

Atmospheric neutrino oscillations with Super-Kamiokande and prospects for SuperK-Gd

*Pablo F. (U. of Liverpool) for the Super-Kamiokande Collaboration
2021/02/24, The Neutrino Telescopes Workshop*

The Super-Kamiokande (SK) Experiment

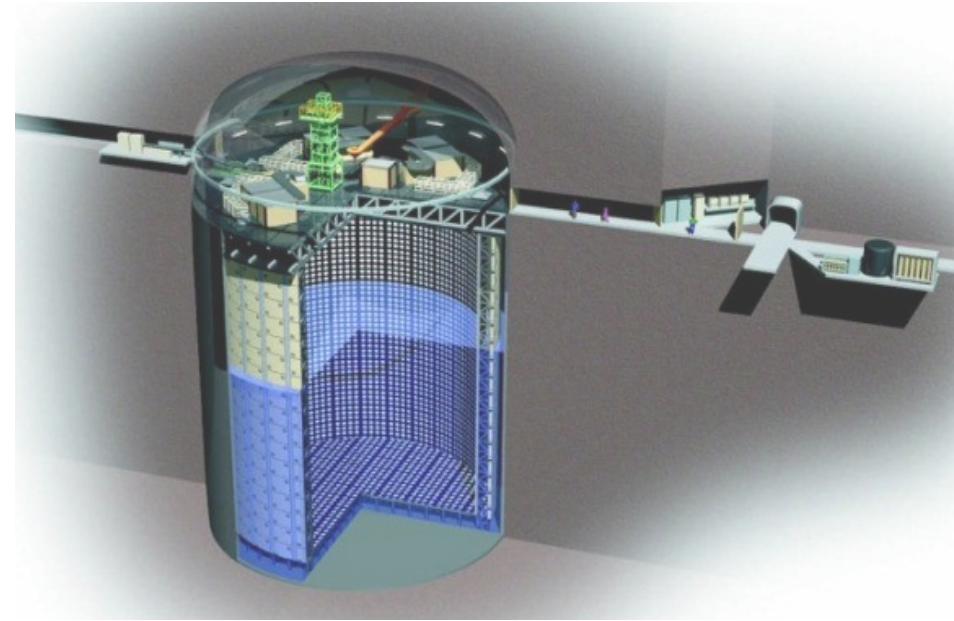
- Water-Cherenkov detector
- Located in Kamioka, Japan
- Under Mt. Ikenoyama
- Overburden 1km of rock
- Total of 50 kton of ultra-pure water, and **22.5 kton fiducial volume**
- Optically divided into inner (ID) and outer (OD) detectors, instrumented with
 - **ID:** ~11000 20"-PMTs → 40% photo-coverage
 - **OD:** ~2000 8"-PMTs mainly used as veto



The Super-Kamiokande Experiment

- More than 20 years of operation and data taking

	Period	Event
SK-I	1996 to 2001	Start
SK-II	2003 to 2005	20% PMT coverage after accident
SK-III	2006 to 2008	Resume 40% PMT coverage
SK-IV	2008 to 2018	Electronics upgrade
SK-V	2019 to 2020	Upgrade for Gd-loading
SK-Gd1	2020 to...	Largest Gd-doped WC detector



- Wide variety of fundamental physics over a wide range of energies:
 - Solar, **atm.**, LBL, SN and astrophysical vs, proton and exotic nucleon decays, dark matter
- Still at the forefront with its latest upgrade, **SuperK-Gd**
 - Will eventually reach a concentration of 0.2% of Gd, capturing 90% of the neutrons
- Even richer physics capabilities:
 - First measurement of DSNB
 - Search for solar antineutrinos
 - Background reduction for proton-decay
 - **Improved distinction of ν s and $\bar{\nu}$ s at higher energies**

The Super-Kamiokande Experiment

e-like

Super-Kamiokande IV

Run 65720 Sub 275 Event 51523895

09-10-19:19:59:15

Inner: 1602 hits, 2951 pe

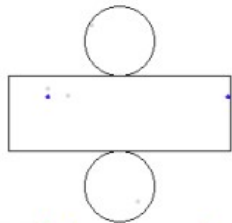
Outer: 2 hits, 2 pe

Trigger: 0x10000007

D_wall: 521.0 cm

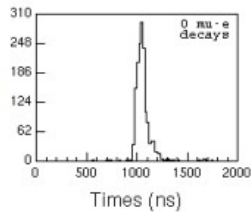
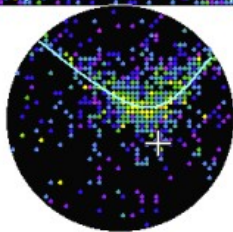
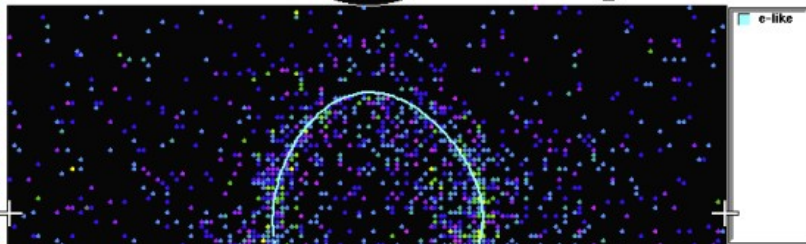
E_vis: 294.0 MeV

e-like, $p = 294.0$ MeV/c



Charge(pe)

- * >26.7
- * 23.3-26.7
- * 20.2-23.3
- * 17.3-20.2
- * 14.7-17.3
- * 12.2-14.7
- * 10.0-12.2
- * 8.0-10.0
- * 6.2-8.0
- * 4.7-6.2
- * 3.3-4.7
- * 2.2-3.3
- * 1.3-2.2
- * 0.7-1.3
- * 0.2-0.7
- * < 0.2



μ -like

Super-Kamiokande IV

Run 65718 Sub 22 Event 3976280

09-10-19:14:35:52

Inner: 3389 hits, 17883 pe

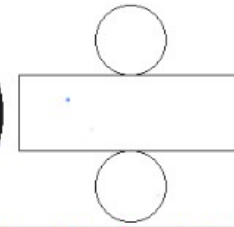
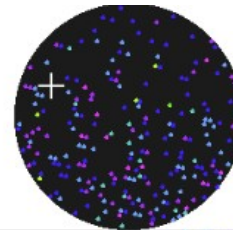
Outer: 1 hits, 1 pe

Trigger: 0x10000007

D_wall: 327.7 cm

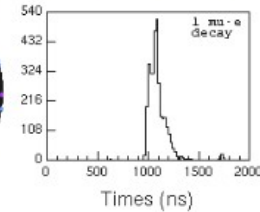
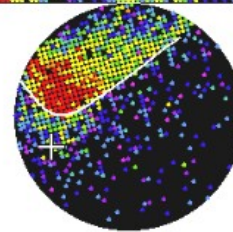
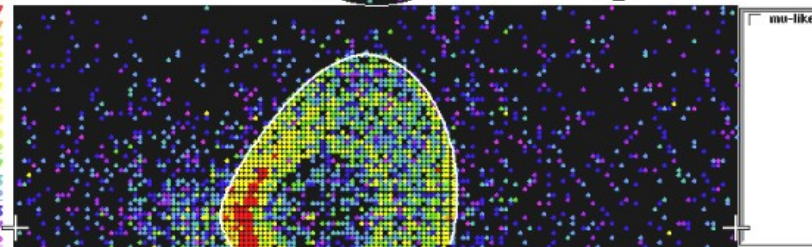
E_vis: 1.9 GeV

μ -like, $p = 1822.0$ MeV/c



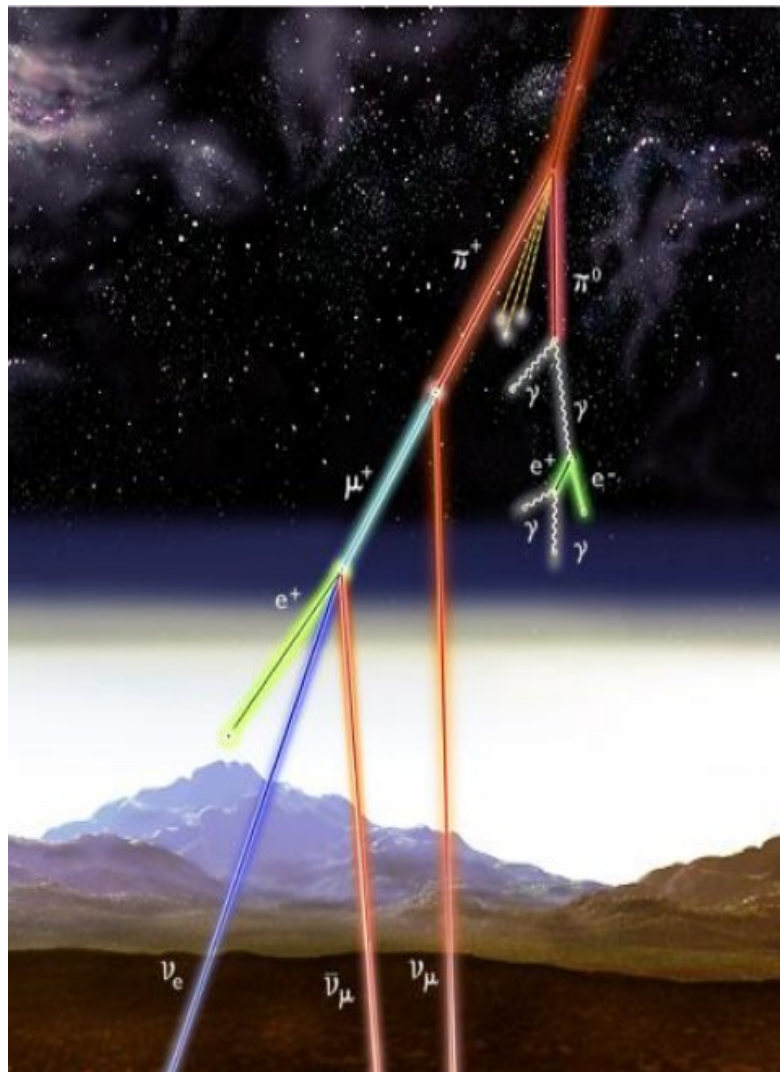
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- * 1.3-2.2
- * 0.7-1.3
- * 0.2-0.7
- * < 0.2



Atmospheric Neutrinos

- Neutrinos produced due to the interaction of cosmic rays and Earth's atmosphere
- Large statistics, but not enough yet
 - **wide range of energies:** $O(10^{-1})$ GeV to $O(10^2)$ GeV
 - **wide range of baselines:** $O(10)$ km to $O(10^4)$ km, half of them through the Earth
- Detailed and specific simulations are required to compute the neutrino flux
- SK discovered the **atmospheric neutrino oscillations** (Nobel 2015)



Atmospheric Neutrino Oscillations

Vacuum oscillations: neutrinos coming from above

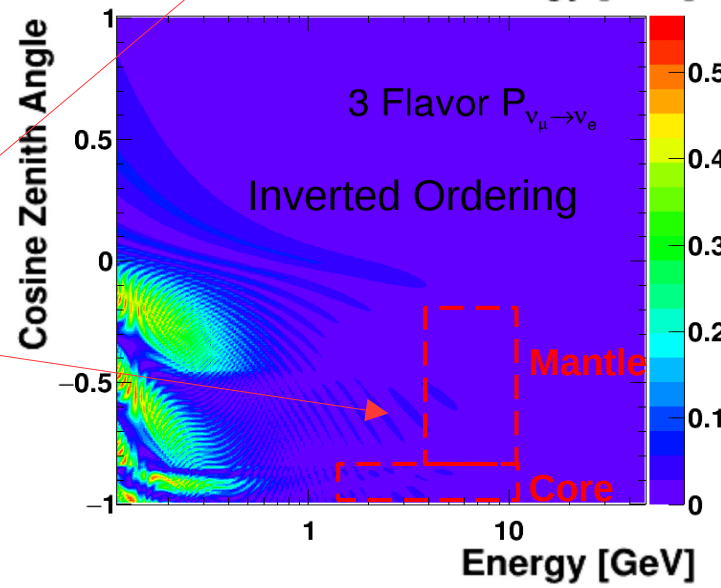
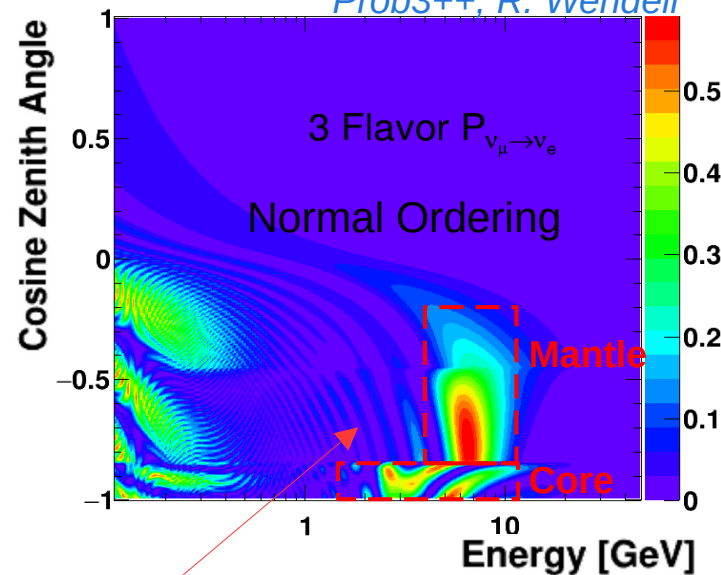
$$|\bar{\nu}_l\rangle = \sum_l U_{PMNS}^{li} |\bar{\nu}_i\rangle$$

$$U_{PMNS} = \begin{pmatrix} c_{13}c_{12} & c_{13}s_{12} & s_{13}e^{-i\delta} \\ -c_{23}s_{12} - s_{23}s_{13}c_{12}e^{i\delta} & c_{23}c_{12} - s_{23}s_{13}s_{12}e^{i\delta} & s_{23}c_{13} \\ s_{23}s_{12} - c_{23}s_{13}c_{12}e^{i\delta} & -s_{23}c_{12} - c_{23}s_{13}s_{12}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

Neutrino coming from below, pass through the Earth changing the effective Hamiltonian and thus, the neutrino propagation

$$H_{eff} = H_0 + H_{CC} = H_0 \pm \sqrt{2}G_F N_e \text{diag}(1, 0, 0)$$

- matter potential is opposite for $\bar{\nu}$ s
- mass ordering changes the matter (MSW) effects between ν s and $\bar{\nu}$ s



SK Atmospheric Neutrino Oscillation Analysis

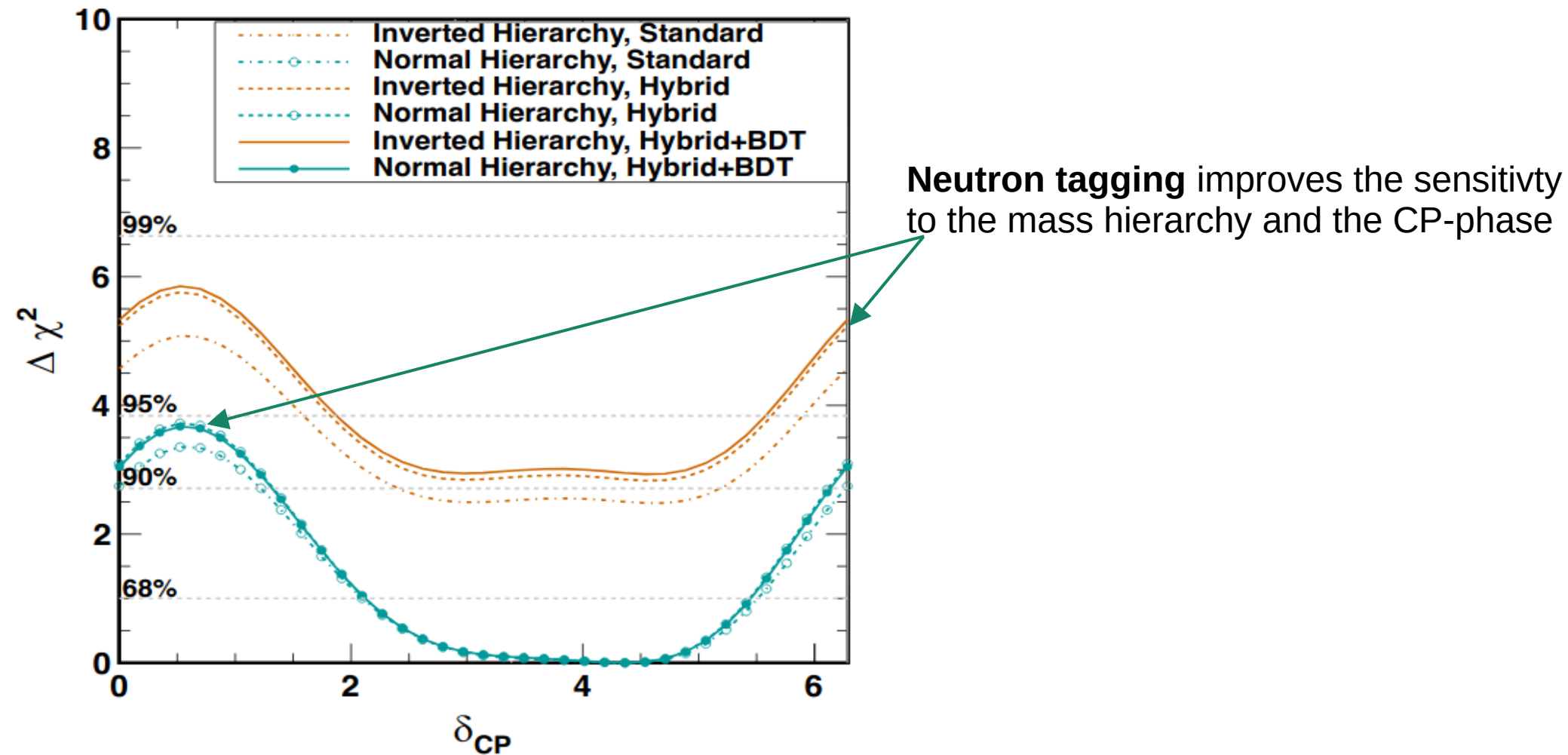
Atmospheric neutrinos are sensitive to crucial oscillation parameters:

- Neutrino mass ordering: e-like, > 1 GeV
 - Leptonic δ_{CP} phase: < 1 GeV
 - Octant of θ_{23} : > 1 GeV
- Separation of ν and $\bar{\nu}$ is crucial*

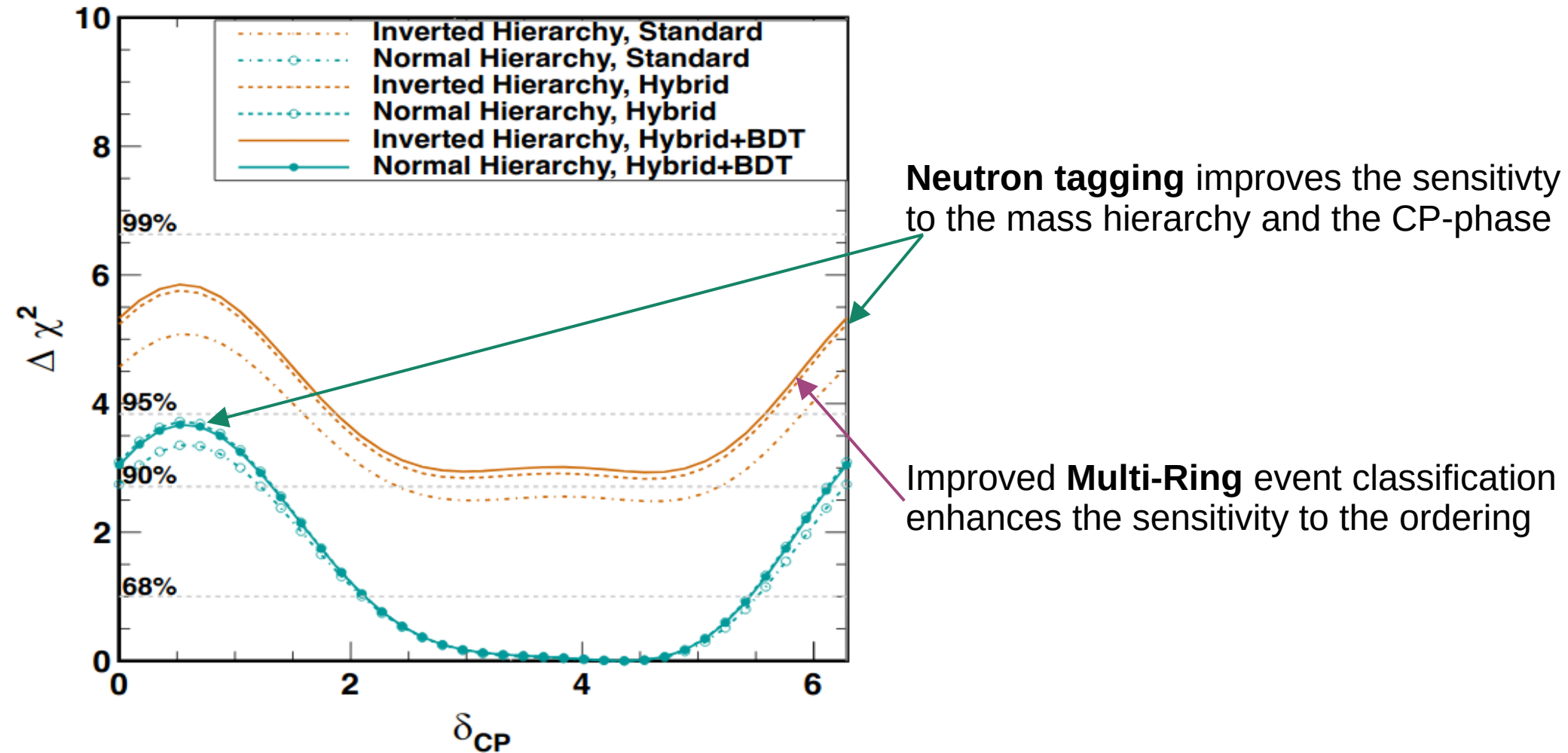
Analysis updates since last publication [*Phys. Rev. D 97, 072001*]

- SK-IV electronics allows for neutron tagging on hydrogen, with an efficiency of 25%
 - New samples in SKIV for better single-ring $\nu/\bar{\nu}$ separation
 - Usual sample scheme for SKI-III
 - **New event selection for Multi-Ring** events based on Boosted Decision Tree
 - New Monte-Carlo using **NEUT 5.4.0** with updated cross-section models and systematics
 - Changed **zenith binning** to better target the **matter effects**
 - Total of 3244.4 days
- Hybrid analysis*

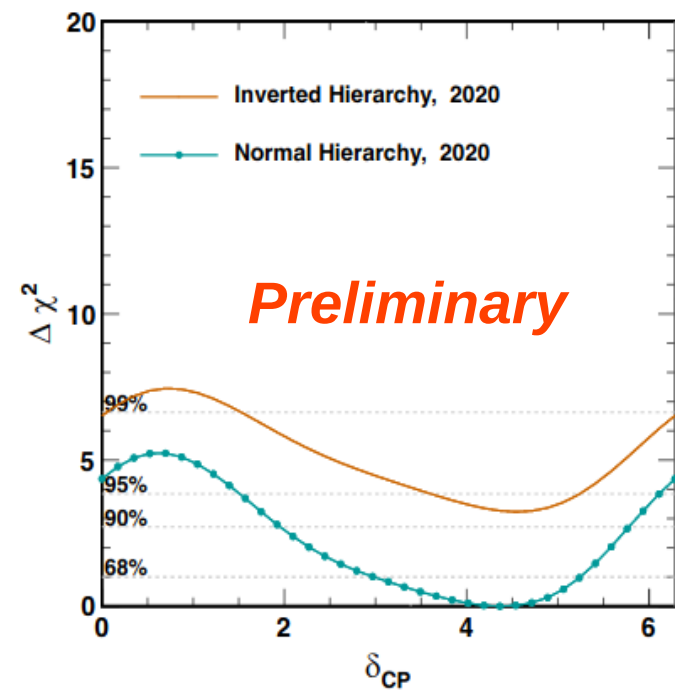
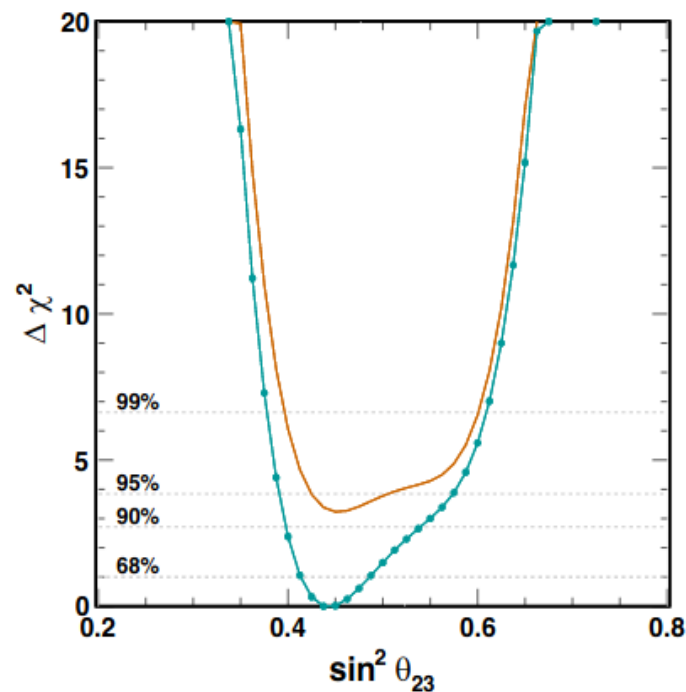
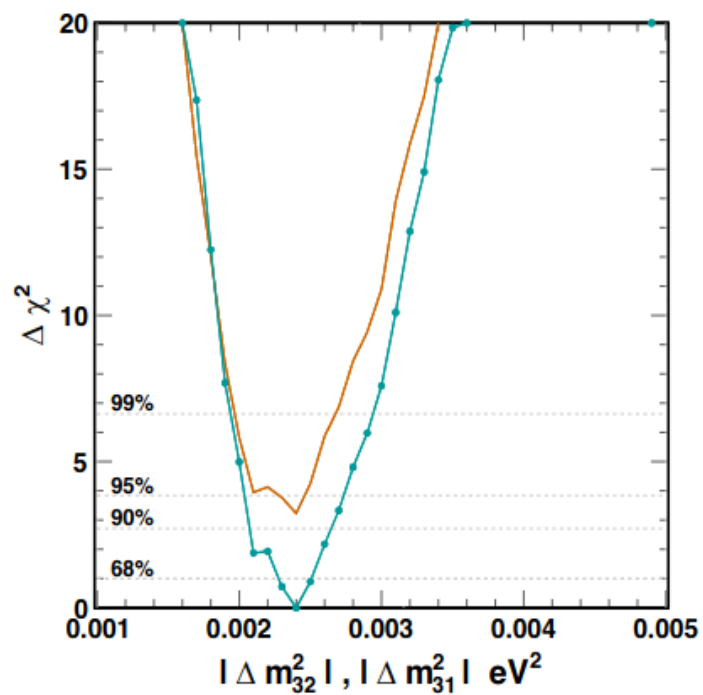
SK Atmospheric Neutrino Oscillation Analysis



SK Atmospheric Neutrino Oscillation Analysis

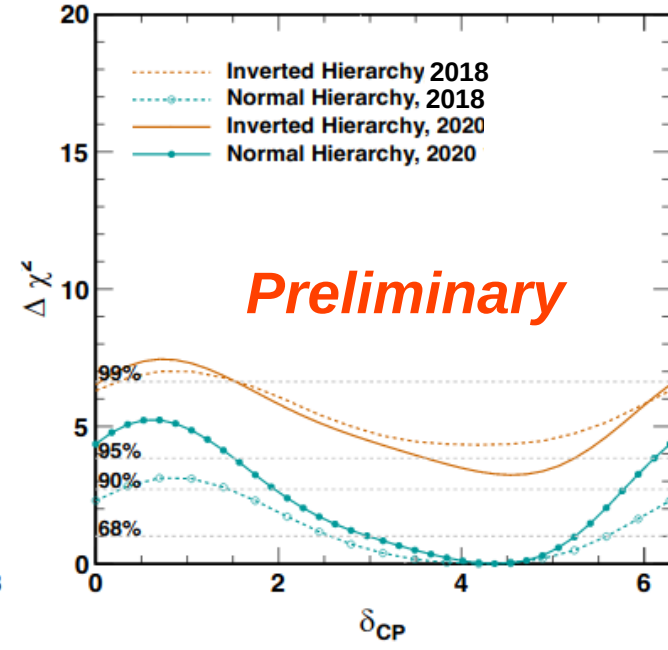
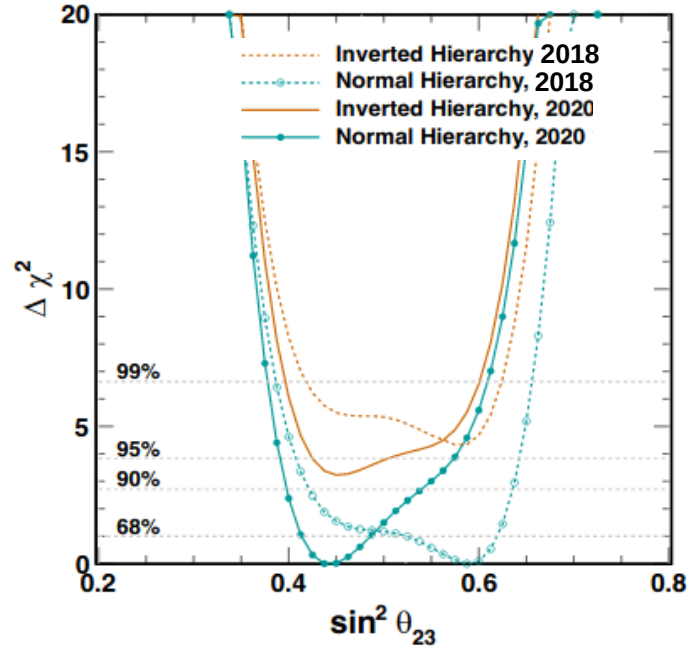
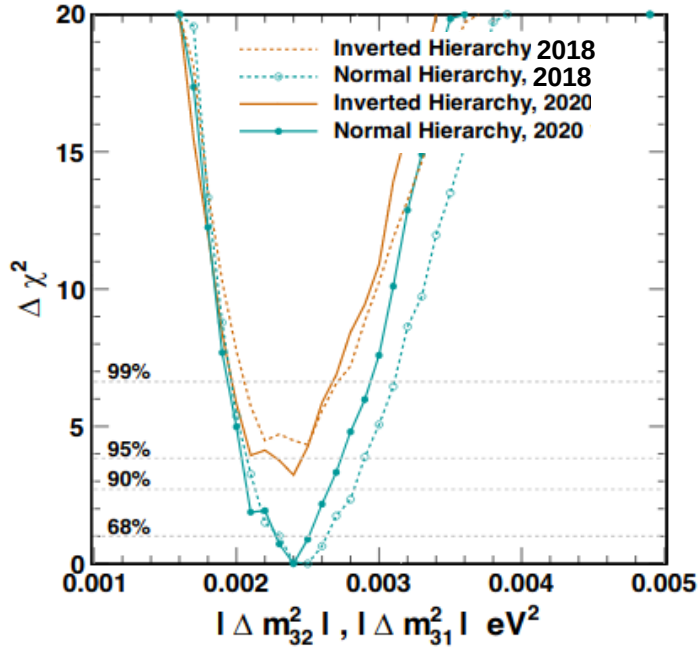


Results from the SK Atm. Neutrino Osc. Analysis



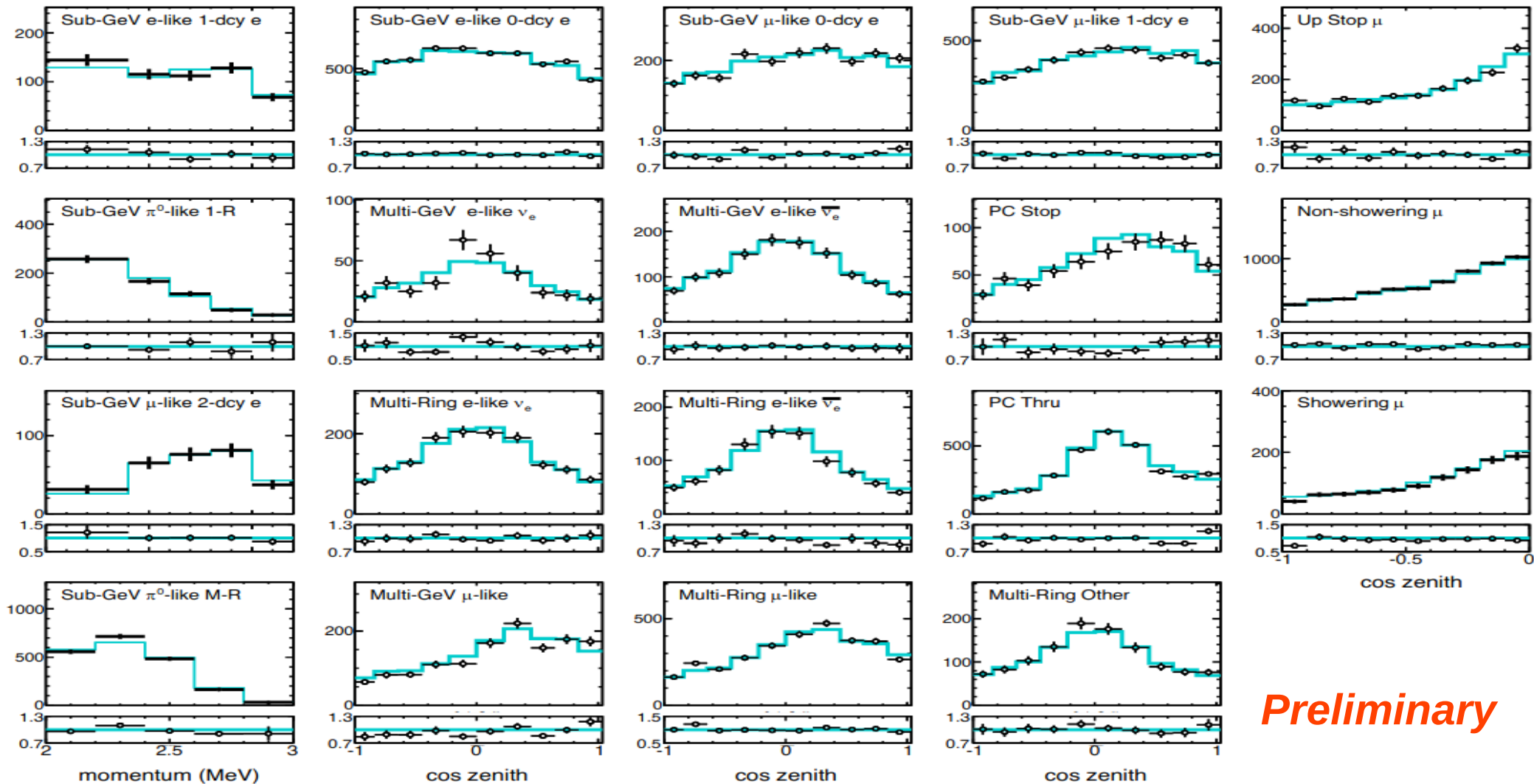
930 bins	χ^2	θ_{13} (reac.)	δ_{CP}	θ_{23}	Δm_{23} ($\times 10^{-3}$)
SK (NH)	1037.5	0.0218	$4.36^{+0.88}_{-1.39}$	$0.44^{+0.05}_{-0.02}$	$2.40^{+0.11}_{-0.12}$
SK (IH)	1040.7	0.0218	$4.54^{+0.88}_{-1.32}$	$0.45^{+0.09}_{-0.03}$	$2.40^{+0.09}_{-0.32}$

Comparison with previous results



- Preference for normal mass ordering at C.L. 86%
- Stronger rejection of small values of CP-phase
- Data prefers first octant and the contours for θ_{23} are significantly more constraining

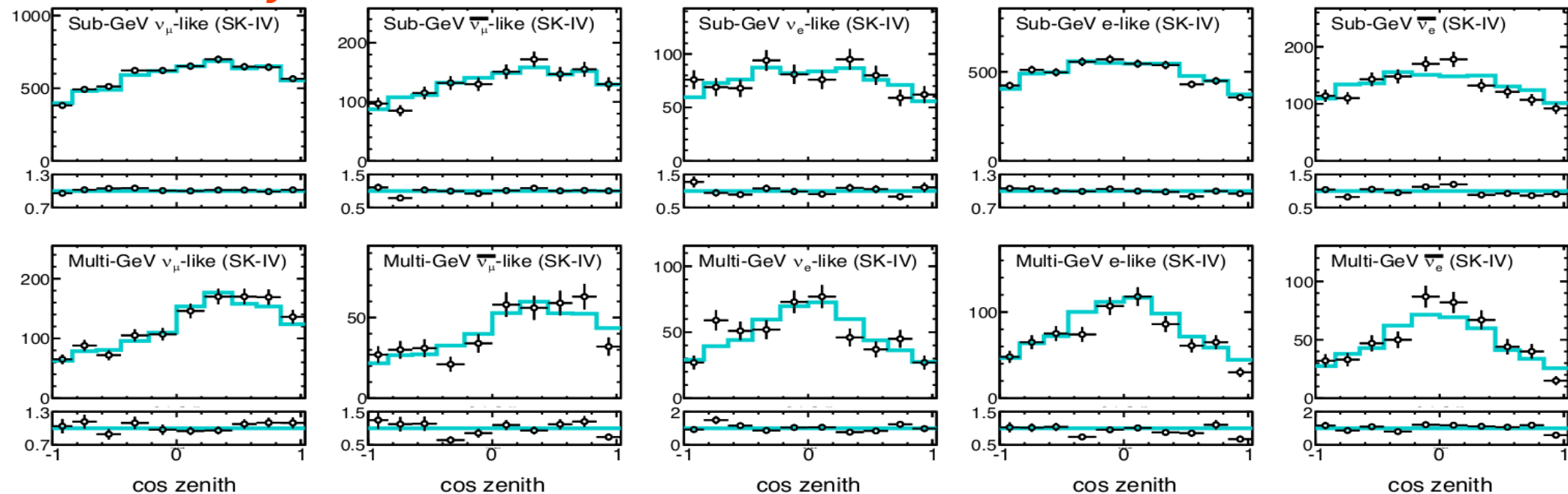
Explicit fit to the data (*traditional SK atm. samples*)



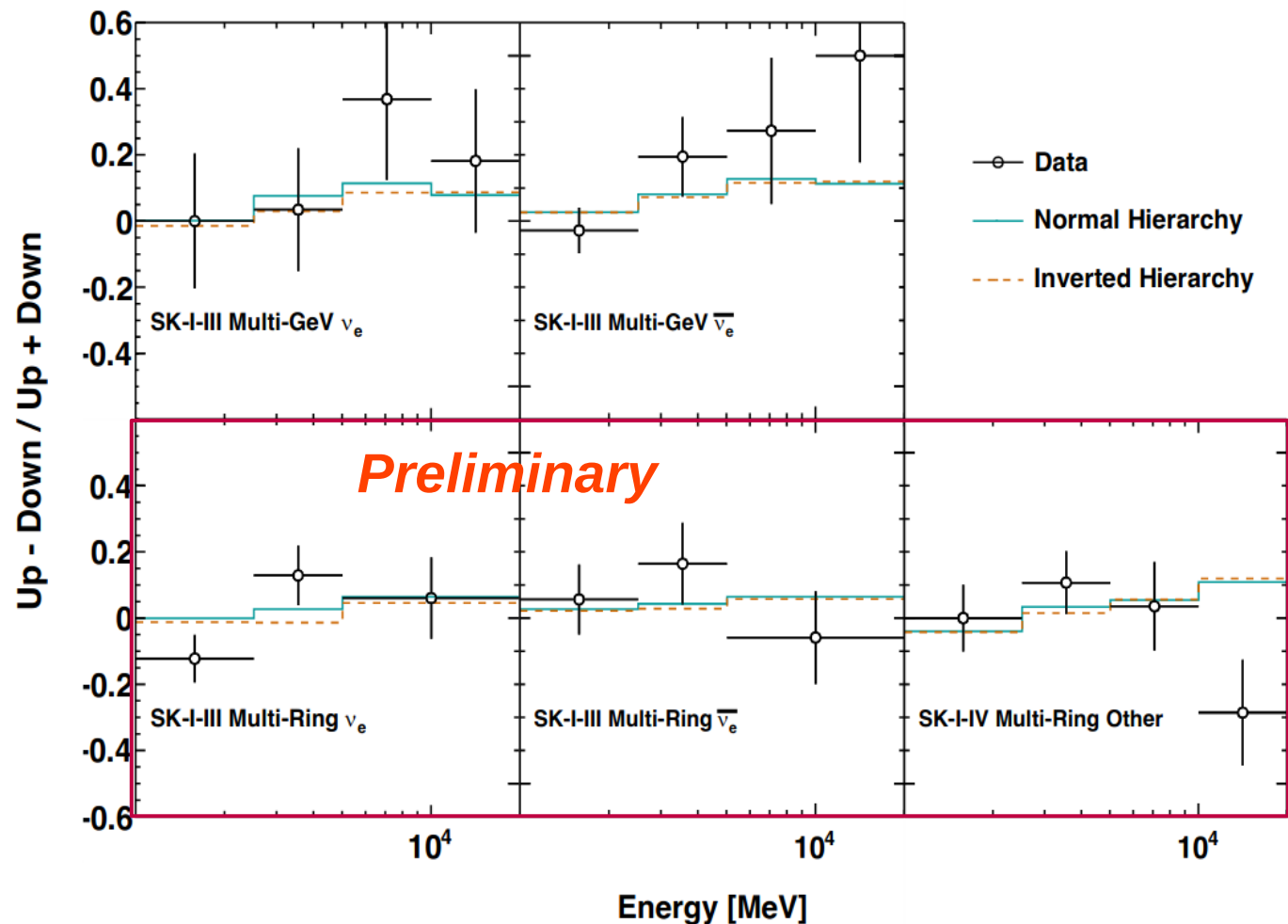
Preliminary

Explicit fit to the data (*neutron-tagging SK atm. samples*)

Preliminary



The hierarchy preference, a closer look

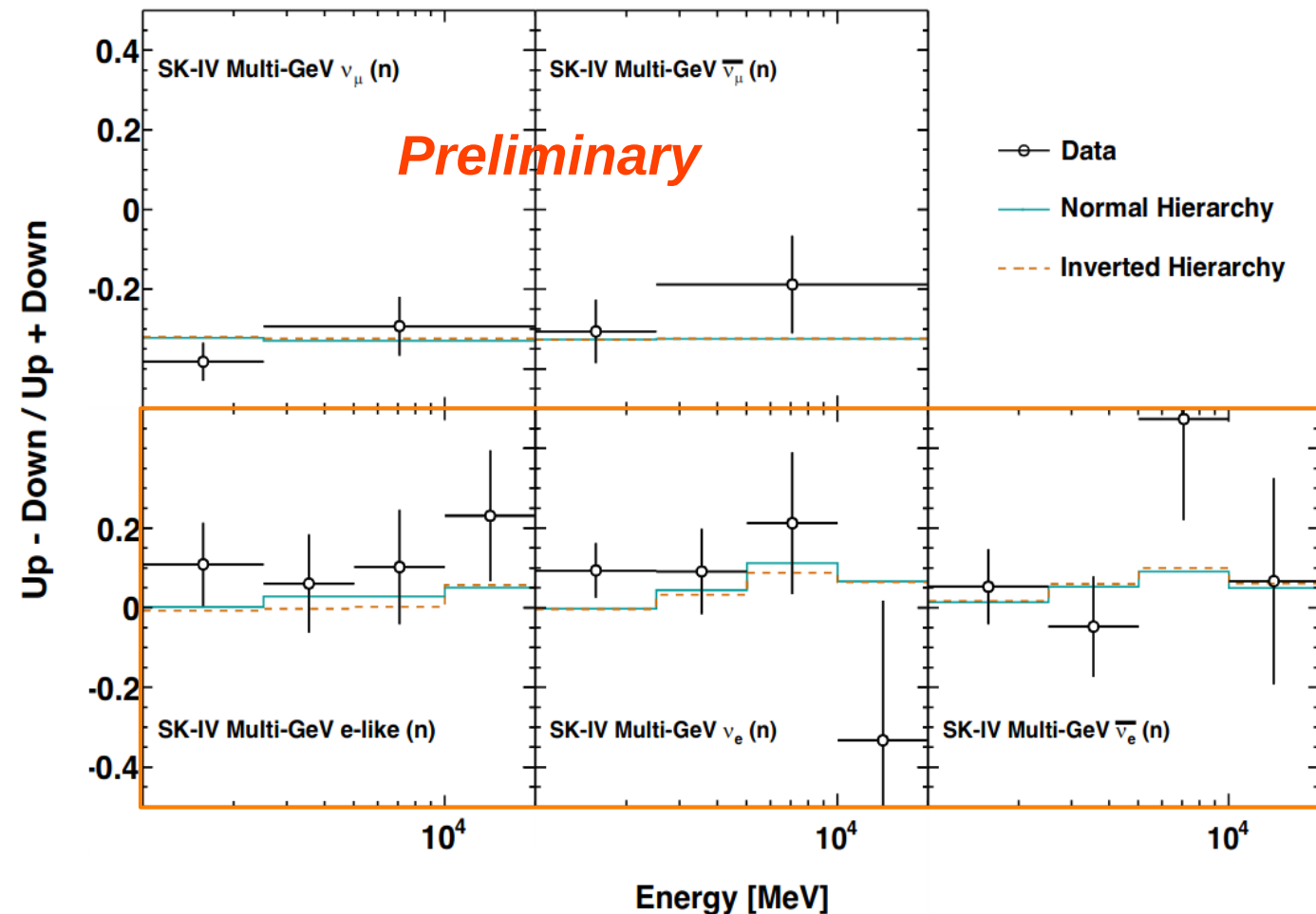


Significance resides mainly in the purity of ν_e -like samples

→ Improved in **Multi-Ring** with new event selection

→ Improved in Single-Ring with new sample definition using neutrons

The hierarchy preference, a closer look

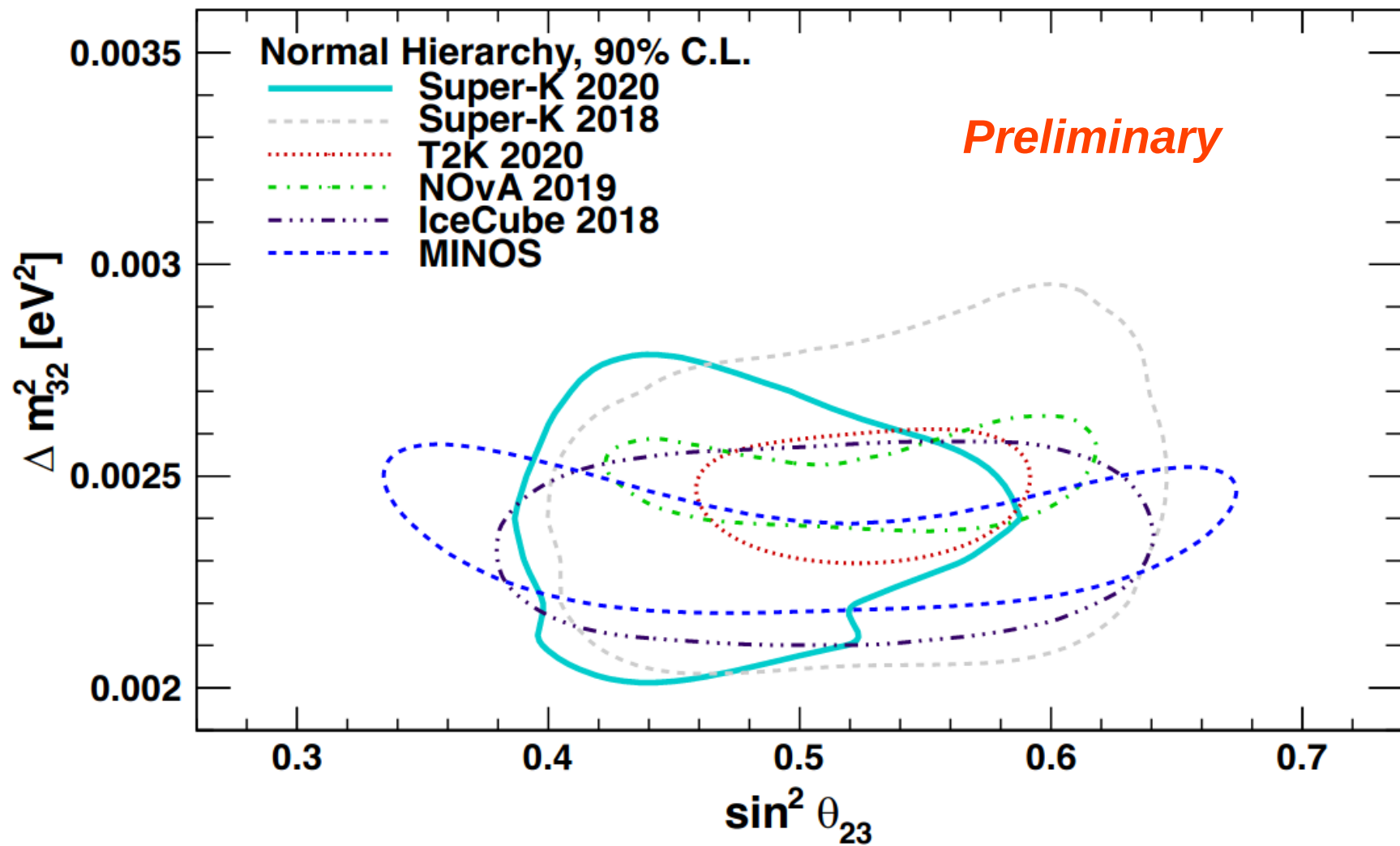


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→ Improved in Multi-Ring with new event selection

→ Improved in Single-Ring with new sample definition using **neutrons**

Results in context



Prospects for the SuperK-Gd Atm. ν Osc. Analysis

The **SuperK-Gd** upgrade has finally started in July 2020

The detector is fully operational and acquiring data with its present Gd concentration. **0.011% Gd**

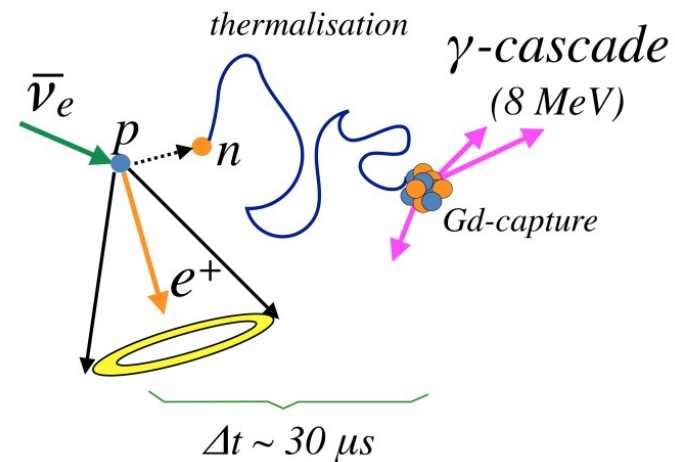
- **~50%** of neutron tagging efficiency (double the H-tagging eff.)

In addition to its impact to low energy physics, we have already seen the **importance of neutron tagging in the atm. ν analysis**

With greater neutron tagging efficiency:

- **Much better classification of neutrinos and antineutrinos**
- **Improve the neutrino energy reconstruction, as they carry information about the invisible energy**
- **Provide additional power in discriminating CC events from NC events**

Complementary studies and measurements will be needed, especially for the two last items as they rely on large neutron multiplicities (>10)



Conclusions

- After 25 years since the beginning, SK has rejuvenated once again to stay at the forefront of ν physics → **SuperK-Gd**
- Discovery of neutrino oscillations in 1998 and currently providing some of the most precise measurements of the oscillation parameters
- **Updated and upgraded atmospheric neutrino analysis**
 - Preference for large ($\sim 3\pi/2$) CP-phase values, agreeing with LBL experiments
 - Preference for normal mass ordering at 86% C.L.
 - Preference for first octant of θ_{23}
- Expecting still improved performance and results from SuperK-Gd data and from the acquired experience over the years