

# Statistical Robustness of the $\beta$ Coupling Parameter in the 3D+3D Discrete Spacetime Theory: A Comprehensive Analysis of SPARC Galaxy Rotation Curves

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## Abstract

We present a comprehensive statistical analysis of the  $\beta$  coupling parameter calibration in the 3D+3D Discrete Spacetime Theory using 122 galaxies from the SPARC (Spitzer Photometry and Accurate Rotation Curves) database. The theory proposes that apparent dark matter effects arise from geometric modifications in a six-dimensional spacetime with two compactified temporal dimensions at characteristic scales  $\lambda_2 = 4.30$  kpc and  $\lambda_3 = 11.7$  kpc. We test both single-mode ( $\beta_2$  only) and two-mode ( $\beta_2 + \beta_3$ ) models through a battery of robustness tests including AIC/BIC analysis, K-fold cross-validation, bootstrap resampling, sensitivity analysis, and hierarchical Bayesian modeling. Our key finding is that while the two-mode model shows superior in-sample fit ( $\Delta\text{AIC} = 1453$ ,  $\Delta\text{BIC} = 1447$ ), it fails to generalize out-of-sample (cross-validation  $p = 0.93$ ). The secondary mode  $\beta_3$  exhibits high sensitivity to methodological choices (50% variation with  $r_{\text{eff}}$  definition), inconsistent behavior across galaxy types ( $\beta_3 = 0$  for LSB galaxies), and a bimodal per-galaxy distribution with 52% of galaxies preferring  $\beta_3 \approx 0$ . We conclude that  $\beta_2 = 0.83 \pm 0.02$  is a robust, universal parameter, while  $\beta_3$  cannot be reliably measured with SPARC data. This does not falsify the  $\lambda_3$  mode but indicates that SPARC operates in the regime dominated by the fundamental Q-field mode. Extended HI surveys (THINGS, HALOGAS, WALLABY) with  $r_{\text{max}} > 25$  kpc are required to probe the secondary temporal scale.

**Keywords:** dark matter alternatives, modified gravity, galaxy rotation curves, extra dimensions, statistical robustness, model selection

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## 1. Introduction

### 1.1 The Dark Matter Problem

The discrepancy between observed galaxy rotation curves and predictions from Newtonian dynamics applied to visible matter remains one of the most compelling puzzles in modern astrophysics. The standard cosmological model ( $\Lambda$ CDM) interprets this as evidence for cold dark matter halos, yet direct detection experiments have not confirmed particle dark matter candidates despite decades of effort.

## 1.2 The 3D+3D Discrete Spacetime Theory

The 3D+3D theory proposes an alternative explanation: spacetime has six dimensions (three spatial, three temporal) with signature  $(-,+,+,+,-,-)$ . Two temporal dimensions ( $\tau_2, \tau_3$ ) are compactified at galactic scales with characteristic lengths:

- $\lambda_2 = 4.30 \text{ kpc}$  (primary compactification scale)
- $\lambda_3 = 11.7 \text{ kpc}$  (secondary compactification scale)

These compactified dimensions generate a scalar Q-field that modifies the effective gravitational potential, producing rotation curve enhancements that mimic dark matter effects without requiring new particles.

## 1.3 The $\beta$ Calibration Problem

The Q-field contribution to galaxy rotation is parameterized by coupling constants  $\beta_2$  and  $\beta_3$  for the two temporal modes. A critical question is whether these parameters are:

1. **Universal constants** derivable from theory, or
2. **Fitted parameters** that vary per galaxy

The theory aspires to the former—zero free parameters per galaxy. This paper systematically tests whether  $\beta$  can be calibrated universally and whether the secondary mode  $\beta_3$  is statistically necessary.

## 1.4 Scope of This Work

We perform a comprehensive robustness analysis following recommendations from independent AI review, testing:

1. Whether the two-mode model generalizes better than single-mode
2. Sensitivity of  $\beta$  to methodological choices
3. Consistency of  $\beta_3$  across galaxy populations
4. Bayesian evidence for model selection
5. Hierarchical structure of per-galaxy  $\beta_3$  estimates

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## 2. Theoretical Framework

### 2.1 Model Equations

The total rotational velocity in the 3D+3D framework is:

$$V_{rot}^2 = V_{bar}^2 + V_{Q2}^2 + V_{Q3}^2$$

where  $V_{\text{bar}}$  is the baryonic contribution and the Q-field contributions are:

$$V_{Qi}^2 = \beta_i^2 \cdot v_{ref}^2 \cdot \left( \frac{M_{bar}}{M_{ref}} \right)^{2\alpha} \cdot F_{total} \cdot f_{shape}(r/\lambda_i)$$

## 2.2 Screening Functions

The screening factor  $F_{\text{total}}$  accounts for disk geometry and gravitational potential:

$$F_{total} = F_{thick}(\chi) \cdot F_{press}(\beta_{press}) \cdot F_{pot}(\psi)$$

with:

$$F_{thick}(\chi) = \frac{1}{\sqrt{1 + (\chi/\chi_0)^2}}$$

$$F_{press}(\beta_{press}) = \frac{1}{1 + \beta_{press}}$$

$$F_{pot}(\psi) = \tanh(\psi/\psi_{crit})$$

## 2.3 Shape Function

The radial profile is governed by:

$$f_{shape}(r/\lambda) = \alpha_{shape} \cdot \tanh(r/\lambda)$$

## 2.4 Fixed Parameters

All parameters except  $\beta$  are fixed from theory:

Parameter	Value	Description
v_ref	90.39 km/s	Reference velocity
M_ref	10 <sup>10</sup> M $\odot$	Reference mass
$\alpha$	0.25	Mass scaling exponent
$\lambda_2$	4.30 kpc	Primary scale
$\lambda_3$	11.7 kpc	Secondary scale
$\chi_0$	0.235	Critical thickness
$\psi_{\text{crit}}$	2.27 $\times 10^{-8}$	Critical potential
$\alpha_{\text{shape}}$	1.5	Shape amplitude

3. Data and Methods

3.1 SPARC Database

We use the SPARC database (Lelli, McGaugh & Schombert 2016), containing 175 galaxies with high-quality rotation curves derived from HI/H $\alpha$  observations and 3.6 $\mu$ m Spitzer photometry.

3.2 Quality Selection

We apply the following cuts:

- Minimum 10 data points per galaxy
- $V_{\text{flat}} \geq 20$  km/s
- $M_{\text{bar}} \geq 10^7$  M $\odot$
- Fraction of points with  $V_{\text{bar}} > 1.1 \cdot V_{\text{obs}} \leq 30\%$

**Result:** 122 galaxies with 3014 data points selected.

3.3 Parameter Estimation

For each galaxy, we estimate:

$$M_{\text{bar}} = r_{\text{eff}} \cdot V_{\text{flat}}^2 / G_N$$

$$\psi = G_N \cdot M_{bar} / (R_{char} \cdot c^2)$$

$$\beta_{press} = (10 \text{ km/s} / V_{flat})^2$$

where  $r_{eff} = \text{median}(r)$  and  $V_{flat} = \text{median of last quarter of velocity points}$ .

### 3.4 Fitting Procedure

We minimize  $\chi^2$  using Nelder-Mead optimization:

$$\chi^2 = \sum_i \sum_j \left( \frac{V_{obs,ij} - V_{model,ij}}{\sigma_{ij}} \right)^2$$

with minimum error floor  $\sigma_{min} = 5 \text{ km/s}$ .

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## 4. Robustness Testing Framework

Following recommendations from independent review, we implement four systematic tests:

### 4.1 Part 1: Information Criteria and Cross-Validation

**Objective:** Compare single-mode vs two-mode using both in-sample (AIC/BIC) and out-of-sample (K-fold CV) metrics.

**Information Criteria:**

$$AIC = 2 \cdot NLL + 2k$$

$$BIC = 2 \cdot NLL + k \cdot \ln(N)$$

**Cross-Validation:** 5-fold CV at galaxy level with paired t-test for significance.

### 4.2 Part 2: Weighting Schemes and Bootstrap

**Objective:** Test sensitivity to data weighting and estimate confidence intervals.

**Standard weighting:** Per-point  $\chi^2$  **Per-galaxy weighting:**  $\chi^2/N$  per galaxy (equal galaxy weight)

**Bootstrap:** Resample galaxies with replacement (N=20 for computational efficiency).

4.3 Part 3: Sensitivity Analysis

**Objective:** Test sensitivity to parameter estimation choices.

**Variations:**

- r\_eff definition: p25, median, mean, p75
- V\_flat definition: last\_quarter, last\_third, p90, max
- Error floor: 3, 5, 7, 10 km/s

4.4 Part 4: Selection Cut Sensitivity

**Objective:** Test impact of quality cuts on results.

**Variations:**

- Fraction cut threshold: 0.1, 0.2, 0.3, 0.4, 0.5, none
- Minimum points: 5, 8, 10, 15, 20, 25

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5. Results: Standard Robustness Tests

5.1 Part 1 Results: AIC/BIC vs Cross-Validation

Information Criteria (In-Sample)

Model	AIC	BIC
Single-mode (k=1)	57,612	57,618
Two-mode (k=2)	56,160	56,172
$\Delta$ AIC	+1,453	—
$\Delta$ BIC	—	+1,447

Both criteria strongly favor the two-mode model ( $\Delta$ AIC,  $\Delta$ BIC >> 10).

Cross-Validation (Out-of-Sample)

Model	Test RMS	Parameter Stability
Single-mode	18.07 ± 2.33 km/s	$\beta = 0.832 \pm 0.018$ (2.1%)
Two-mode	18.09 ± 2.04 km/s	$\beta_2 = 0.675 \pm 0.036$ (5.3%)

**Out-of-sample improvement: -0.02 km/s, p = 0.93 (NOT SIGNIFICANT)**

**The AIC/BIC vs CV Paradox**

This is a critical finding: the two-mode model shows dramatically better in-sample fit but **zero improvement** on new galaxies. This is the textbook signature of overfitting—the additional parameter  $\beta_3$  is capturing noise specific to the training set rather than generalizable physics.

**5.2 Part 2 Results: Weighting and Bootstrap**

**Weighting Comparison**

Scheme	$\beta_2$	$\beta_3$	Change in $\beta_3$
Standard (per-point)	0.6748	0.5472	—
Per-galaxy weighted	0.7052	0.4031	<b>-26.3%</b>

$\beta_3$  shows high sensitivity to weighting scheme—a physical parameter should be more robust.

**Bootstrap Results (N=20)**

Parameter	Mean	Std	95% CI
$\beta_2$	0.697	0.055 (8%)	[0.585, 0.773]
$\beta_3$	0.401	0.092 (23%)	[0.263, 0.589]

$\beta_3$  has nearly 3× higher relative uncertainty than  $\beta_2$ .

**5.3 Part 3 Results: Parameter Estimation Sensitivity**

Variation	$\beta_2$ Spread	$\beta_3$ Spread
r_eff definition	<b>24.9%</b>	<b>48.9%</b>
V_flat definition	1.8%	18.6%

Variation	$\beta_2$ Spread	$\beta_3$ Spread
Error threshold	4.6%	10.8%
<b>TOTAL</b>	<b>24.8%</b>	<b>50.4%</b>

**Critical finding:**  $\beta_3$  varies by ~50% simply by changing how  $r_{\text{eff}}$  is calculated. A physical parameter cannot depend this strongly on arbitrary methodological choices.

### 5.4 Part 4 Results: Selection Cut Sensitivity

Variation	$\beta_2$ Spread	$\beta_3$ Spread
Fraction cut	2.7%	4.9%
Minimum points	10.8%	30.7%
<b>TOTAL</b>	<b>11.0%</b>	<b>32.2%</b>

Selection cuts have moderate impact, with  $\beta_3$  again showing higher sensitivity.

## 6. Results: Extended Tests (Vega Protocol)

Following recommendations from a second independent AI review (Vega), we implement four additional diagnostic tests.

### 6.1 Test A: Removal Test (Per-Galaxy Likelihood Drop)

**Objective:** For each galaxy, compute the likelihood improvement from adding  $\beta_3$ . If  $\beta_3$  is physical, expect correlation with  $r_{\text{max}}/\lambda_3$ .

**Results:**

- Galaxies improved by  $\beta_3$ : 63 (51.6%)
- Galaxies worsened by  $\beta_3$ : 59 (48.4%)
- Significantly improved ( $\Delta\chi^2 > 3.84$ ): 58 (47.5%)

**Correlations with  $\Delta\text{NLL}$ :**



Property	r	p-value
r_max/ $\lambda_3$	+0.237	0.009
V_flat	+0.237	0.009
log(M)	+0.185	0.042
N_pts	+0.323	0.0003

**Interpretation:** Weak positive correlation with r\_max/ $\lambda_3$  ( $r = 0.24$ ) suggests that some extended galaxies do benefit from  $\beta_3$ , but the effect is not strong enough to establish  $\beta_3$  as universally necessary.

### 6.2 Test B: Galaxies by Type

**Objective:** Separate galaxies by morphological type and test if  $\beta_3$  emerges consistently.

**Classification:**

- LSB (Low Surface Brightness): 24 galaxies
- Normal spirals: 38 galaxies
- HSB (High Surface Brightness): 49 galaxies
- Massive ( $M > 10^{10.5} M_\odot$ ): 7 galaxies

**Results:**

Type	N	$\beta_3$	p-value	$\beta_3$ Significant?
LSB	24	<b>0.000</b>	0.998	NO
Normal	38	0.654	<0.001	Yes (but $\Delta$ RMS negative!)
HSB	49	0.626	<0.001	Yes
Massive	7	0.591	<0.001	Yes

**Critical findings:**

1. **LSB galaxies:**  $\beta_3$  converges to exactly zero
2. **Normal galaxies:**  $\beta_3$  is significant but *worsens* RMS by 0.49 km/s
3. **Coefficient of Variation:** CV = 58% across types

**Interpretation:**  $\beta_3$  is NOT a universal parameter—it varies from 0 to 0.65 depending on galaxy type. This inconsistency is incompatible with a fundamental physical constant.

### 6.3 Test C: Bayesian Evidence

**Objective:** Compute Bayes Factor using Laplace approximation.

**Results:**

- $\Delta \ln(Z) = 722.8$
- Bayes Factor  $B = \infty$  (numerical overflow)
- Interpretation: "Decisive evidence" for two-mode (Jeffreys scale)

**Critical caveat:** This result is *misleading* because the Bayes Factor is computed from in-sample likelihood. The cross-validation showed no out-of-sample improvement. High Bayes Factor with failed CV is the classic signature of overfitting.

### 6.4 Test D: Hierarchical Bayesian Model

**Objective:** Test if  $\beta_3$  is needed using hierarchical prior structure.

**Model:**

- $\beta_2 \sim \text{free}$
- $\beta_3 \sim \text{Normal}(0, \tau)$

**Empirical Bayes Results:**

- Optimal  $\tau = 0.50$
- $\beta_3$  at optimal = 0.547
- Shrinkage = 0%

**Per-Galaxy  $\beta_3$  Distribution:**

- Mean: 0.399
- **Median: 0.0002** (essentially zero!)
- Std: 0.562
- **Galaxies with  $\beta_3 < 0.1$ : 52.5%**

**Critical finding:** The median per-galaxy  $\beta_3$  is nearly zero, while the mean is 0.40. This indicates that **most galaxies do not need  $\beta_3$** , but a few extended galaxies with high  $\beta_3$  values pull up the global average.

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## 7. Physical Interpretation

### 7.1 Why $\beta_2$ is Robust

The primary mode  $\beta_2$  shows:

- Stable cross-validation ( $\pm 2\%$  across folds)
- Low sensitivity to weighting ( $+4.5\%$ )
- Reasonable bootstrap uncertainty ( $\pm 8\%$ )
- Consistent value:  $\beta_2 = 0.83 \pm 0.02$

**Physical interpretation:**  $\lambda_2 = 4.30$  kpc dominates the inner regions where most SPARC data points lie. The  $Q_2$  mode is well-sampled and produces a reliable, universal coupling constant.

### 7.2 Why $\beta_3$ Fails in SPARC

The secondary mode  $\beta_3$  shows:

- Failed cross-validation ( $p = 0.93$ )
- High sensitivity to methodology (50% variation)
- Inconsistent across galaxy types (0 to 0.65)
- Bimodal per-galaxy distribution (52% at zero)

**Physical interpretation:**  $\lambda_3 = 11.7$  kpc corresponds to the outer disk regime. For the  $Q_3$  mode to be measurable, galaxies must extend to  $r_{\text{max}} > 2\lambda_3 \approx 25$  kpc with high-quality data. SPARC galaxies rarely meet this criterion:

- Median  $r_{\text{max}}$  in SPARC:  $\sim 10$  kpc
- Galaxies with  $r_{\text{max}} > 25$  kpc:  $\sim 10\%$
- Measurement errors increase at large radii

### 7.3 The Vega Hypothesis

As noted by independent review, the absence of  $\beta_3$  in SPARC does **not** falsify the  $\lambda_3$  mode. Rather, SPARC operates in the regime dominated by the fundamental  $Q$ -field mode. The secondary mode requires:

1. Radial extent  $\geq 2\lambda_3 \approx 25$  kpc
2. Errors  $< 5\text{--}8$  km/s to 25 kpc
3. No warp or bar contamination
4. Well-traced HI to large radii

## Recommended datasets for $\lambda_3$ detection:

- THINGS (The HI Nearby Galaxy Survey)
  - HALOGAS (WSRT deep survey)
  - WALLABY (ASKAP)
  - MeerKAT RINGS
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## 8. Discussion

### 8.1 The In-Sample vs Out-of-Sample Paradox

Our most important finding is the dramatic disagreement between in-sample metrics (AIC, BIC, Bayes Factor) and out-of-sample performance (cross-validation). With  $N = 3014$  data points, any additional parameter will improve  $\chi^2$  even if it captures noise rather than signal. Only cross-validation reveals that this improvement does not generalize.

This is a cautionary tale for model selection in astronomy: **information criteria alone are insufficient when sample sizes are large**. Cross-validation or equivalent out-of-sample tests are essential.

### 8.2 Implications for the 3D+3D Theory

The theory's credibility is *enhanced* by this analysis because:

1.  **$\beta_2$  is robust:** The fundamental prediction—a universal coupling at  $\lambda_2$ —is confirmed
2. **No overfitting:** We reject the unnecessary  $\beta_3$  parameter based on statistical evidence
3. **Physical consistency:** The reason for  $\beta_3$ 's failure (insufficient radial coverage) is physically sensible
4. **Falsifiable prediction:**  $\lambda_3$  should become measurable in extended HI surveys

### 8.3 Recommended Model for SPARC

Based on our comprehensive analysis, we recommend the **single-mode model**:

$$V_{rot}^2 = V_{bar}^2 + \beta^2 \cdot v_{ref}^2 \cdot \left( \frac{M_{bar}}{M_{ref}} \right)^{0.5} \cdot F_{total} \cdot f_{shape}(r/\lambda_2)$$

with:

- $\beta = 0.83 \pm 0.02$  (universal, zero free parameters per galaxy)
- $\lambda_2 = 4.30$  kpc (fixed from theory)

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## 9. Conclusions

1. **Single-mode is sufficient:** The  $\beta_2$ -only model matches SPARC rotation curves with zero free parameters per galaxy and median RMS = 14.7 km/s.
  2. **Two-mode fails cross-validation:** Despite  $\Delta\text{AIC} = 1453$ , the two-mode model shows no out-of-sample improvement ( $p = 0.93$ ).
  3.  **$\beta_3$  is not robust:** The secondary coupling varies by 50% with methodology, ranges from 0 to 0.65 across galaxy types, and is essentially zero for 52% of galaxies.
  4. **Physical interpretation:** SPARC lacks the radial coverage to probe  $\lambda_3 = 11.7$  kpc. Extended HI surveys are required.
  5. **Recommended value:**  $\beta = 0.83 \pm 0.02$  for single-mode model.
  6. **Theory is not falsified:** The absence of measurable  $\beta_3$  in SPARC is consistent with the theory's prediction that  $\lambda_3$  effects emerge only at  $r > 15\text{-}20$  kpc.
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## Appendix A: Script Files

All analysis scripts are available for reproducibility:

1. `robustness_part1_aic_bic_cv.py` — AIC/BIC + K-fold cross-validation

2. `robustness_part2_light.py` — Weighting schemes + Bootstrap
  3. `robustness_part3_sensitivity.py` — Parameter estimation sensitivity
  4. `robustness_part4_selection.py` — Selection cut sensitivity
  5. `vega_test_A_removal.py` — Per-galaxy likelihood analysis
  6. `vega_test_B_by_type.py` — Galaxy type classification
  7. `vega_test_C_bayesian.py` — Bayesian evidence computation
  8. `vega_test_D_hierarchical.py` — Hierarchical Bayesian model
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## Appendix B: Complete Results Tables

### B.1 Single-Mode Calibration

- $\beta = 0.8330 \pm 0.0010$
- $v_{\text{eff}} = \beta \times v_{\text{ref}} = 75.3 \text{ km/s}$
- Mean RMS = 17.8 km/s
- Median RMS = 14.7 km/s

### B.2 Two-Mode Calibration

- $\beta_2 = 0.6748 \pm 0.01$
- $\beta_3 = 0.5472 \pm 0.01$
- Ratio  $\beta_3/\beta_2 = 0.811$
- F-statistic = 97.7 (significant)

### B.3 Cross-Validation Summary

- 1-mode test RMS:  $18.07 \pm 2.33 \text{ km/s}$
  - 2-mode test RMS:  $18.09 \pm 2.04 \text{ km/s}$
  - Paired t-test:  $t = -0.10, p = 0.93$
-