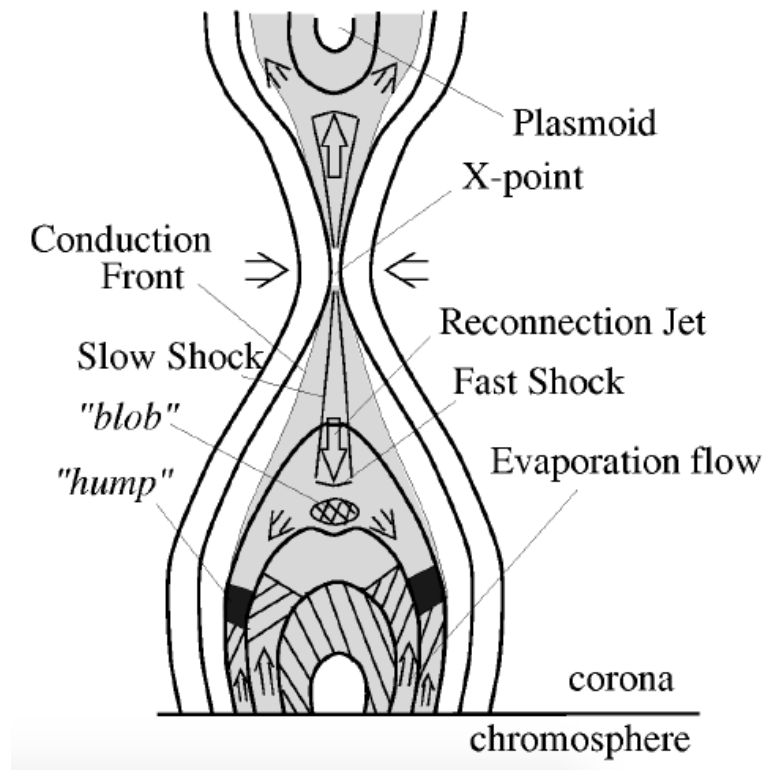
- One of the most energetic events in the sun
- Energy release via *magnetic reconnection*
  - Magnetic → thermal, kinetic
- Energy transport
  - Radiation
  - Mass ejection
  - Shock formation
  - Wave propagation
  - Particle acceleration



*Benz, Living Review in Solar Physics (2008)*

# Reconnection and MHD shocks

- Reconnection site
  - Above loop tops
- Shocks, inflows, reconnection outflows
  - Magnetic tension snaps newly reconnected field lines away from reconnection point
  - Slow-mode MHD shock heating
    - Cool coronal Inflows → hot outflows, i.e. reconnection jets
  - **Termination shock (TS)**
    - Fast-mode MHD shock
    - Supersonic reconnection jet collides with downstream plasma at loop tops



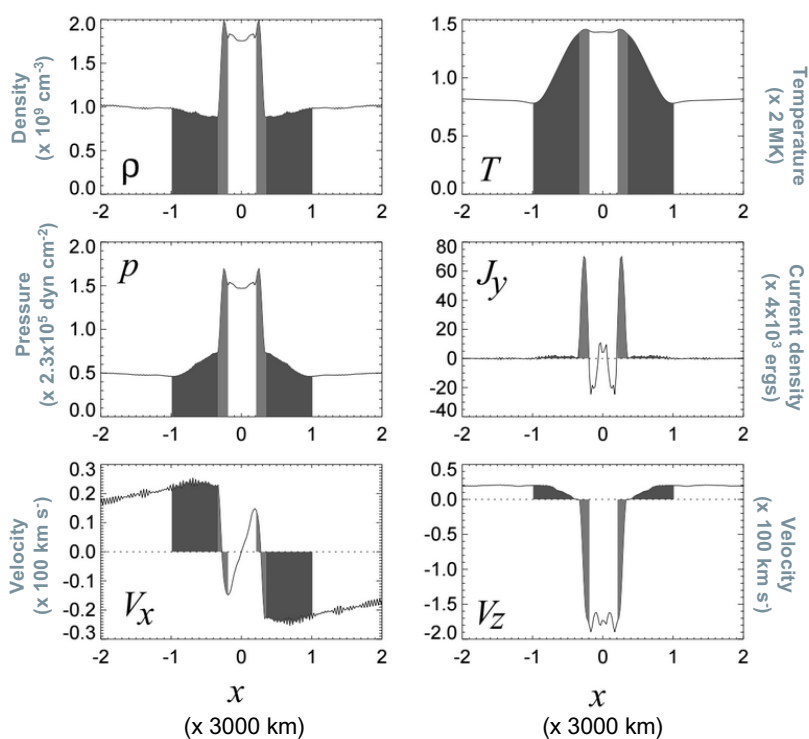
Yokoyama and Shibata, ApJ (2008)



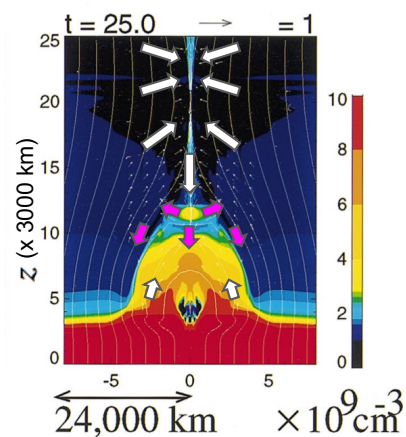
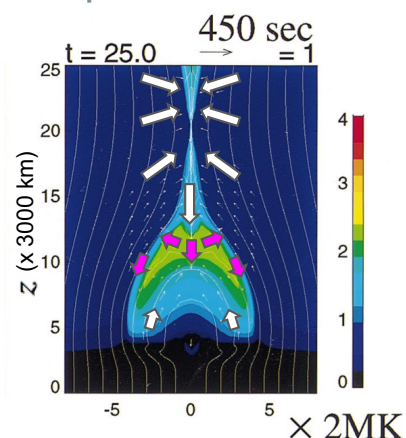
# Signs of a TS

1-dimensional plots along the x-axis  
at  $z=15$ ,  $t = 25$

MHD SIMULATION OF SOLAR FLARE



Temperature distribution

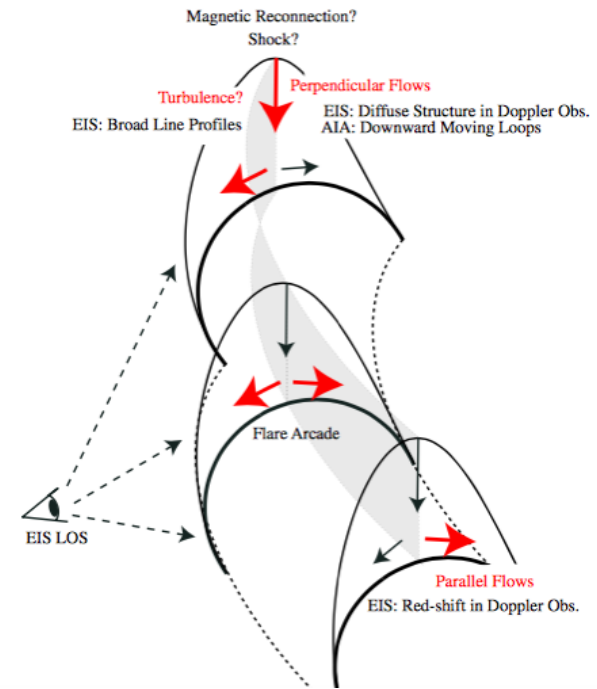
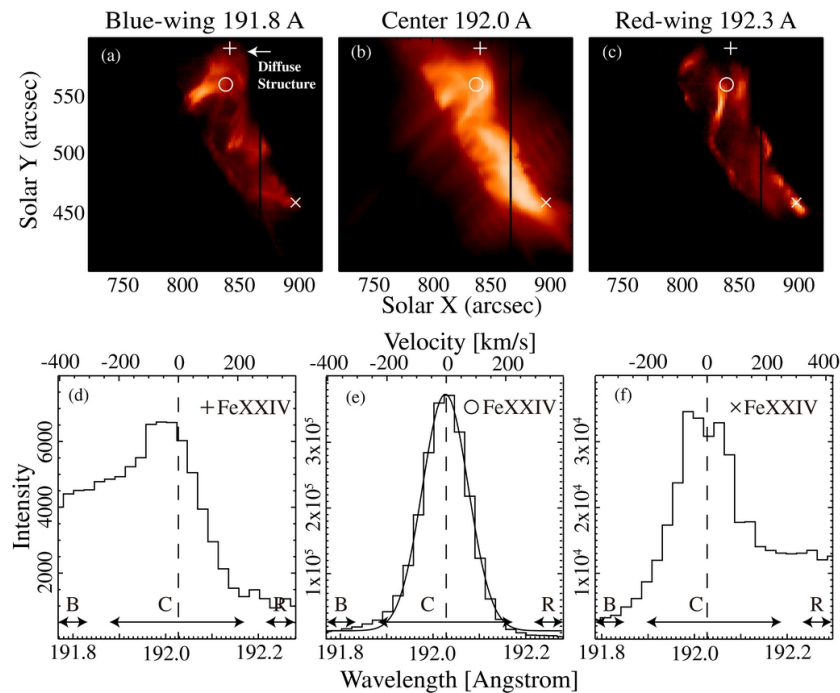


Density distribution

- Speed discontinuity of the fast downward jet accompanied by bidirectional flows away from TS site
- Formation of bright, dense blob
  - Compressed reconnection jet plasma due to shock
  - Slow downward velocity

# Termination shock - Previous related findings

- Fast downward-moving structure above LT
- Bidirectional horizontal flows on LT



*Imada et. al, ApJ, 2013*

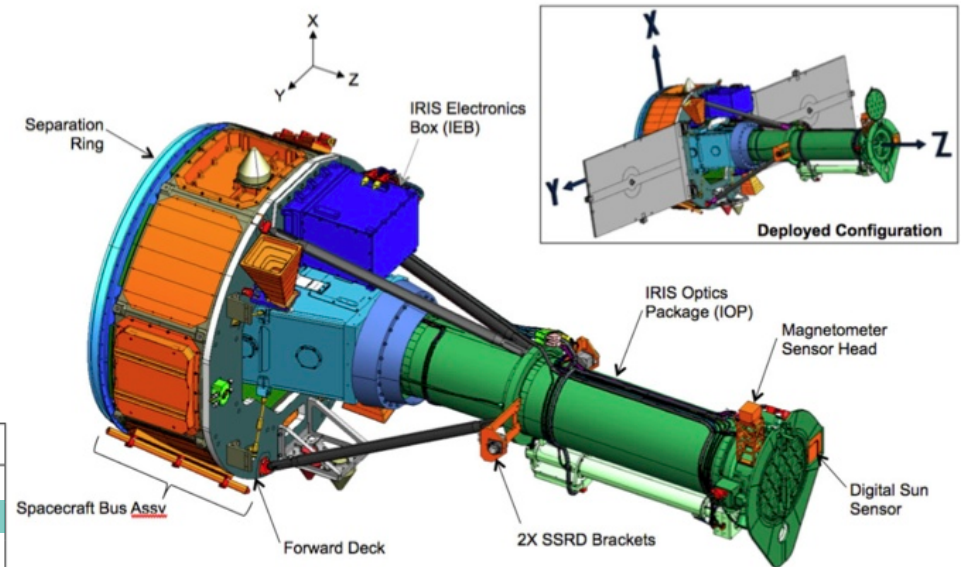
# IRIS - The Interface Region Imaging Spectrograph

- UV Spectrograph

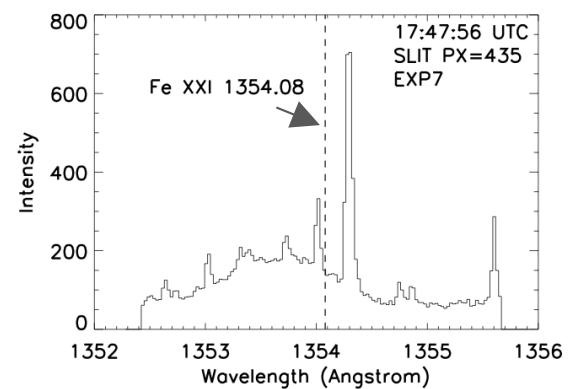
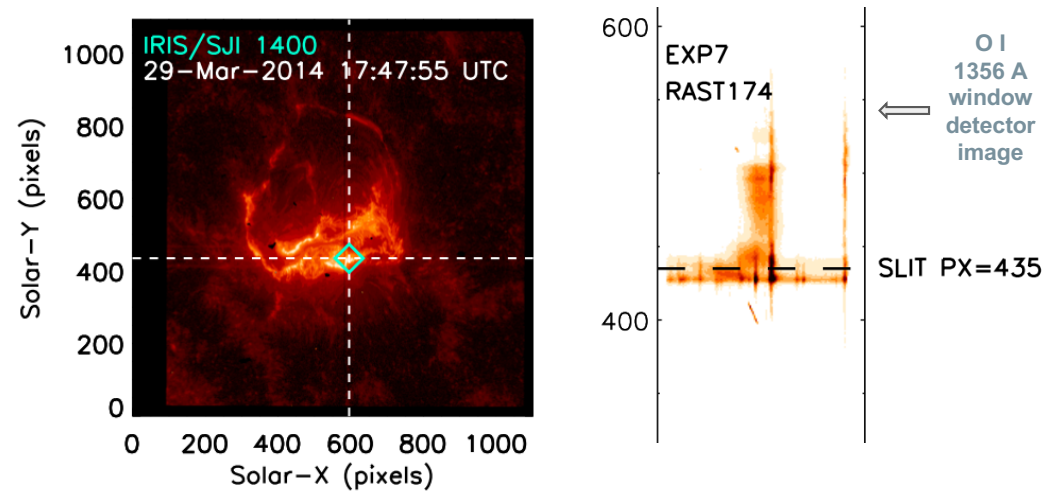
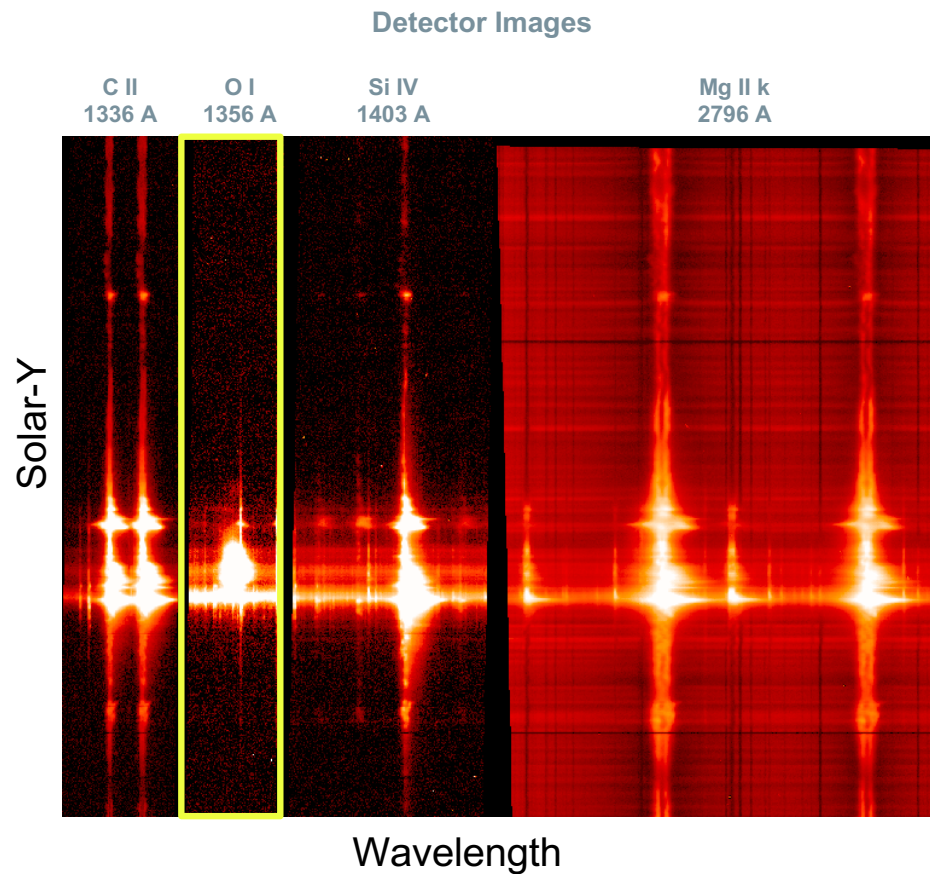
Band	Wavelength (Å)	Temperature (log T)
FUV 1	1331.7–1358.4	3.7–7.0
FUV 2	1389.0–1407.0	3.7–5.2
NUV	2782.7–2851.1	3.7–4.2

- Slit-jaw imager

SJI Passband	Type	Wavelength (Å)	FWHM (Å)	log T
Glass	T	5000	2000	–
C II	M	1330	40	3.7–7.0
Si IV	M	1400	40	3.7–5.2
Mg II h/k	T	2796	4	3.7–4.2
Mg II wing	T	2832	4	3.7–3.8
Broad-band	M	1600	400	–

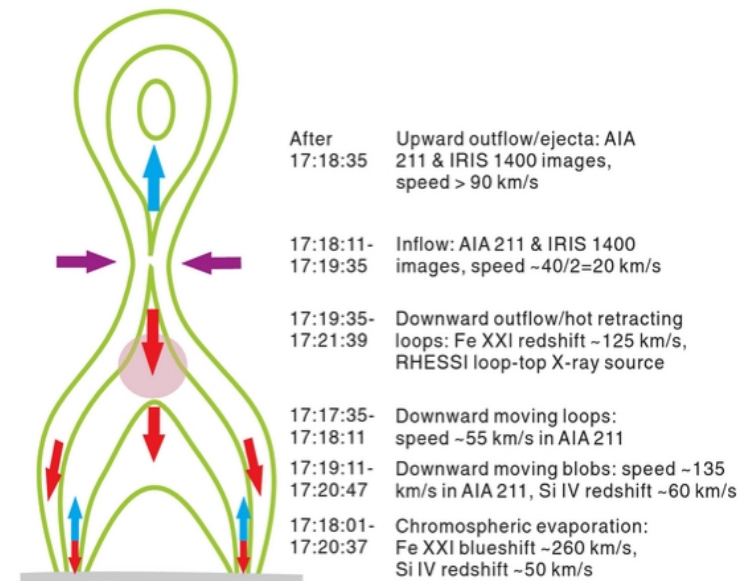
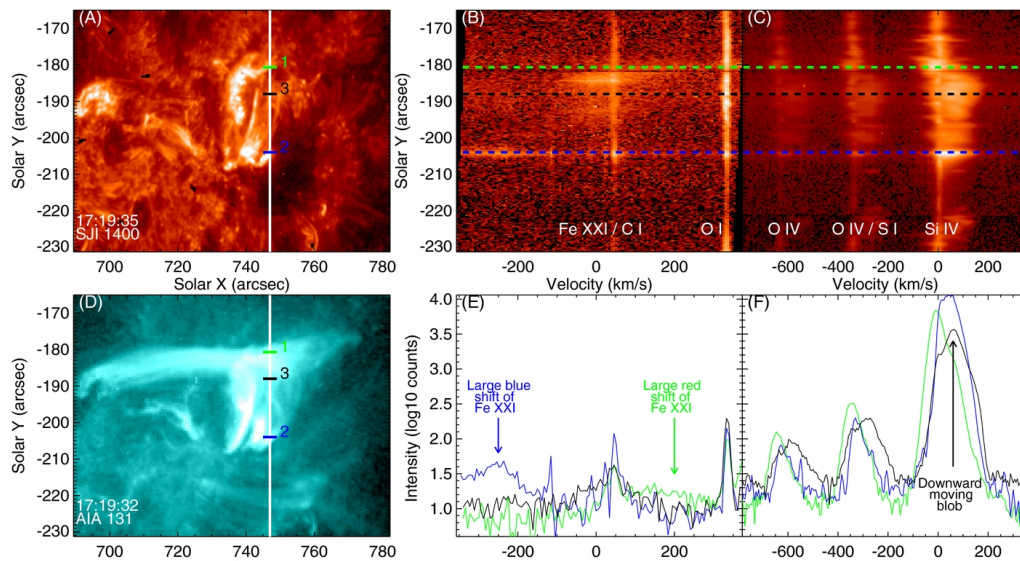


# IRIS - The Interface Region Imaging Spectrograph



# Termination shock - Previous related findings

- Fast downward outflow on LTs,  $v \sim 125$  km/s
- Downward moving blob along loops,  $v \sim 135$  km/s



Images from Tian et. al, ApJ, 2014

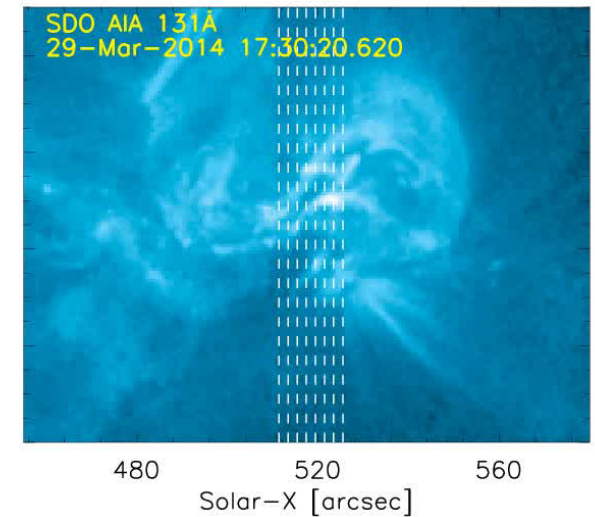
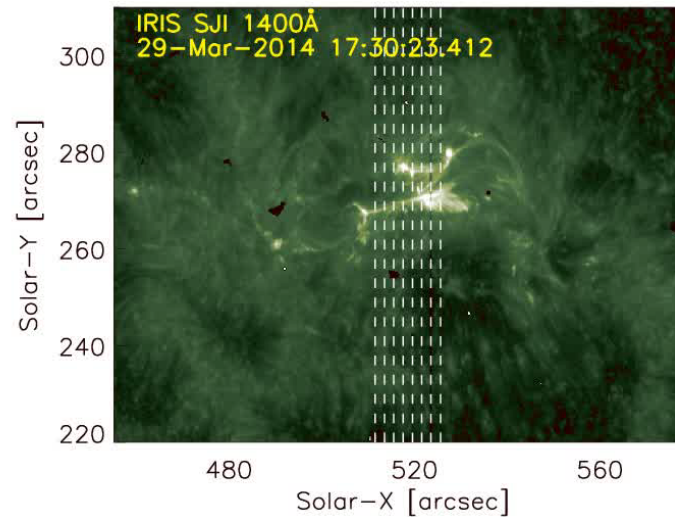


# Methodology

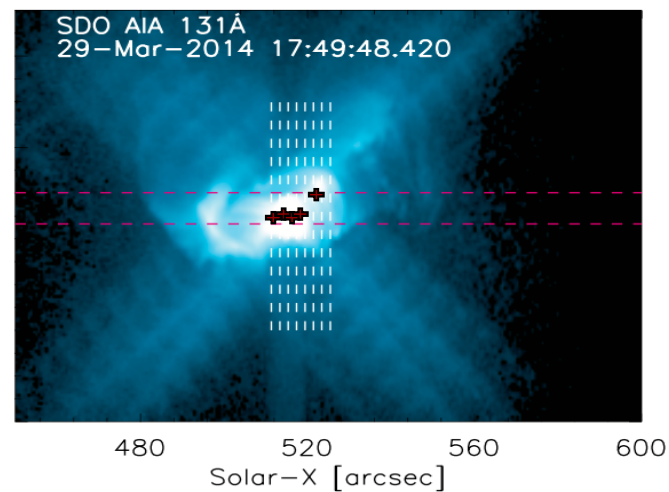
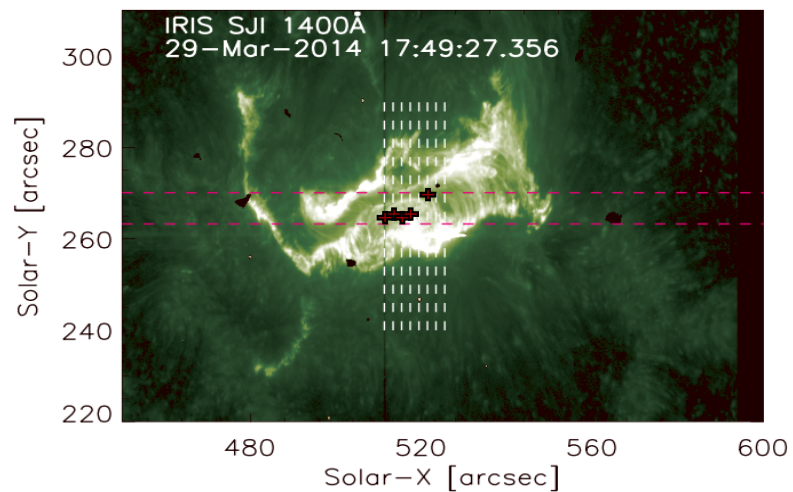
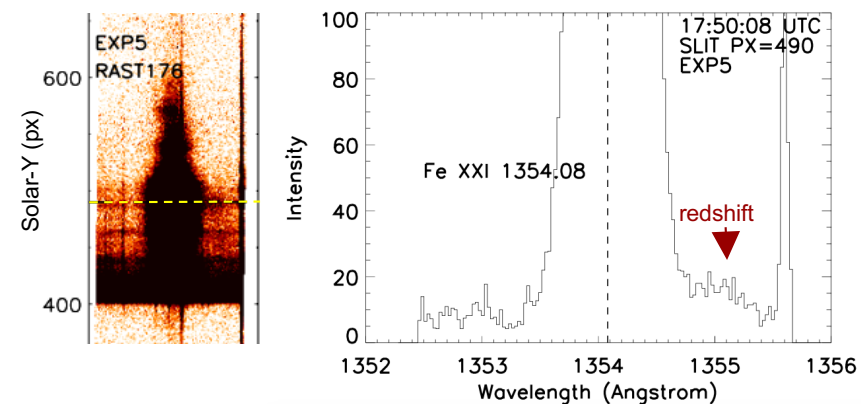
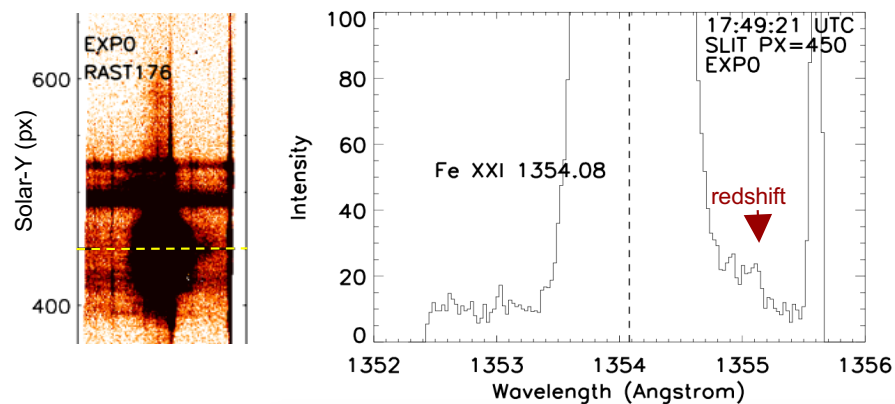
- Analysis of IRIS UV imaging and spectral data
  - Spectral observations
    - Looked for redshifted Fe XXI (1354.08 Å) emissions in LT region
      - Reconnection downflows, plasma motions away from TS site
    - Formation of Fe XXI line only during flare, plasma  $T = > \sim 10$  MK plasma
  - Imaging observations
    - Spatial context
    - Coterporal 1330 SJI imaging
    - Co-alignment with AIA imaging observations
      - AIA 131 Å passband - dominated by Fe XXI emissions during flares

# Results: 2014-03-29 Flare (OBS 3860258481)

- 180 rasters
- Very large coarse 8-step raster
  - Step = 2''
- Exposure time: ~8s



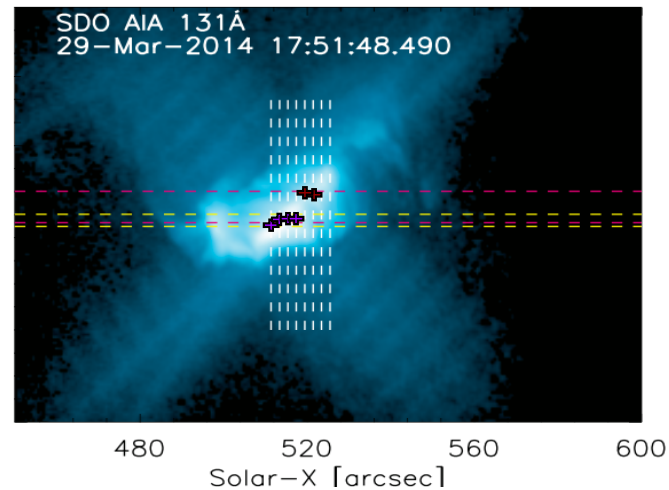
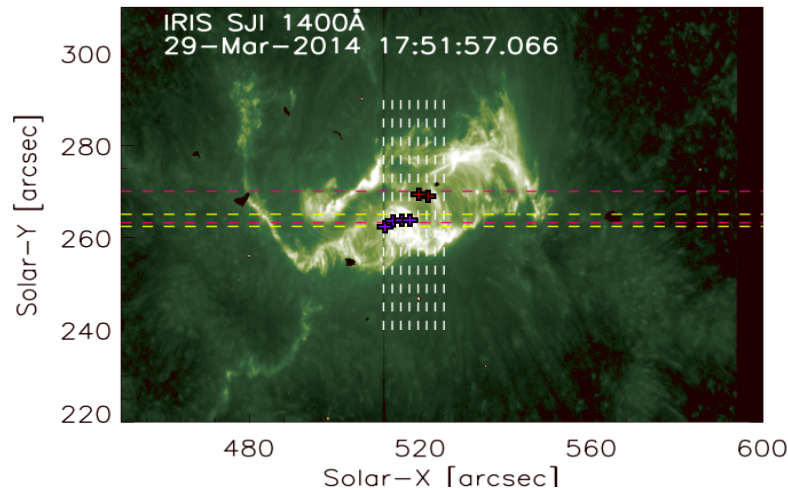
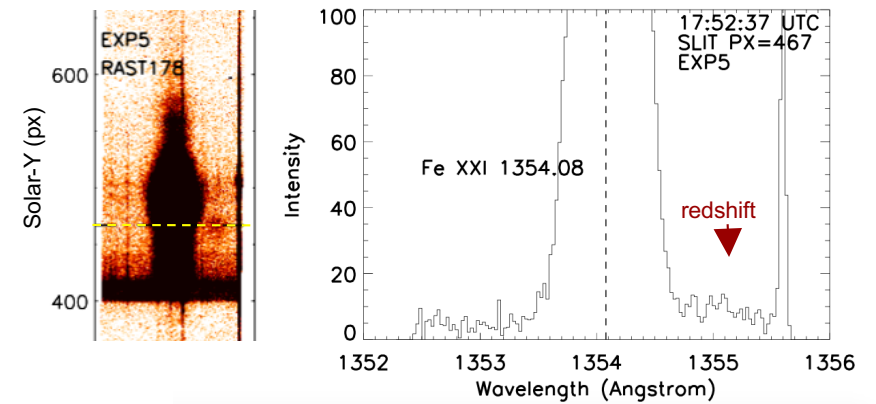
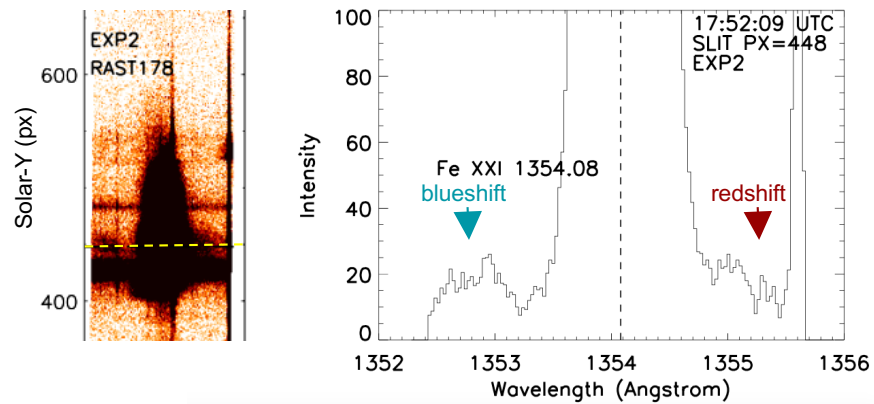
# Results: 2014-03-29 Flare (OBS 3860258481)



⊕ = redshift only

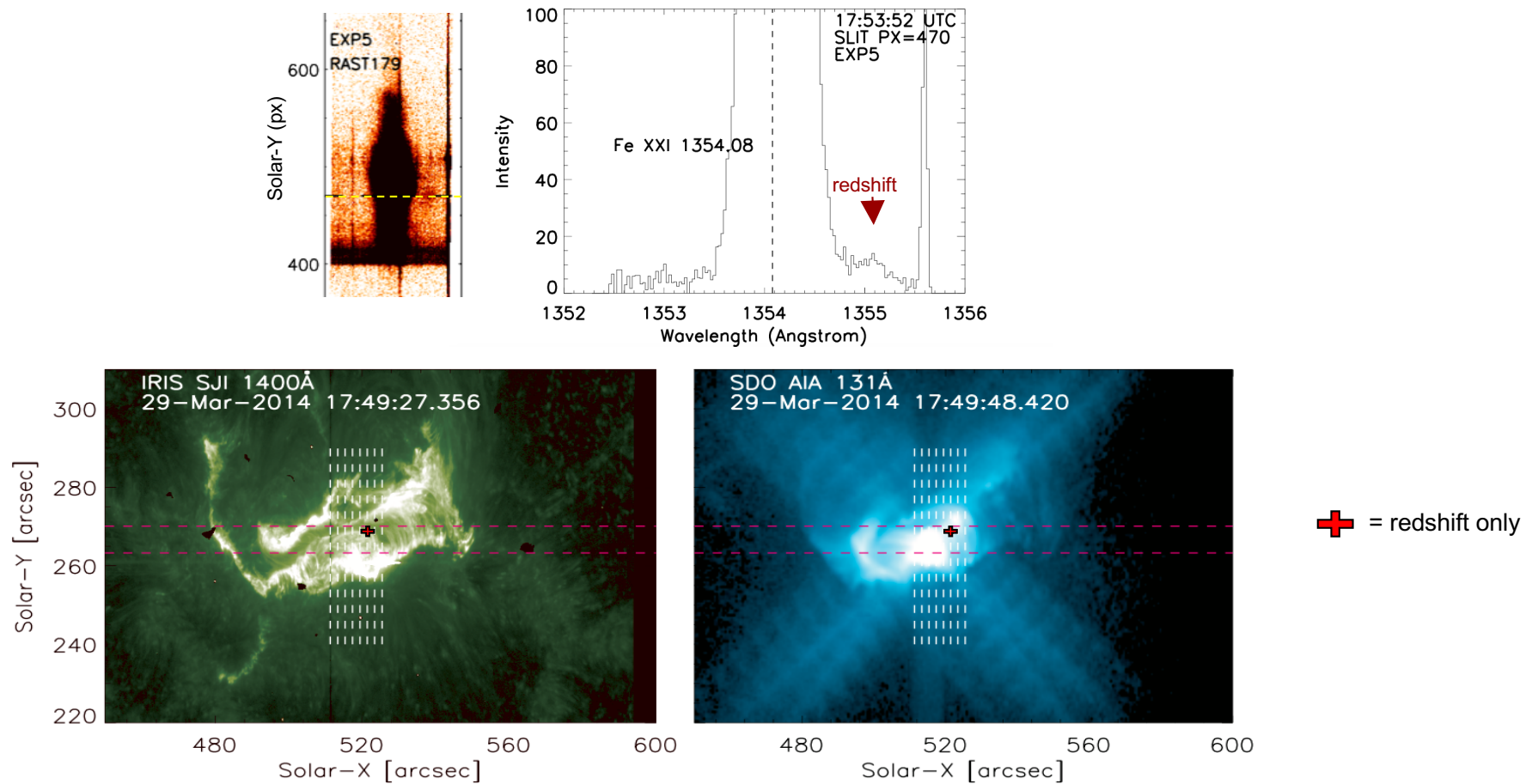


# Results: 2014-03-29 Flare (OBS 3860258481)



⊕ = redshift only  
⊕ = redshift and blueshift

# Results: 2014-03-29 Flare (OBS 3860258481)



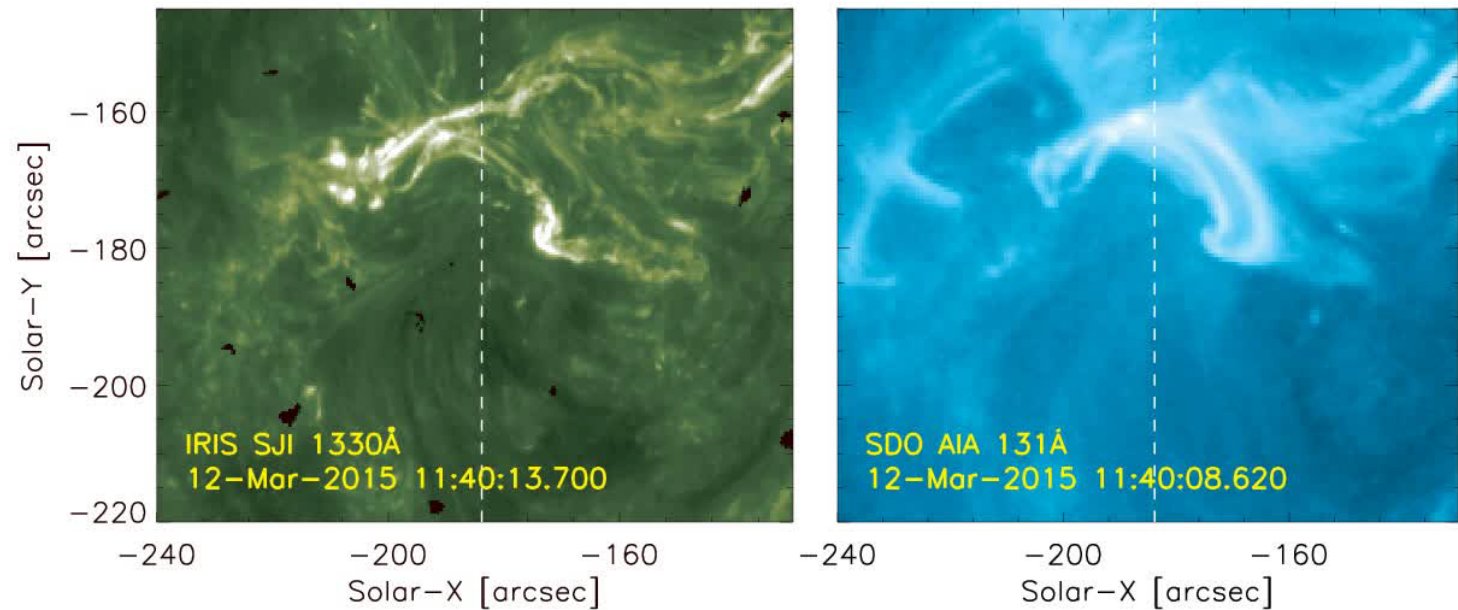
# Results: 2014-03-29 Flare

- Summary

- Redshifts on bright loop top bar,  $v = \sim 100\text{-}200 \text{ km/s}$ 
  - Possible downward flows along the LOS on the bright loop top bar region
- Cospatial blueshifts and redshifts,  $v = \sim 100\text{-}200 \text{ km/s}$ 
  - Redshifts: Possible downward flows along the LOS on bright loop bar region
  - Blueshifts: Chromospheric evaporation from ribbon

## Results: 2015-03-12 Flare (OBS 3860107053)

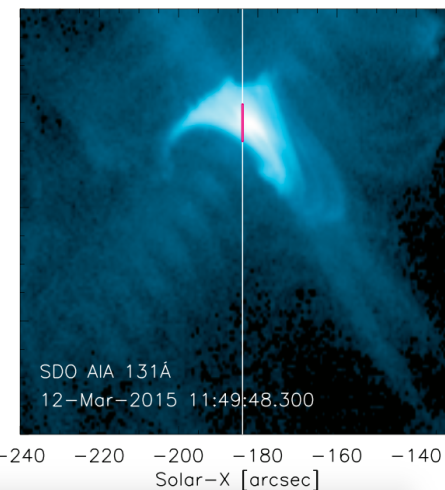
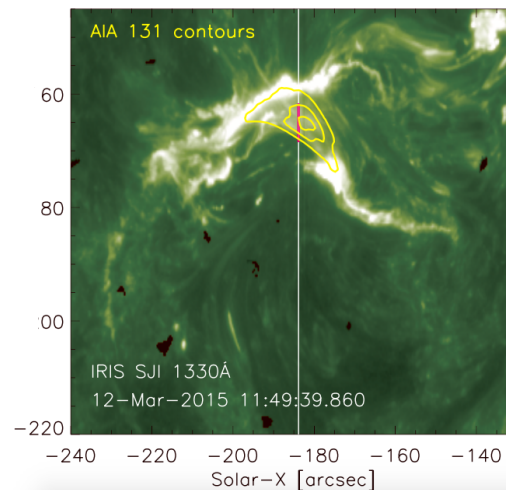
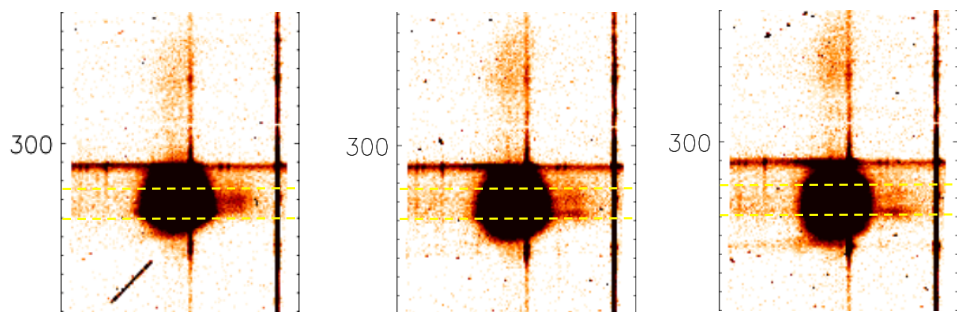
- 11:48:11-12:00:40 UTC
- Large sit-and-stare
- Exposure time:  $\sim 5$ s



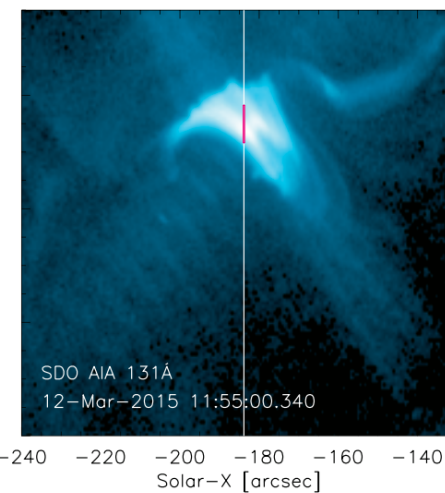
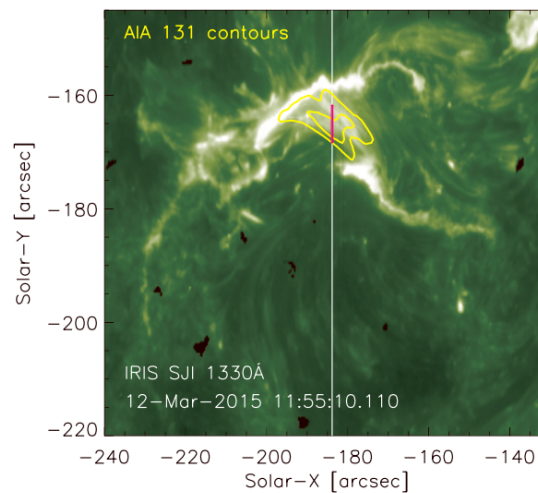
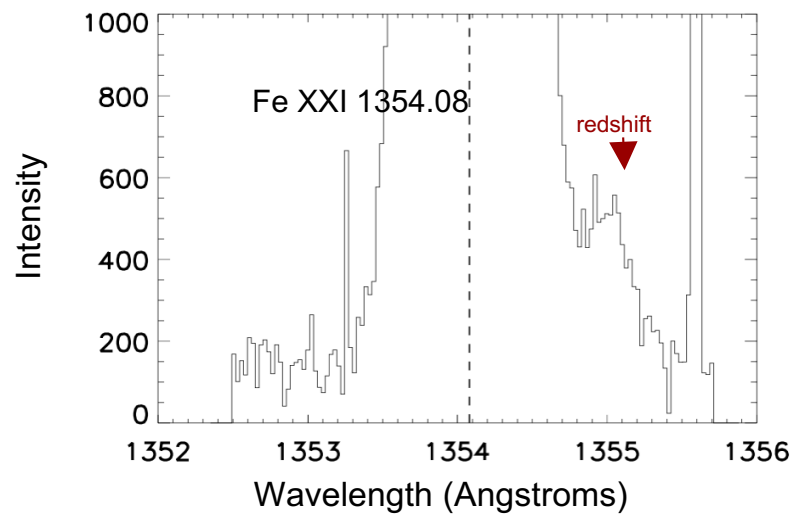
11:50:16-11:52:17

11:52:22-11:54:22

11:54:28-11:56:28



Total intensity over  $y = 260-265$ px  
from 11:50:16-11:52:17



# Results: 2015-03-12 Flare (OBS 3860107053)

- Summary

- Redshifts on bright loop top bar region,  $v = \sim 100\text{-}200$  km/s
  - Possible downward flows along the LOS on the bright loop top bar region

# Conclusion

- Cannot conclusively determine that termination shocks occurred
- Identified flows indicative of a TS were identified in the form of redshifted Fe XXI emissions in two separate flares
- Similarities
  - Location: bright loop top bar
  - Comparable speeds; and emission intensities and widths
  - Appeared in gradual phase, right after the flare peak
  - Redshifts persisted for ~5 and ~10 minutes, respectively
- Further investigation
  - Redshifts appeared at later time (gradual phase) than expected (impulsive phase)



# Acknowledgements

*Thank you to my mentors **Vanessa Polito** and **Katherine Reeves** for their guidance and patience. Thank you also to the **NSF-REU** solar physics program at SAO, grant number AGS-1560313.*