

European  
Innovation  
Council



UK Research  
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C L O U D

**experiment's first release**

**“Neutrino Telescopes” Conference**

25 October 2023 — Venezia (Italia)

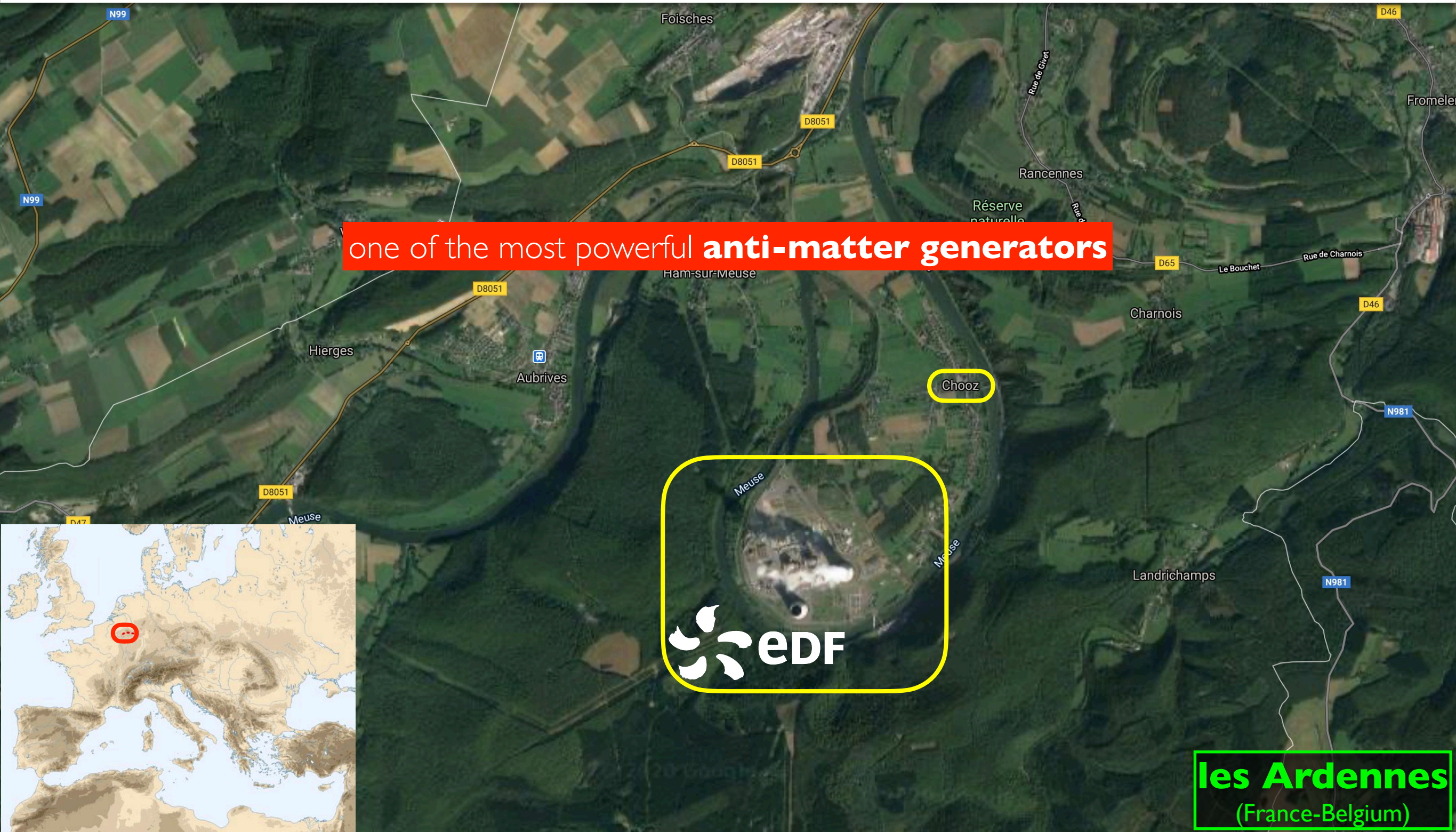
**Anatael Cabrera**

IJCLab / LNCA - Université Paris-Saclay / CNRS  
Orsay, France

(co)spokesperson:  
• DoubleChooz  
• LiquidO  
• CLOUD — AM-OTech (EIC)  
• SuperChooz Pathfinder



in the **middle of central Europe** (between France-Belgium): **Chooz** [meeting point with Germany, Luxembourg, Netherlands]

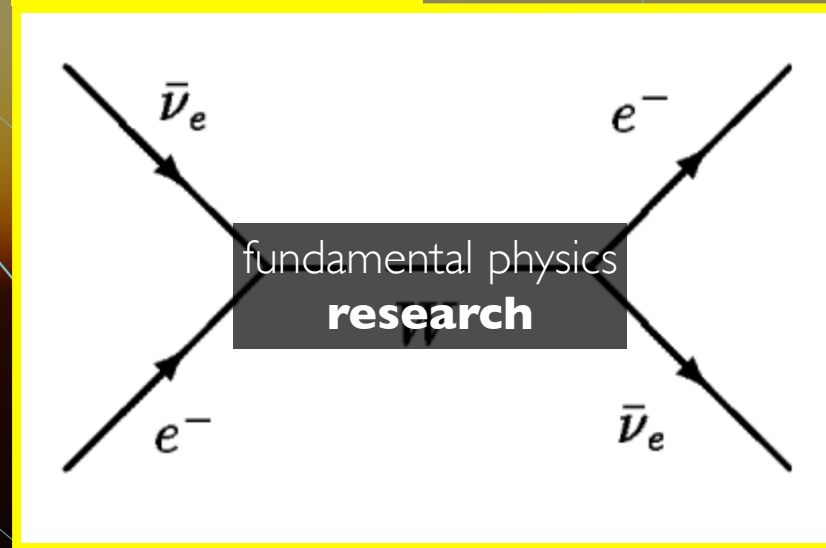


Europe's most powerful reactor site...

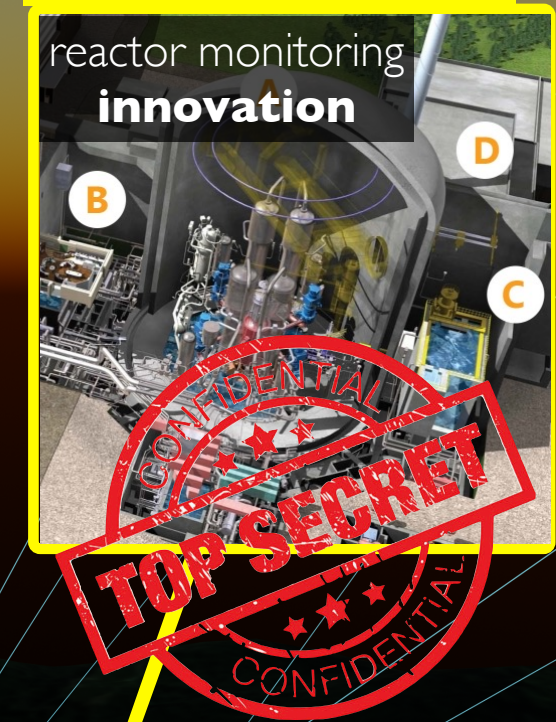
3<sup>rd</sup> generation of reactor neutrino experiments @ Chooz



## CLOUD's view...



## AM-OTech's view...



antineutrino (neutrino?):  $\sim 10^{21}$  v/s per core

**Chooz-B:** Nuclear Reactor Cores

## experimental setup...

- Detector Mass:  $\sim [5, 10]$  ton — **LiquidO** technology
- Overburden:  $\sim 3$  mwe
- Baseline:  $\geq 30$  m (**Ultra Near Detector** site @ Chooz)
- Rate:  $\sim 25,000$  anti- $\nu$  per day —  $\sim 10M$  anti- $\nu$  per year

# CLOUD vs AntiMatter-OTech...



today, **most experiments** bypass (whenever possible) the **absolute flux knowledge** — complex!  
**relative knowledge** (ex. multi-detector, etc.) well suited to **extract “known model” parameters**



confront **absolute flux knowledge** for **new neutrino physics** via “flux bias explorations”

- **extreme signal** (to BG) rates — unprecedented
- *must*: the **best-known cross-section**(s) today ( $\leq 1\%$ )
- **extreme energy control** ( $\leq 1\%$ ) — avoid spectral distortions ( $\rightarrow$  flux biases)
- **much redundancy** — as much as Nature kindly allows...



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on the shoulders of giants...



# breakthroughs with reactor antineutrinos...

- [1950s;  $L \approx 10\text{m}$ ] **electron-antineutrino discovery** by **Poltergeist** [Nobel Prize 1995]
- [1980s;  $L \approx 10\text{m}$ ] **Bugey3** (shape) & **Bugey4** (rate): reactor flux understood  $\leq 3\%$ ? [ILL data: **prediction**]
- [1990s;  $L \approx 1\text{km}$ ] **Chooz** & **Palo Verde** absence of oscillation  $\Delta m^2$  — **limit in  $\theta_{13}$** 
  - corroborate **Kamiokande's oscillation  $\nu_\mu \rightarrow \nu_\tau$  dominant transition** [Nobel Prize 2015]
- [2000s;  $L \approx 180\text{km}$ ] **KamLAND** **favoured solar “LMA”** — **SNO** complementary [Nobel Prize 2015]
- [2010s;  $L \approx 1\text{km}$ ] **Daya Bay, Double Chooz, RENO**: observe **predicted  $\theta_{13}$**   $\Rightarrow$  consolidation!!
  - **Double Chooz sub-team**: rate prediction bias  $\rightarrow$  new physics vs prediction? [biased prediction]
  - **Double Chooz et al.**: **spectral distortion** — contradiction to **Bugey3** [what's wrong?]
- [2020s;  $L \approx 50\text{km}$ ] **JUNO** will **measure  $\theta_{12}$ ,  $|\delta m^2|$ ,  $|\Delta m^2|$  to  $\leq 1\%$**  — first “bi-oscillation energy spectrum”
  - **mass ordering** ( $\geq 5\sigma$ ) need synergy with accelerator experiments [**backup**]
- [ $\geq 2030\text{s}$ ;  $L \approx 1\text{km}$ ] **SuperChooz** will **measure  $\theta_{13}$ - $|\Delta m^2|$  &  $\theta_{12}$ - $|\delta m^2|$  to  $\leq 1\%$**  — **exploring more!**

**reactors** leading most of **world's neutrino knowledge**  $\Rightarrow$  **NO new physics (so far)**



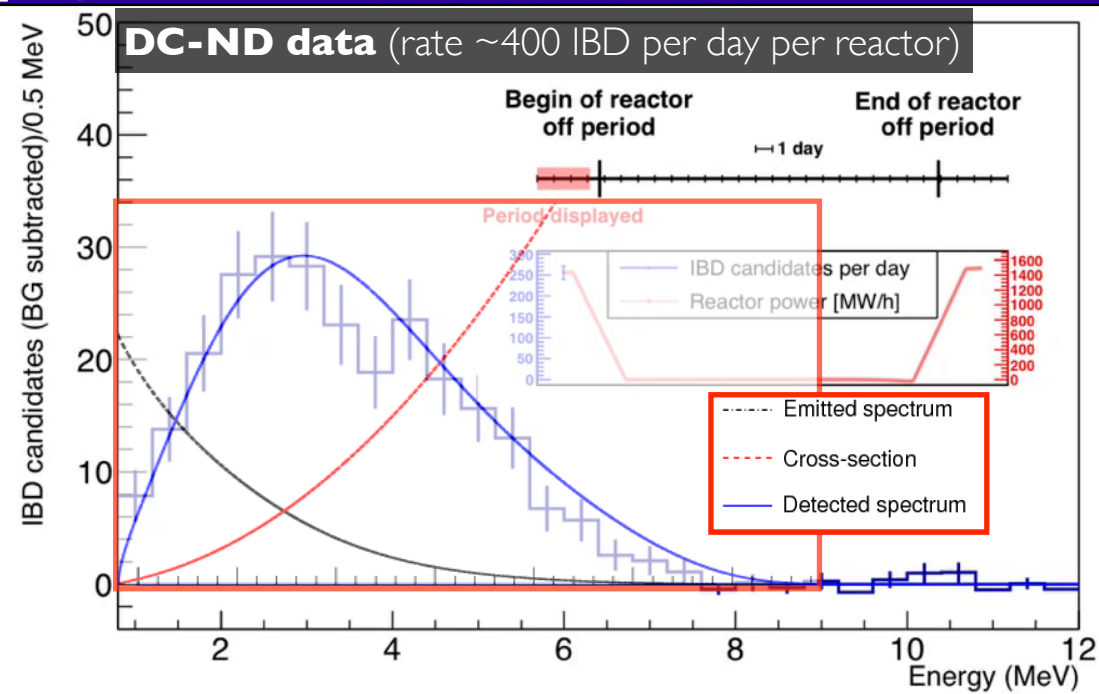
# reactor neutrinos...



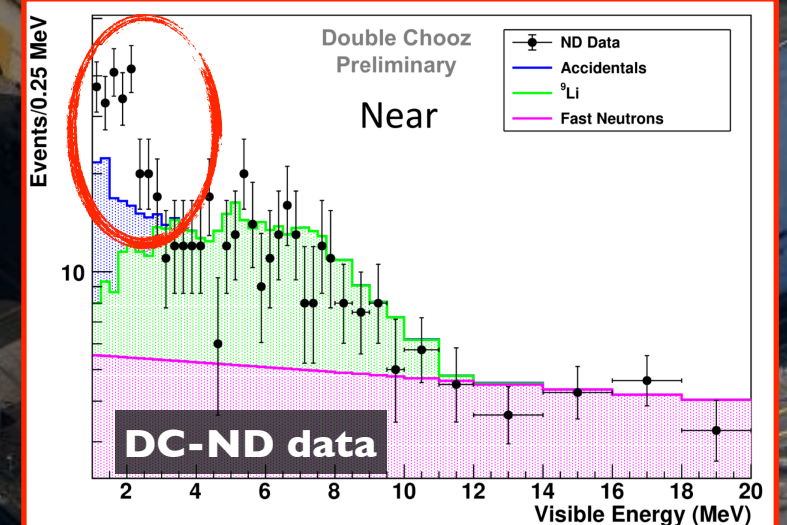
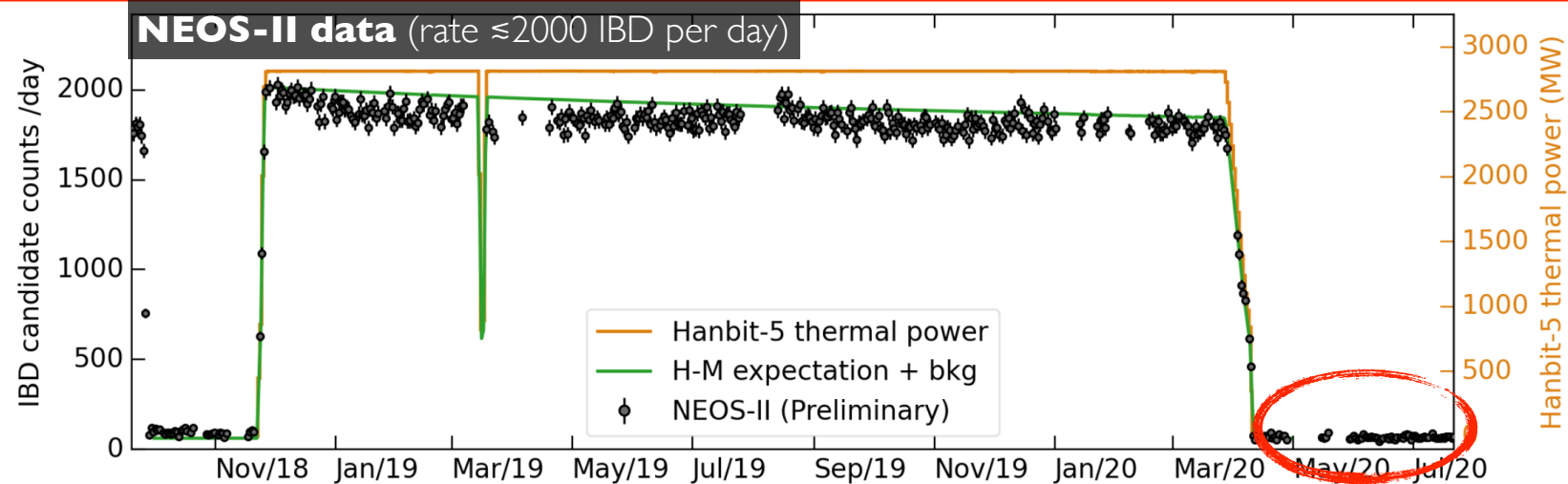
nature  
physics

ARTICLE

First Double Chooz  $\theta_{13}$  Measurement via Total Neutron Capture Detection



**DC-ND overburden  $\sim 30\text{m}$**





# today's reactor neutrino methodology...

**the most powerful source on Earth**

- **interaction: IBD@p** [inverse-beta-decay on proton]
  - $\sigma(\text{IBD@p})$ : CC, high and known to  $\sim 0.2\%$  ( $\rightarrow$  neutron lifetime) with threshold  $\geq 1.8\text{MeV}$
  - no other interactions — few attempts to “**electron elastic scattering**” (past)
- **flux:  $\sim 10^{20}$  antineutrino per second per GW** (thermal)
  - **experimental precision  $\leq 1\%$**  [world's precision by Double Chooz]
  - **prediction precision  $[2\sim 6]\%$**  [ILL-based  $\oplus$  approximations  $\oplus$  bias correction by Kopeikin et al.]
    - **URGENT: new accurate reactor predictions** — how to **ensure reliable precision?**
- **signal (IBD@p) features (typically underground):**
  - **reactor modulation** (up to 100%):  $\text{rate(ON)} / \text{rate(OFF)} \approx 100$  [residual flux during rate(OFF)]
    - high precision **spectral reactor-OFF information** [DoubleChooz]
  - **Signal-to-BG order 10** (GW reactors — commercial) [ $\geq 10\times$  loss with research reactors]
    - BG dominated by cosmogenic  $\rightarrow$  **irreducible** [unless **e+ ID** was possible]
      - **e+ ID** via ortho-positronium [a la Borexino  $\rightarrow$  Double Chooz] — [impractical for reactor physics](#)
    - **monolithic & hermetic detectors** — segmentation limited gain & risk radiogenic-BG issues

reactor neutrinos **experimental methodology largely similar** for the last  **$\sim 70$  years** (Reines et al.)  
**powerful framework** so far, but **good enough for discoveries  $\geq 2025$ ?**



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**WARNING:** from now we should talk about **neutrinos** and **antineutrino**...

reactor (anti)neutrino future... (?)



# future: discoveries?

today's BG: *standard neutrino oscillation*

**neutrino $\oplus$ weak-interaction** remains **bizarre** (Majorana, etc)...

- new neutrino **phenomenology**? [ex. mixing and masses]
- new neutrino **interactions**?
- new neutrino **states**? [assume: “3+1 sterile” is largely ruled out]



# the future reactor neutrino...

- **interaction:** go beyond the **IBD@p** (antineutrino-CC)
  - **precise ES@e** [*elastic scattering on electrons*]: **CC $\oplus$ NC** &  **$\theta_w$  @ 1 MeV** (renormalisation running)
    - combined with **IBC@p**  $\Rightarrow$  isolate **NC-only** component? [à la SNO]
  - **reactor neutrinos** ( $\beta^+/\text{EC}$  at the reactor)? [ $\rightarrow$  the “**missing MeV neutrino source**”?]
  - **IBD@X < 1.8 MeV:** geoneutrino  $^{40}\text{K}$  & (non-intrusive) direct reactor-fuel monitoring (pool, etc.)
- **flux:** measure **all known** & **unknown(s)** possible emissions [**discovery potential**]
  - **$\phi(\text{anti-}\nu_e; \text{CC})$ :** ultimate precision  $\sim 0.5\%$ <sub>thermal-power</sub>  $\Rightarrow$  **unitarity violation?** — **new physics?**
    - **novel reactor predictions methodology?** probe & demonstrate accuracy
  - **$\phi(\nu_e; \text{CC})$ :** first observation ever (**surprises?**) & complementary to  $\phi(\text{anti-}\nu_e; \text{CC})$  prediction
  - **$\phi(\nu_x; \text{NC})$ :** NC validation: **agreement to CC?** [à la SNO] — **new physics?**
- **signal** features:
  - **(IBD@p) S-to-BG  $\geq 100$**  (GW reactors)  $\Rightarrow$  address also low power reactors — **the future?**
    - **empowered coincidences & PID**; namely **topology e+ ID** (but not only)
  - other improvements elsewhere (especially for **ES@e**) — radiopurity, etc.
- **neutrino-based innovation?**  $\rightarrow$  exploring in **AntiMatter-OTech** [**CONFIDENTIAL**]

**worthy challenges  $\Rightarrow$  major breakthrough(s) ahead ( $\geq 2025$ ) & probing new discovery potential**



C L O U D

**a probe to the future?**

the experiment...



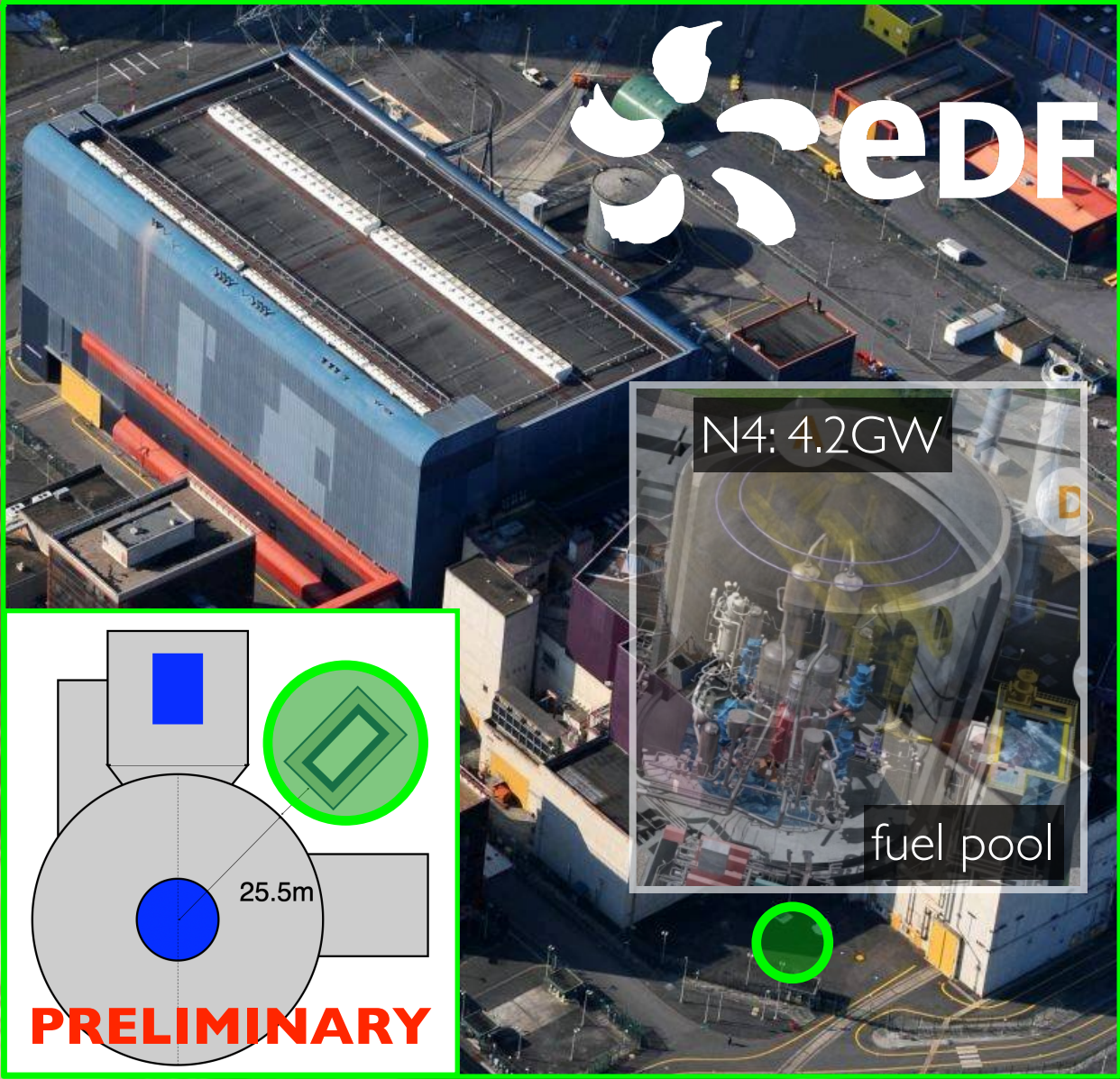
**Chooz-B Power Station**

- facility: EDF CNPE
- location: Chooz (France)
- reactor cores: 2x PWR AREVA-N4
- thermal power: 8.4GW (total)

**Double Chooz**  
Near Detector

**LNCA-Hall** (CNRS)

**Ultra Near Detector (UND) sites**



N4: 4.2GW

fuel pool

25.5m

**PRELIMINARY**

ON

OFF

due to global warm → more frequent reactor-OFF (2022: several months)

**CLOUD** = "Chooz Liquid **U**ltraneur **D**etector"

**Double Chooz**  
Far Detector

# Europe's best reactor-V site...



**Water Pool** [20,40]cm thick

- $4\pi$  shield & neutron moderator
- controllable thermal-bath

**IGLOO** [ $\sim 3$ mwe]

- concrete **bunker** (with boron?)
- DC's iron steel shield (15cm thick)

**Redundant “surface neutron” layers...**

- **IGLOO** (absorption) — passive
- **Water** (moderator $\oplus$ absorption) — passive
- **Armour** (veto $\oplus$ moderator $\oplus$ absorption) — **OD**
- **Tracker** (PID $\oplus$ moderator) — **ID**

cross-section view

side-section view

**ID Physics Volume:** [5,10]tons

**LiquidO-Tracker** (or **inner-detector**) [ $\leq 10$ tons fiducial]

- opaque scintillator — new formulation(s) [more on this soon]
- **$\sim 10,000$  fibres $\oplus$ SiPM** readout channels (GHz waveforms)
- designed light level:  **$\geq 200$ pe/MeV**

**ARMOUR** (or **outer-detector**) [ **$\sim 0.5$ m thickness**]

- transparent scintillator (LAB $\oplus$ PPO $\oplus$ Bis-MSB)
- **$\leq 180$  DC-PMTs** & highly reflecting walls
- designed light yield  **$\geq 400$ pe/MeV**

**experimental setup...**

- Detector Mass:  $\sim [5,10]$  ton — **LiquidO** technology
- Overburden:  $\sim 3$  mwe
- Baseline:  $\geq 30$  m (**Ultra Near Detector** site @ Chooz)
- Rate:  $\sim 25,000$  anti- $\nu$  per day —  $\sim 10$ M anti- $\nu$  per year

# CLOUD detector...



**CLOUD is powered by...**

L I Q U I D

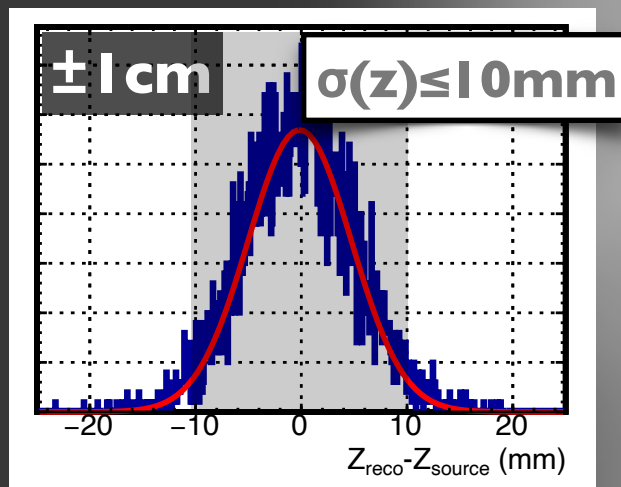
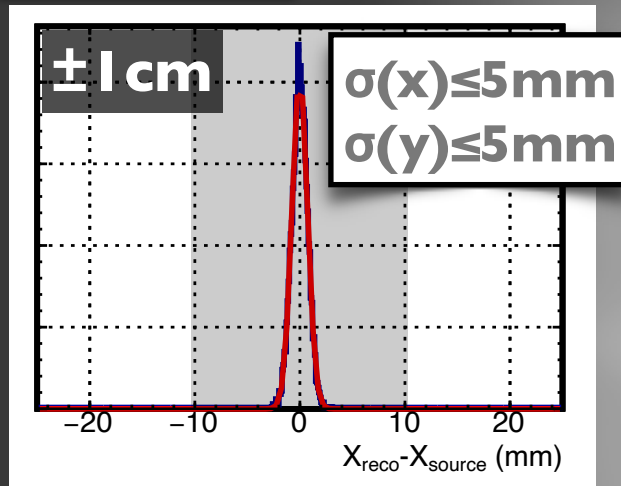




# LiquidO $\leftrightarrow$ stochastic light confinement

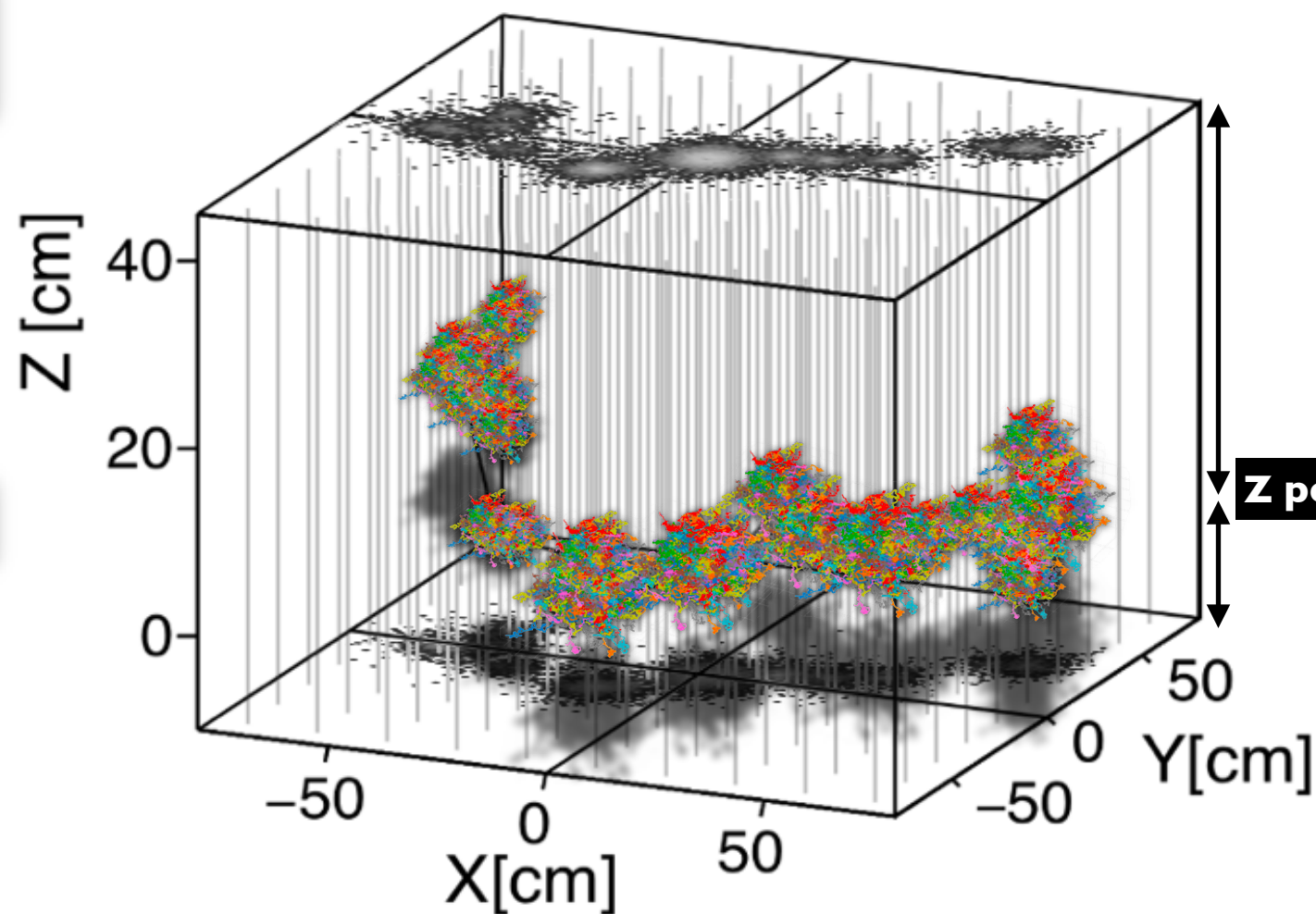
Topology (X,Y) direct & native (PID)  $\rightarrow$  possible sub-mm vertex precision

**$\sim 1.0\text{MeV}$**



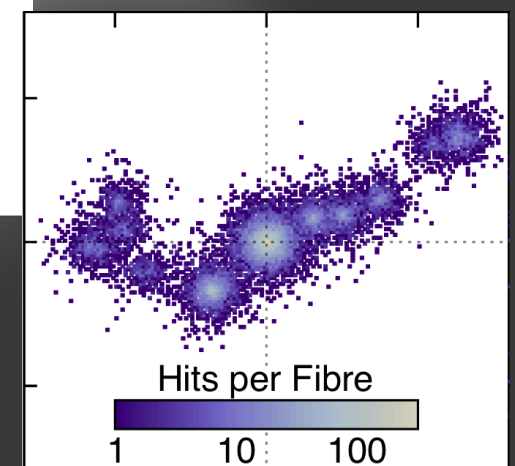
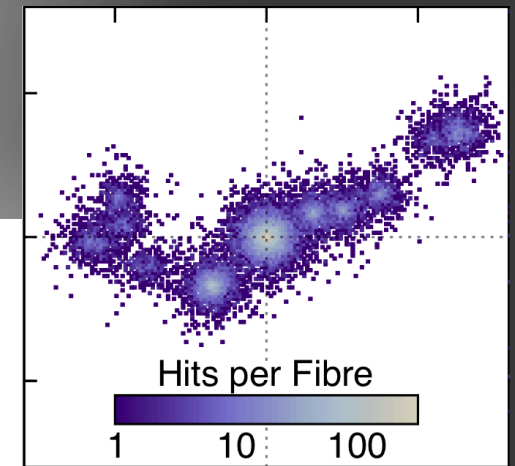
**Vanilla LiquidO: 1D lattice** (fibres along Z-axis only)

**TOP VIEW: (X,Y) Projection  $\rightarrow$  direct readout**



**Z position:  $\Delta t$  (time difference)**

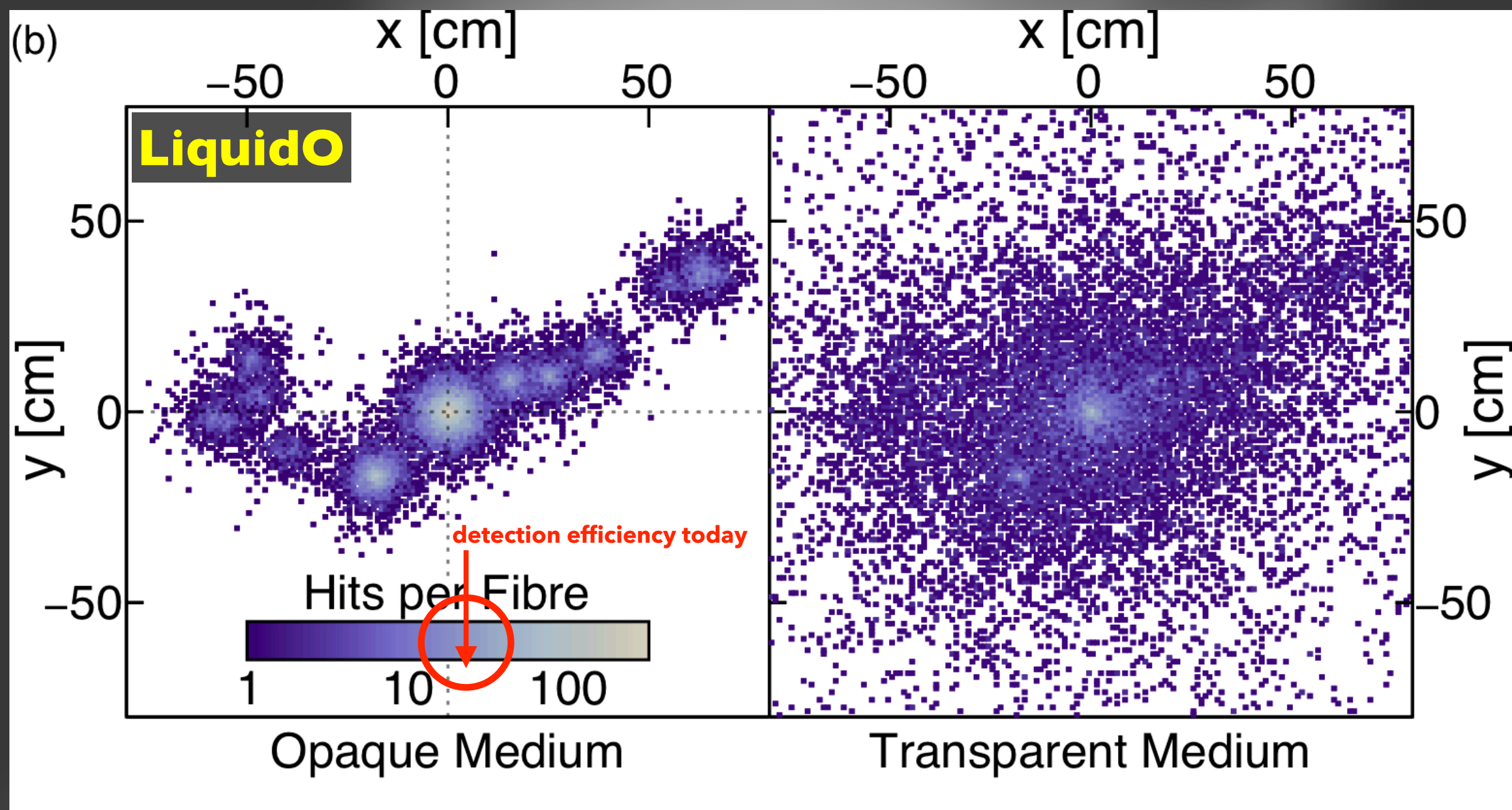
**BOTTOM VIEW: (X,Y) Projection  $\rightarrow$  direct readout**



**LiquidO can have up to 3 orthogonal fibre lattice orientations (3D)**



assume conventional liquid scintillators's yield  $\sim 10,000 \gamma/\text{MeV}$



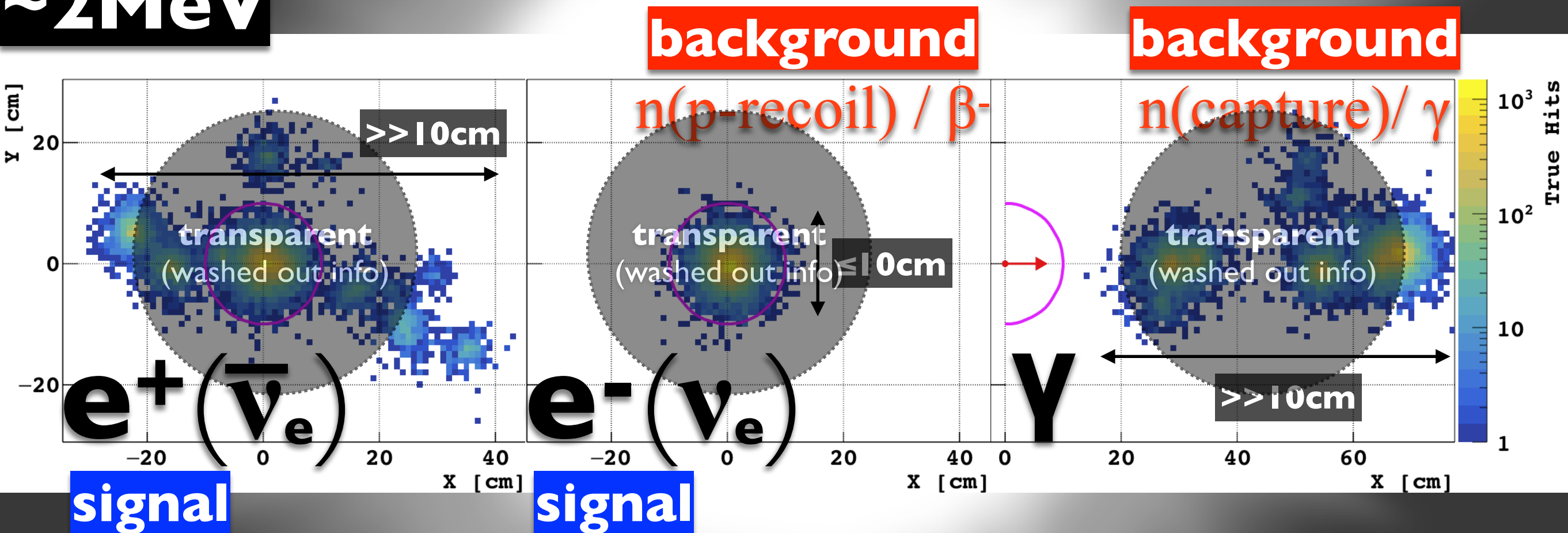
opacity implies neat images...



# unprecedented MeV imaging...

reduce overburden/shielding

**~2MeV**



**LiquidO: stochastic confinement** (NO segmentation)



# topology's PID (no timing)...

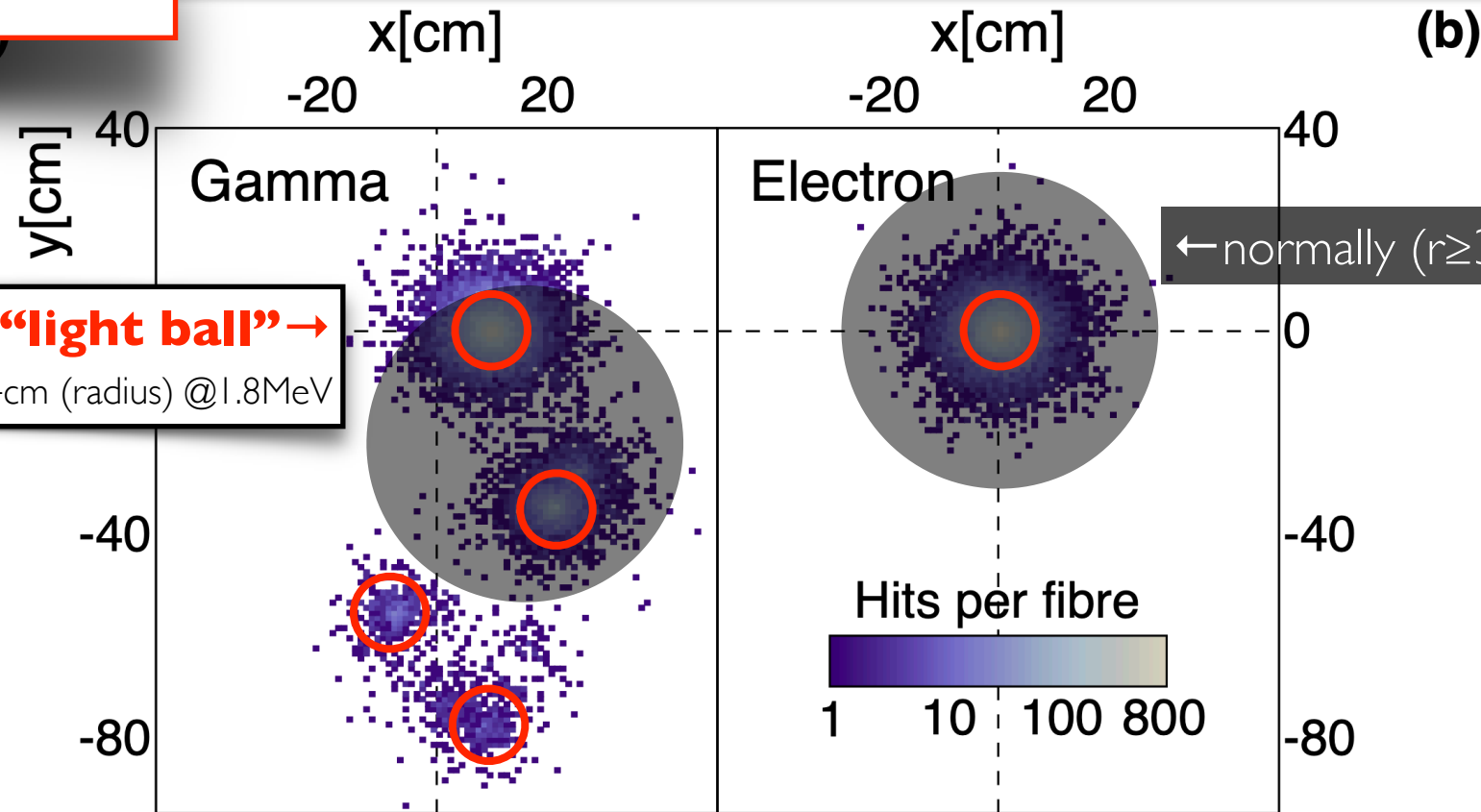
Neutrino physics with an opaque detector

LiquidO Consortium

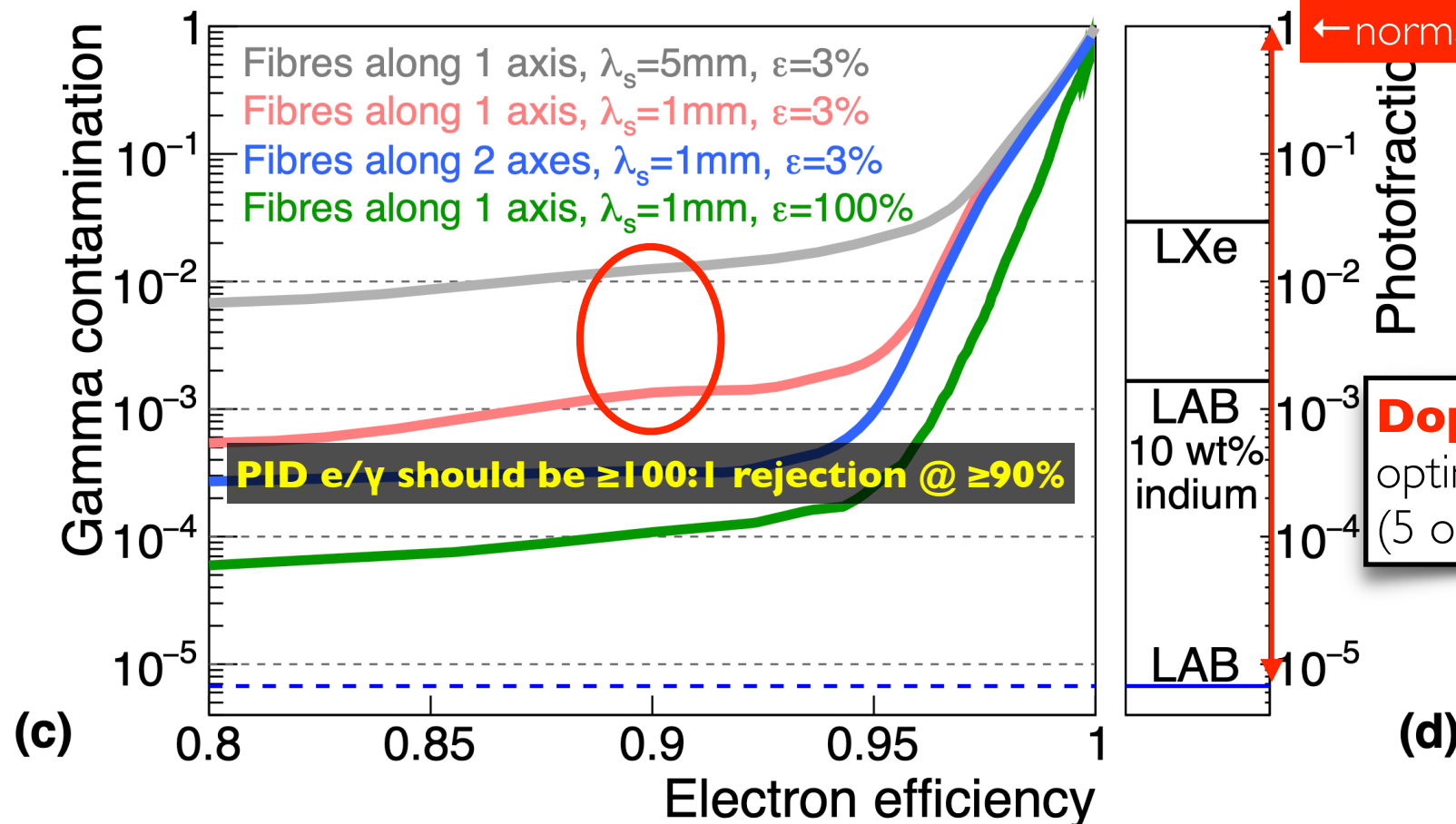
Communications Physics 4, Article number: 273 (2021) | Cite this article

**LiquidO prototype “light ball”** →

~80% light within ~4cm (radius) @1.8MeV



← normally ( $r \geq 30\text{cm} \Leftrightarrow \geq 1.5\text{ns}$  for  $\sigma^{\text{PMT}}$ )



← normally here: **NO e/γ PID**

**Doping Impact**  
optimisation **PID vs doping**  
(5 orders of magnitude)



# LiquidO's prototype MINI-II (upgrade)

data taking since 2021



**overall view**

**3" PMT**

(test transparency)

**single electrons**

[0.4, 1.8] MeV mono-energetic

**~10L multi-media**

- water (transparent)
- scintillator (transparent)
- scintillator (transparent ↔ opaque)

**64 channels readout**

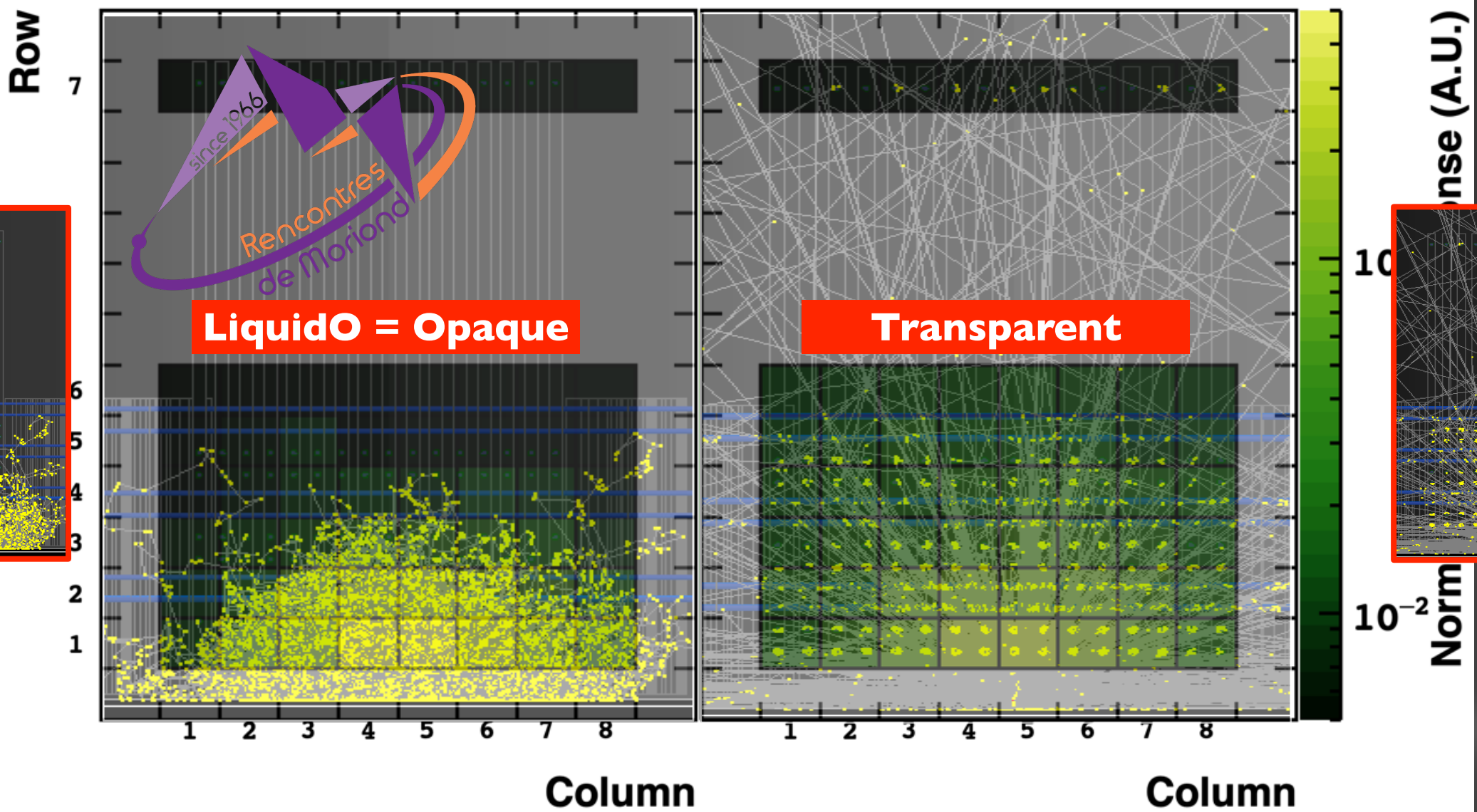
(pitch  $\xi \approx 1.5\text{cm}$ )

**top view**

**T control**

radiator ⊕ chiller: [5, 40] °C

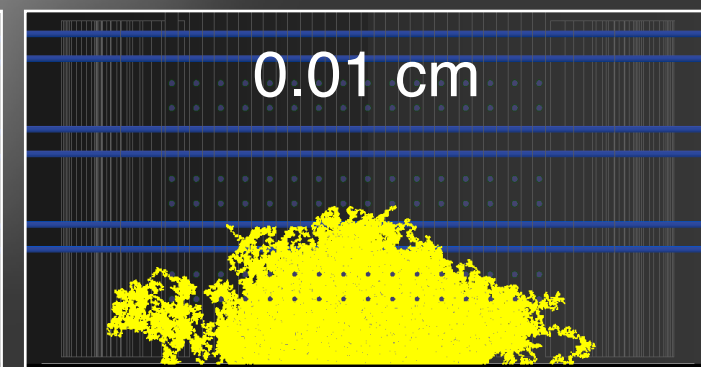
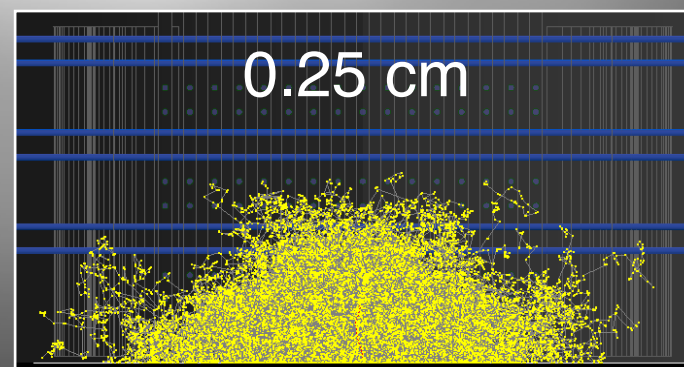
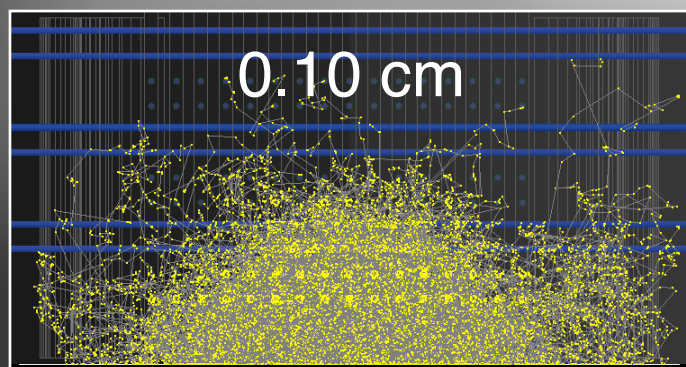
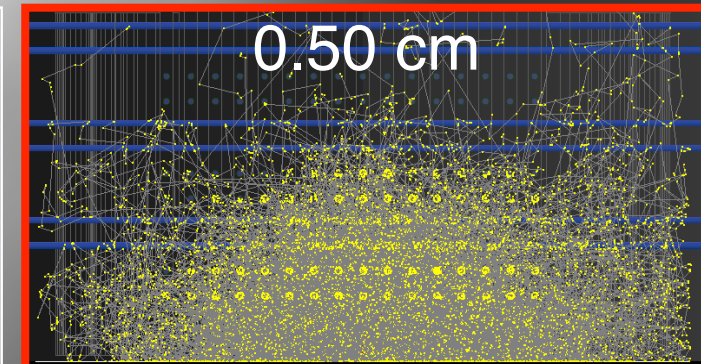
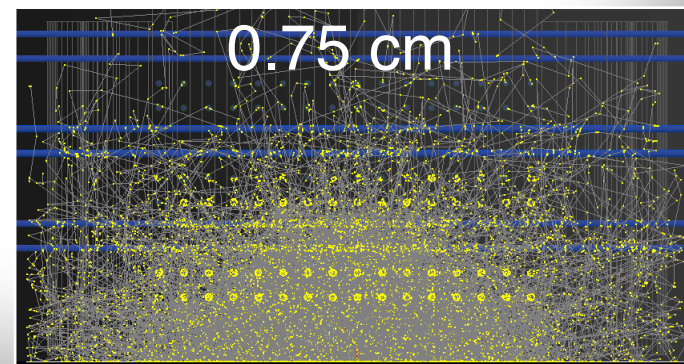
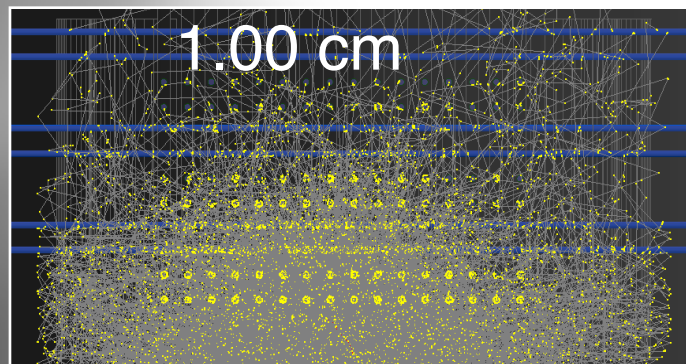




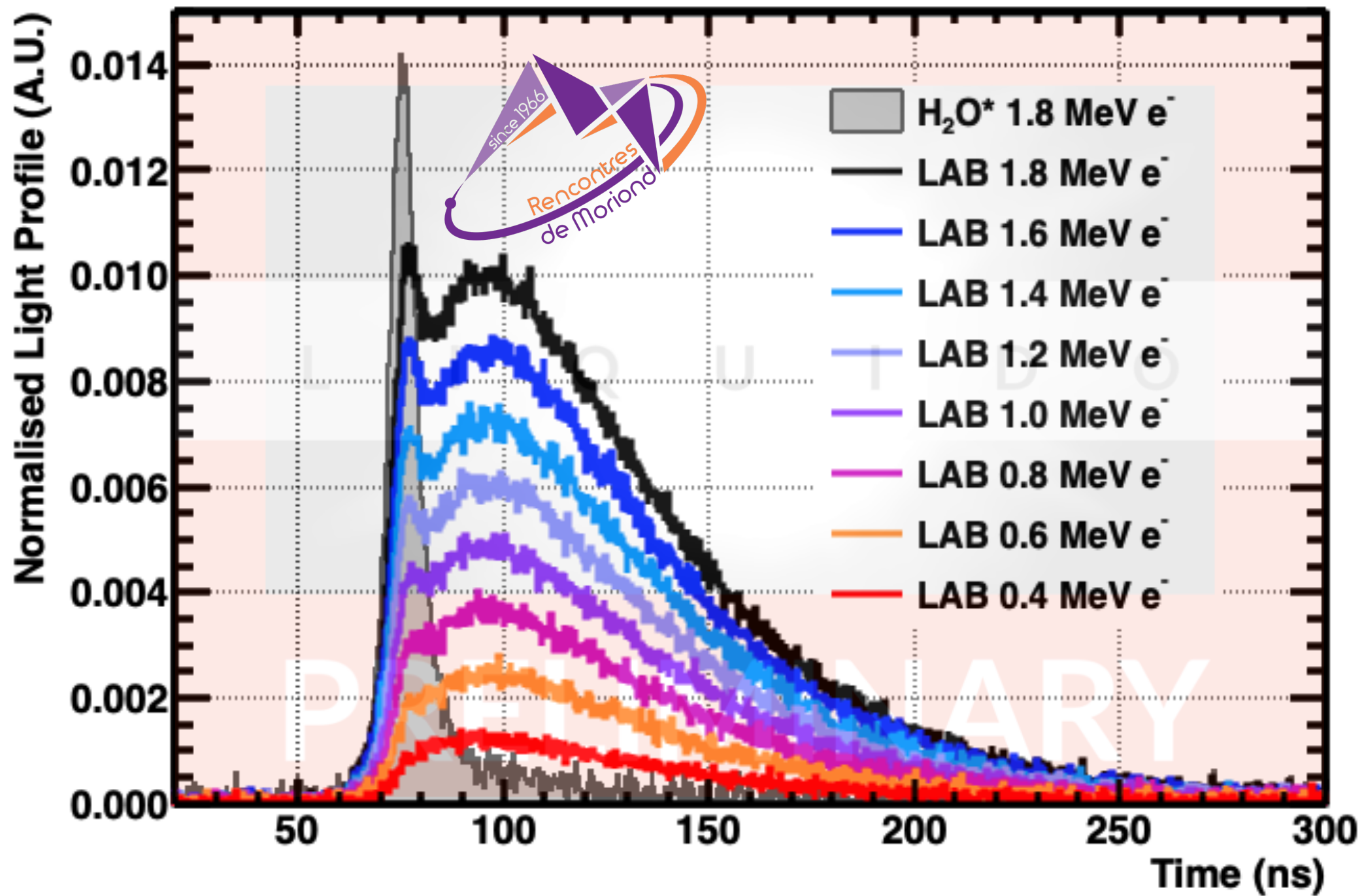
## Geant4 Simulation (under tuning)

“light ball” size:

- scattering:  $\lambda_s$
- # fibres
- absorption?







**ANY light detection: Cherenkov / Scintillation / anything!**  
(ensure the opaque medium is granted)



**First Release at CERN** July 2019 (detector seminar)

<https://indico.cern.ch/event/823865/>

# nature communications physics

**Neutrino 2022**  
(June 2022)

Article | [Open access](#) | Published: 21 December 2021

## Neutrino physics with an opaque detector

[LiquidO Consortium](#)

*Communications Physics* **4**, Article number: 273 (2021) | [Cite this article](#)

5131 Accesses | 9 Citations | 23 Altmetric | [Metrics](#)

### Abstract

**COVID delayed**

In 1956 Reines & Cowan discovered the neutrino using a liquid scintillator detector. The neutrinos interacted with the scintillator, producing light that propagated across transparent volumes to surrounding photo-sensors. This approach has remained one of the most widespread and successful neutrino detection technologies used since. This article introduces a concept that breaks with the conventional paradigm of transparency by confining and collecting light near its creation point with an opaque scintillator and a dense array of optical fibres. This technique, called LiquidO, can provide high-resolution imaging to enable efficient identification of individual particles event-by-event. A natural affinity for adding dopants at high concentrations is provided by the use of an opaque medium. With these and other capabilities, the potential of our detector concept to unlock opportunities in neutrino physics is presented here, alongside the results of the first experimental validation.

[www.nature.com/articles/s42005-021-00763-5](https://www.nature.com/articles/s42005-021-00763-5)



**FNAL Seminar 2023**  
(May 2023)



**LiquidO Official WEB:** <https://liquido.ijclab.in2p3.fr/>



## LiquidO Consortium<sup>(a-z)\*</sup>

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Contact: **LiquidO-Contact-L@in2p3.fr**

Web: **<https://liquido.ijclab.in2p3.fr/>**



**Chooz** (most powerful reactor)  $\oplus$  **UND** ( $\geq 30\text{m}$  baseline)  $\oplus$  **LiquidO** (BG rejection)  
[**EDF** within the team — unprecedented]

C L  U D  
I - II - III

CLOUD's sequence...

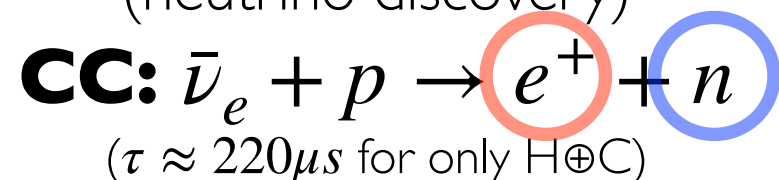


# the power of coincidences

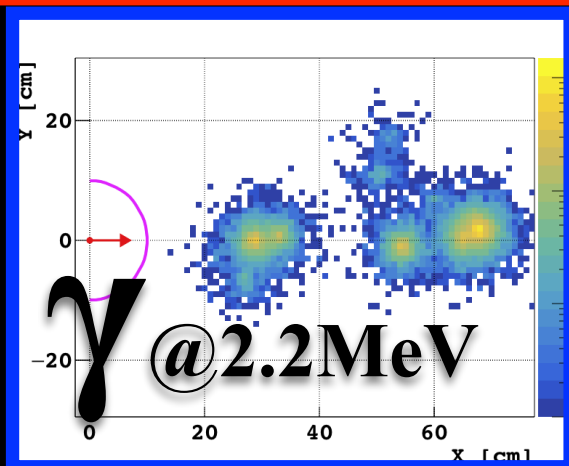
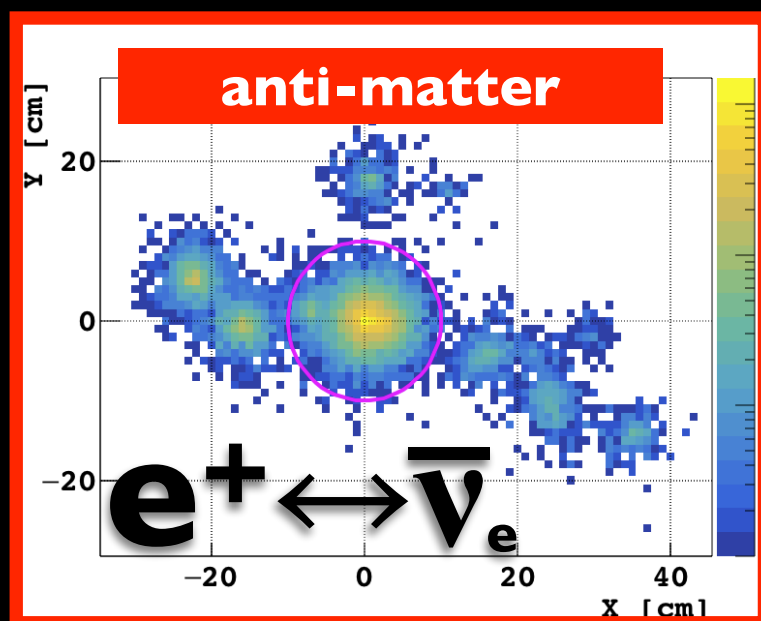
being at **the right “place & time & energy & PID” — huge rejection(s)**

## Reines et al ‘50s

(neutrino discovery)

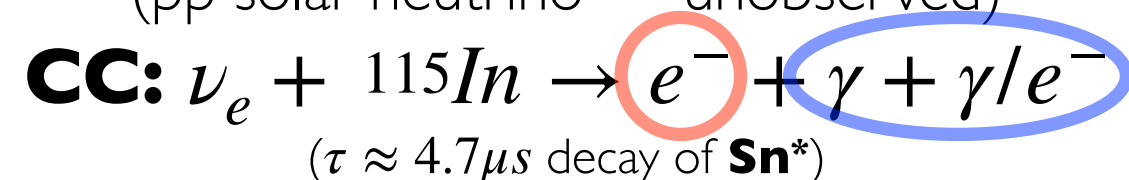


(anti)neutrino **discovery** [ $\tau_n$  &  $\Delta m_{p\sim n}$ ]

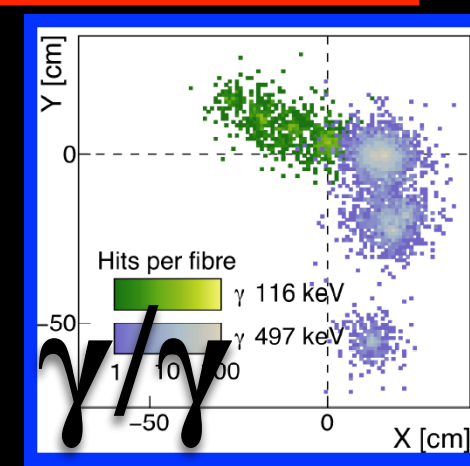
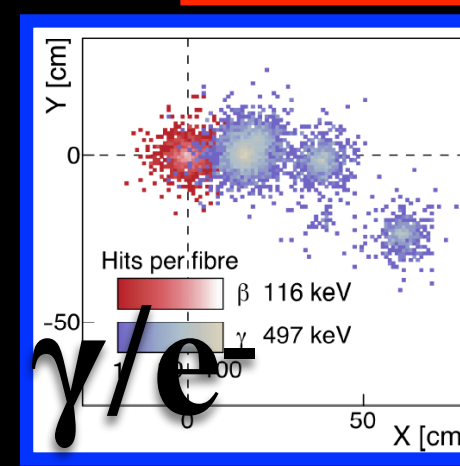
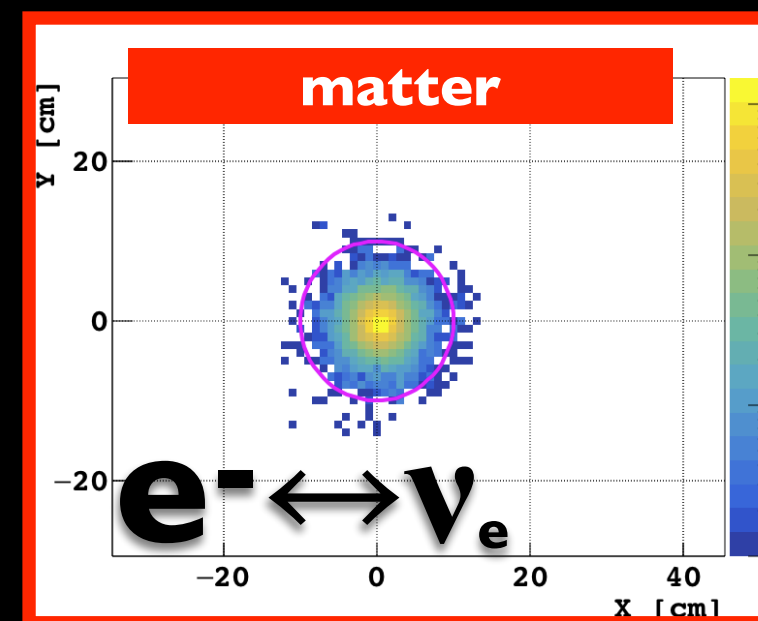


## Raghavan et al ‘70s

(pp solar neutrino — unobserved)



major **R&D** [ $\sim 2$  decades] by **LENS** *et al.*

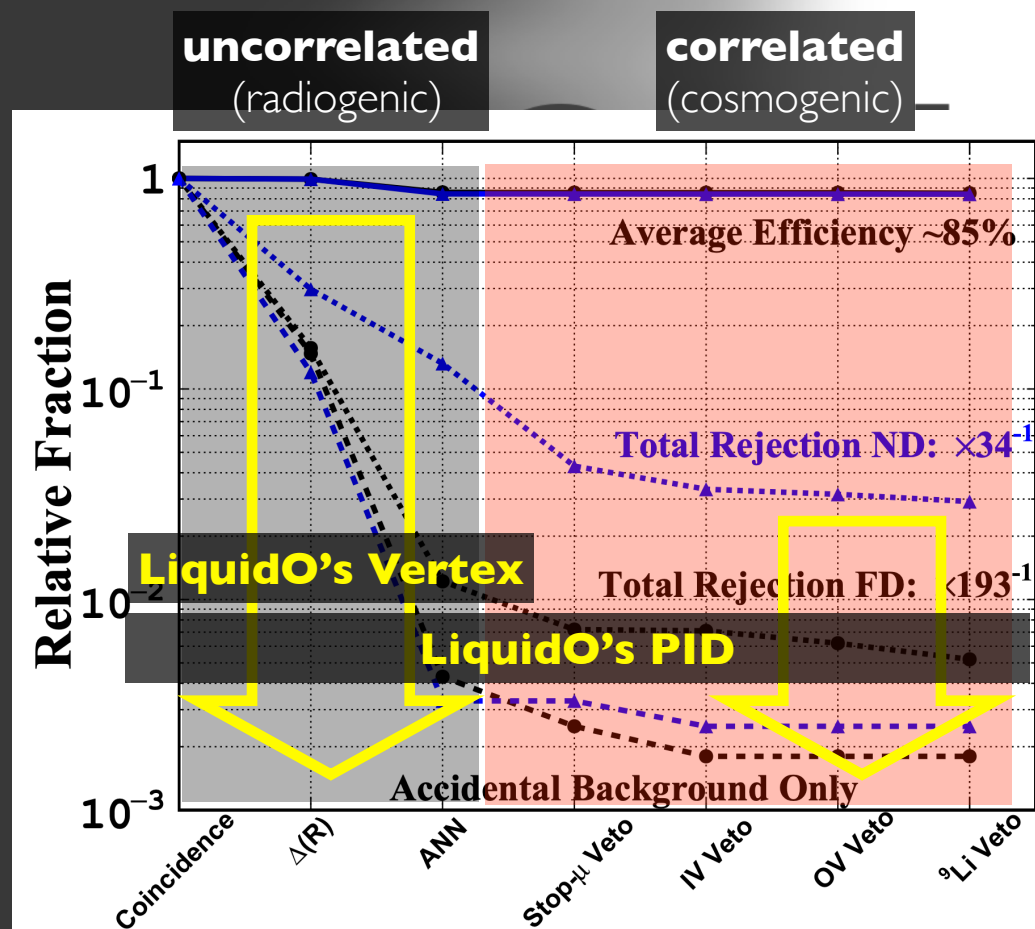




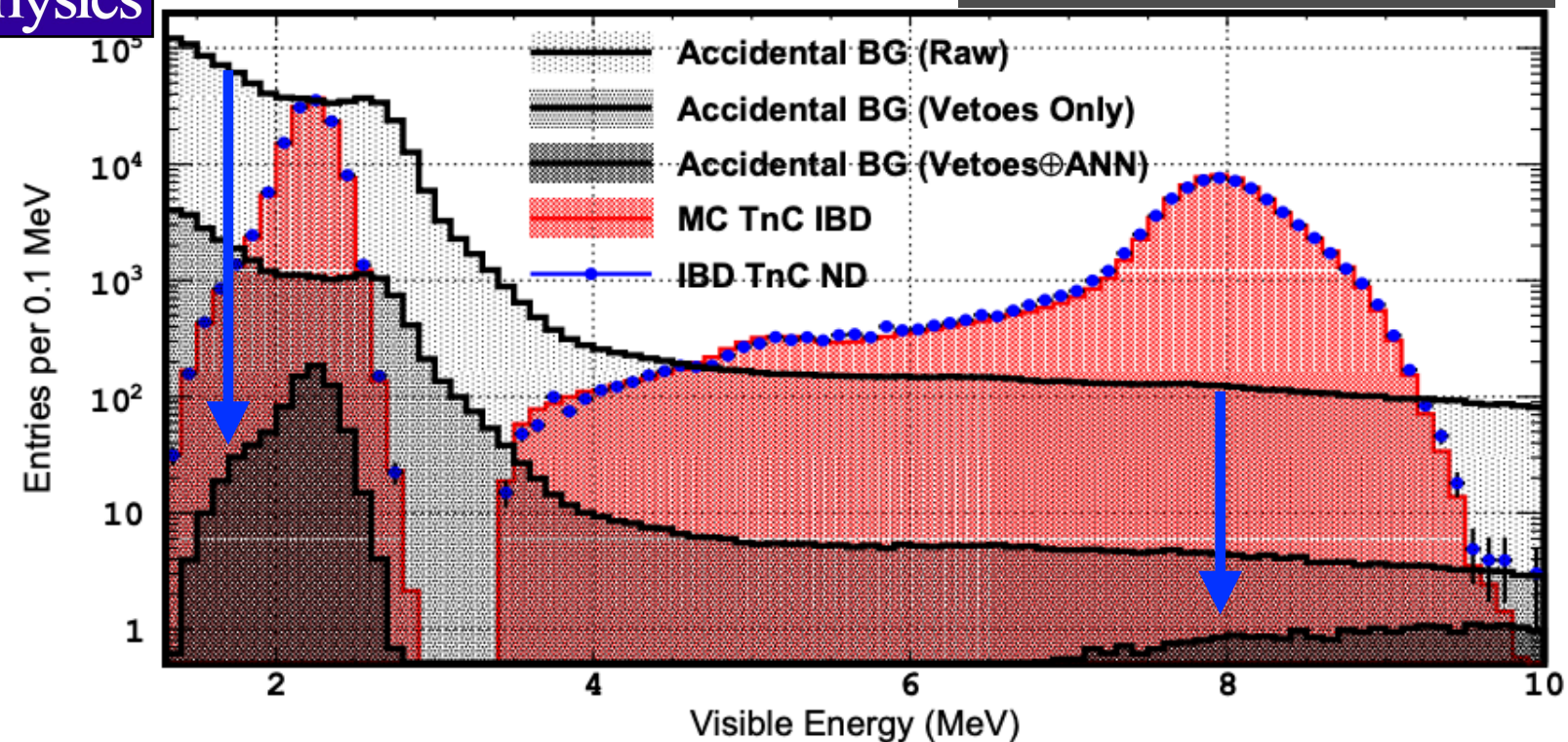
# active BG rejection and control...

- detection using **coincidence-signal** (ex. **IBD@p**)  $\Rightarrow$  prompt-delayed correlations
  - **combinatory (uncorrelated) BG(s)**: **5D-coincidence** ( $\Delta t \oplus \Delta r \oplus \Delta E$ ) — **LiquidO's** mm-vertex
  - **cosmogenic (correlated) BG(s)**: **particle-ID** — **LiquidO's** imaging [impossible so far]
- **active rejection**  $\rightarrow$  rejected-BG as **data-driven BG input** (high accuracy physics extraction)
- **radiogenic control**: in-situ radiogenic BG model tuning (radiopurity control order  $\leq 10^{-14}$ g/g)

at right **place $\oplus$ time $\oplus$ energy $\oplus$ PID** — many orders of magnitude



nature  
physics



**DC: no need for Gd?**  
(cosmogenic BG)

easier to lower **combinatory-BG** ( $\sim 3$  orders of magnitude) than **cosmogenic-BG** ( $\sim 1$  order of magnitude)



# C L I U D

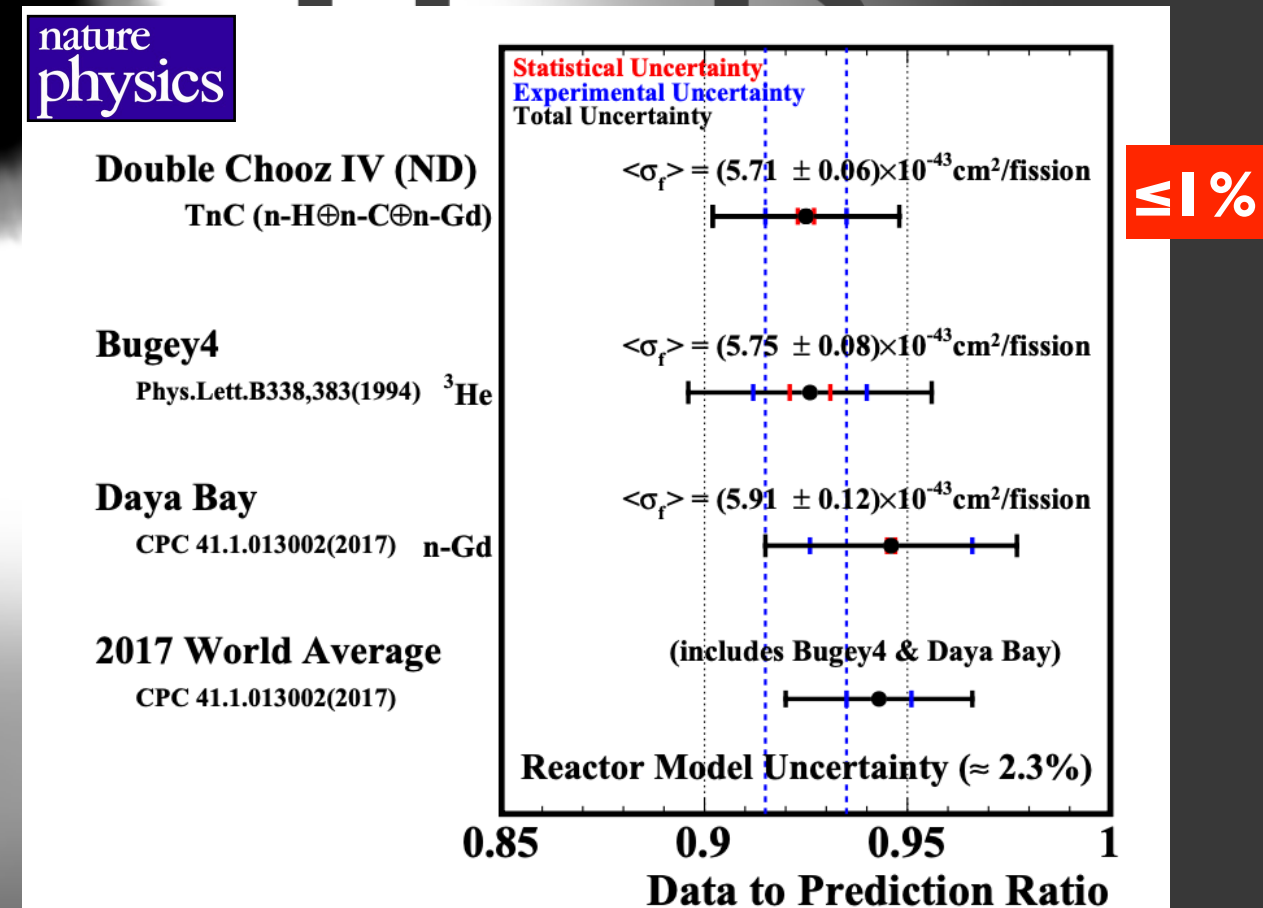
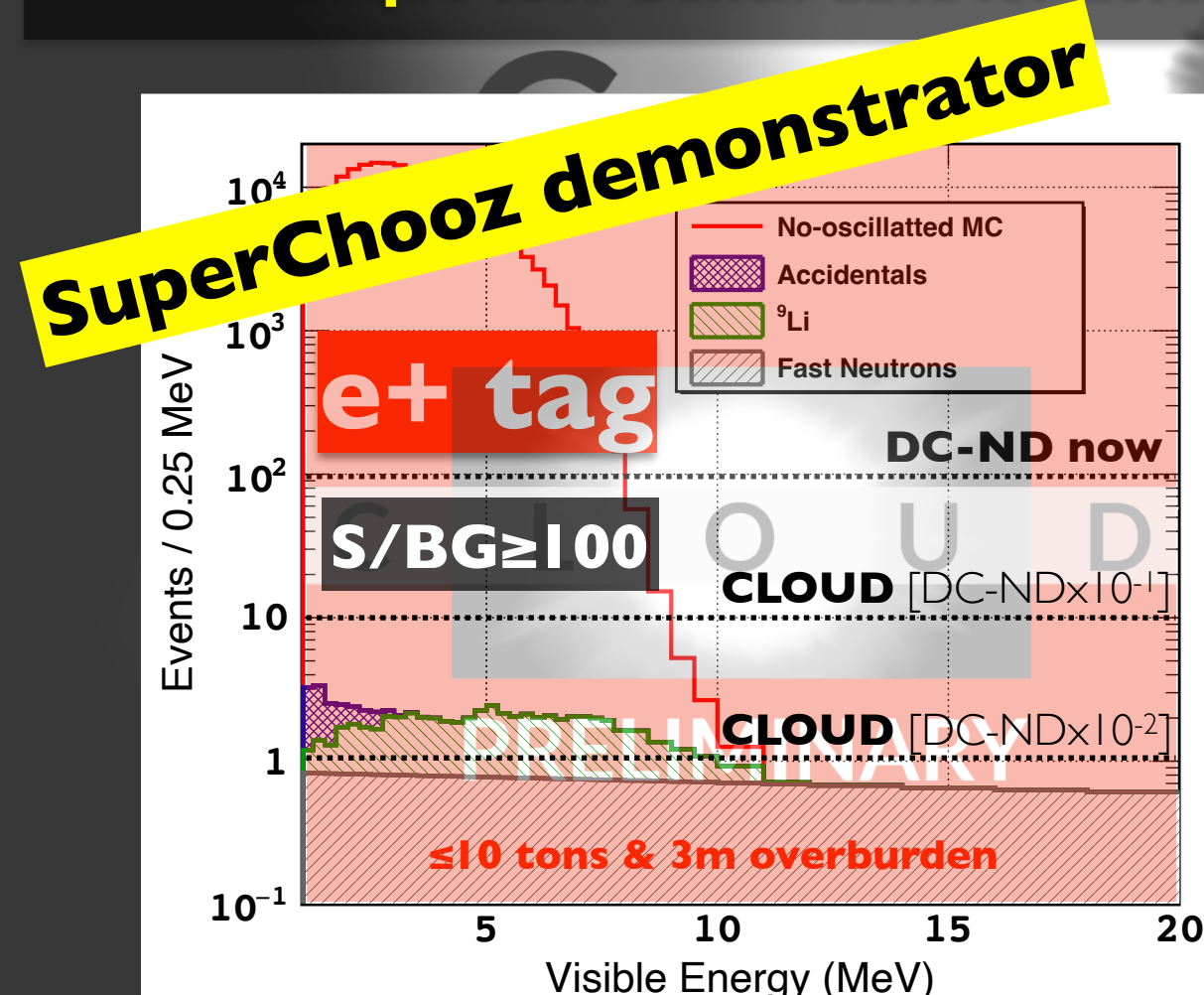
**AntiMatter-OTech** — synergy  
**SuperChooz's** antineutrino golden channel **demonstration** — byproduct

antineutrino CC & NC? (~~doping~~)...



# CLOUD-I physics programme: **IBD@p...**

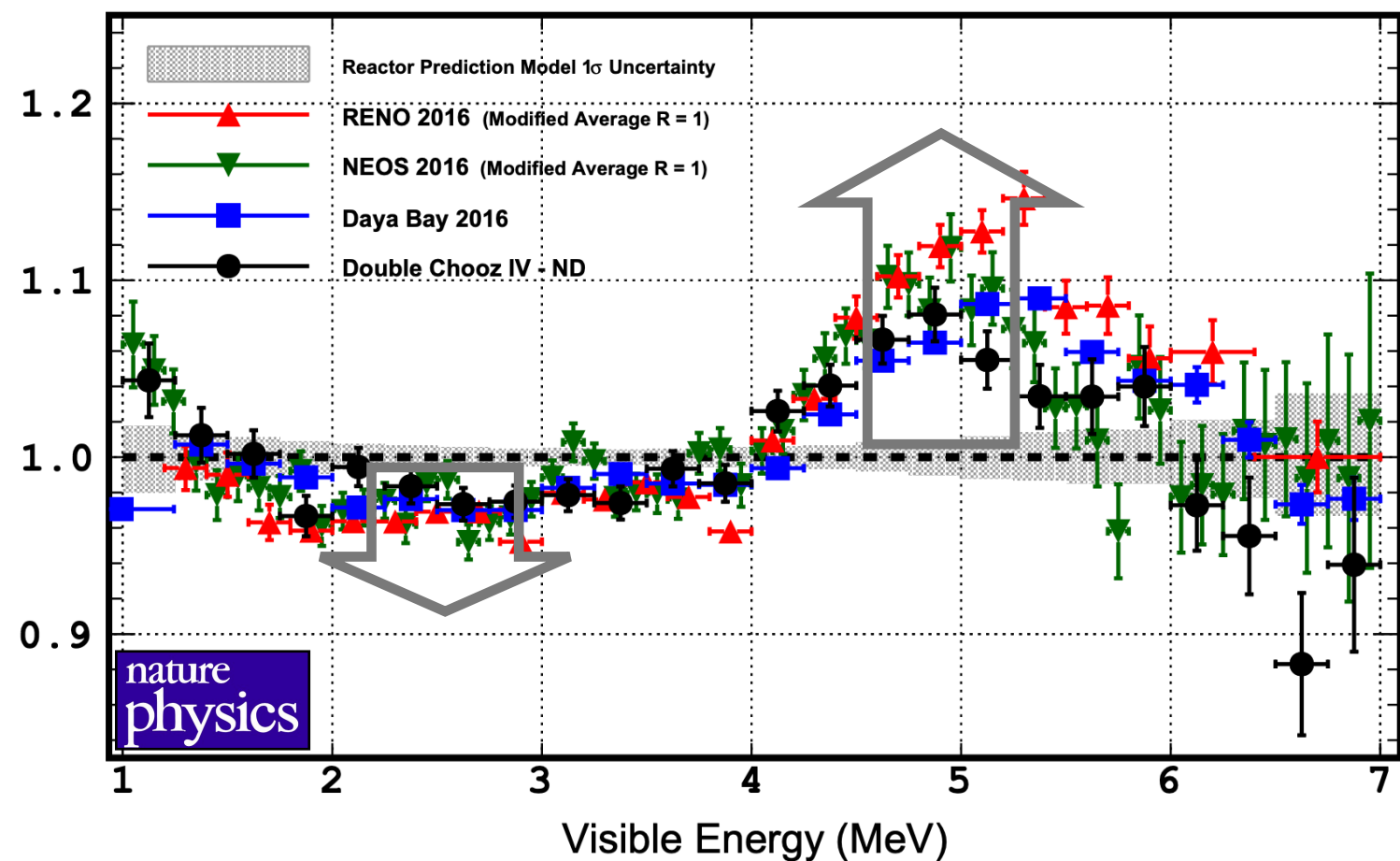
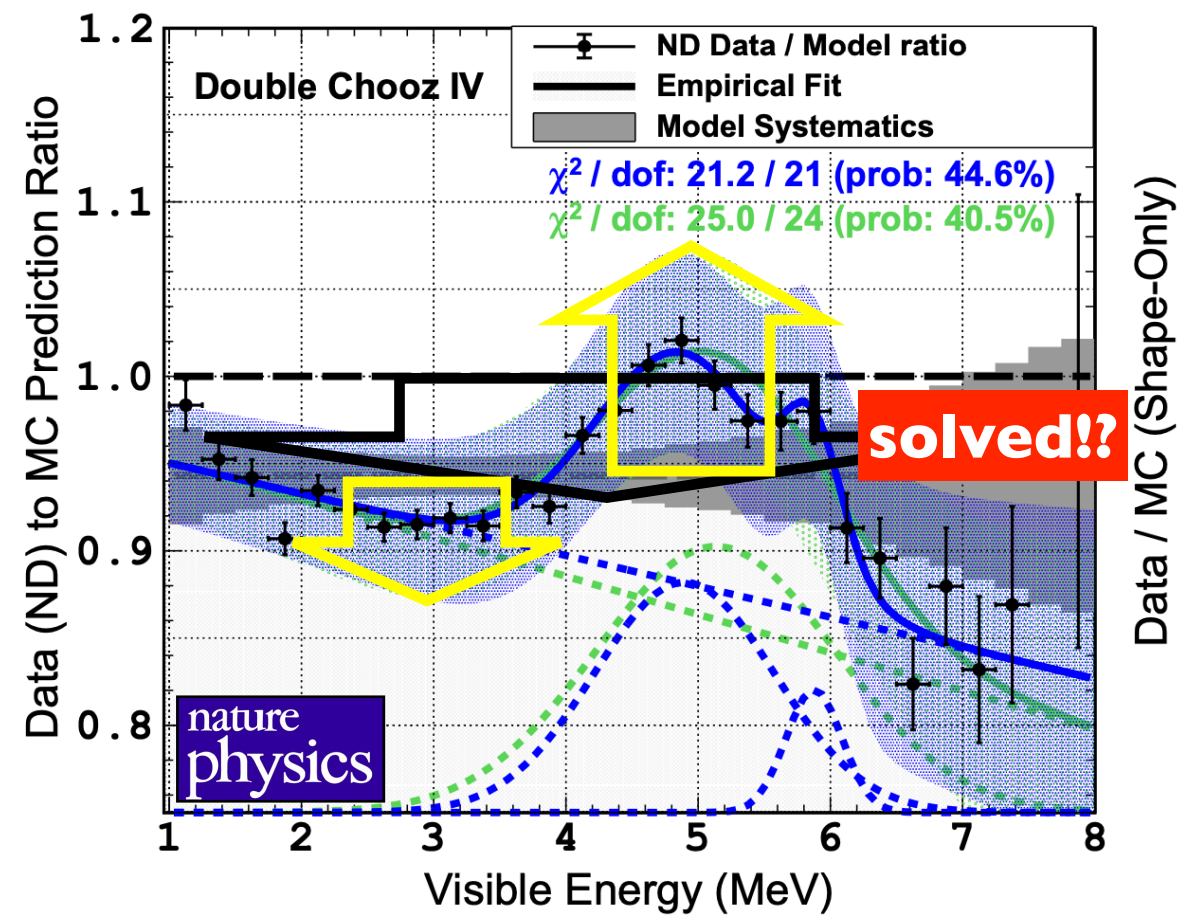
- **IBD@p** (anti- $\nu$  CC):  $\geq 10,000$  interaction per day for **10tons ID** [ $\geq 3M$  interactions per year]
- **LiquidO** reach a **background-less** regime — improve  $\geq 3x$  today's BG control (ex. DC-ND)
- **Signal(ON)-to-BG  $\geq 100$**  — unprecedented high precision reactor characterisation
  - dominant  **$\sim 0.5$ (thermal power) uncertainty** & accurate **U/Pu composition**
- **Signal(OFF)-to-BG  $\geq 1$**  — unprecedented **reactor-fuel monitoring**
  - accurate monitoring of **transitions OFF-ON-OFF** — some interesting physics
- **unique test-bench data for accurate prediction** — validate uncertainties, too?



**CLOUD precision  $\geq 0.6\% \Rightarrow$  Unitarity Violation? (if predictions are improved)!**



all experiments consistent — except **Bugey3!!**



**ABSOLUTE FLUX: the future of reactor-antineutrino physics**

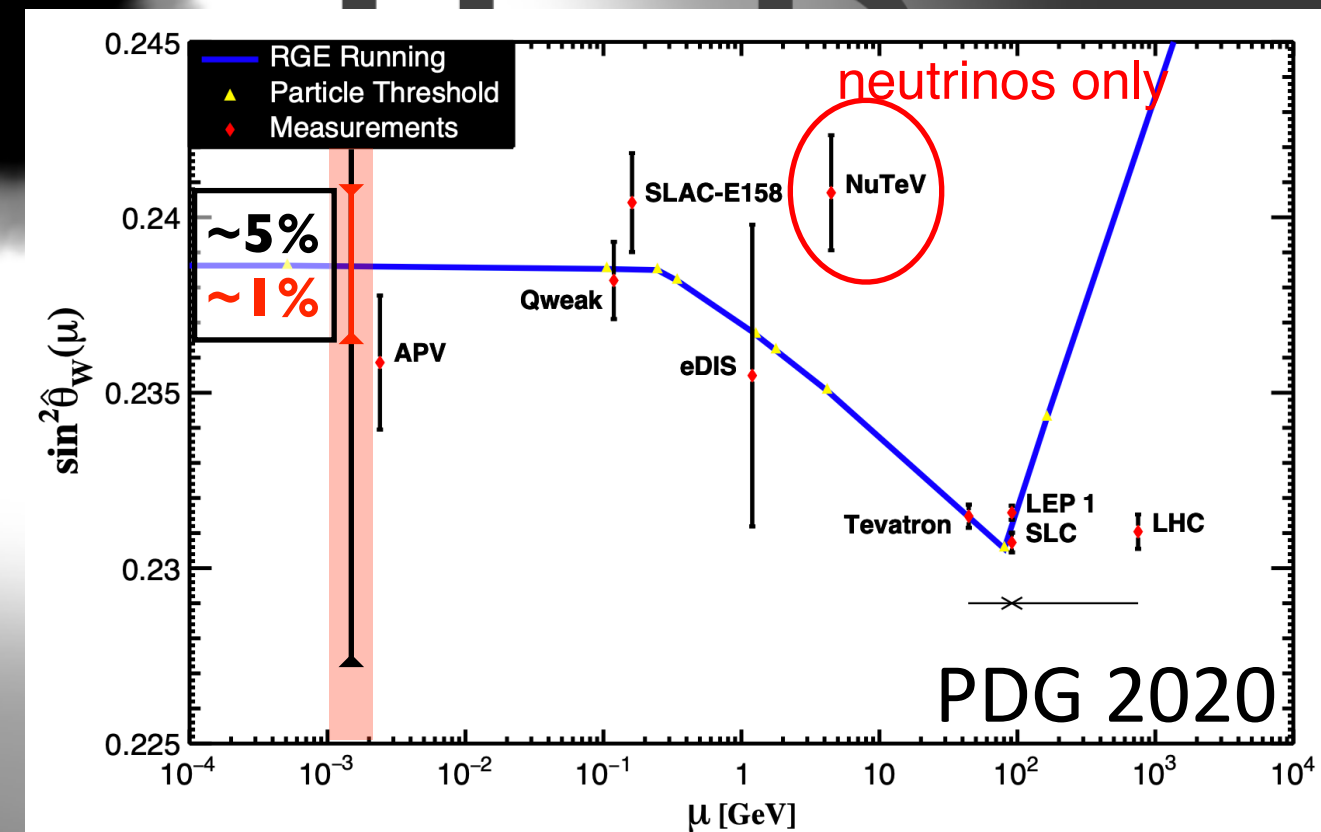
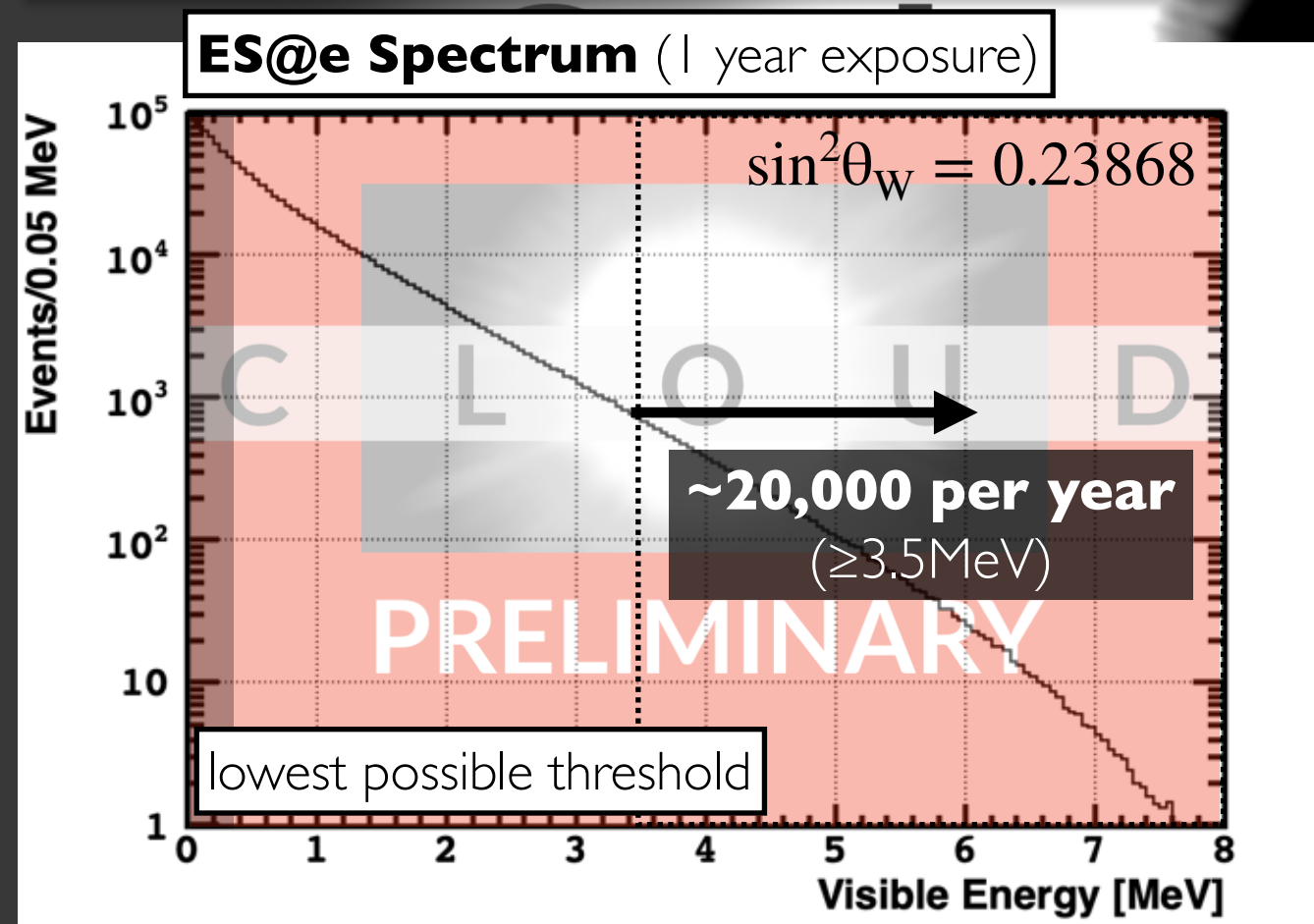
**we must solve this “mess”**  $\Rightarrow$  the reward **possible new physics!**  
(if so, prediction should not use neutrino input  $\Rightarrow$  **no** new physics)

must understand flux  $\leq 1\%$ ...



# CLOUD-I physics programme: **ES@e**...

- **eES** (anti- $\nu$  CC $\oplus$ NC):  **$\leq 5,000$**  interactions per day for **10tons ID** [ $\approx 2M$  interactions per year]
  - interference CC & NC — different for neutrino (easier) and antineutrino (harder)
  - measure  $\theta_w$  or use to **decompose the NC flux** component
  - PDG-2022's  $\sin^2\theta_w = [0.231??, 0.23868]$  — running due to SM's renormalisation
- **major challenge: LiquidO** isolate “e-like” **PID** and exploit **high-rate reactor modulation**
  - likely strong **fiducial volume** & **higher energies** — reduce detected rate drastically
  - **$\leq 10\%$  precision** ( $\geq 5\sigma$  observation) tolerates much BG but  **$\leq 1\% \Rightarrow S/BG \geq 2(!!)$  impossible?**



R.L. Workman et al. (Particle Data Group), Prog. Theor. Exp. Phys. 2022, 083C01 (2022)



# CLIU D

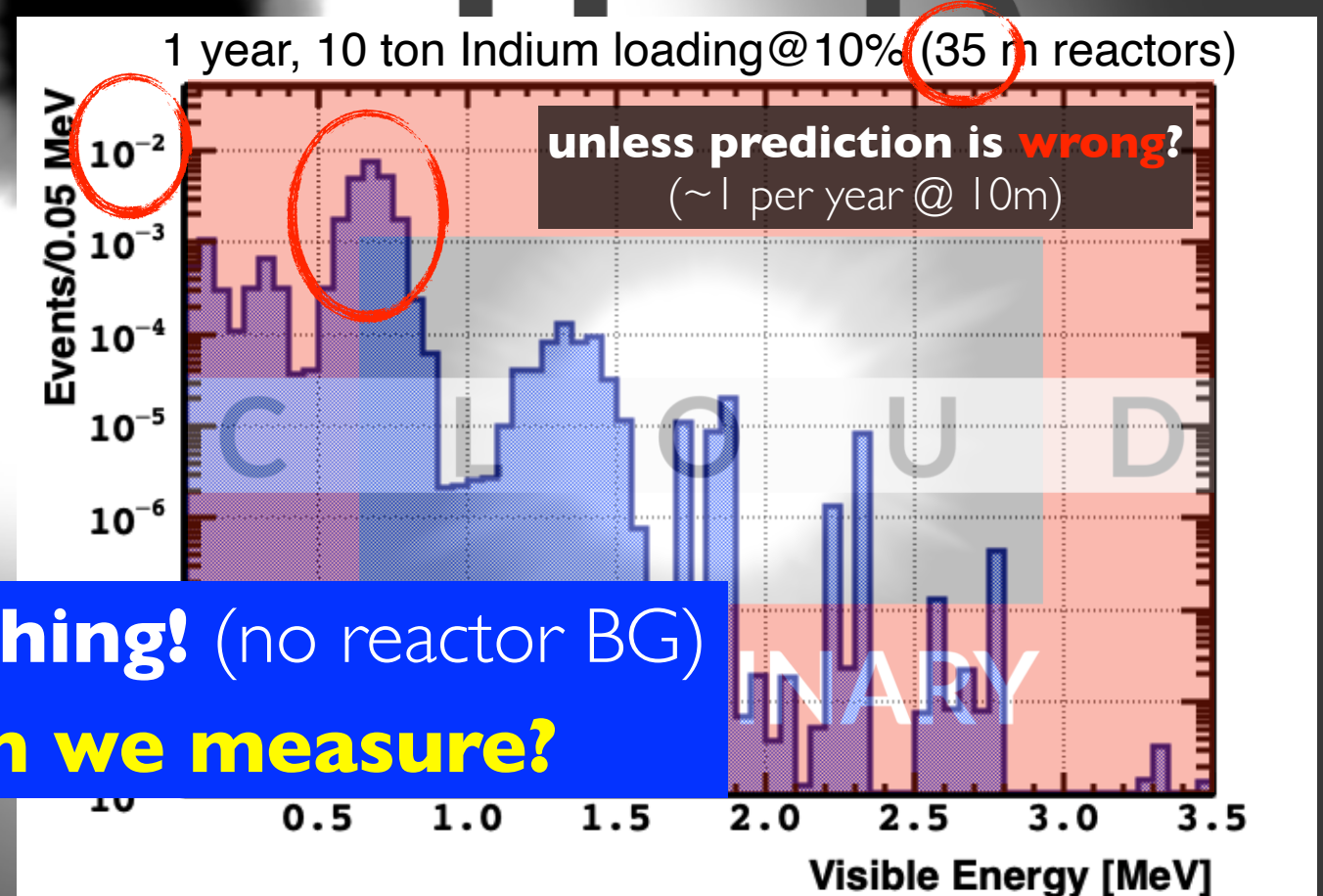
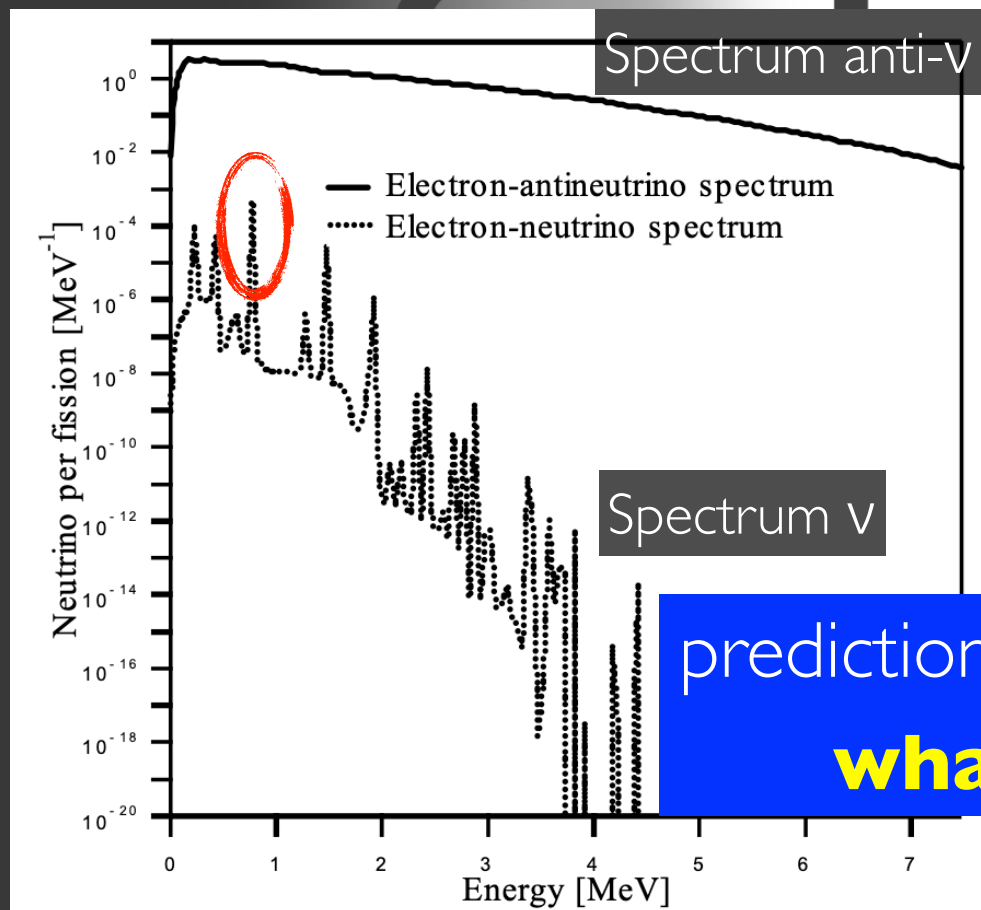
**SuperChooz's** neutrino golden channel **demonstration** — byproduct

neutrino CC (doping)...



# CLOUD-II physics programme: **neutrino**...

- loading **indium** on the detector — **unique strong coincidence  $\geq 114\text{keV}$**  (“solar-pp” in mind)
  - low threshold  $\oplus$  high natural-abundance  $\oplus$  high-ish cross-section  $\oplus$  BG-killer (coincidence)
  - CC interaction:  $\nu_e + {}^{115}\text{In} \rightarrow e^- + {}^{115}\text{Sn}^*$  [ $\tau: 4.8\mu\text{s}$  decay:  $\gamma/e(116\text{keV}) + \gamma(496\text{keV})$ ]
  - reactor neutrino **modulate with the reactor power** — no ambiguity whatsoever
- why to detector neutrinos close to a reactor?
  - **reactor neutrinos** (from  $\beta^+/\text{EC}$ ): **rate( $\nu$ )  $\approx 10^{-5}$  rate(anti- $\nu$ )** — prediction (both correlated)
  - could **reactors be the missing MeV neutrino source?** [otherwise impractical]

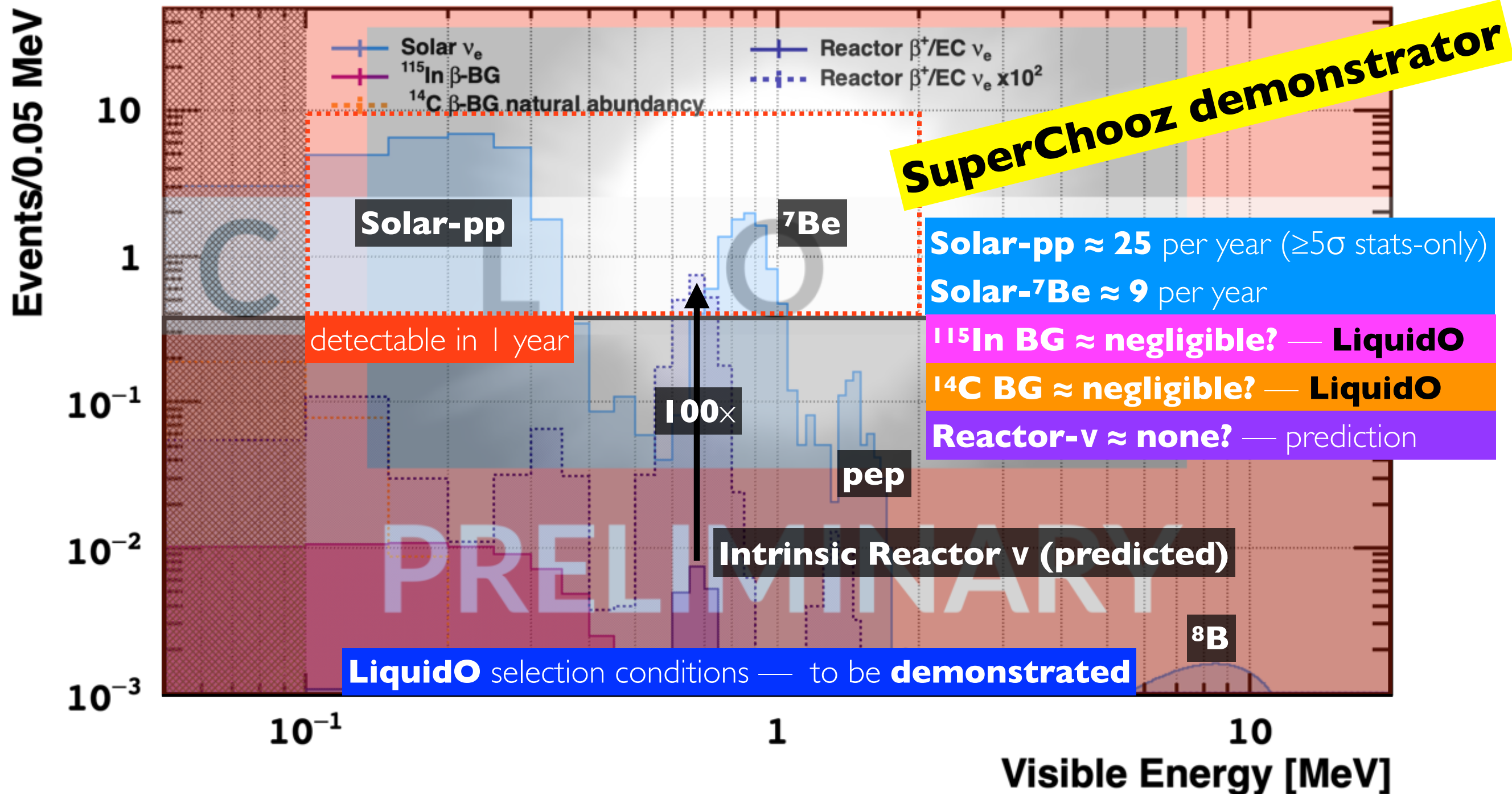




# the big picture of neutrinos @CLOUD...

assuming the **LENS “BG model”** — valid at **overburden ~3m?** (to be demonstrated)

1 year, 10 ton Indium loading @10% (35 m reactors)



detection **solar-pp neutrinos** on a **10ton** detector **almost on surface** right **next a nuclear reactor?**



# CLIU D

**R&D for low energy reactor-fuel monitoring & geoneutrino  $^{40}\text{K}$  discovery** — demonstration

new antineutrino CC (doping)...



arXiv:2308.04154

# Probing Earth's Missing Potassium using the Unique Antimatter Signature of Geoneutrinos

A. Cabrera<sup>\*12 $\alpha$ ,2,a</sup>, M. Chen<sup>†6</sup>, F. Mantovani<sup>‡3 $\alpha$ ,3 $\beta$</sup> , A. Serafini<sup>§3 $\alpha$ ,3 $\beta$ ,13 $\alpha$ ,13 $\beta$</sup> , V. Strati<sup>¶3 $\alpha$ ,3 $\beta$</sup> , J. Apilluelo<sup>18</sup>, L. Asquith<sup>1</sup>, J.L. Beney<sup>11</sup>, T.J.C. Bezerra<sup>1</sup>, M. Bongrand<sup>11</sup>, C. Bourgeois<sup>12 $\alpha$</sup> , D. Breton<sup>12 $\alpha$</sup> , M. Briere<sup>12 $\alpha$</sup> , J. Busto<sup>10</sup>, A. Cadiou<sup>11</sup>, E. Calvo<sup>8</sup>, V. Chaumat<sup>12 $\alpha$</sup> , E. Chauveau<sup>4</sup>, B.J. Cattermole<sup>1</sup>, P. Chimenti<sup>7</sup>, C. Delafosse<sup>12 $\alpha$</sup> , H. de Kerret<sup>1a</sup>, S. Dusini<sup>13 $\alpha$</sup> , A. Earle<sup>1</sup>, C. Frigerio-Martins<sup>7</sup>, J. Galán<sup>18</sup>, J. A. García<sup>18</sup>, R. Gazzini<sup>12 $\alpha$</sup> , A. Gibson-Foster<sup>1</sup>, A. Gallas<sup>12 $\alpha$</sup> , C. Girard-Carillo<sup>9 $\alpha$</sup> , W.C. Griffith<sup>1</sup>, F. Haddad<sup>11</sup>, J. Hartnell<sup>1</sup>, A. Hourlier<sup>17</sup>, G. Hull<sup>12 $\alpha$</sup> , I. G. Irastorza<sup>18</sup>, L. Koch<sup>9 $\alpha$</sup> , P. Lanièce<sup>12 $\alpha$ ,12 $\beta$</sup> , J.F. Le Du<sup>12 $\alpha$ ,2</sup>, C. Lefebvre<sup>6</sup>, F. Lefevre<sup>11</sup>, F. Legrand<sup>12 $\alpha$</sup> , P. Loaiza<sup>12 $\alpha$</sup> , J. A. Lock<sup>1</sup>, G. Luzón<sup>18</sup>, J. Maalmi<sup>12 $\alpha$</sup> , C. Marquet<sup>4</sup>, M. Martínez<sup>18</sup>, B. Mathon<sup>12 $\alpha$</sup> , L. Ménard<sup>12 $\alpha$ ,12 $\beta$</sup> , D. Navas-Nicolás<sup>12 $\alpha$</sup> , H. Nunokawa<sup>15</sup>, J.P. Ochoa-Ricoux<sup>5</sup>, M. Obolensky<sup>a</sup>, C. Palomares<sup>8</sup>, P. Pillot<sup>11</sup>, J.C.C. Porter<sup>1</sup>, M.S. Pravikoff<sup>4</sup>, H. Ramarijaona<sup>12 $\alpha$</sup> , M. Roche<sup>4</sup>, P. Rosier<sup>12 $\alpha$</sup> , B. Roskovec<sup>14</sup>, M.L. Sarsa<sup>18</sup>, S. Schoppmann<sup>9 $\beta$</sup> , W. Shorrocks<sup>1</sup>, L. Simard<sup>12 $\alpha$</sup> , H.Th.J. Steiger<sup>9 $\alpha$ ,9 $\beta$</sup> , D. Stocco<sup>11</sup>, J.S. Stutzmann<sup>11</sup>, F. Suekane<sup>16,a</sup>, A. Tunc<sup>9 $\alpha$</sup> , M.-A. Verdier<sup>12 $\alpha$ ,12 $\beta$</sup> , A. Verdugo<sup>8</sup>, B. Viaud<sup>11</sup>, S. M. Wakely<sup>9 $\alpha$</sup> , A. Weber<sup>9 $\alpha$</sup> , and F. Yermia<sup>11</sup>

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(LiquidO Consortium)

**<sup>40</sup>K geoneutrino new methodology → good enough for discovery?**

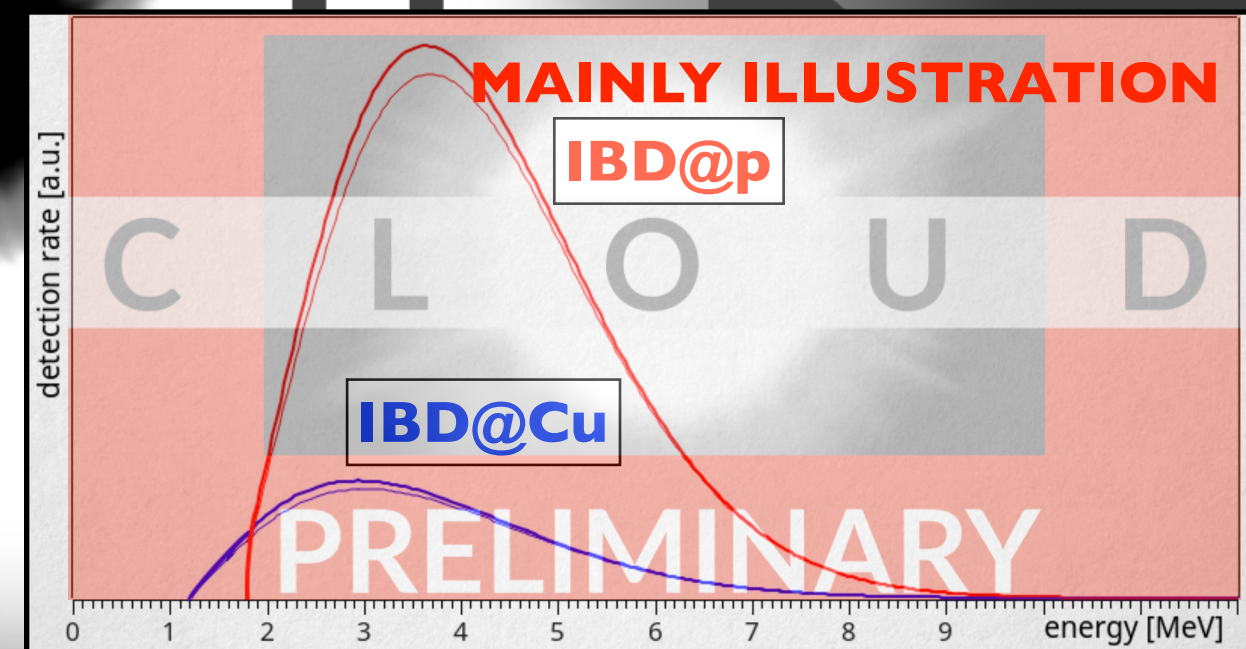
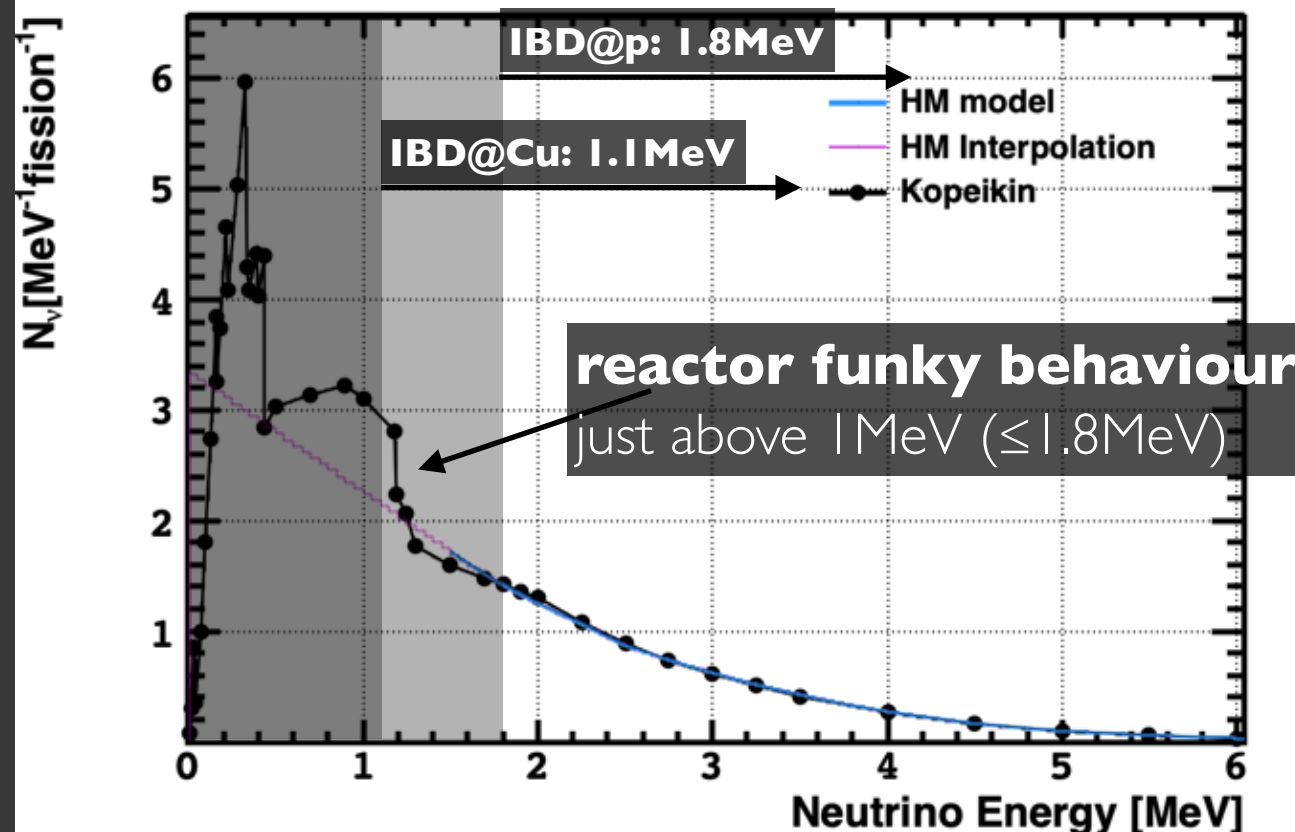


# CLOUD-II physics programme...

## Probing Earth's Missing Potassium using the Unique Antimatter Signature of Geoneutrinos

- **unique IBD-like** interaction on **Cu** [ $E(\text{threshold})$ :  $\sim 1.176 \text{ MeV}$ ] — **unique in the Universe**
  - low threshold  $\oplus$  high natural-abundance  $\oplus$  high-ish cross-section  $\oplus$  BG-resilience? (even coincidence?)
  - CC interaction: **anti- $\nu_e$**  +  **$^{63}\text{Cu}$**   $\rightarrow$   **$e^+$**  +  $^{63}\text{Ni}$  + [if  $^{63}\text{Ni}$  was excited:  **$\gamma$** (87keV;  $\tau \approx 1.67 \mu\text{s}$ )]
- **possible applications:**
  - **direct reactor-fuel monitoring?** — fuel-storage systems?
  - **$^{40}\text{K}$  geoneutrino exploration** (discovery) — extremely challenging

### Reactor Flux Prediction (Kopeikin et al.)



IBC@Cu: net increase of events — to be demonstrated

- **detection feasibility**
- **cross-section** measurement — relative to IBD@p
- branching-ratio for **Cu\*** (**tagging**) versus Cu

**note:** geoneutrino slide in **backup**



C L O U D

a long story short...

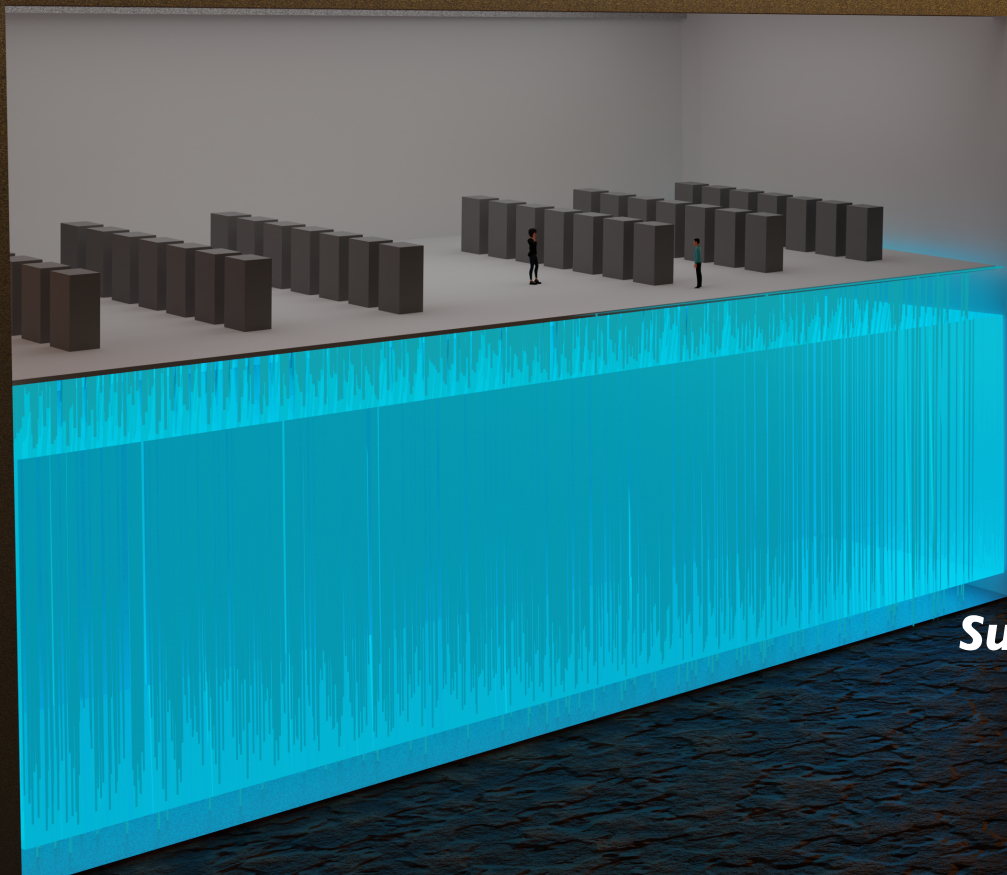


# SuperChooz exploration...

**flagship neutrino oscillation experiment in Europe?**

the Ardennes mountains

**Chooz-A:** Cavern Reactor Core



European  
Innovation  
Council



UK Research  
and Innovation

**AM-OTech** project [EIC-UKRI]  
**CLOUD** experiment

1 Dec 2022

**Chooz-B:** Reactor Cores

**Ultra Near Detectors @ Chooz-B:**

- LiquidO technology
- Mass:  $\leq 5$  tons
- Overburden:  $\leq 5$  m
- Baseline:  $\leq 30$  m

**Super Far Detector @ Chooz-A**

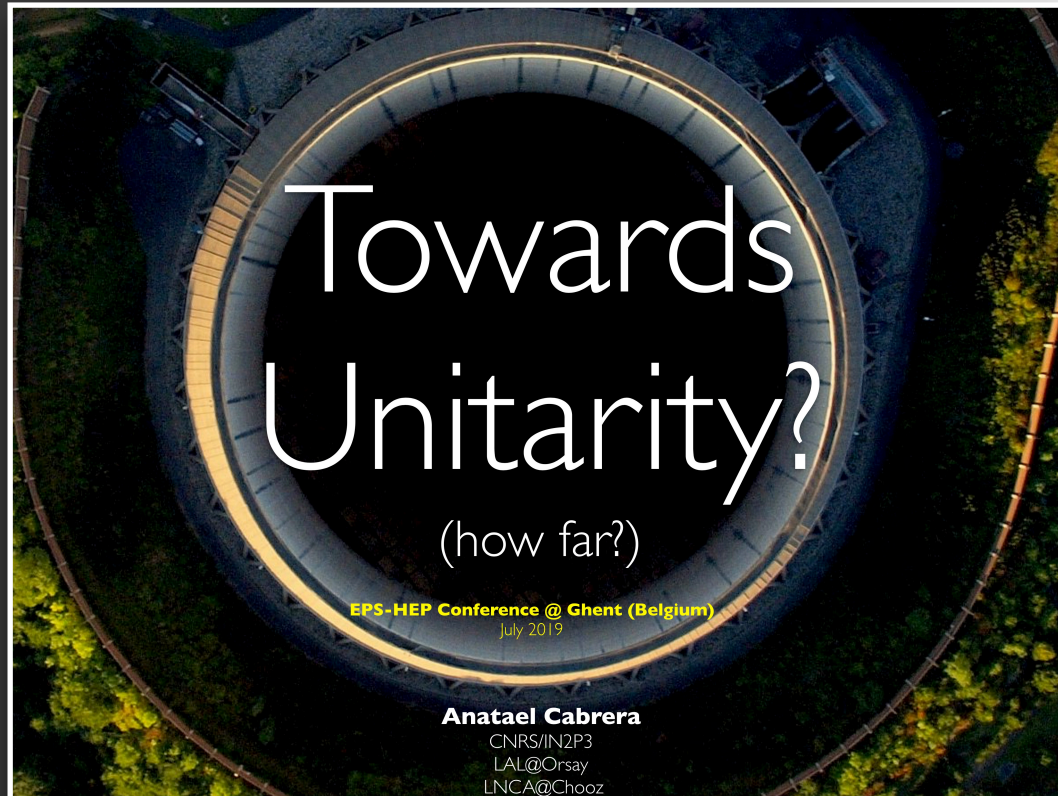
- LiquidO technology
- Mass:  $\sim 10,000$  tons
- Overburden:  $\leq 100$  m
- Baseline:  $\sim 1$  km

the Meuse river





HEP-European Physics Society  
(July 2019 @ Ghent Belgium)



<https://indico.cern.ch/event/577856/contributions/3421609/>

EP Seminar

## The SuperChooz Experiment: Unveiling the Opportunity

by Dr Anatael CABRERA (IJCLab - IN2P3/CNRS)

Tuesday 29 Nov 2022, 11:00 → 12:00 Europe/Zurich

222/R-001 (CERN)



<https://indico.cern.ch/event/1215214/>

<https://zenodo.org/record/7504162>

<https://liquido.ijclab.in2p3.fr/>

# exploring since 2018...



- **CLOUD** demonstrator for **LiquidO's (anti)neutrino detection capabilities** — **a revolution?**
  - byproduct to **new reactor monitoring capability?** [a dream since '70s]
- **CLOUD-I:** **approved & funded [AM-OTech]** **plan: data by 2025**
  - **most precise absolute CC-antineutrino flux** — **new physics?**
  - possible **first NC-(anti)neutrino flux** — **new physics?**
- **CLOUD-II:** **under feasibility study** (→ new **indium**-loaded opaque scintillator)
  - (first) **absolute CC-neutrino flux reactor** — **new physics?**
  - **measure solar-pp ( $\geq 5\sigma$ )** in a tiny detector almost on the surface? ⇒ **a major breakthrough**
    - [backup] ⇒ explore **new physics?**
- **CLOUD-III:** **under feasibility study** (→ new **copper**-loaded opaque scintillator)
  - probe **reactor flux at low energies?** — **surprises?** [first time ever below 1.8MeV]
  - demonstration for  **$^{40}\text{K}$  detection methodology** — **a discovery one day?**

**an even vaster future of reactor (anti)neutrinos ahead?**

conclusions...

**SuperChooz demonstrator**



# our collaboration...

European  
Innovation  
Council



UK Research  
and Innovation

## CLOUD

### CLOUD International collaboration

- **EDF** (France) — **first time in neutrino science**
- **Brookhaven National Laboratory** (USA)
- **Charles University** (Czechia)
- **CIEMAT** (Spain)
- **IJCLab** / Université Paris-Saclay (France)
- **Imperial College London** (UK)
- **INFN-Padova** (Italy)
- **Instituto Superior Técnico** (Portugal)
- **Johannes Gutenberg Universität Mainz** (Germany)
- **Pennsylvania State University** (USA)
- **Pontifícia Universidade Católica do Rio de Janeiro** (Brazil)
- **Queen's University** (Canada)
- **Subatech / Nantes Université** (France)
- **Tohoku University / RCNS** (Japan)
- **Universidad de Zaragoza** (Spain)
- **Universidade Estadual de Londrina** (Brazil)
- **University of California Irvine** (USA)
- **University of Michigan** (USA)
- **University of Sussex** (UK)

⇒ 19 institutions in 11 countries

### Spokespersons:

- A. Cabrera — IJCLab / Université Paris-Saclay (France)
- J. Hartnell — Sussex University (UK)

### IB Chair:

- M. Chen — Queen's University (Canada)

### Webs:

<https://antimatter-otech.ijclab.in2p3.fr/> [AMOTech]

<https://liquido.ijclab.in2p3.fr/nucloud> [via LiquidO]