### Atmospheric Chemistry Experiment (ACE): Recent Results

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## **ACE Satellite**

#### See http://www.ace.uwaterloo.ca



#### Advantages:

Radiance of sun gives higher S/N than emission

Limb view gives longer path length  $\sim$ 500 km (lower detection limits) than nadir

"Self-calibrating" so excellent long-term accuracy and precision

#### **Disadvantages**:

Modest global coverage Samples only free troposphere

## Timeline

-	Jan.	1998	Proposal to CSA
•	Feb.	2001	FTS and Imager CDR
•	Mar.	2001	MAESTRO CDR
•	Jun.	2001	Bus CDR
•	Sept.	2002	S/C integration & test
•	Mar.	2003	Instrument test (Toronto)
•	Мау	2003	Final integration (DFL)
•	Aug.	2003	Launch
•	Sept.	2003	Commissioning
•	Feb.	2004	Routine operations
•	Mar.	2004	Arctic campaign
•	Feb.	2005	Arctic campaign
•	Feb.	2006	Arctic campaign
•	Feb.	2007	Arctic campaign
•	Feb	2008	Arctic campaign

First ACE data Feb. 2004, mission currently approved to March 2010. Mission had a 2year lifetime – fifth anniversary Aug. 2008.

## Instruments

- Infrared Fourier Transform Spectrometer operating between 2 and 13 microns with a resolution of 0.02 cm<sup>-1</sup>
- 2-channel visible/near infrared Imagers, operating at 0.525 and 1.02 microns (cf., SAGE II)
- Suntracker keeps the instruments pointed at the sun's radiometric center.
- UV / Visible spectrometer (MAESTRO) 0.4 to 1.03 microns, resolution ~1-2 nm
- Startracker

Bernath et al. GRL, 32, L15S01 (2005)

### SCISAT-1 Spacecraft (Bristol)



## **Optical Layout (ABB-Bomem)**



MAESTRO PI: T. McElroy, MSC



UV 270-540 nm

MAESTRO

Vis 520-1040 nm

**Dump/Alignment** 

Port

90/10 Beamsplitter

Dual concave grating spectrograph, 1-2 nm resolution

70/30 Beamsplitter

Input

Port

# **Global Occultation Distribution**

#### 2005-10-08 18:00:00+00

## FTS – Decontamination Results



**Ryan Hughes** 

### **Occultation sequence**



Atmospheric Transmittance

# **ACE-FTS Species Measured**

• Baseline species (version 2.2):

 $H_2O$ ,  $O_3$ ,  $N_2O$ , CO,  $CH_4$ , NO,  $NO_2$ ,  $HNO_3$ , HF, HCI,  $N_2O_5$ ,  $CIONO_2$ ,  $CCI_2F_2$ ,  $CCI_3F$ , as well as pressure and temperature from  $CO_2$  lines

- Other routine species: COF<sub>2</sub>, CHF<sub>2</sub>Cl, CF<sub>4</sub>, CH<sub>3</sub>Cl, C<sub>2</sub>H<sub>6</sub>, SF<sub>6</sub>, OCS, HCN
- *Research* species:

CCl<sub>4</sub>, HOCl, H<sub>2</sub>O<sub>2</sub>, HO<sub>2</sub>NO<sub>2</sub>, CCl<sub>2</sub>FCClF<sub>2</sub>, CH<sub>3</sub>CClF<sub>2</sub>, ClO, C<sub>2</sub>H<sub>2</sub>, C<sub>2</sub>H<sub>6</sub>, COFCl, COCl<sub>2</sub>, CH<sub>3</sub>OH, HCOOH, H<sub>2</sub>CO, N<sub>2</sub> and additional isotopologues

## **CO<sub>2</sub> line near 61 km**

Typical ACE-FTS fitting results

Note: results are plotted on the raw measurement grid

Boone et al. Appl. Opt. 44, 7218 (2005)







CFC-11, CFC-12, HCFC-22, CCl<sub>4</sub>, CH<sub>3</sub>Cl, CF<sub>4</sub>, CFC-113, HCFC-142b, HFC134a, F<sub>2</sub>CO, CIFCO, Cl<sub>2</sub>CO



50

0

C,H, C,H, CO HCN

CH<sub>3</sub>Cl

OCS

species - lifetime (years)

HCFC22

CFC11CFC12



## **Global Distribution of Phosgene,** Cl<sub>2</sub>CO

C

Fu et al. GRL, 34, L17815 (2007) (Toon et al.'s calculated linelist for  $v_5$  near 850 cm<sup>-1</sup>)





# **Distribution of COCIF**

- Carbonyl chlorofluoride is a product of chlorofluorocarbon (CFC-11 mainly) decomposition
- Previously studied by aircraft (5 - 12 km)
- First global picture obtained from ACE-FTS
- Spectroscopy based on Brown's ATMOS linelist created from Kitt Peak spectra, with rough intensities.



D. Fu et al., to be submitted

## **Effect of Lightning (HNO<sub>3</sub>)**

Lightning produces NO, which is oxidized to HNO<sub>3</sub>.

Need to have 6 Tg N/yr from lightning to match ACE observations of tropospheric HNO<sub>3</sub> (Martin et al. JGR, 112, D09309 (2007))











ACE solar spectrum (F. Hase): 224782 spectra added, improvement over ATMOS, no telluric lines, but 0.02 cm<sup>-1</sup> vs 0.01 cm<sup>-1</sup> resolution (resolution largely determined by width of solar lines) and 750-4400 cm<sup>-1</sup> vs 600-4800 cm<sup>-1</sup>.

New atomic and molecular assignments (ACE linelist) by L. Wallace (NOAO); improved spectroscopic data for CH, NH and OH.

For OH, Reg Colin (ULB) finds v=4 can be improved.



### **Air Quality and Biomass Burning**



The Global Impact of Biomass Burning

An Interview with EPA'S Robert Huggett

Wanted: Environmental Chemistry Gradua<u>tes</u>

## **Biomass Burning in Brazil**

NOAA HYSPLIT MODEL Backward trajectories ending at 15 UTC 02 Oct 04 CDC1 Meteorological Data



### H<sub>3</sub>OH contribution to the spectrum



### **Global Methanol**

ACE is an upper tropospheric "air quality" mission measuring global  $CH_4$ ,  $CH_3OH$ , HCN,  $C_2H_2$ ,  $C_2H_6$ ,  $H_2O_2$ , HCOOH,  $H_2CO$ , plus likely PAN and acetone.



Dufour et al. ACP, 7, 6119, 2007 LDMz-INCA model (D. Hauglustaine)

## **Young Biomass Burning Plume**



### PAN, Peroxyacetyl nitrate, etc.

Coheur et al. (Brussels), PAN from a biomass plume near East Africa



## **HCHO spectroscopy: new linelist**

#### Use of HCHO line intensities calculated by A. Perrin

Dufour et al.



Perrin, Jacquemart, Kwabia-Tchana & Lacome

## ICHO contribution to the spectrum

6 spectral windows selected in the range 2735 - 2830 cm<sup>-1</sup> :

2739.85 ; 2765.65 ; 2778.4 ; 2781.2 ; 2812.25 ; 2826.67





#### years of HCHO measurements with ACE-FTS

ACE-FTS - 8.5 km





## Preliminary comparisons with CTMs



Comparison with 2 state-of-theart models (LMDz-INCA and GEOS-Chem) that use different emissions inventories.

#### North Hemisphere:

seasonality of UT HCHO well reproduced intensity of the maximum not

always reproduced

#### **South Hemisphere**:

LMDz-INCA systematically smaller

#### South Tropics:

small impact of biomass burning larger variability in the models

## **ACE** Partners (Selected)

- Canada- K. Walker, J. Drummond, K. Strong, J. McConnell, W. Evans, T. McElroy, I. Folkins, R. Martin, J. Sloan, T. Shepherd, etc.
- USA- NASA launched ACE: C. Rinsland, L. Thomason (NASA-Langley), C. Randall (U. Colorado), B. Bojkov (NASA-Goddard), M. Santee, L. Froidevaux, G. Manney (JPL), etc.
- Belgium- supplied CMOS imager chips: R. Colin, P.-F. Coheur, M. Carleer (ULB), D. Fussen, M. DeMaziere (IASB), M. Mahieu, R. Zander (Liege), etc.
- UK- J. Remedios (Leicester), P. Palmer (Edinburgh), M. Chipperfield (Leeds)
- France- C. Camy-Peyret, C. Clerbaux, C. Brogniez, G. Dufour, D. Hauglustaine (Paris)
- Japan- M. Suzuki, Y. Kasai (JAXA)
- Sweden- G. Witt (Stockholm)



## Sunset over Kitt Peak, AZ