



Can AI help users find data?

The experience with the ESO Science Archives

Martino Romaniello (ESO)

Nima Sedaghat , Felix Stoehr, Jon Carrick, FX Pineau

Vojtech Cvrcek, Wolfram Freudling, Pascal Ballester, Radim Sara

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CONTEXT: THE PROBLEM



The context: The ESO Science Archives

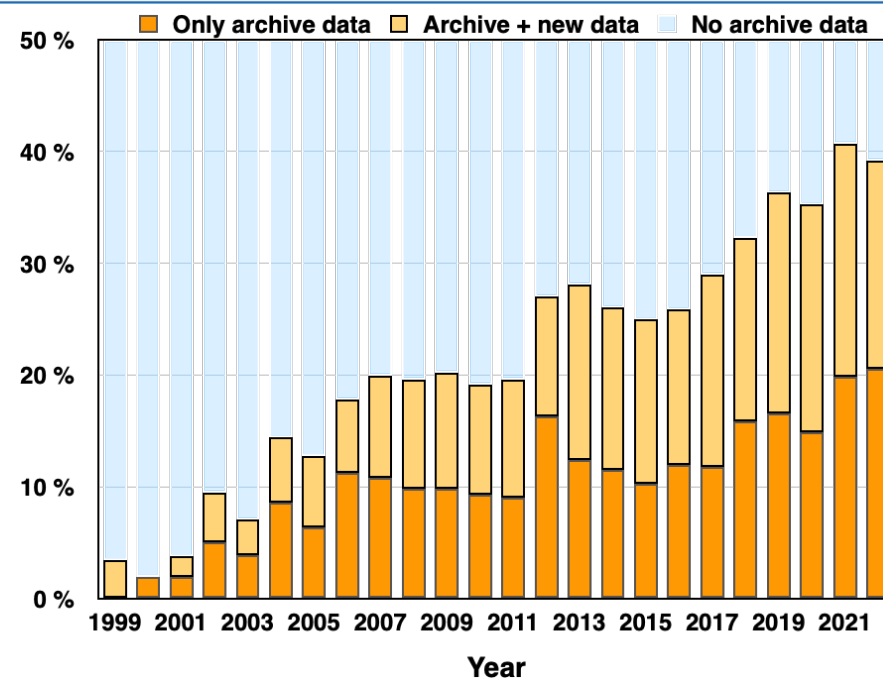
- Very substantial contributors to ESO's science output: e.g. 40% of VLT publications

- Millions of science files, tens of millions of metadata items

- It is key to present and characterize the data in a language that speaks to users

- Sky position
- Instrument description (setup, ...)
- Data description (SNR, resolution, depth, ...)

- The next step: characterization by source properties (object type, redshift, chemical composition, ...) and/or by similarity



Source: telbib.eso.org



Deep Learning on the ESO Science Archives - I

Goals

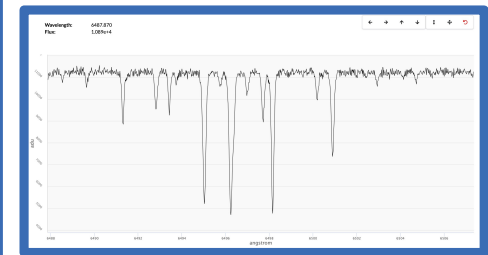
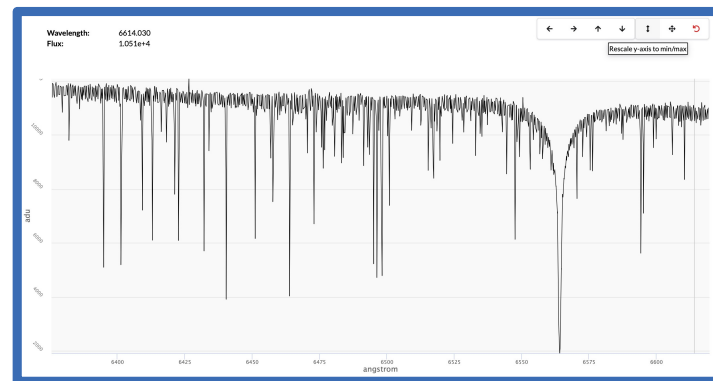
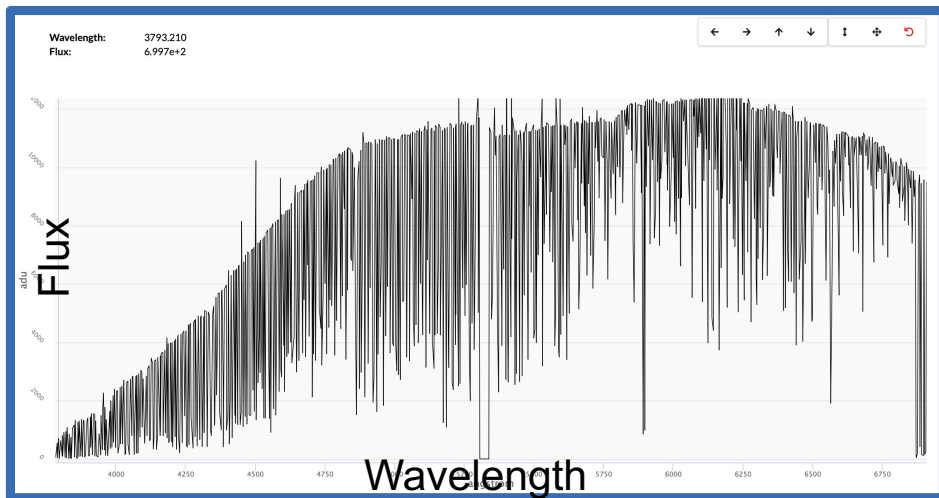
- Assess whether AI is useful in providing users with novel ways to identify data in the ESO Science Archives
 - Starting point: processed data
- Data is very heterogeneous, e.g. La Silla Paranal processed data:
 - 3.6 million files
 - 28 instruments
 - 71 data collections, 56 data providers
 - 3000 PIs, 9000 individual programmes
- Large and varied user base
 - More than half of professional astronomers worldwide
- Strive to limit the imposition of preconceived categories and criteria
- Results should be robust, understood, reproducible and user-friendly



Deep Learning on the ESO Science Archives – II

The HARPS experiment

- Deep Learning analysis of the entire HARPS archive
 - High-resolution, high-stability spectrograph
 - Relatively clean sample: mostly stars in the solar neighborhood
 - Data readily available
 - 1D spectra, processed in physical units (wavelength vs flux) to high accuracy and uniformity
 - ~270k spectra, ~300k wavelengths channels each





Sedaghat, MR, Carrick, Pineau 2021, MNRAS, 501, 6026

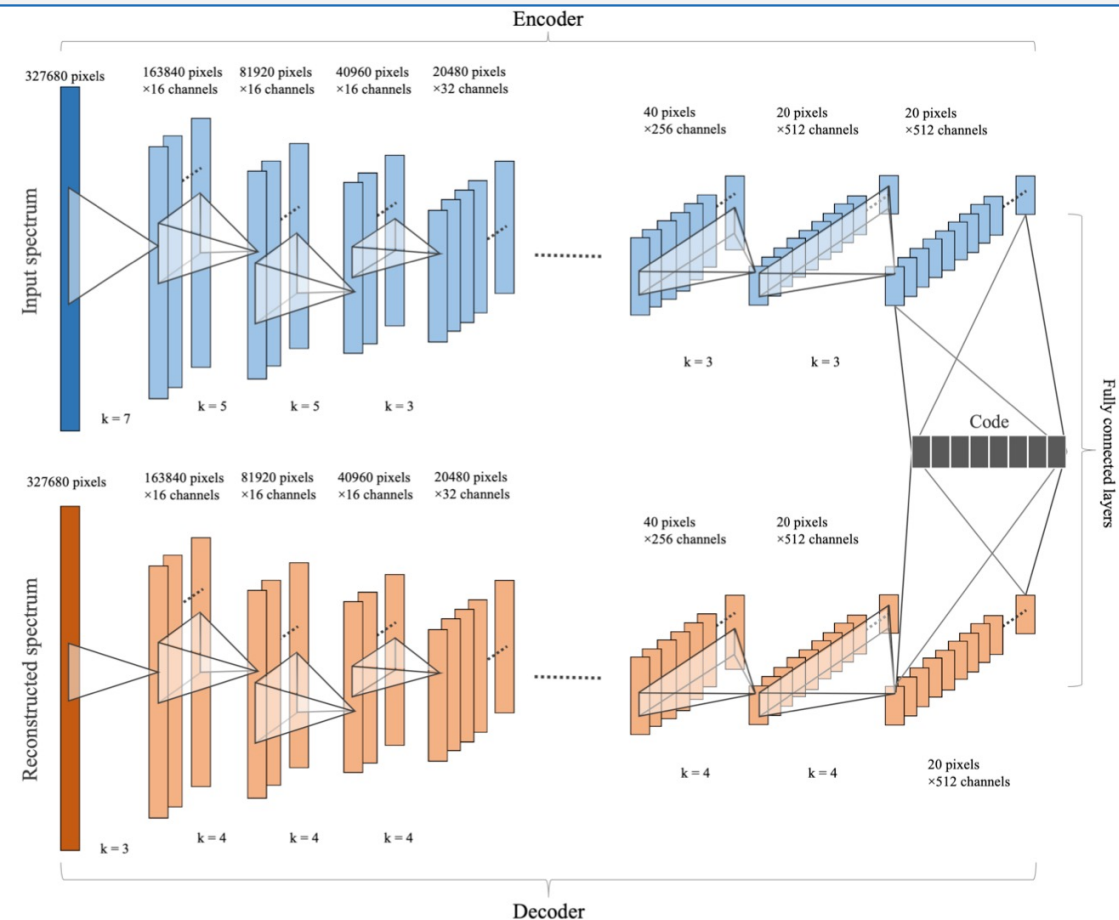


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APPROACH 1: UNSUPERVISED LEARNING

Unsupervised learning: Network schematics

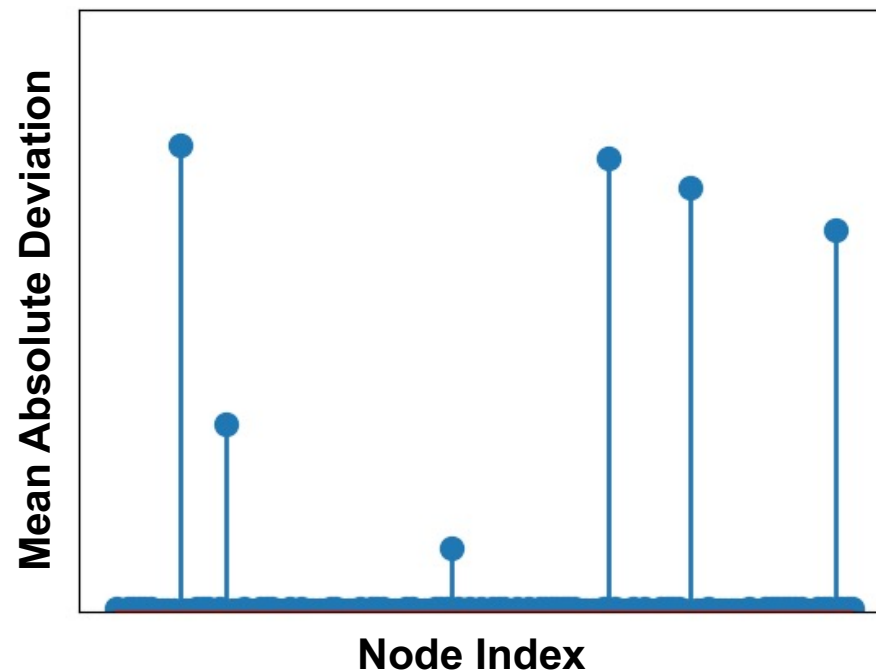
- Variational AutoEncoder
- Loss function: L1 norm
- Disentanglement for interpretation of latent space dimensions
- Understand what the networks "learned" in latent space (Code)
- Distances in the latent space for searches based on similarity



Sedaghat, MR, Carrick, Pineau 2021, MNRAS, 501, 6026

Unsupervised learning: Results - I

- 128 latent dimensions needed for good reconstruction
- Not all of them carry significant information
 - In fact, only 6 out of 128 do
- So, what are they? Do they have a physical interpretation?

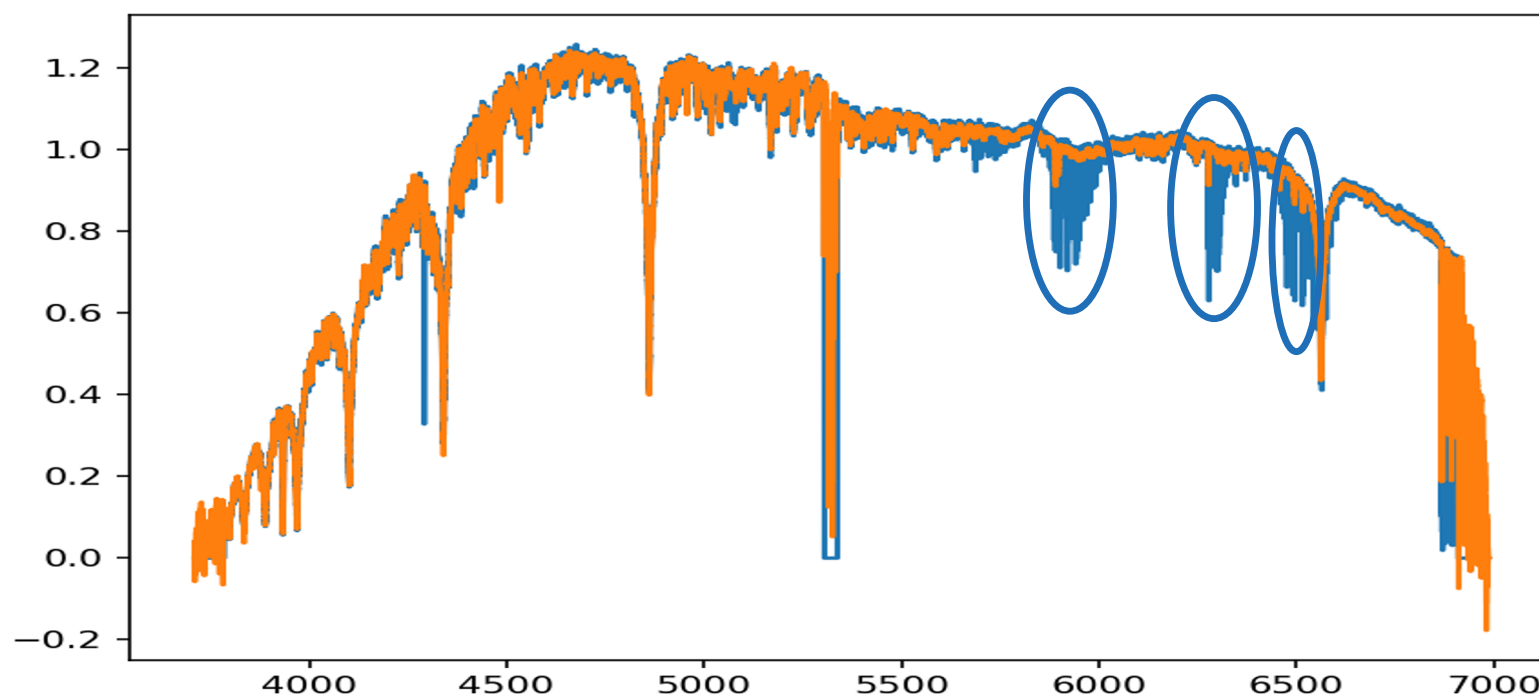


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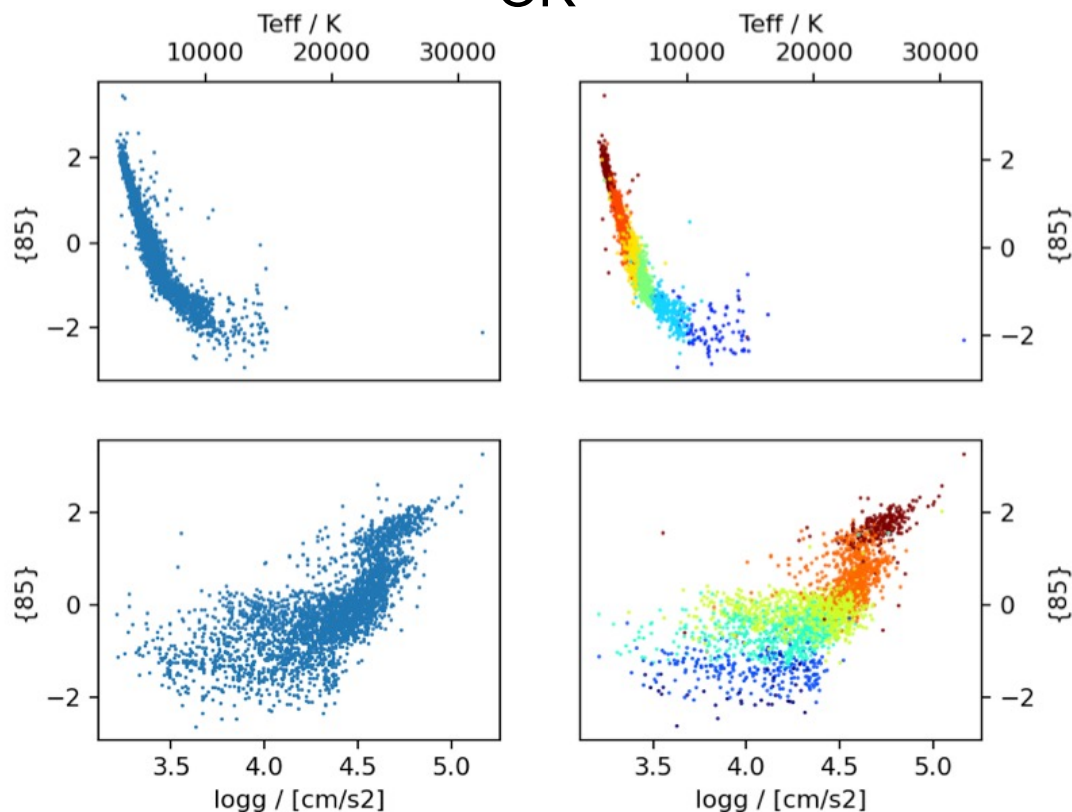
Unsupervised learning: Tantalizing hints

- Tendency to separate stellar vs Earth atmosphere features
 - In any reference frame, except topocentric

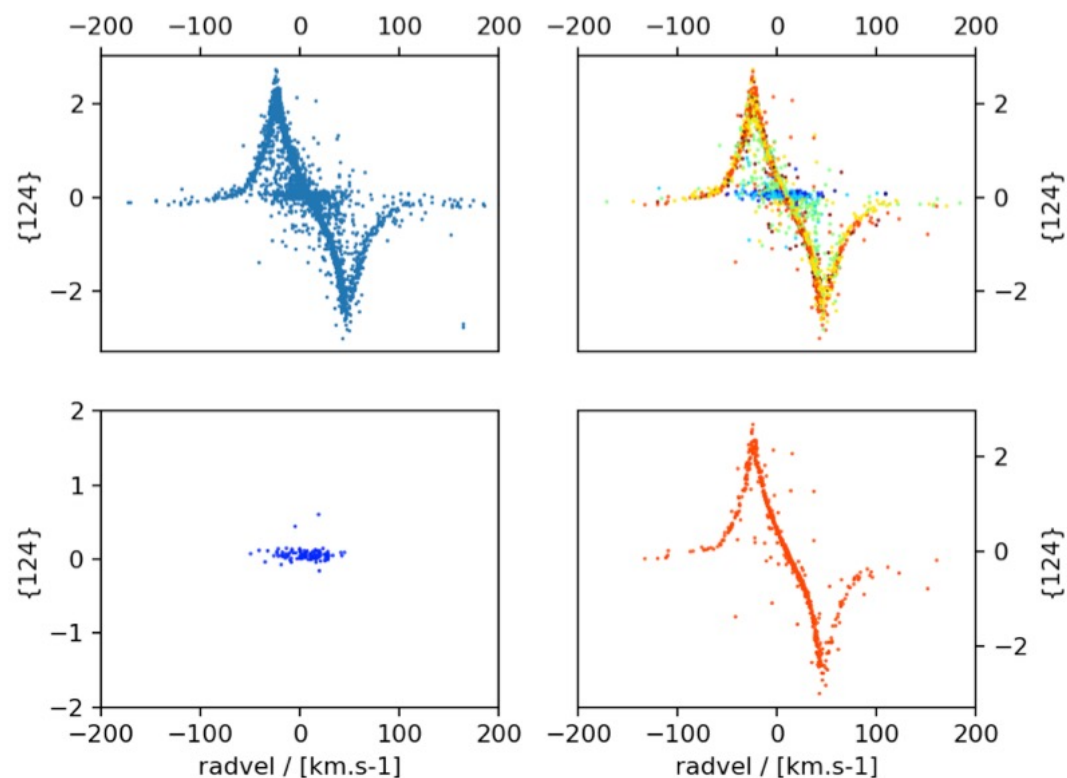


Unsupervised learning: Puzzling results - I

OK



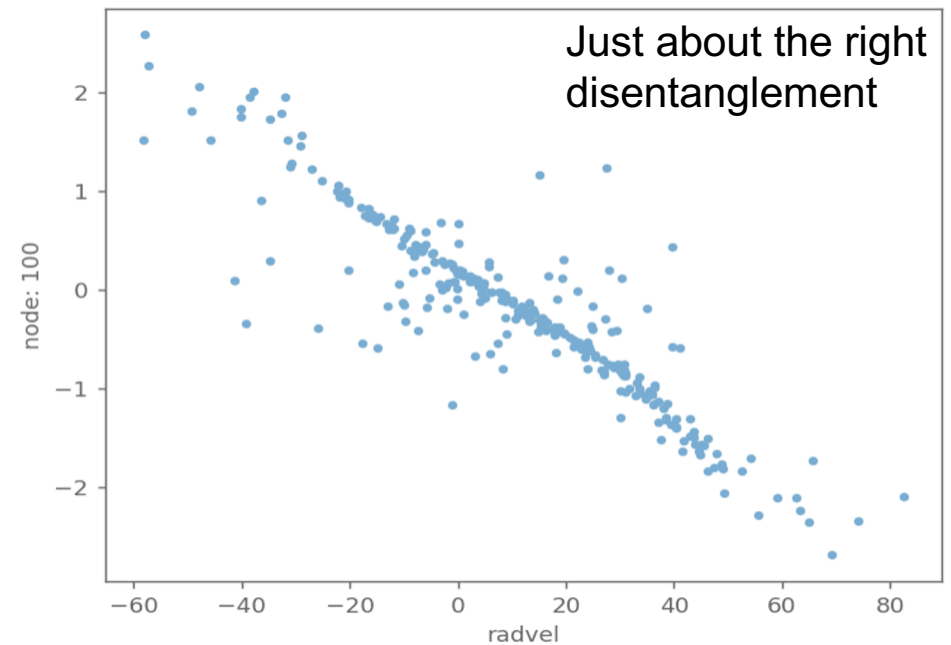
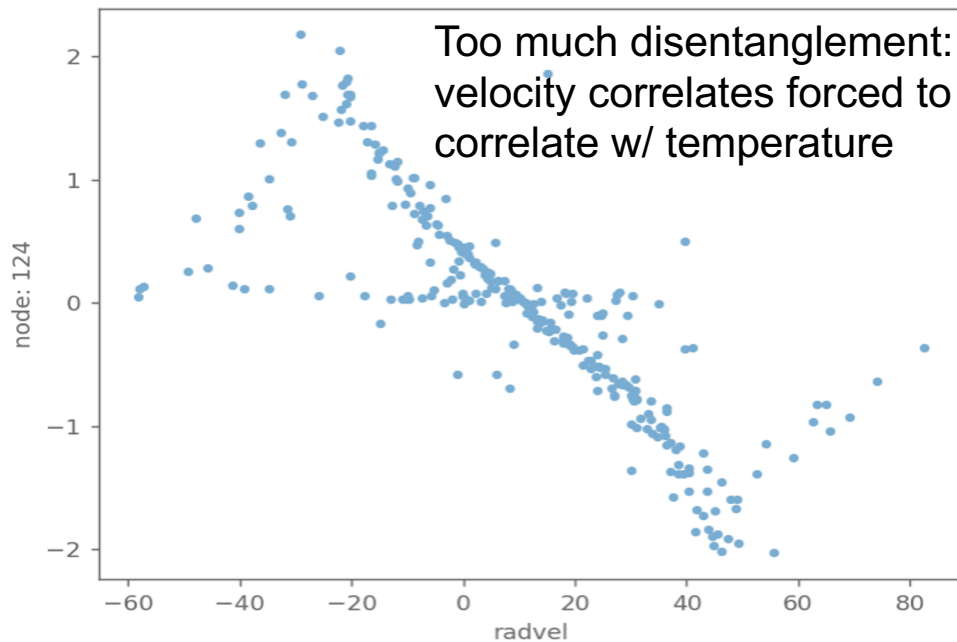
Non-monotonic?!?



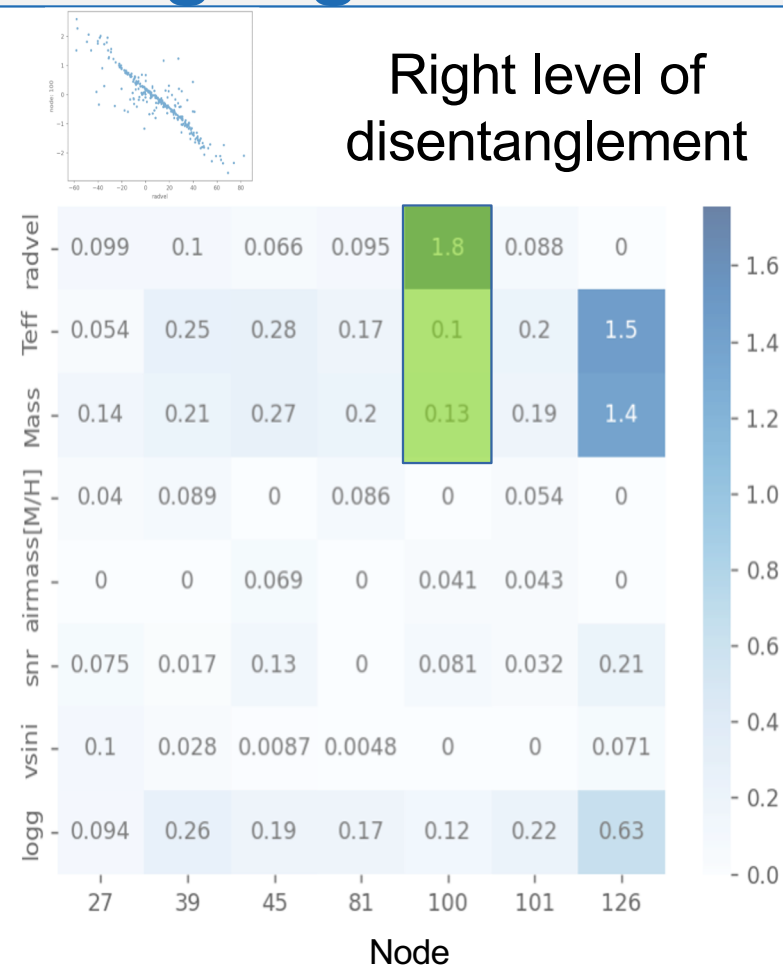
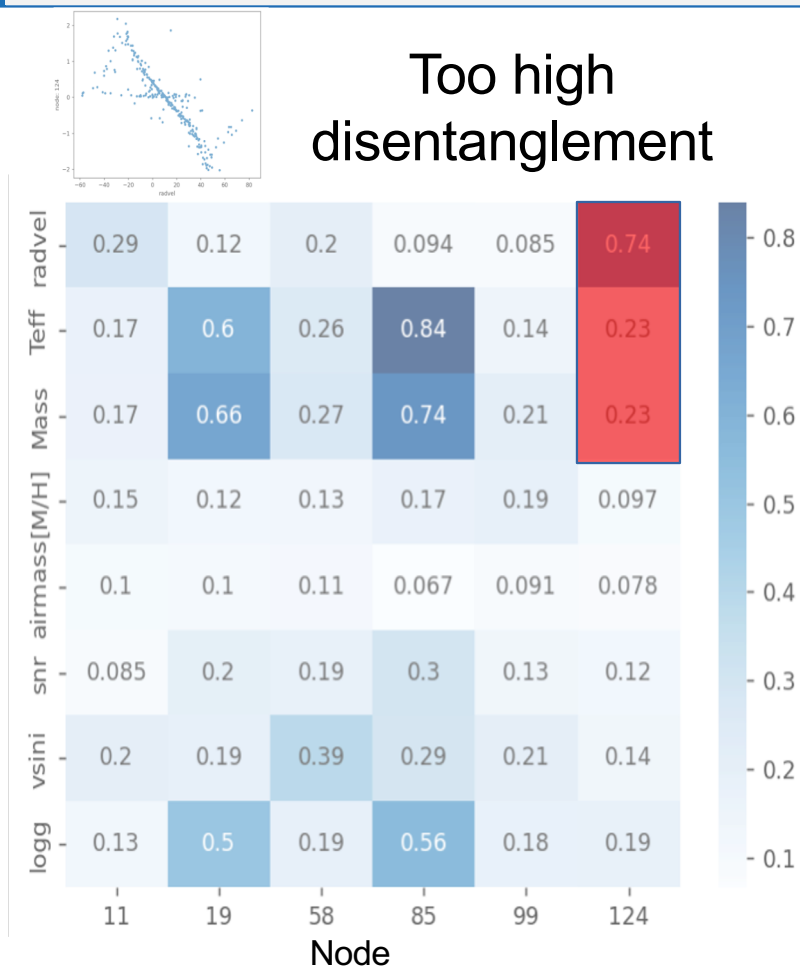
Sedaghat, MR, Carrick, Pineau 2021, MNRAS, 501, 6026

Unsupervised learning: Fine tuning disentangling - I

- Too high disentanglement penalty: latent nodes forced to correlate with multiple uncorrelated physical variables
- Too low disentanglement penalty: nodes are entangled



Unsupervised learning: Fine tuning disentangling - II



Unsupervised learning: Provisional summary

- No labels used in the training, checked a-posteriori for interpretability
- Only a handful of latent space dimensions carry significant information
 - Some of them relate directly to physical parameters of the stars ...
 - Effective Temperature, surface gravity, radial velocity
 - ... but NOT ALL
 - No Mutual Information between chemical composition and nodes; some nodes unexplained
- Disentanglement needs tuning to be effective
- Physically correlated quantities remain so in the latent space (e.g., effective temperature, surface gravity, mass)
 - Problem for interpretability of, e.g., archive queries based on similarity
- Interesting features, but not quite ready for primetime



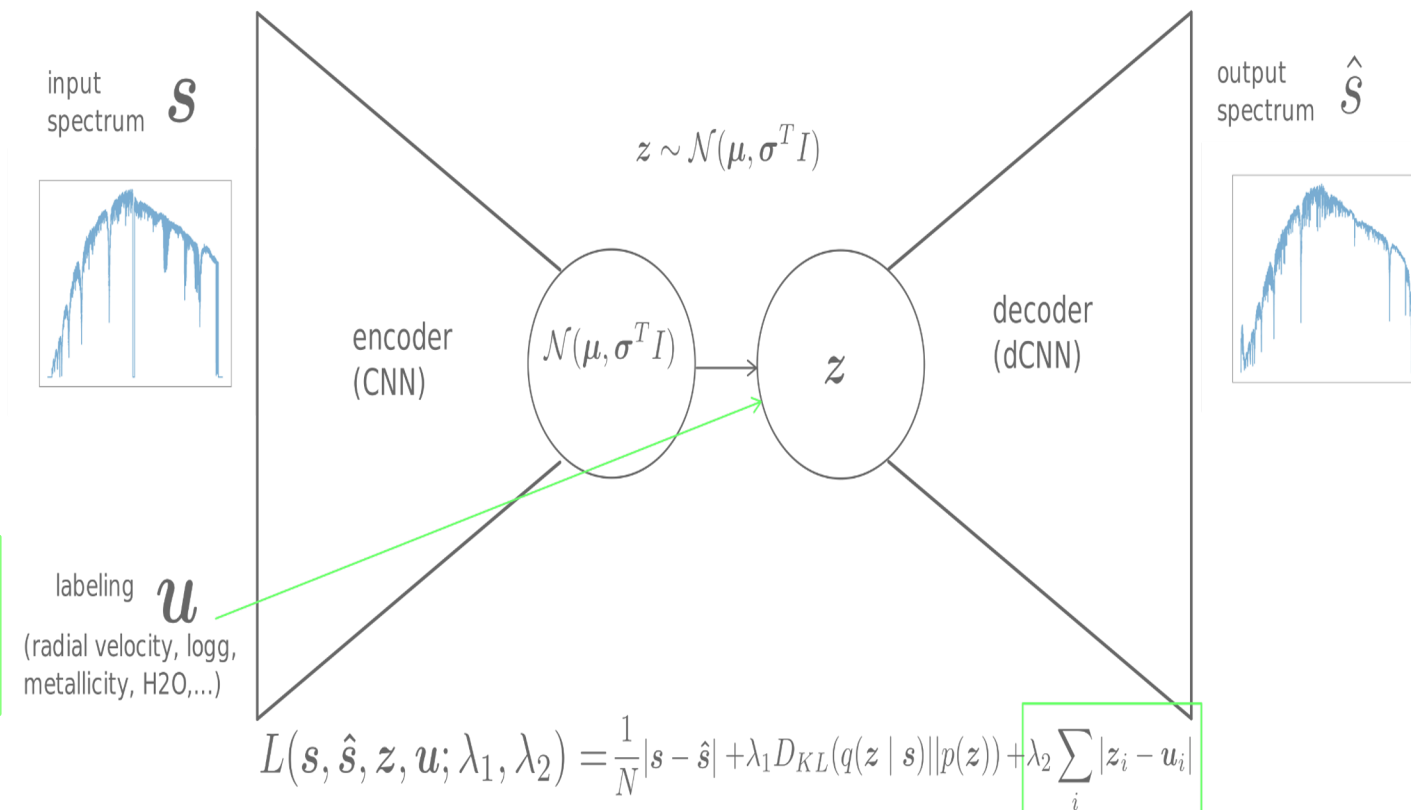
Very much work in progress ...

APPROACH 2: WEAKLY SUPERVISED LEARNING

Weakly supervised learning: Network schematics

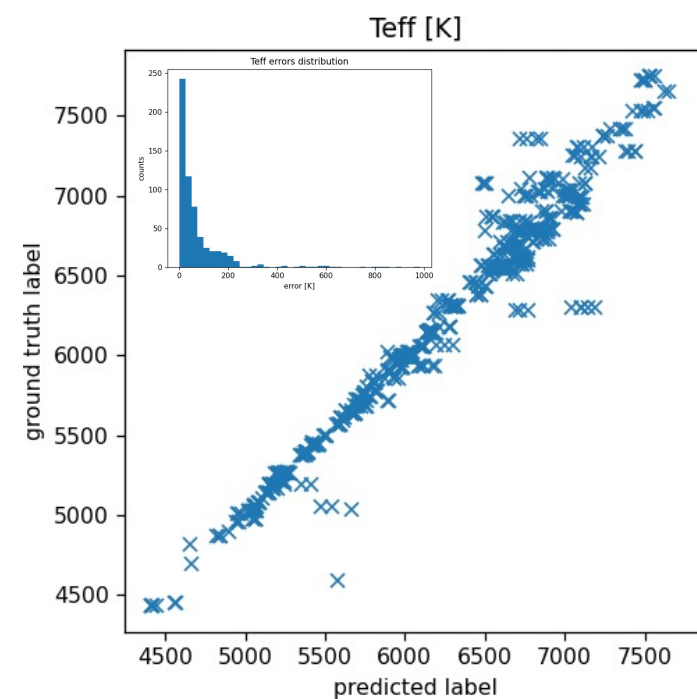
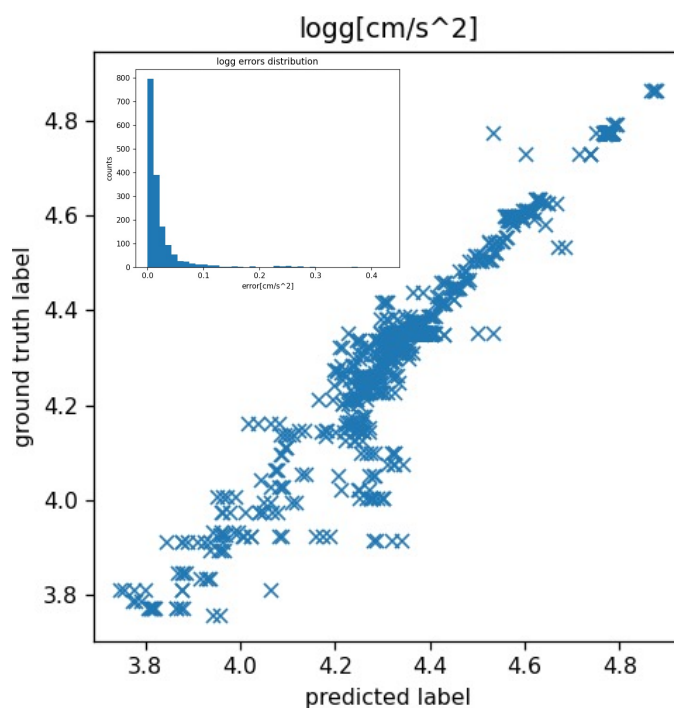
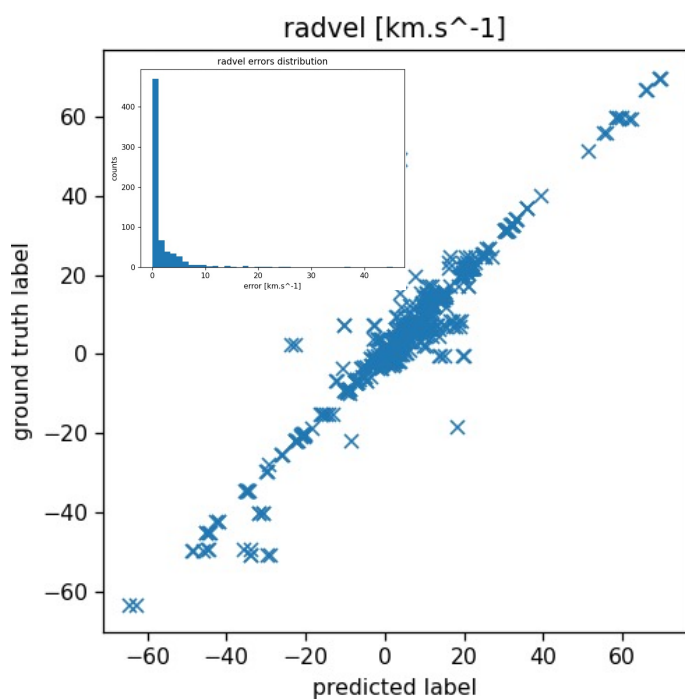
- Labels used for training, alongside reconstruction losses

Labels include source and Earth atmosphere parameters



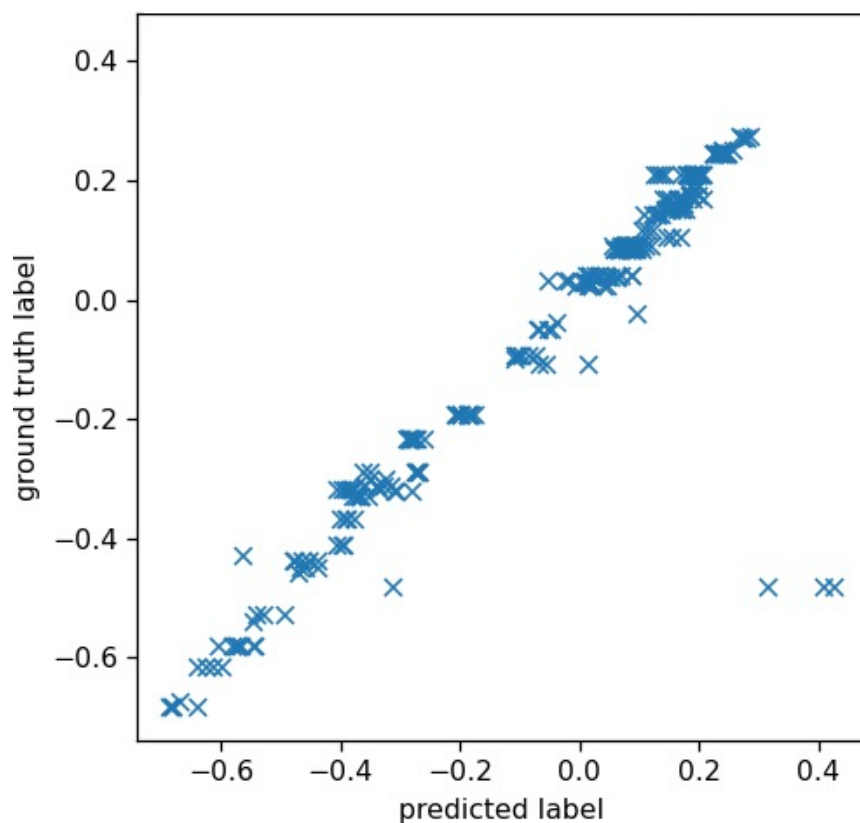
Weakly supervised learning: Results - I

■ Rather good reconstruction of the labels

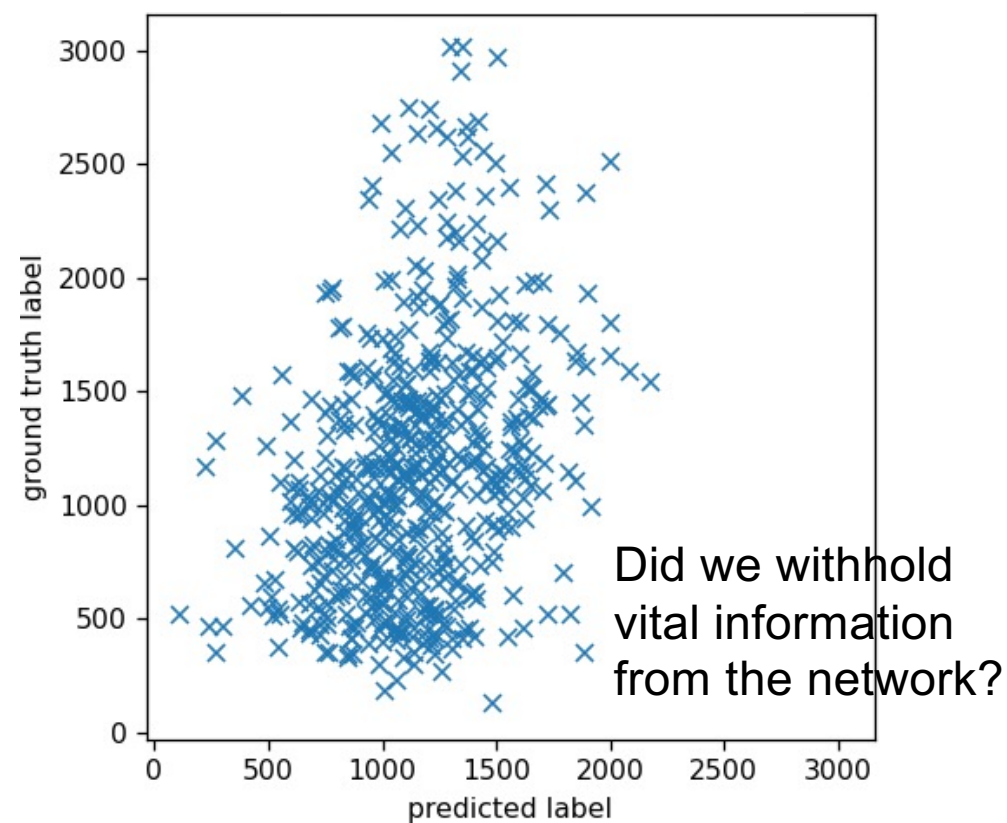


Weakly supervised learning: Results - II

Stars' chemical composition

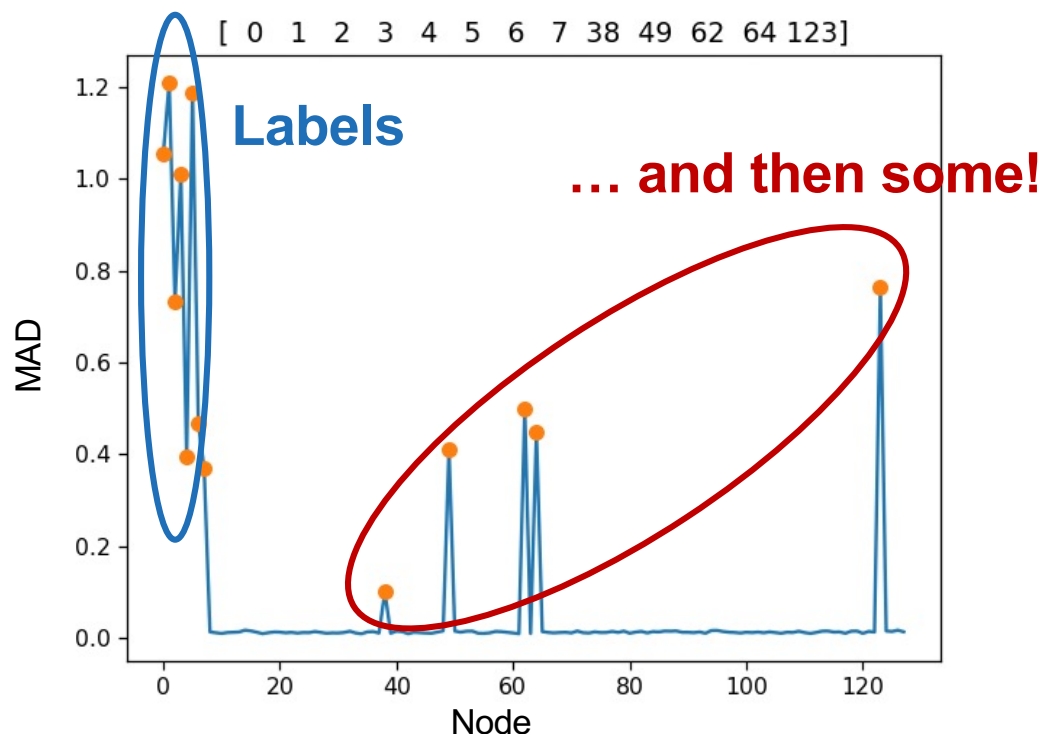


Earth's atmosphere

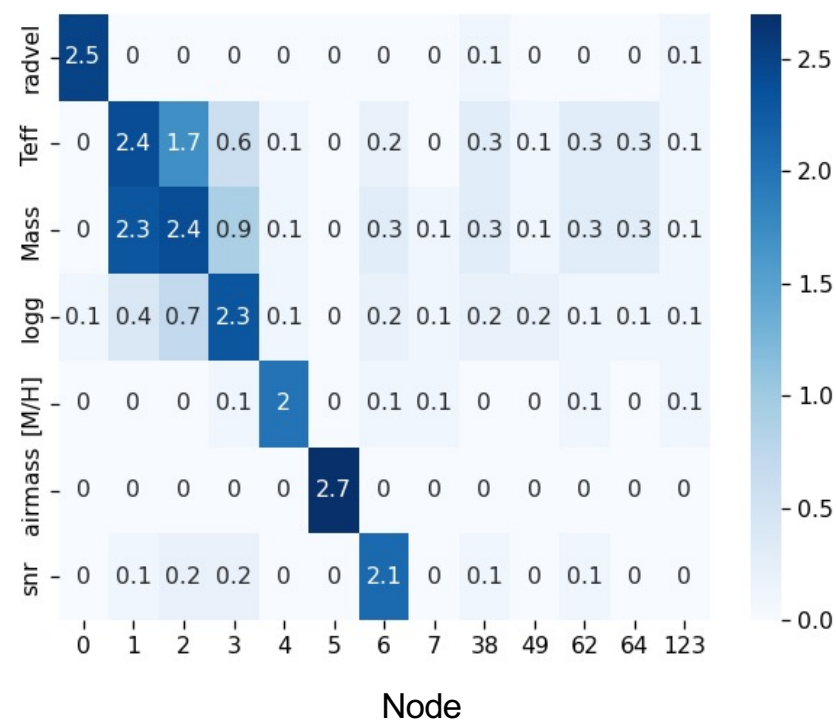


Weakly supervised learning: Results - III

■ The network learns the labels,
and then some



■ Label nodes are well
disentangled



Weakly supervised learning: Provisional summary

- Labels use in training alongside reconstruction losses
- Reconstruction of labels is solid (“supervised is easy”, ©Maggie Lieu)
 - BUT, several significant dimensions in addition: what are they?
- Do we have reliable sources of labels for all the cases?
- What to do for diverse samples with disjoint sets of parameters, e.g., stars and galaxies and QSOs and ...,?
 - Pre-classification? (Cf. Caroline Heneka)
 - Spars(er) label matrix?
- Simulations may help
 - We are after physics, after all
 - Reinforcement learning (Cf. Maxime Quesnel’s talk with simulator as decoder)
 - Domain adaptation

Provisional conclusions

- We are running an experiment to extract physical parameters from massive dataset to build new query capabilities for archive research
 - Still very much work in progress
- The purely unsupervised approach has issues if interpretability in terms of the object's physical parameters is desired
 - Interpretability is important to present results to the intended broad and diverse audience of archive users
- The weakly supervised approach is promising in that sense, but brings the question of quality and availability of labels
 - Which anyhow affects the unsupervised approach, where labels are needed to validate the interpretability
- Simulations may help both approaches
 - WIP ...

