

Evidence for Gravitational Screening in Massive Elliptical Galaxies: A 25σ Detection from SLACS Strong Lensing Data

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Abstract

We present a statistically robust detection of gravitational screening in massive elliptical galaxies using 66 strong gravitational lenses from the Sloan Lens ACS (SLACS) survey. Comparing lensing masses (M_{Ein}) with dynamical masses (M_{dyn}) derived from stellar velocity dispersions, we find $M_{\text{Ein}}/M_{\text{dyn}} = 0.603 \pm 0.043$, indicating a 39.7% screening effect at 9.2σ significance. Within the framework of the 3D+3D Discrete Spacetime Theory, we introduce the density-dependent screening scale $\lambda_{\text{core}}(\sigma) = \lambda_2/\sqrt{1 + \psi/\psi_{\text{crit}}}$, derived from the non-linear term $\gamma(\nabla^2 Q)^2$ in the 6D quantum field theory. At resonance ($R_E \approx \lambda_{\text{core}}$), screening reaches 61.8% with 25.7σ significance. The Spearman correlation $\rho = 0.675$ ($p = 5 \times 10^{-10}$) between $M_{\text{Ein}}/M_{\text{dyn}}$ and distance from resonance confirms the predicted V-shaped pattern. Bootstrap analysis ($N = 5000$) yields 95% confidence intervals of $[19\sigma, 37\sigma]$ for the resonance detection. These results provide strong observational evidence for gravitational screening in the non-linear regime of modified gravity.

Keywords: gravitational lensing, modified gravity, screening mechanisms, elliptical galaxies, dark matter alternative

1. Introduction

Strong gravitational lensing by massive elliptical galaxies provides a unique laboratory for testing theories of gravity on kiloparsec scales. The SLACS survey (Bolton et al. 2006, 2008; Auger et al. 2009) has assembled the largest homogeneous sample of galaxy-scale strong lenses, enabling precision comparisons between lensing and dynamical mass estimates.

In standard General Relativity (GR), the lensing mass within the Einstein radius should equal the dynamical mass inferred from stellar velocity dispersions. Any systematic deviation would indicate either: (a) systematic errors in mass modeling, (b) the presence of dark matter with a different spatial distribution than baryons, or (c) modifications to gravity.

The 3D+3D Discrete Spacetime Theory (Calzighetti 2025a,b,c) proposes a six-dimensional spacetime with signature $(-, +, +, +, -, -)$, where two temporal dimensions (τ_2, τ_3) are compactified at galactic scales. The breathing modes of these compact dimensions generate a scalar field Q that couples to baryonic matter, producing effects observationally similar to dark matter. Crucially, the non-linear self-interaction term $\gamma(\nabla^2 Q)^2$ leads to screening: the effective gravitational mass is reduced relative to the dynamical mass in dense environments.

In this paper, we test the screening prediction using SLACS data, comparing lensing masses (M_{Ein}) with dynamical masses (M_{dyn}). We find strong evidence for $M_{\text{Ein}} < M_{\text{dyn}}$ at 9.2σ significance, with maximum screening at the predicted resonance scale.

2. Theoretical Framework

2.1 The 3D+3D Theory

The 3D+3D theory posits that spacetime has six dimensions with metric signature $(-, +, +, +, -, -)$. After dimensional reduction to 4D, the effective Lagrangian contains a scalar field Q representing the breathing mode of the compact temporal dimensions:

$$\mathcal{L}_{4D} = \mathcal{L}_{GR} + \frac{1}{2}(\partial_\mu Q)^2 - \frac{1}{2}m_Q^2 Q^2 - \gamma(\nabla^2 Q)^2 + \beta Q T^{\text{bar}} \quad (1)$$

where m_Q is the Q -field mass (determining the characteristic scale $\lambda_2 = 4.30$ kpc from SPARC rotation curves), γ is the non-linear coupling constant, β is the matter coupling, and T^{bar} is the baryonic stress-energy trace.

2.2 The Screening Mechanism

The non-linear term $\gamma(\nabla^2 Q)^2$ causes the effective screening scale to depend on the local gravitational potential. Solving the modified Klein-Gordon equation in the presence of a massive elliptical galaxy with velocity dispersion σ , we obtain:

$$\lambda_{\text{core}}(\sigma) = \lambda_2 / \sqrt{1 + \psi/\psi_{\text{crit}}} \quad (2)$$

where $\psi = (\sigma/c)^2$ is the dimensionless gravitational potential and $\psi_{\text{crit}} = (\sigma_{\text{crit}}/c)^2$ defines the threshold where non-linear effects become important. From optimization against SLACS data, we find $\sigma_{\text{crit}} \approx 150$ km/s.

For SLACS ellipticals with $\sigma \sim 200$ -300 km/s, this yields $\lambda_{\text{core}} \approx 2$ -3 kpc, significantly smaller than $\lambda_2 = 4.30$ kpc applicable to spiral galaxies.

2.3 Resonance Condition

Screening is maximized when the Einstein radius R_E matches the screening scale λ_{core} . This 'resonance condition' $R_E \approx \lambda_{\text{core}}$ leads to the prediction of a V-shaped pattern: screening should decrease monotonically as $|\log(R_E/\lambda_{\text{core}})|$ increases from zero.

3. Data and Methods

3.1 SLACS Sample

We analyze 66 galaxy-scale strong lenses from the SLACS survey (Auger et al. 2009). The sample spans lens redshifts $z_l = 0.06$ -0.43 and velocity dispersions $\sigma = 164$ -404 km/s. For each lens we have:

1. Einstein radius θ_E from HST imaging
2. Stellar velocity dispersion σ from SDSS spectroscopy
3. Stellar mass M_{star} from photometry (Chabrier IMF)
4. Lensing mass $M_{\text{Ein}} = \pi\theta_E^2 \Sigma_{\text{crit}}$ within Einstein radius

3.2 Mass Estimates

The dynamical mass is computed from the virial theorem:

$$M_{\text{dyn}} = k(n) \times \sigma^2 \times R_{\text{eff}} / G \quad (3)$$

where $k(n)$ is the virial coefficient depending on the Sérsic index. The lensing mass M_{Ein} is derived directly from the observed Einstein radius using the thin lens equation.

3.3 Statistical Methods

We test the null hypothesis $H_0: M_{\text{Ein}}/M_{\text{dyn}} = 1$ (no screening) using: (a) Student's t-test for the mean deviation from unity; (b) Wilcoxon signed-rank test as a non-parametric alternative; (c) Spearman rank correlation to test the V-shape pattern; (d) Bootstrap resampling ($N = 5000$) to estimate confidence intervals.

4. Results

4.1 Main Result: $M_{\text{Ein}} < M_{\text{dyn}}$

The key finding is presented in Table 1. Across the full sample of 66 lenses:

Table 1: Main Results

Quantity	Value	Significance
Sample size	66	—
$\langle M_{\text{Ein}}/M_{\text{dyn}} \rangle$	0.603 ± 0.043	9.2σ
Screening fraction	39.7%	$p = 2.5 \times 10^{-13}$
Wilcoxon test	$W = 143$	$p = 7.8 \times 10^{-10}$

The lensing mass is systematically lower than the dynamical mass by $\sim 40\%$. Both parametric (t-test) and non-parametric (Wilcoxon) tests reject the null hypothesis at extreme significance levels.

4.2 Resonance Analysis

Using the λ_{core} theory with $\sigma_{\text{crit}} = 150$ km/s, we identify 22 galaxies at resonance ($0.7 < R_E/\lambda_{\text{core}} < 1.5$). The results are presented in Table 2.

Table 2: Resonance vs Far-from-Resonance Comparison

Region	N	$\langle M_{\text{Ein}}/M_{\text{dyn}} \rangle$	Signif.
At resonance	22	0.382 ± 0.024	25.7σ
Far from resonance	28	0.790 ± 0.081	2.6σ

At resonance, screening reaches 61.8% ($M_{\text{Ein}}/M_{\text{dyn}} = 0.382$) with 25.7σ significance. Far from resonance, screening is only 21% and barely significant. This dramatic difference confirms the resonance prediction.

4.3 V-Shape Pattern

The theory predicts screening should decrease monotonically with distance from resonance. Table 3 shows the binned analysis.

Table 3: V-Shape Pattern Analysis

Distance [dex]	N	$\langle M_{\text{Ein}}/M_{\text{dyn}} \rangle$	Screening
0.00 - 0.10	13	0.366 ± 0.030	63.4%
0.10 - 0.20	13	0.461 ± 0.064	53.9%
0.20 - 0.35	22	0.587 ± 0.046	41.3%

Distance [dex]	N	$\langle M_{\text{Ein}}/M_{\text{dyn}} \rangle$	Screening
0.35 - 0.60	17	0.921 ± 0.114	7.9%

The Spearman rank correlation between $M_{\text{Ein}}/M_{\text{dyn}}$ and distance from resonance is $\rho = 0.675$ ($p = 5.2 \times 10^{-10}$), confirming the predicted positive correlation (screening decreases with distance). This provides strong support for the λ_{core} theory.

4.4 Bootstrap Robustness

Bootstrap resampling ($N = 5000$) yields the following 95% confidence intervals:

5. Overall $M_{\text{Ein}}/M_{\text{dyn}}$: [0.524, 0.692]
6. Resonance $M_{\text{Ein}}/M_{\text{dyn}}$: [0.335, 0.430]
7. Resonance significance: [19σ , 37σ]
8. Spearman ρ : [0.514, 0.793]

All confidence intervals exclude the null hypothesis values (1.0 for mass ratios, 0 for correlation), demonstrating the robustness of our results.

5. Discussion

5.1 Physical Interpretation

The finding that $M_{\text{lensing}} < M_{\text{dynamical}}$ has a natural interpretation in the 3D+3D framework. The dynamical mass, measured from stellar motions within the galaxy, includes the full gravitational effect of the Q-field. However, gravitational lensing probes scales comparable to or larger than the Einstein radius ($R_E \sim 3\text{--}5$ kpc), where screening reduces the effective mass.

The 61.8% screening at resonance ($R_E \approx \lambda_{\text{core}}$) implies that nearly two-thirds of the gravitational mass is 'hidden' from long-range interactions when the probing scale matches the screening wavelength. This is consistent with the non-linear saturation predicted by the $\gamma(\nabla^2 Q)^2$ term.

5.2 Comparison with Standard Dark Matter

In the standard Λ CDM paradigm with dark matter halos, one would expect $M_{\text{lensing}} \approx M_{\text{dynamical}}$ within the Einstein radius, where both baryons and dark matter contribute. The systematic deficit we observe (40% on average, 62% at resonance) cannot be explained by standard NFW halos.

We note that Auger et al. (2010) previously reported $M_{\text{Ein}} < M_{\text{dyn}}$ for some SLACS lenses and attributed this to IMF variations or stellar mass uncertainties. However, these explanations cannot account for the strong correlation with $R_E/\lambda_{\text{core}}$ that we find, which is a specific prediction of the screening theory.

5.3 Outliers

We identify 4 galaxies with anomalously low $M_{\text{Ein}}/M_{\text{dyn}} < 0.25$ that cause a minor bump in the V-shape pattern at 0.08–0.15 dex. These outliers may represent: (a) measurement errors in θ_E or σ ; (b) unusually strong local screening from environmental effects; or (c) line-of-sight contamination. Removing these outliers increases the clean sample significance to 13.5σ without qualitatively changing our conclusions.

6. Conclusions

Using 66 strong gravitational lenses from the SLACS survey, we have detected gravitational screening at high statistical significance:

9. **Main result:** $M_{\text{Ein}}/M_{\text{dyn}} = 0.603 \pm 0.043$, representing 39.7% screening at 9.2σ significance.
10. **Resonance enhancement:** At $R_E \approx \lambda_{\text{core}}$, screening reaches 61.8% at 25.7σ .
11. **V-shape confirmation:** Spearman $\rho = 0.675$ ($p = 5 \times 10^{-10}$) confirms the predicted pattern.
12. **Bootstrap robustness:** 95% CI for resonance significance is $[19\sigma, 37\sigma]$.

These results provide the strongest observational evidence to date for gravitational screening in the non-linear regime, consistent with predictions of the 3D+3D Discrete Spacetime Theory. The density-dependent screening scale $\lambda_{\text{core}}(\sigma)$ explains why massive ellipticals show different screening behavior than spiral galaxies.

Future observations with Euclid and LSST will dramatically increase the sample of strong lenses, enabling even more stringent tests of the resonance prediction across a wider range of galaxy properties.

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