

Parameter Dependence of Ambipolar Diffusion Effects in Alfvén-Wave-Driven Stellar Winds

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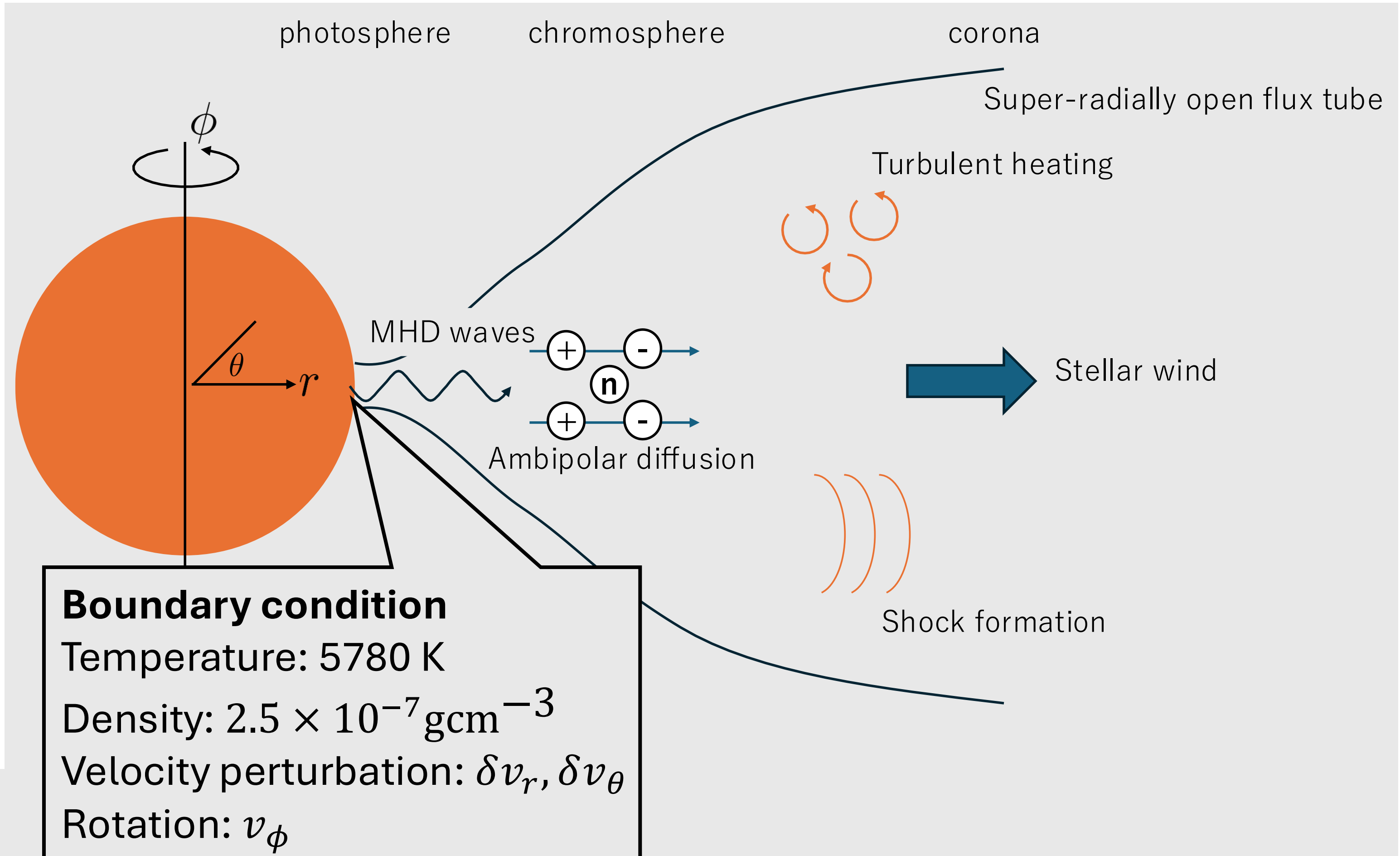
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Background & Motivations

- Mass-loss rate and torque are key quantities in stellar wind.
- Shoda et al. (2020) investigated how stellar rotation affects the mass-loss rate and torque with Alfvén-wave-driven model.
- However, their model does not include ambipolar diffusion, which possibly significant role for solar-like stellar winds (Matsuoka et al. (2024)).
- In this study, we perform stellar wind simulations that include both rotation and ambipolar diffusion, and investigate the parameter dependence of the mass-loss rate and torque over a wide parameter range.

Methods

- Equations: nonideal 1D-MHD simulation
 - Ohm & ambipolar diffusion included
 - Flux tube model (Kopp & Holzer 1976, Suzuki et al. 2013)
 - Heat conduction, radiation (Suzuki 2018), phenomenological turbulent dissipation (Hossain et al. 1995, Shoda et al. 2018) included
- Simulation region:
Photosphere to wind region (outermost: $r \sim 30\text{-}70 R_{\odot}$)
- Changing input parameters:
 B_0 : Radial magnetic field
 f_0 : Filling factor
 $\delta v_r, \delta v_{\theta}$: Velocity perturbation
 Ω_* : Rotation rate



Ohmic and ambipolar diffusion

$$\eta_O = \frac{c^2 m_e \nu_{en}}{4\pi e_c^2 n_e} \simeq 2.3 \times 10^2 \frac{\max((1-x_e), 0)}{x_e} \sqrt{\frac{T}{K}} \text{ cm}^2 \text{ s}^{-1}$$

$$\eta_{AD} = \frac{B^2 (\rho_n / \rho)^2}{4\pi \chi \rho_i \rho_n} \simeq 2.1 \times 10^{-16} \frac{(B/\text{G})^2 \max((1-x_e), 0)^2}{[\rho/(\text{g cm}^{-3})]^2 x_e} \text{ cm}^2 \text{ s}^{-1}$$

m_e : electron mass, e_c : elementary charge
 ν_{en} : collision frequency between electron and neutral species
 $\chi = \overline{\sigma_{in} \nu_{in}} / (m_i + m_n)$

Model examples		
Model	diffusion	B_0 (kG)
I1	off	1.41
D1	on	1.41
I2	off	2.82
D2	on	2.82

Other parameters:
 $f_0 = 1/1265$
 $\delta v_r = \delta v_{\theta} = 1.0 \text{ km/s}$
 $\Omega_* = 2.9 \times 10^{-6} \text{ rad/s}$
 (Period $\sim 25 \text{ day}$)

* Other 184 cases are also simulated.

Results & Discussions

Scaling laws of mass loss rate and torque

$$\dot{M}_{\text{ideal}} = 2.09 \times 10^{-14} \times \left(\frac{\delta v}{1 \text{ km s}^{-1}} \right)^{3.23} \times \left(\frac{B_0}{1 \text{ kG}} \right)^{-0.0763} \times \left(\frac{f_0}{7.91 \times 10^{-4}} \right)^{0.468} \times \left(\frac{\Omega}{\Omega_{\odot}} \right)^{0.0171} M_{\odot} \text{ yr}^{-1}$$

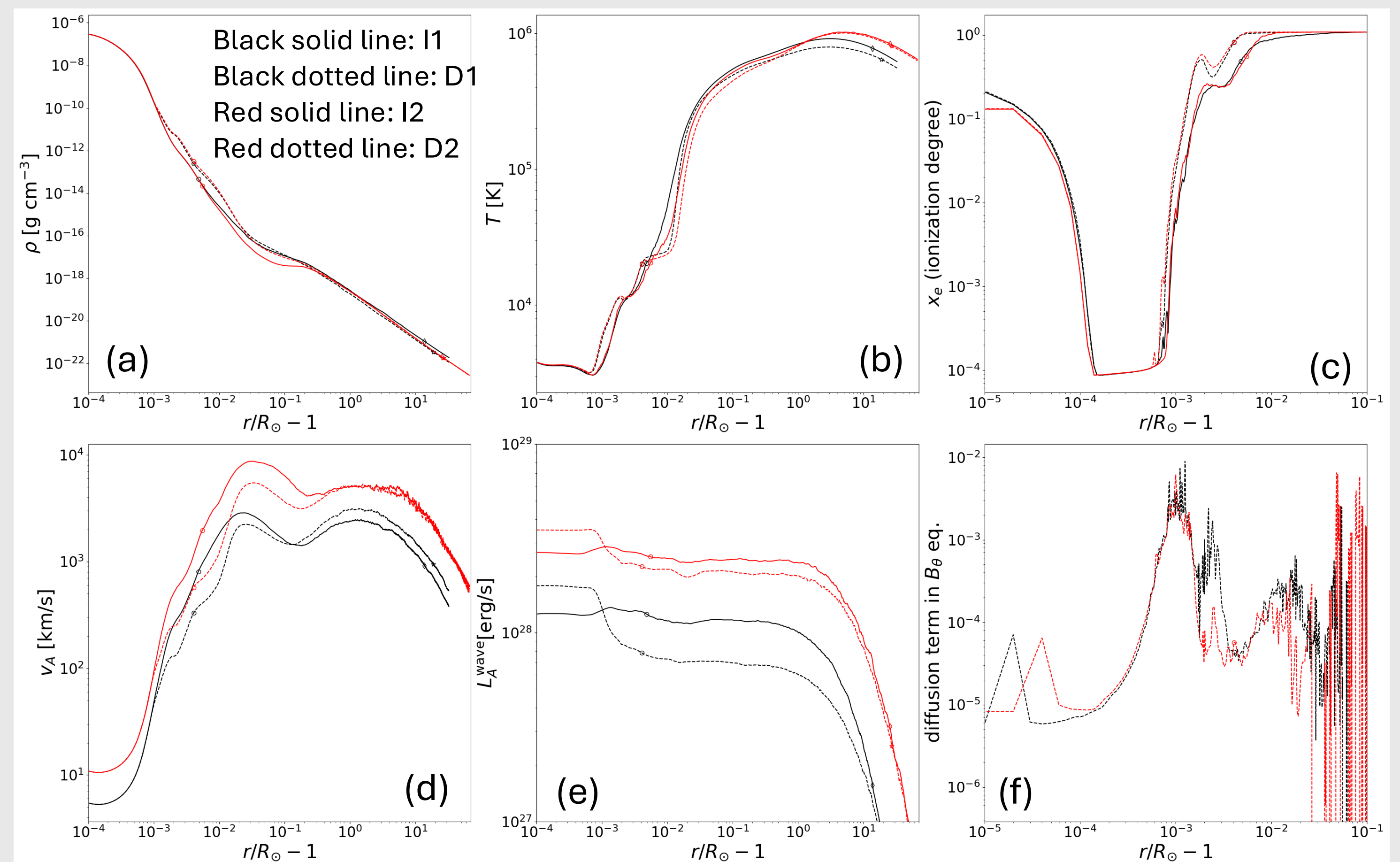
$$\dot{M}_{\text{diff}} = 0.917 \times 10^{-14} \times \left(\frac{\delta v}{1 \text{ km s}^{-1}} \right)^{3.87} \times \left(\frac{B_0}{1 \text{ kG}} \right)^{0.564} \times \left(\frac{f_0}{7.91 \times 10^{-4}} \right)^{0.742} \times \left(\frac{\Omega}{\Omega_{\odot}} \right)^{0.0280} M_{\odot} \text{ yr}^{-1}$$

$$\tau_{\text{ideal}} = 1.88 \times 10^{30} \times \left(\frac{\delta v}{1 \text{ km s}^{-1}} \right)^{0.220} \times \left(\frac{B_0}{1 \text{ kG}} \times \frac{f_0}{7.91 \times 10^{-4}} \right)^{1.47} \times \left(\frac{\Omega}{\Omega_{\odot}} \right)^{0.921} \text{ erg}$$

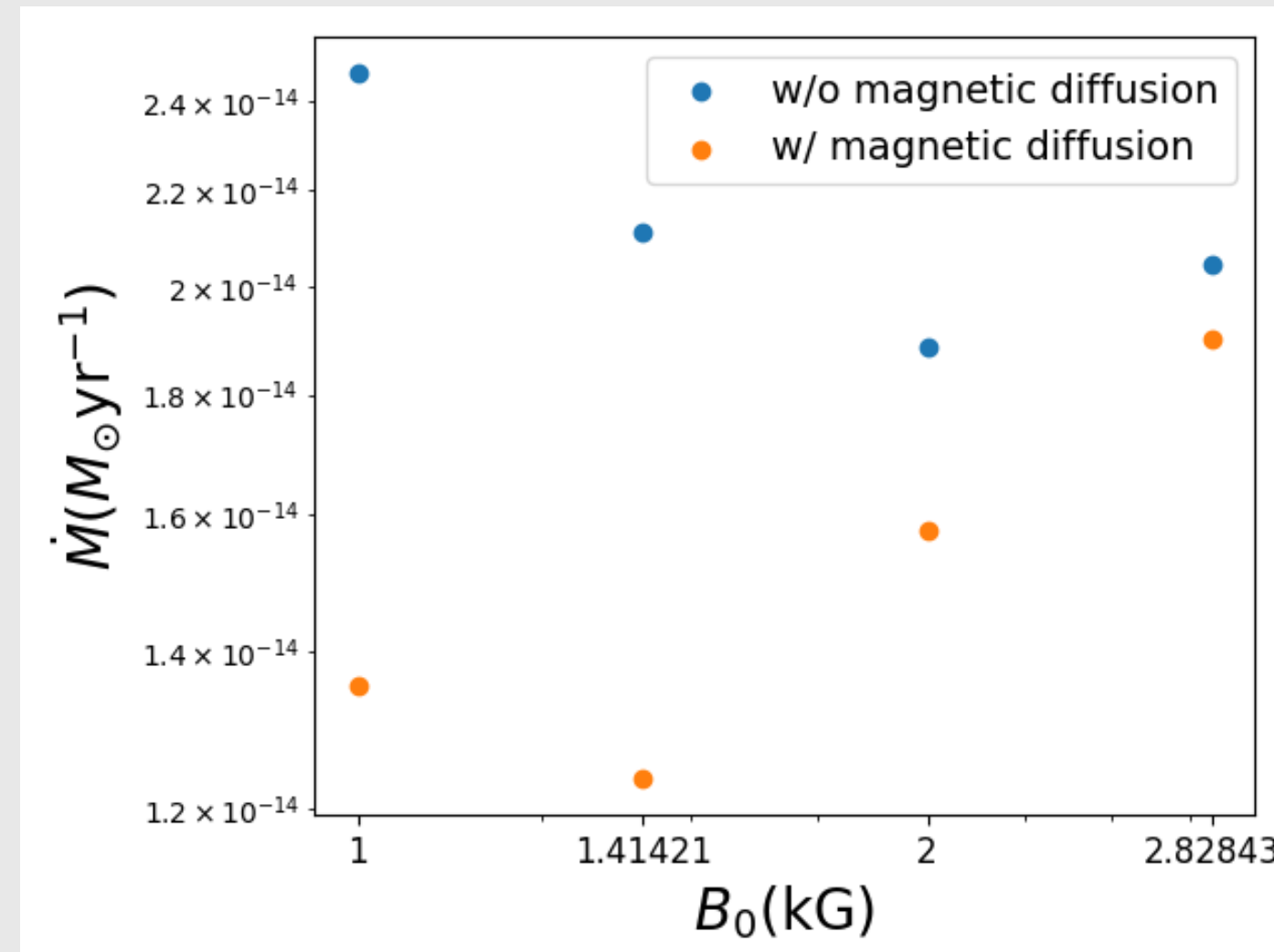
$$\tau_{\text{diff}} = 1.97 \times 10^{30} \times \left(\frac{\delta v}{1 \text{ km s}^{-1}} \right)^{0.176} \times \left(\frac{B_0}{1 \text{ kG}} \times \frac{f_0}{7.91 \times 10^{-4}} \right)^{1.44} \times \left(\frac{\Omega}{\Omega_{\odot}} \right)^{0.910} \text{ erg}$$

- Mass-loss rate:
 w/o ambipolar diffusion \rightarrow independent of B_0
 w/ ambipolar diffusion \rightarrow dependent on B_0
 - $B_0 \uparrow \rightarrow l_{AW} \uparrow \rightarrow$ ambipolar \downarrow
 (l_{AW} : the spatial scale of Alfvén waves)
 - $B_0 \uparrow \rightarrow v_A \uparrow \rightarrow \tau_A \downarrow$
 (τ_A : time that waves are exposed to ambipolar diffusion)
- The magnitude of the diffusion term depends only weakly on the magnetic field strength.
- Torque: independent of ambipolar diffusion
 - Ambipolar diffusion $\rightarrow \dot{M} \downarrow, r_A \uparrow$
 \rightarrow Ambipolar diffusion little change the torque

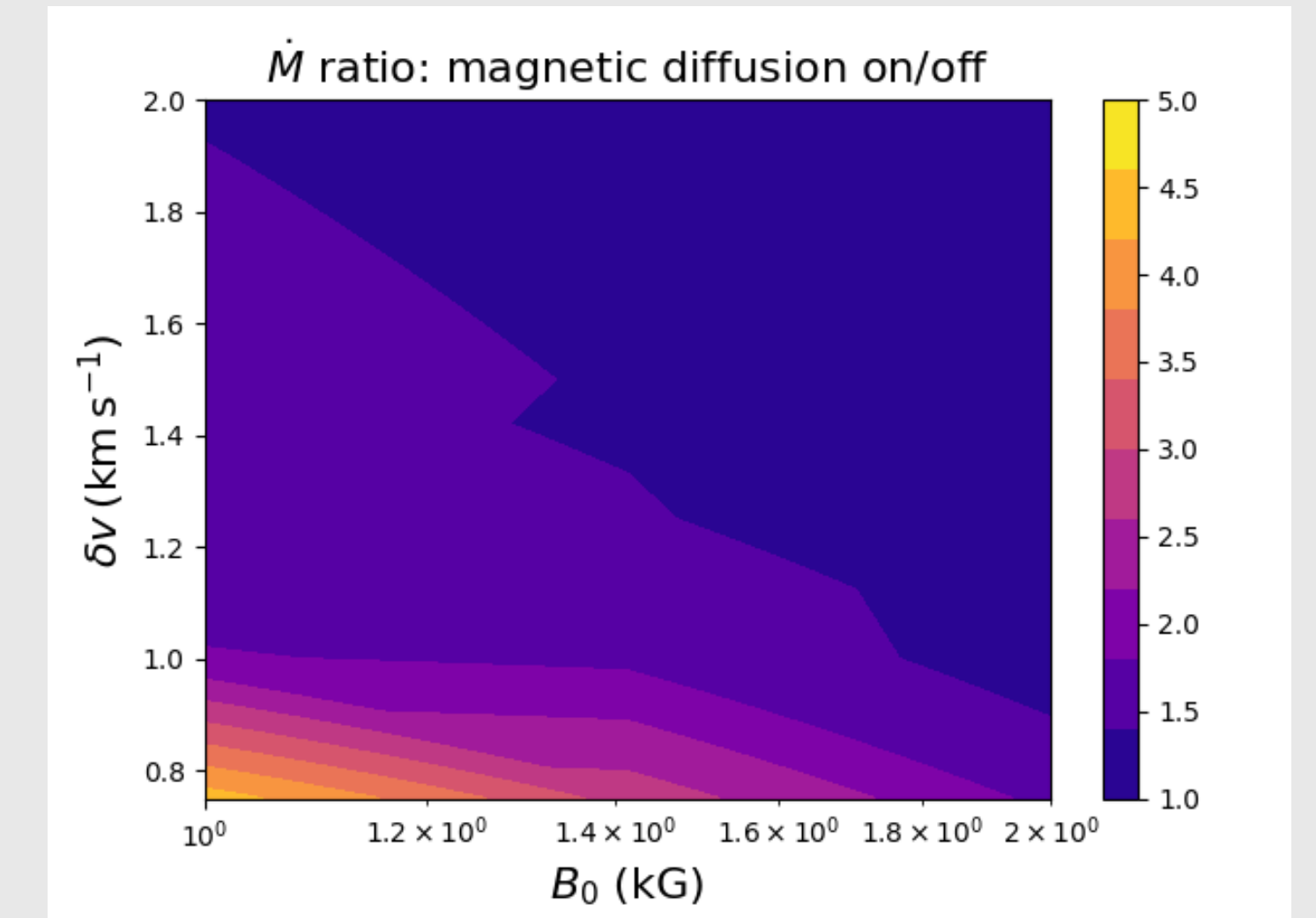
$$\tau_w \sim \dot{M} r_A^2 \Omega_* \quad (\tau_w : \text{torque})$$



Radial profiles of (a) density, (b) temperature, (c) ionization degree, (d) Alfvén velocity, (e) Alfvén luminosity, (f) rms of diffusion term in induction equation in θ direction



Mass loss rates in cases with/without magnetic diffusion under various magnitude of the magnetic field



Ratio of the mass-loss rate with magnetic diffusion to that without magnetic diffusion

Summary

- We simulated 1D-MHD simulation over wide parameter range.
- Our results suggest that ambipolar diffusion plays an important role only when magnetic field is weak.
- Scaling laws for the mass-loss rate and torque are also derived for both the ideal and ambipolar-diffusion cases.

References

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