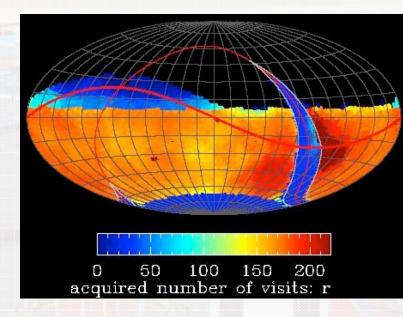
#### LSST Observing Strategy Željko Ivezić, University of Washington LSST Project Scientist AAS 227, January 2016







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An optical/near-IR survey of half the sky in ugrizy bands to r~27.5 based on ~1000 visits over a 10-year period:

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More information at www.lsst.org and arXiv:0805.2366

LSST: a digital color movie of the Universe...

A catalog of 20 billion stars and 20 billion galaxies with exquisite photometry, astrometry and image quality!

### **Main points**



- 1) The ultimate deliverable of LSST is not just the telescope, nor the camera, but the fully reduced science-ready data as well. No classical TAC.
- 2) The fundamental basis of the LSST observing strategy is to scan the sky wide, fast, and deep, and to obtain a dataset that simultaneously satisfies the majority of the science goals.
- 3) The LSST Science Advisory Council (with community input) is advising the Project on cadence-related questions. Also, a community white paper on observing strategy is in prep.

### **Observing Strategy: baseline**



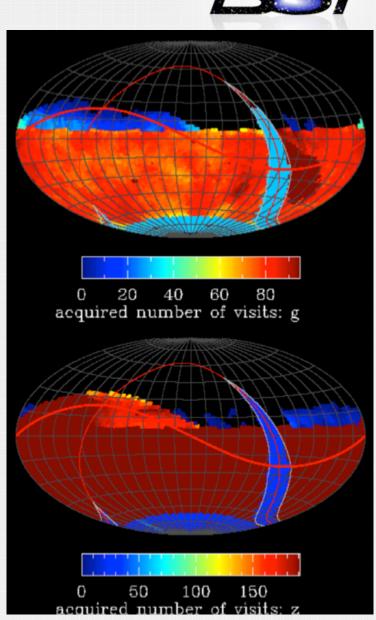
- 1) The main wide-fast-deep survey will use **about 90%** of the observing time and will be simultaneously optimized for the homogeneity of depth and number of visits, and for timedomain science (e.g., asteroids, supernovae, variable stars).
- 2) The remaining observing time will be used to obtain **improved coverage of parameter space** such as very deep observations, observations with very short revisit times, and observations of "special" regions (e.g., the Ecliptic, the Large and Small Magellanic Clouds).
  - For more details, see arXiv:0805.2366

## **Operations Simulations**

Observing constraints are provided by the astrophysical properties of the site (e.g. sky background), engineering models (e.g. settle time) and science requirements.

Operations simulator generates sequences of LSST observations together with their properties (seeing, sky brightness, depth, filter). About 2.5 million visits

(30 sec exposures) over 10 years.
For more technical details about
observing strategy, simulations and
cadence optimization, see talks by
Connolly, Yoachim and Ivezic at
http://ls.st/kaq



# Simulated surveys are public



Setup	Simulation Name	Description of the Survey Setup
0	enigma_1189	<u>Modern Version of the Baseline Cadence</u> A candidate replacement simulation for the current Baseline Cadence (opsim3.61) produced with the latest version (v3.2.1) of the Operations Simulation (OpSim) code. The following adjustments have been made: includes Science Council approved Deep Drilling fields; Wide-Fast-Deep (WFD) design specification for areal coverage (18,000 deg) & WFD "boosted visits" = 75, 105, 240, 240, 210, 210 for u, g, r, i, z, & y filters where g, r, i and z visits are collected in pairs separated by about 30 minutes; includes revised scheduled downtime as well as random downtime; minAlt = 20 deg; MinDistance2Moon = 30 deg. Note that SRD design visits = 56, 80, 184, 184, 160, 160 for u, g, r, i, z, & y filters.
1	<u>9ps2_1098</u>	Uniform cadence (WFD), which asks for visits in pairs, and no other proposal.
2	9 <u>ps2</u> 1093	Only uniform cadence (WFD), but does not require pairs of visits.
3	kraken_1033	As the baseline cadence (Setup 0), but does not require pairs of visits.
4	enigma_1271 enigma_1266	As the baseline cadence, but requests 3 visits per Wide-Fast-Deep field chosen instead of 2 visits, using the same window function for both 1-2 visits and 2-3 visits. As the baseline cadence, but requests 4 visits per Wide-Fast-Deep field.
5	kraken_1034	As the baseline cadence, except that the u-band exposure time is 60 sec instead of 30 sec.; Nyisit for the u-band remains the same.
6	kraken_1035	As the baseline cadence, except that the u-band exposure time is 60 sec instead of 30 sec; Nyisit for the u-band is decreased by a factor of 2.
7	kraken_1036	As the baseline cadence, except for a shorter visit exposure time: 20 sec instead of 30 sec. Deep drilling proposal has visits based on 30sec exposure due to code issues.
8	kraken_1037	As the baseline cadence, except for a longer visit exposure time: 60 sec instead of 30 sec.
9	<u>9252_</u> 1092	Pan-STARRS-like Cadence This is the uniform cadence, and no other proposal, keeping pairs of visits, but increase the area to include everything with Dec <+15 deg (about 27,400 deg2), and keeping the default <u>airmass</u> limit of 1.5.
10	kraken_1038	As the baseline cadence, except for the more relaxed airmass limit of 2.0 instead of 1.5.
11	<u>9252_1096</u>	As Setup1 (uniform cadence with no other proposal), except for the more relaxed airmass limit of 2.0 instead of 1.5.
12	ops2_1097	As Setup 1 (uniform cadence with no other proposal), except for the more stringent airmass limit of 1.3 instead of 1.5.

For each visit: RA,Dec Filter **MJD** ExposureTime Seeing Airmass Sky brightness 5-sigma depth and ~100 more...



## **Ongoing optimization work**



- minimizing the impact of read-out noise (mostly in u band)
- optimizing sky coverage (Galactic plane, south celestial pole, LMC/SMC, Ecliptic)
- temporal sampling (SNe, variable stars, asteroids)
- interplay between sky coverage and temporal sampling
- deep drilling fields
- dynamic cadence (in response to expected SNR)
- evolving cadence (in response to changing science drivers)
   LSST Science Advisory Council (SAC)
- the main mechanism for officially collecting and delivering community (your!) input to the Project.
- For existing input from the community, see <a href="http://ls.st/smg">http://ls.st/smg</a>

A white paper by the community: on arXiv by August 2016 Go to community.lsst.org to participate!