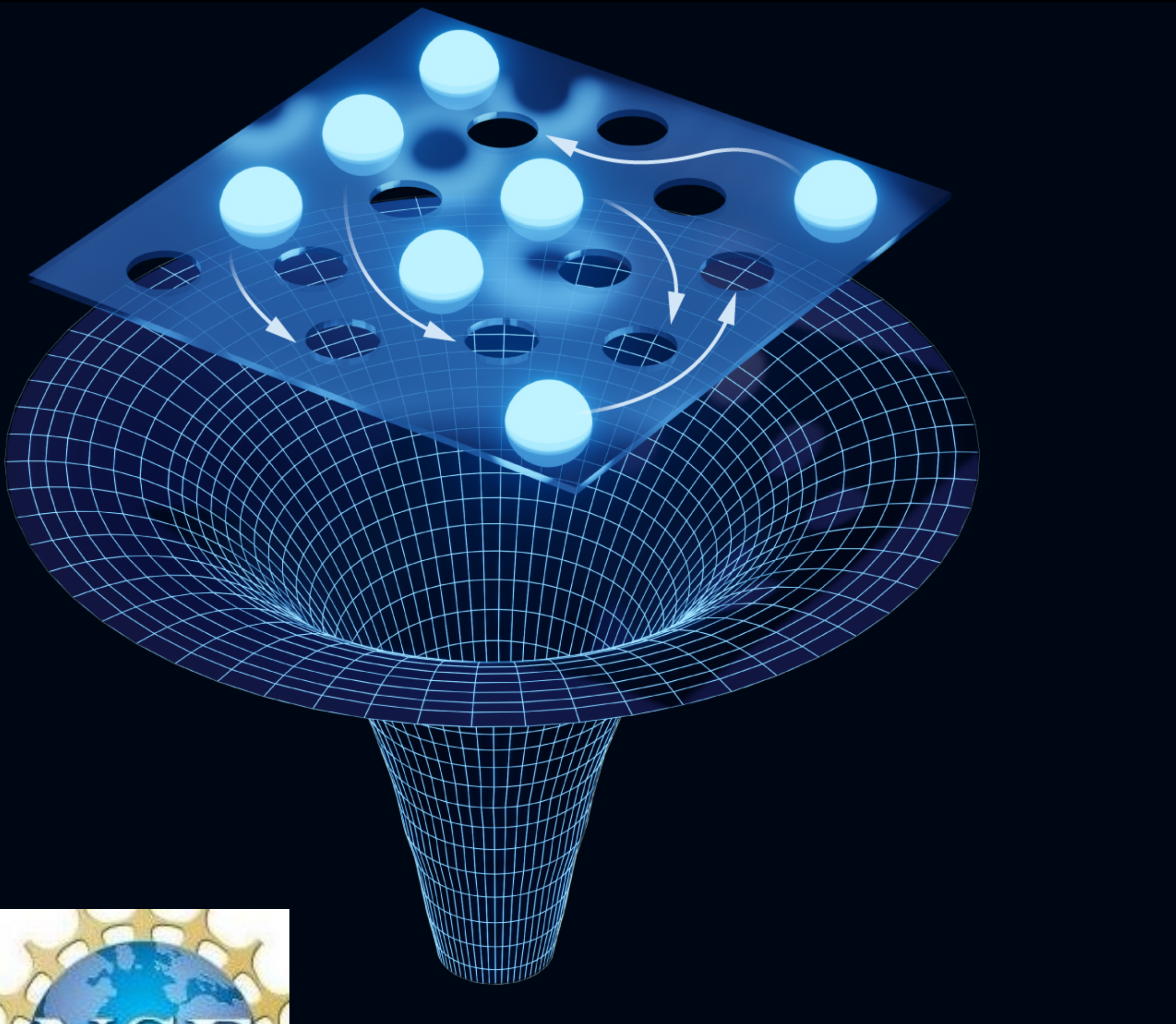


Quantum entanglement in nature: high temperature superconductors and black holes



New York University
October 3, 2024

Subir Sachdev



Talk online: sachdev.physics.harvard.edu



Boltzmann-Landau theory of metals

Statistical interpretation of entropy (1870)

$$S = k_B \log W$$

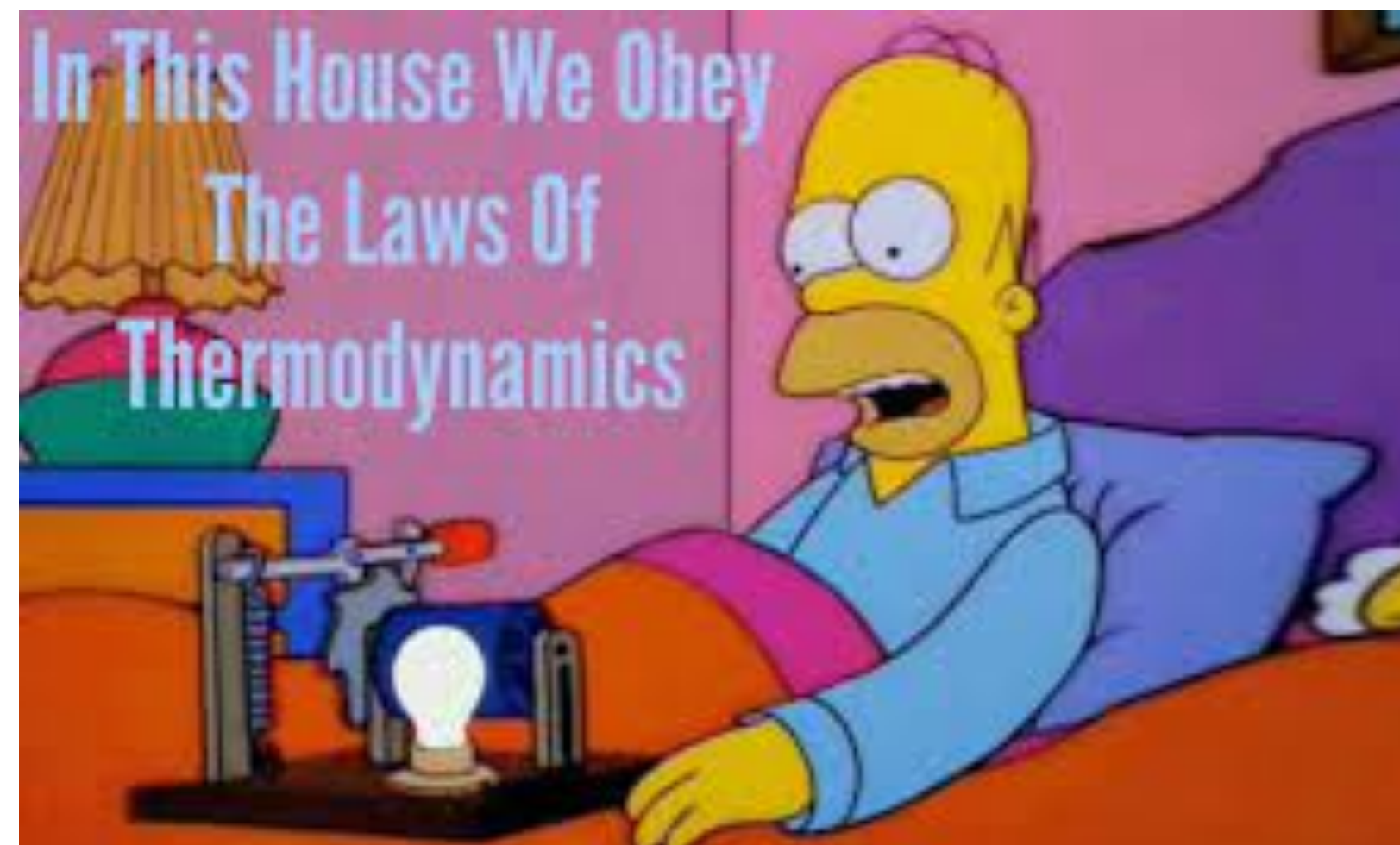
Density of quantum states $D(E) = \exp(S(E)/k_B)$

$$\frac{1}{T} = \frac{dS}{dE}$$



Ludwig Boltzmann

20 February 1844 - September 5, 1906
Vienna, Austria



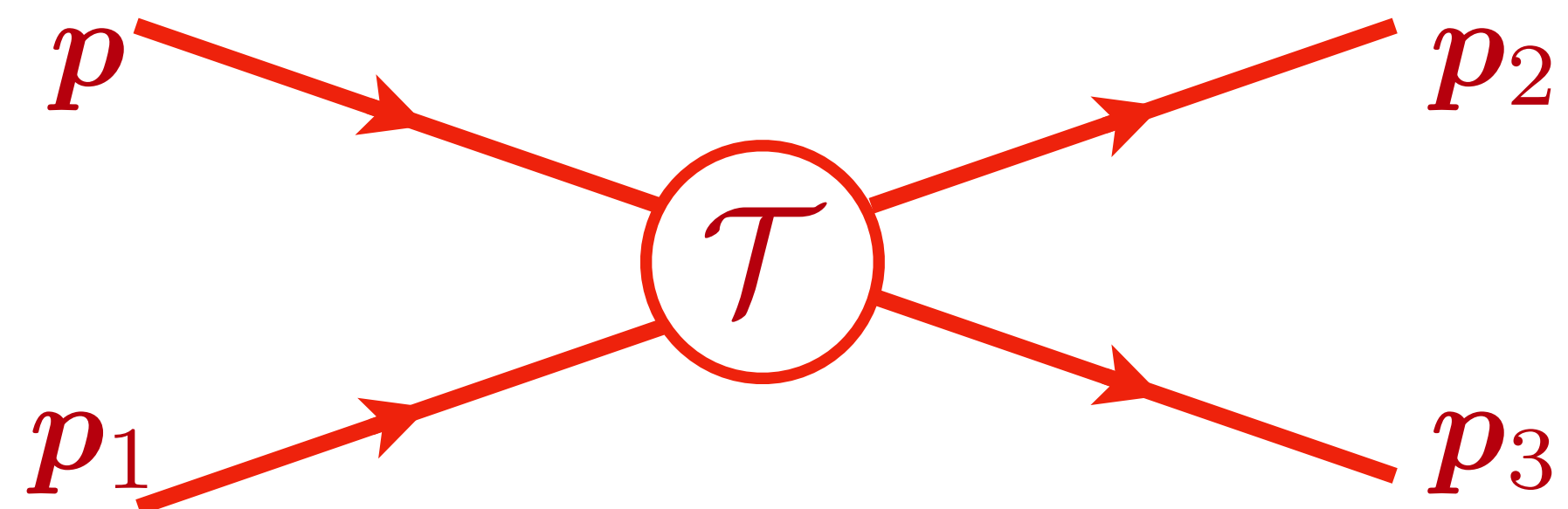
No perpetual
motion machines!

Boltzmann equation (1872)

Dilute classical gas

Molecular chaos: successive collisions are statistically independent

$$\frac{\partial f_{\mathbf{p}}}{\partial t} + \frac{\partial \varepsilon_{\mathbf{p}}}{\partial \mathbf{p}} \cdot \nabla_{\mathbf{r}} f_{\mathbf{p}} + \mathbf{F} \cdot \nabla_{\mathbf{p}} f_{\mathbf{p}} =$$
$$- 2\pi \int_{\mathbf{p}_{1,2,3}} |\mathcal{T}|^2 \delta(\varepsilon_{\mathbf{p}} + \varepsilon_{\mathbf{p}_1} - \varepsilon_{\mathbf{p}_2} - \varepsilon_{\mathbf{p}_3}) \delta(\mathbf{p} + \mathbf{p}_1 - \mathbf{p}_2 - \mathbf{p}_3)$$
$$\times [f_{\mathbf{p}} f_{\mathbf{p}_1} - f_{\mathbf{p}_2} f_{\mathbf{p}_3}]$$



Ludwig Boltzmann

20 February 1844 - September 5, 1906
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Ludwig Boltzmann
20 February 1844 - September 5, 1906
Vienna, Austria

Quantum Boltzmann equation (Landau)

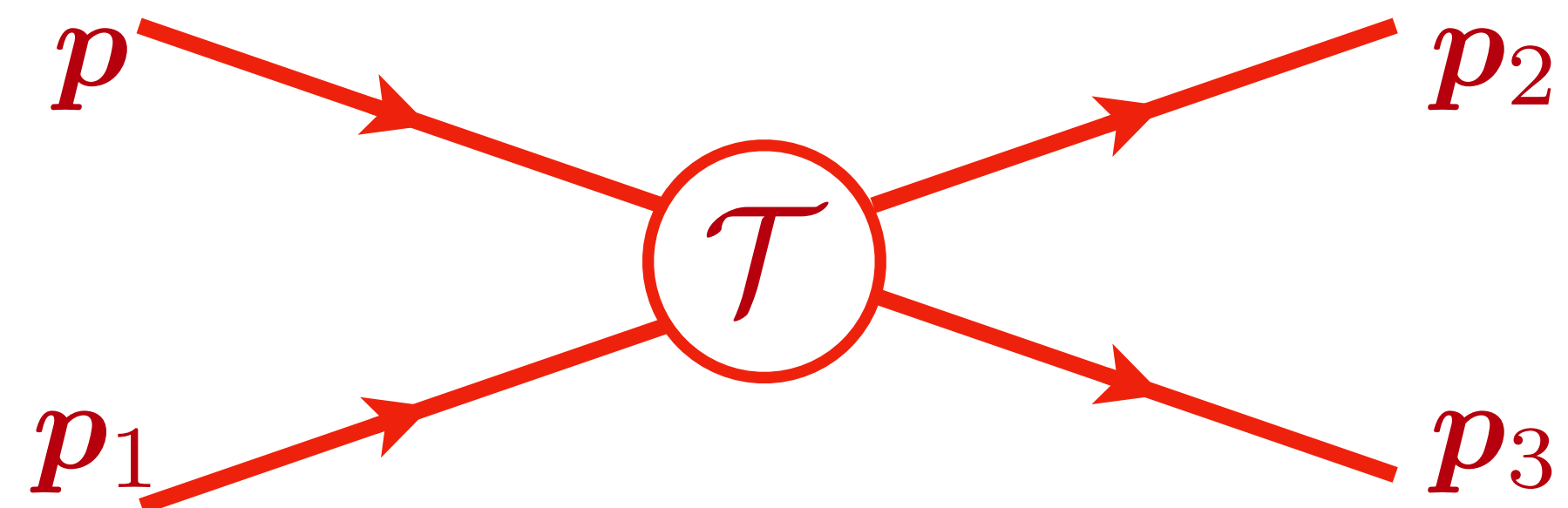
Dense gas of electrons

Neglects quantum interference (entanglement)
between successive collisions

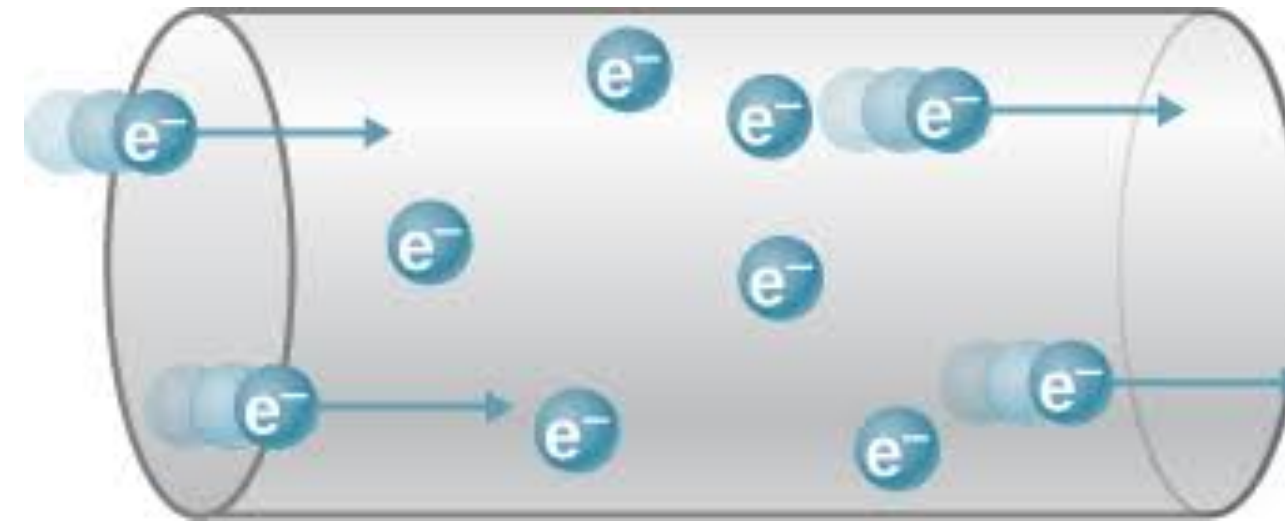
$$\frac{\partial f_{\mathbf{p}}}{\partial t} + \frac{\partial \varepsilon_{\mathbf{p}}}{\partial \mathbf{p}} \cdot \nabla_{\mathbf{r}} f_{\mathbf{p}} + \mathbf{F} \cdot \nabla_{\mathbf{p}} f_{\mathbf{p}} =$$

$$- 2\pi \int_{\mathbf{p}_{1,2,3}} |\mathcal{T}|^2 \delta(\varepsilon_{\mathbf{p}} + \varepsilon_{\mathbf{p}_1} - \varepsilon_{\mathbf{p}_2} - \varepsilon_{\mathbf{p}_3}) \delta(\mathbf{p} + \mathbf{p}_1 - \mathbf{p}_2 - \mathbf{p}_3)$$

$$\times [f_{\mathbf{p}} f_{\mathbf{p}_1} (1 - f_{\mathbf{p}_2}) (1 - f_{\mathbf{p}_3}) - f_{\mathbf{p}_2} f_{\mathbf{p}_3} (1 - f_{\mathbf{p}}) (1 - f_{\mathbf{p}_1})]$$



Current flow with electrons in ordinary metals



Flow of electrons described by Boltzmann equation \Rightarrow
typical scattering time $\tau \sim 1/(UT)^2$ (U is the strength of interactions),
resistivity $\rho(T) = \rho(0) + AT^2$

The time τ is much longer than a limiting ‘Planckian time’ $\frac{\hbar}{k_B T}$.

The long scattering time implies that individual electrons are well-defined.

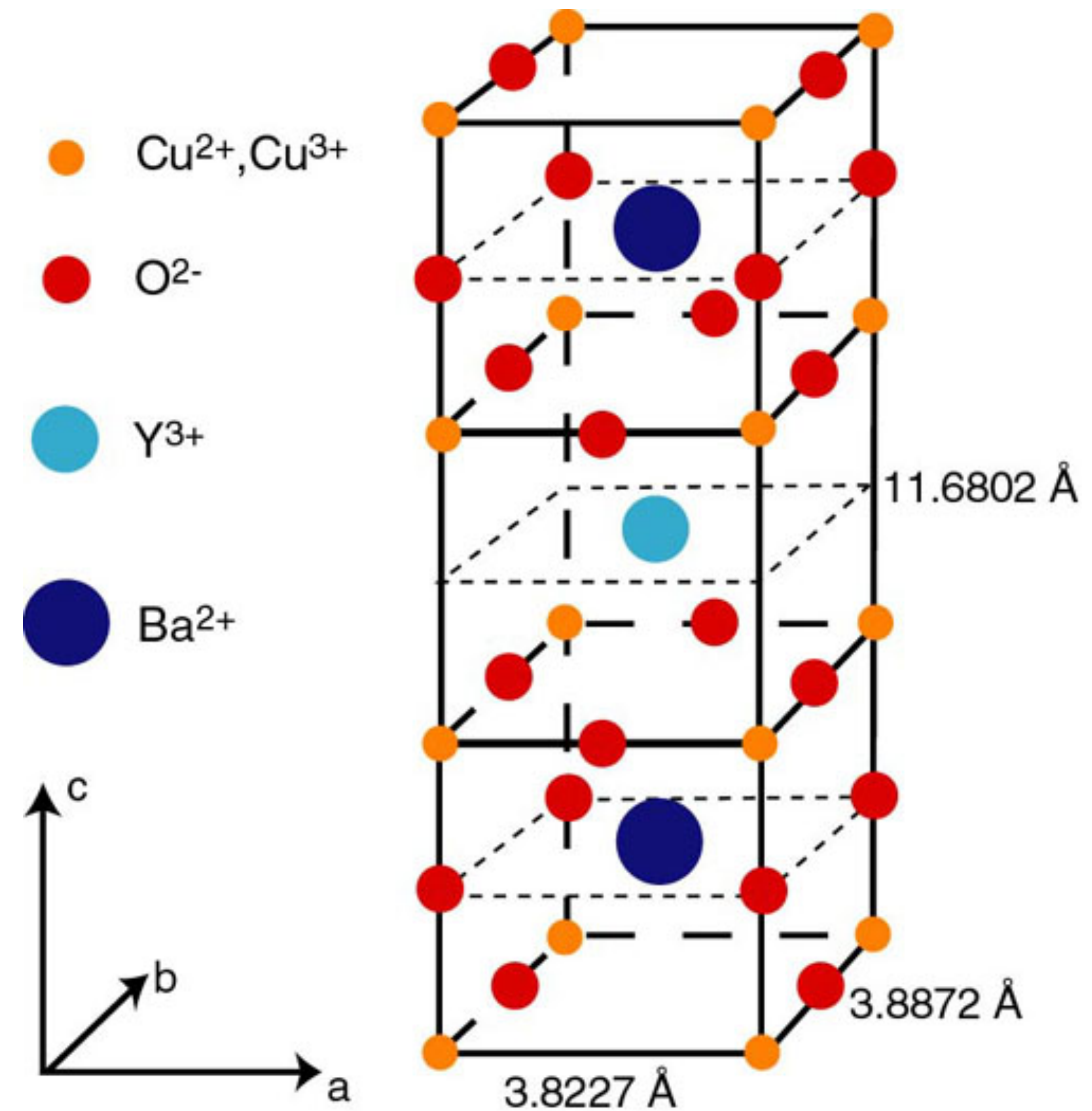
The motion of electrons is ‘ballistic’ or ‘integrable’
up to the long time τ , after which it is chaotic.

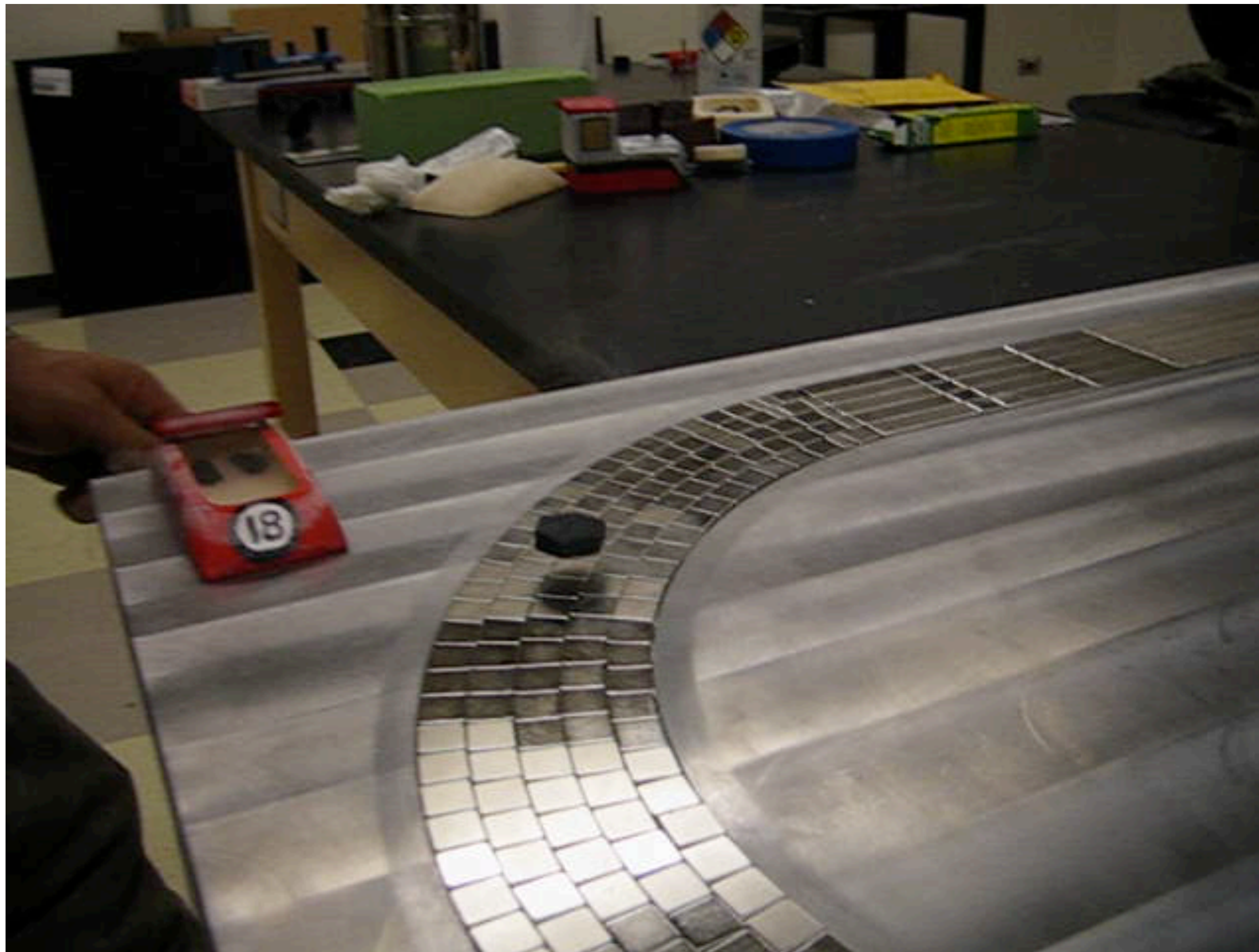
*High temperature
superconductivity*



Kamerlingh Onnes 1911:
Mercury is a superconductor below $-269\text{ }^{\circ}\text{C}$

Cuprate high temperature superconductors





Nd-Fe-B magnets, YBaCuO superconductor

Julian Hetel and Nandini Trivedi, Ohio State University

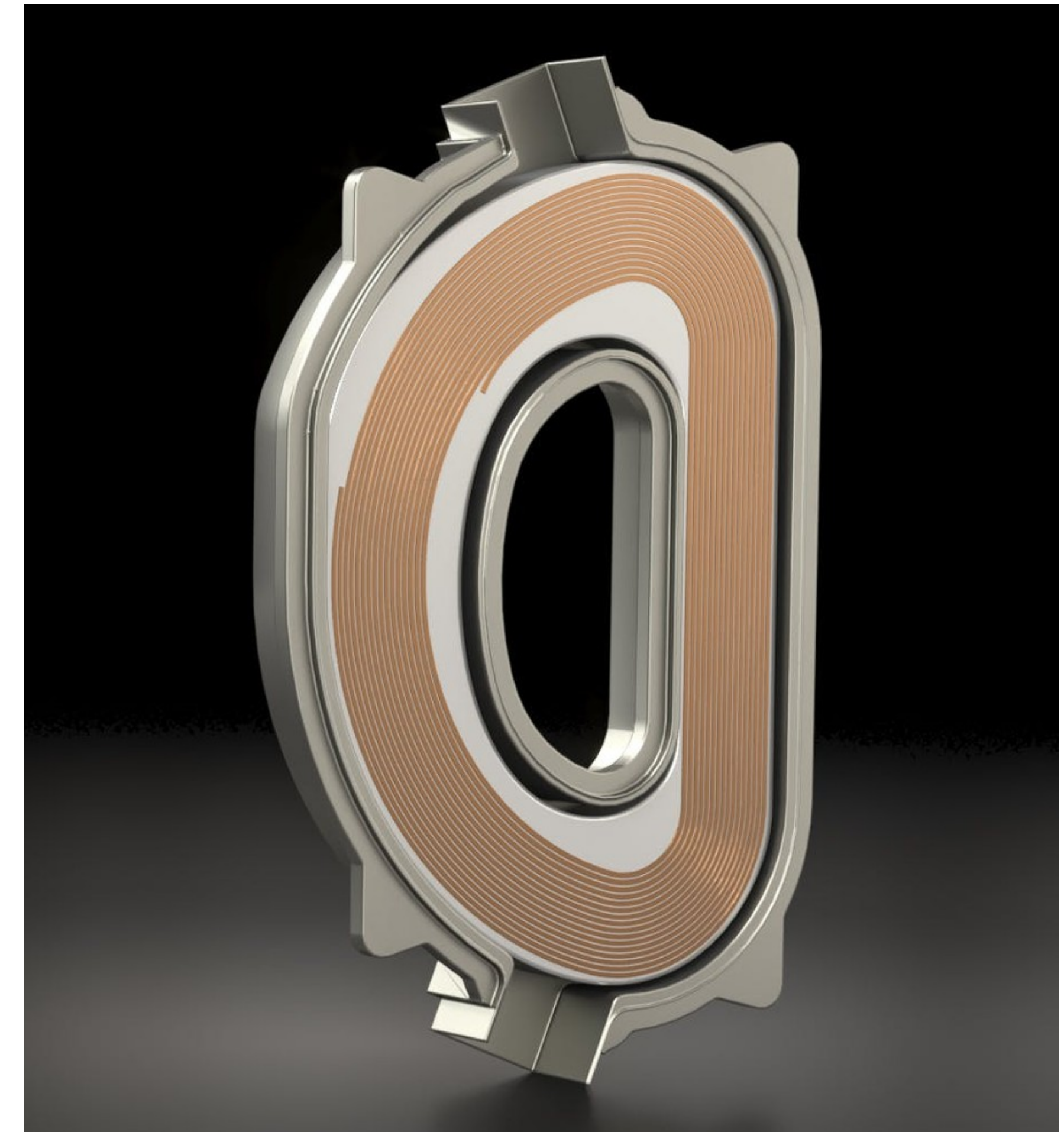
HTS Magnets: Enabling Technology

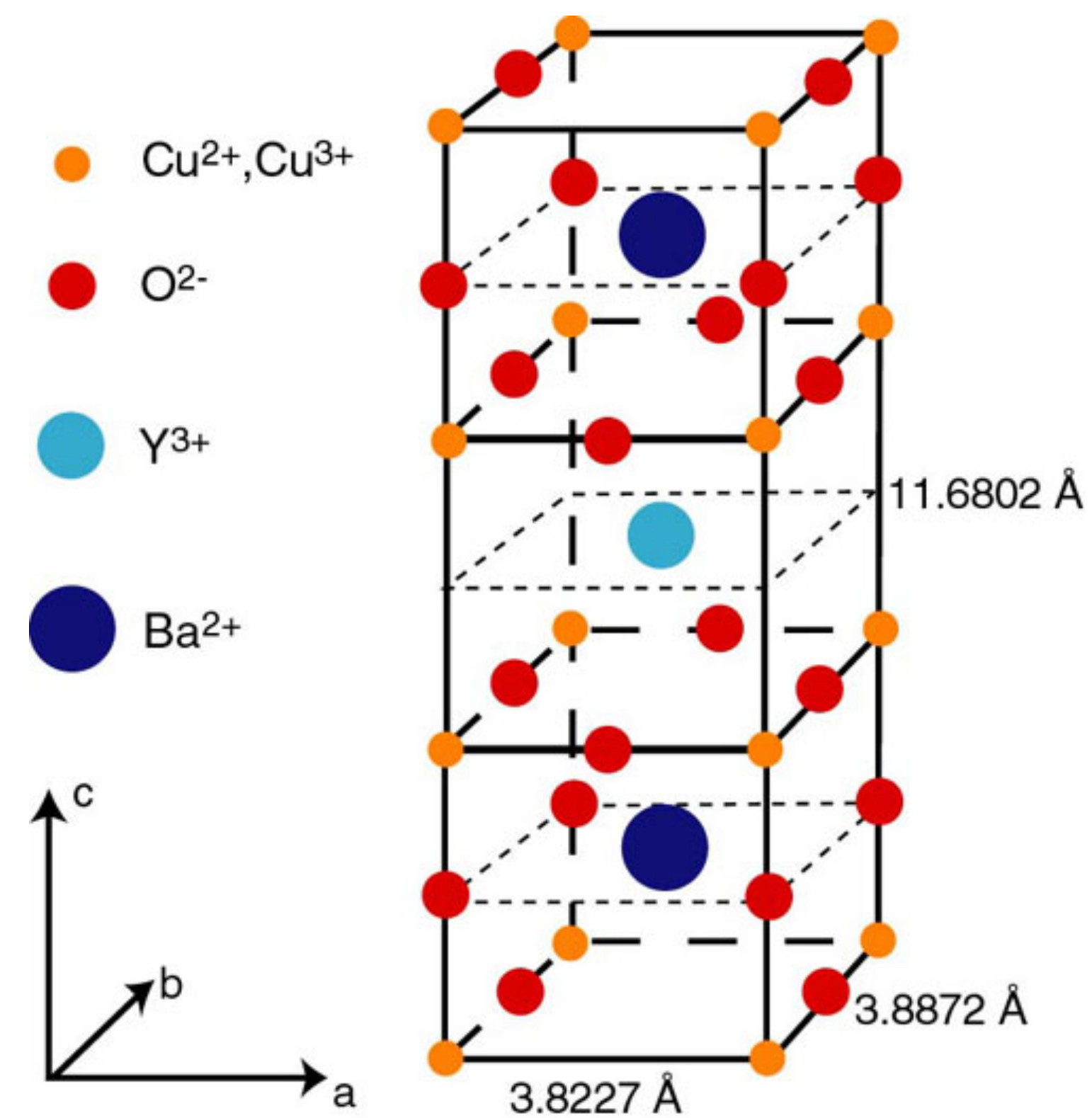
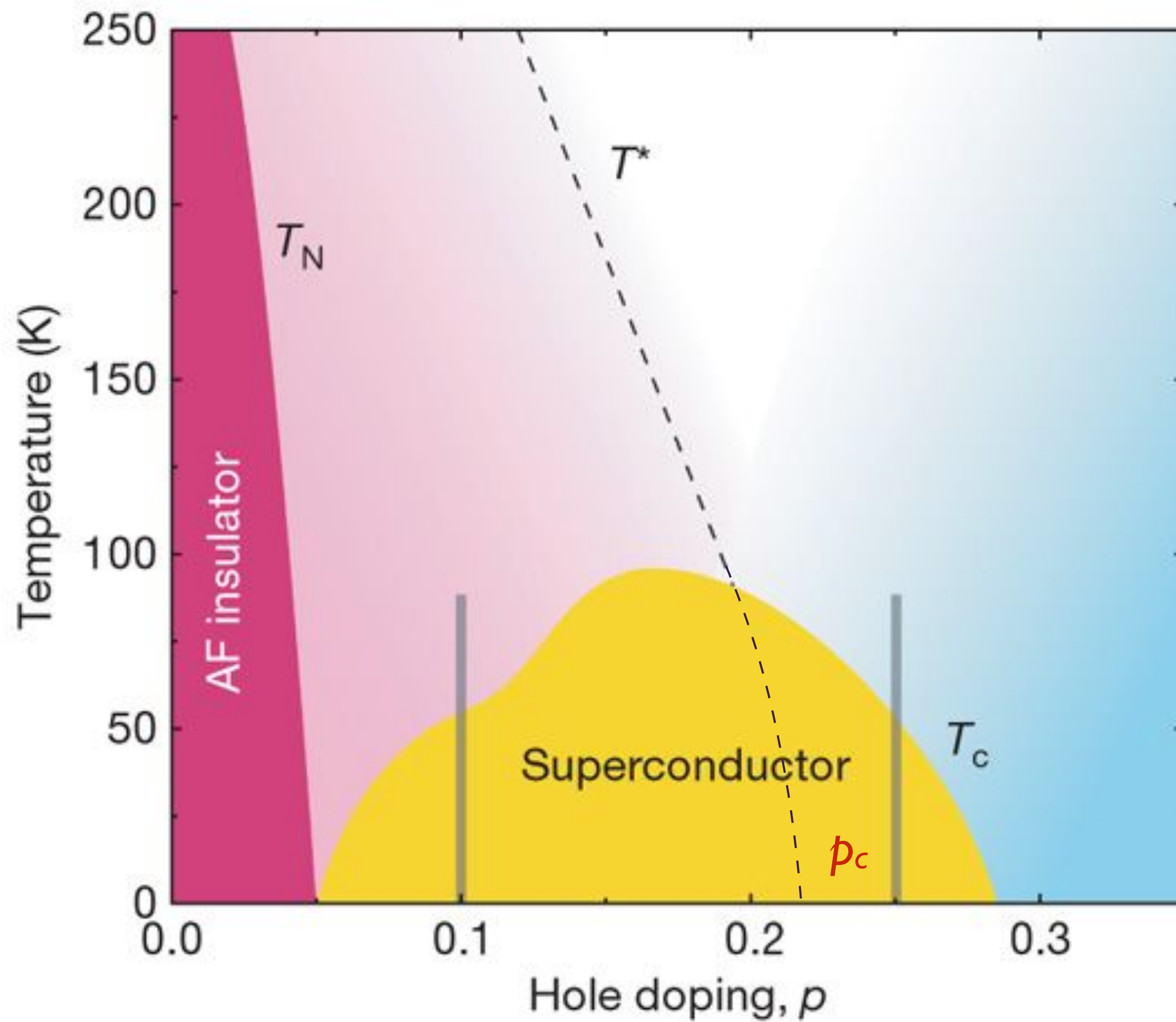
The surest path to limitless,
clean, fusion energy

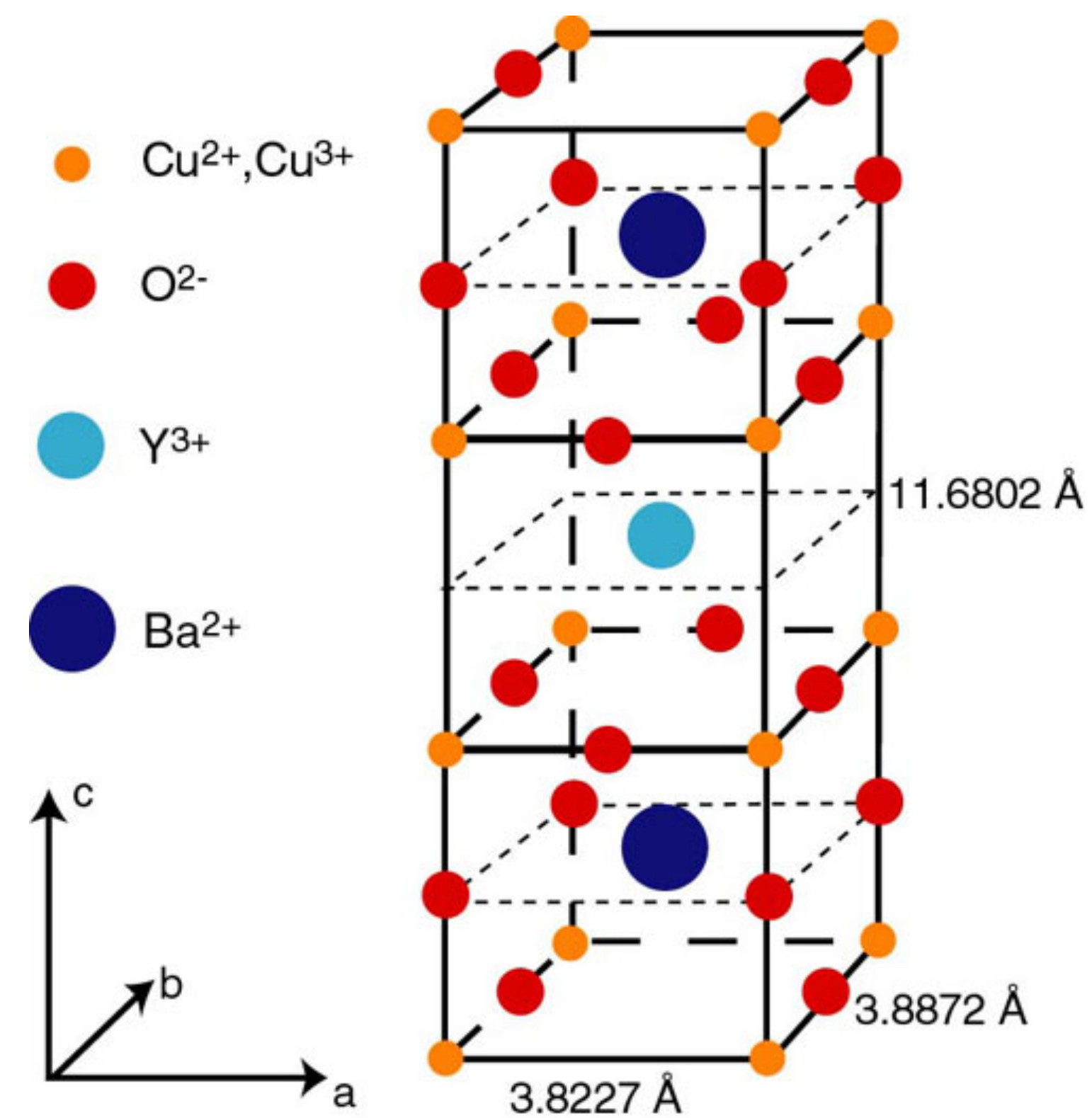
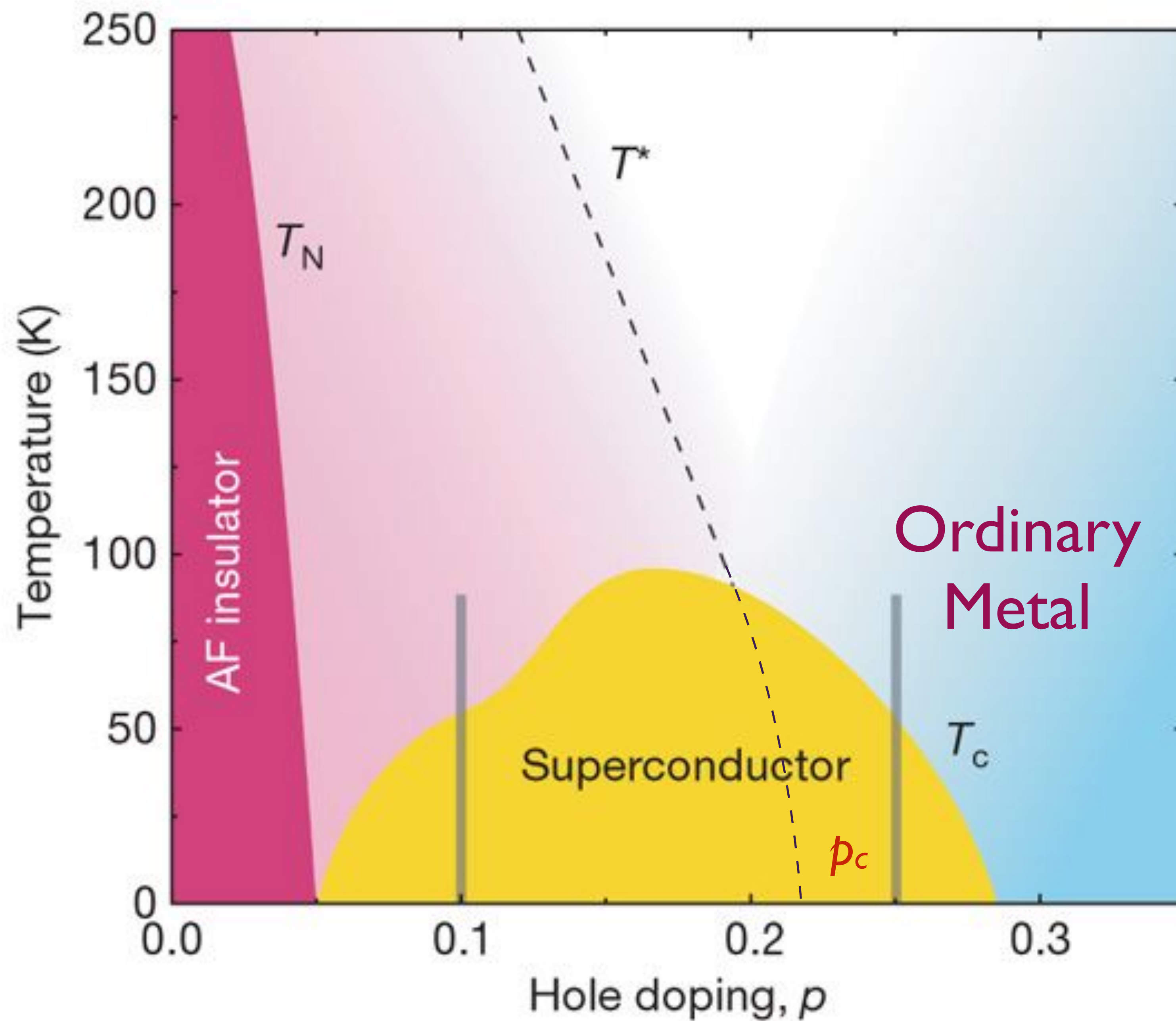
YBCO magnets allow for smaller,
faster, and less expensive
tokamaks for plasma fusion

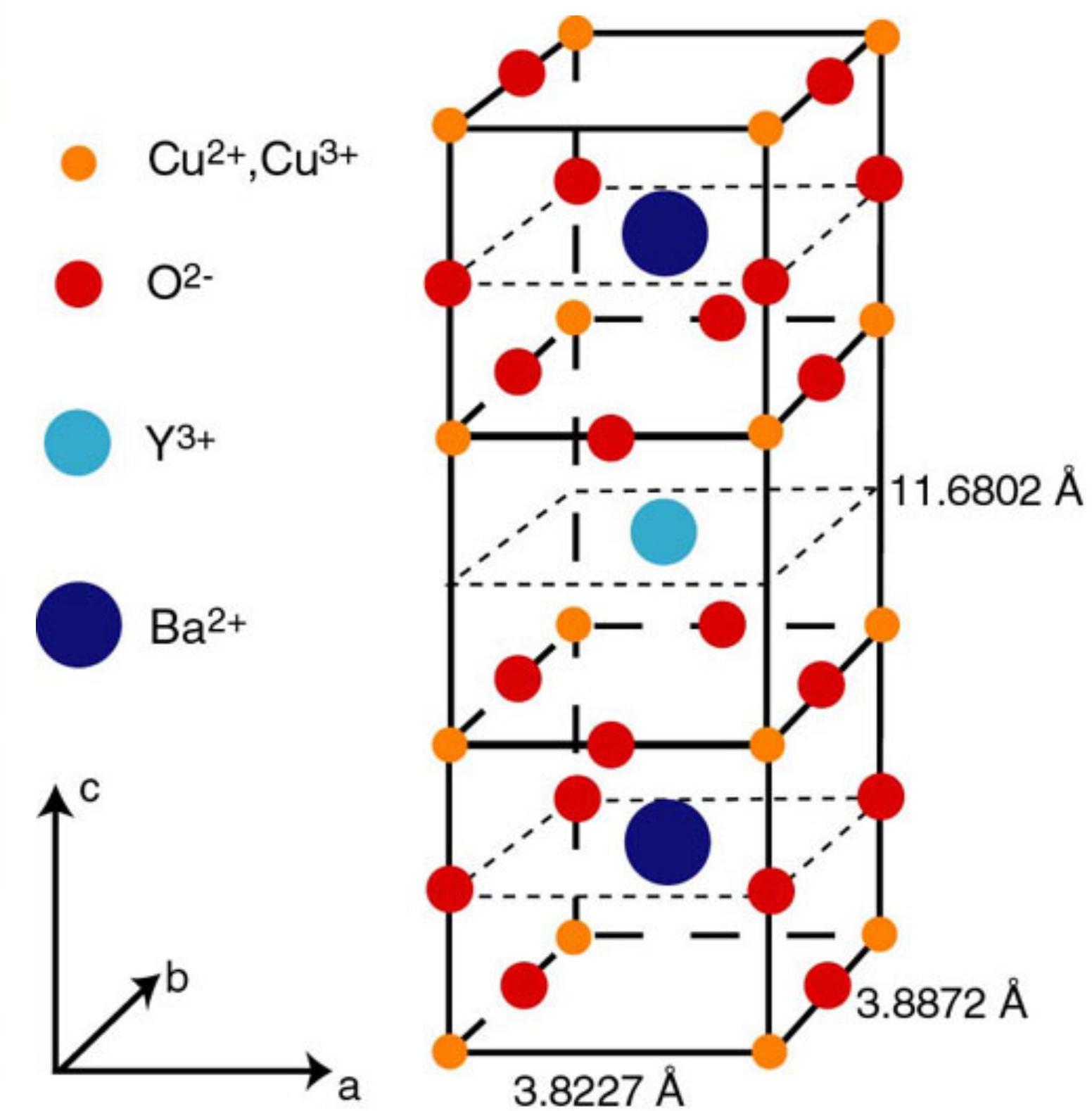
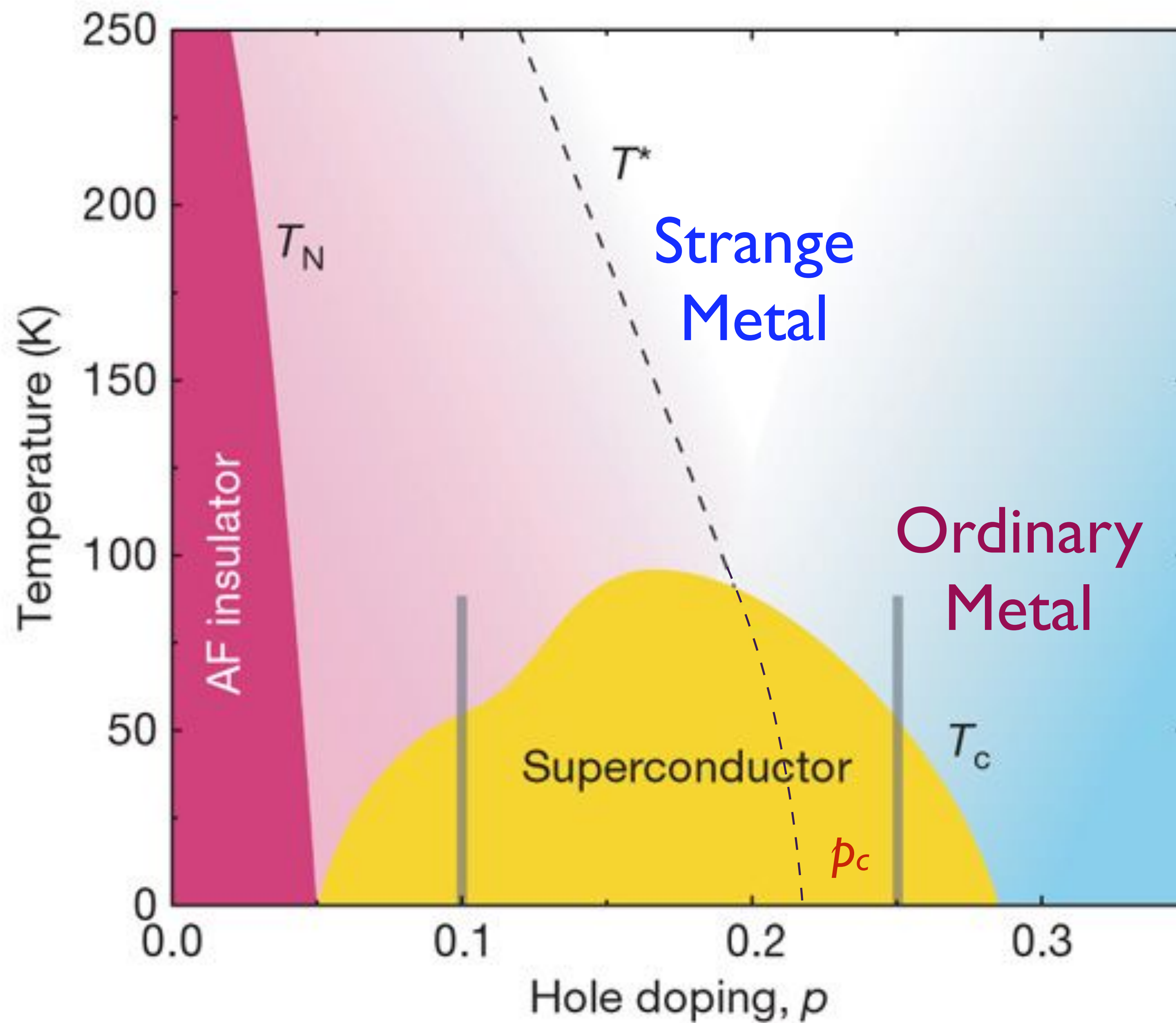


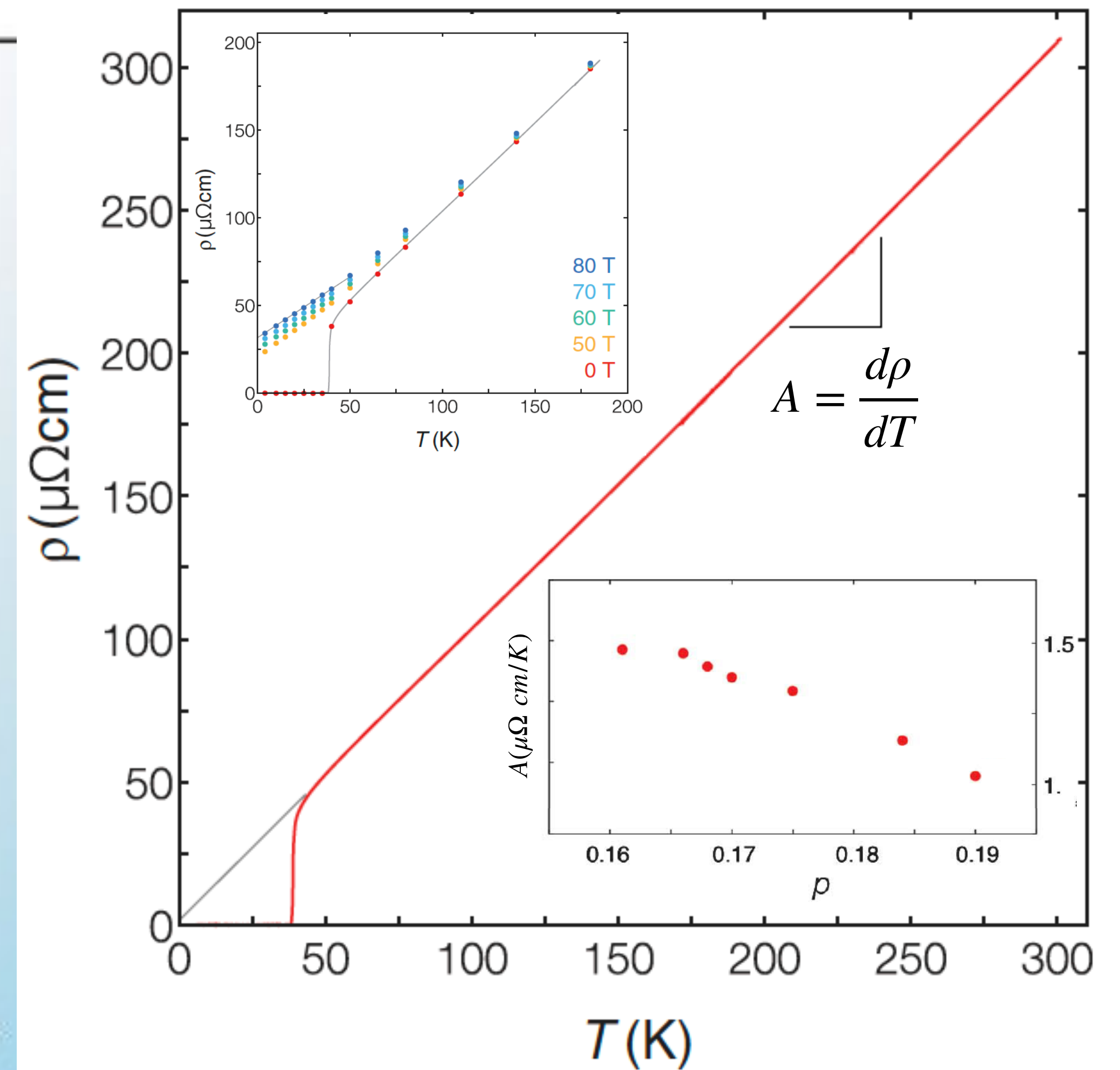
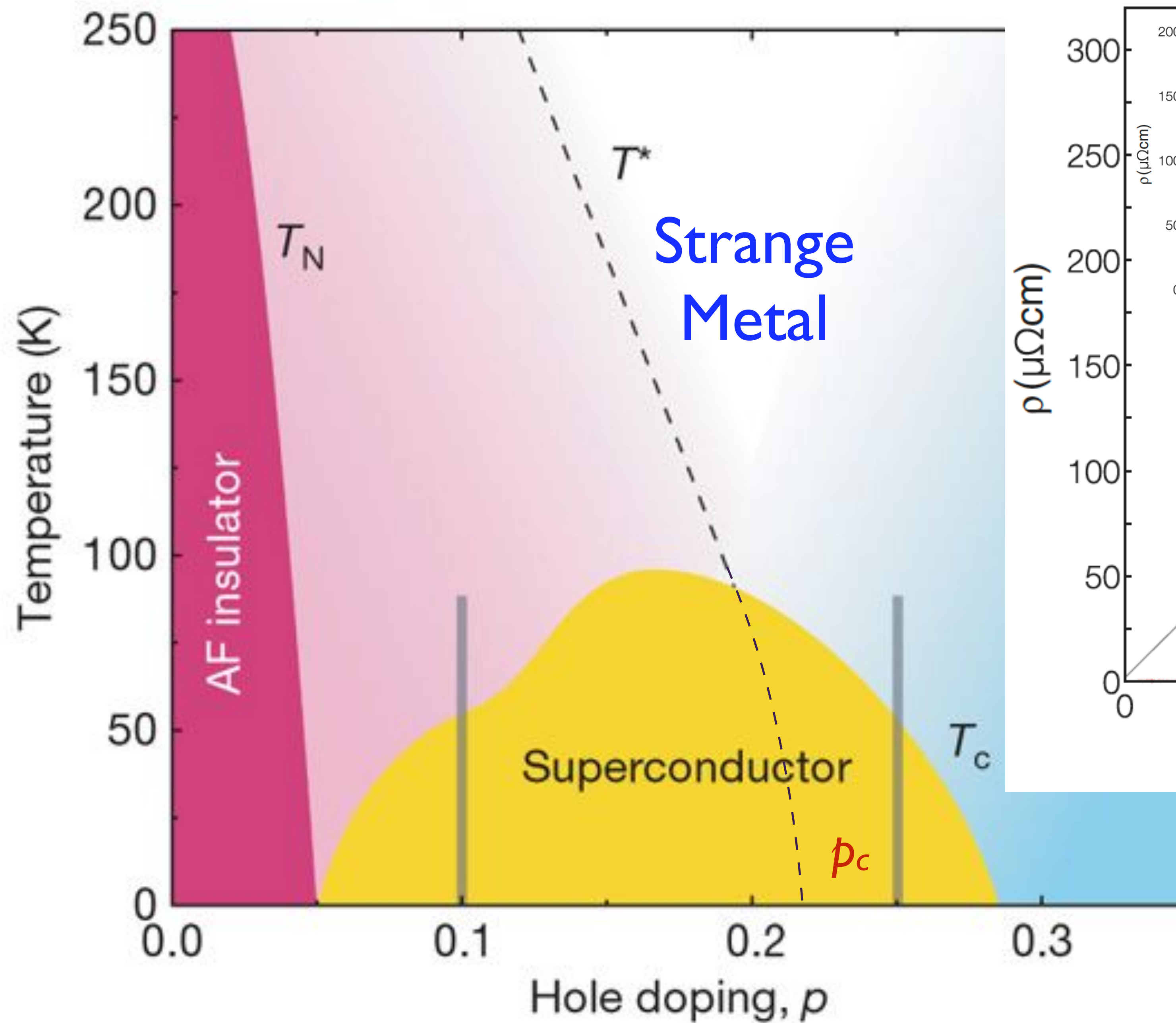
Commonwealth
Fusion Systems











LSCO: Giraldo-Gallo et al. 2018

Reconciling scaling of the optical conductivity of cuprate superconductors with Planckian resistivity and specific heat

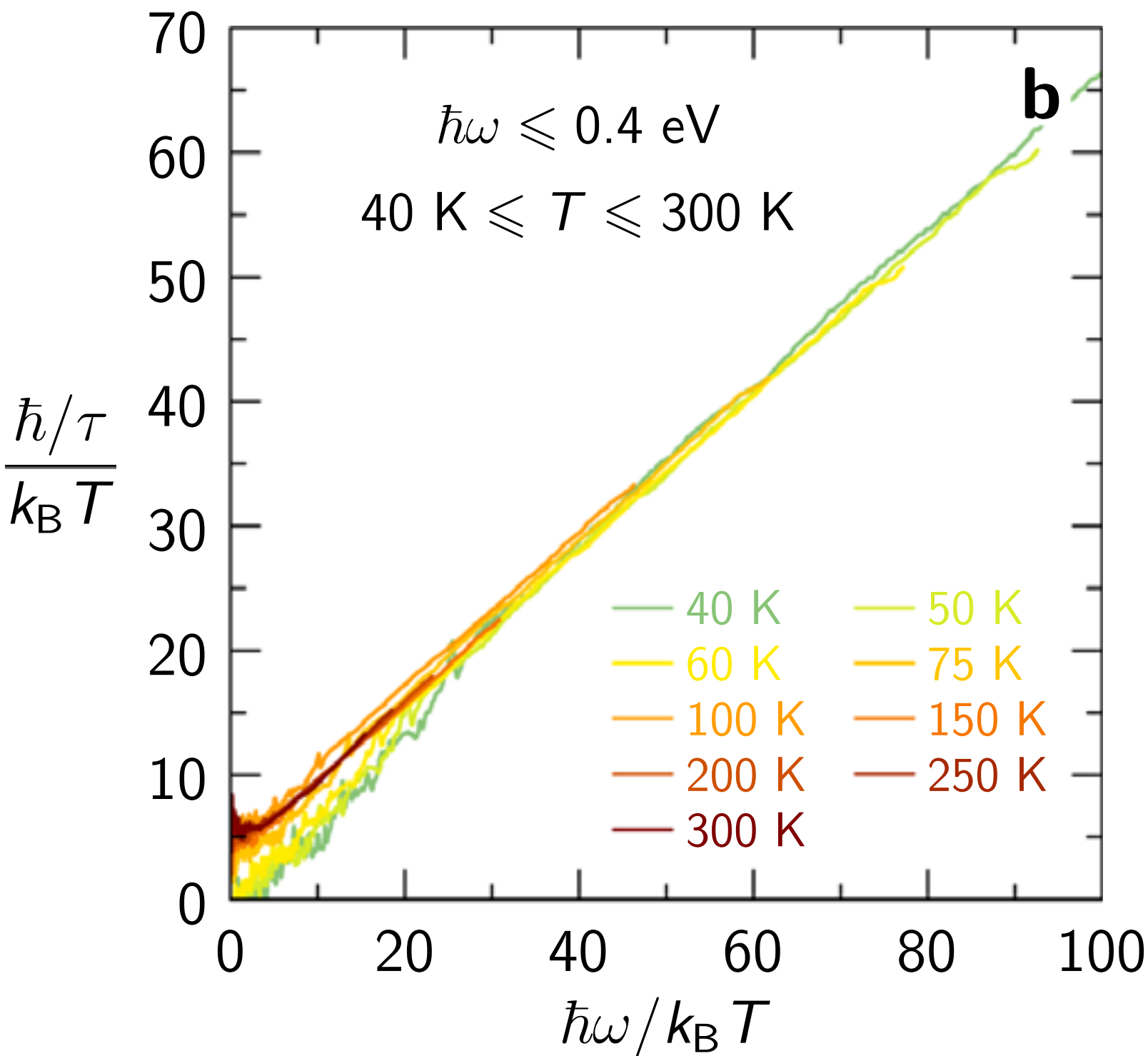
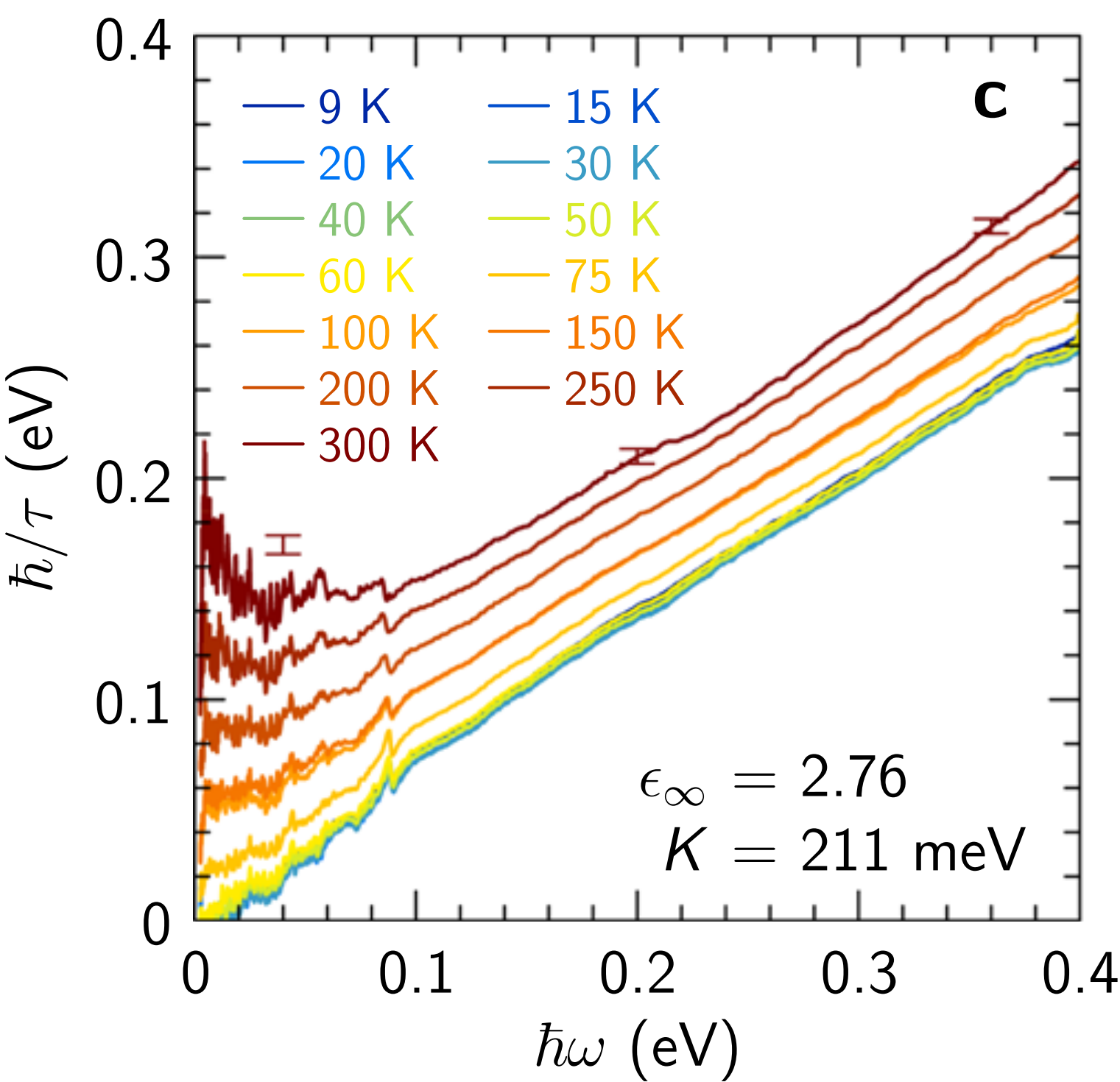
B. Michon, C. Berthod, C. W. Rischau, A. Ataei, L. Chen, S. Komiya, S. Ono, L. Taillefer, D. van der Marel, A. Georges

Nature Communications **14**, Article number: 3033 (2023)

$$\sigma(\omega) = i \frac{e^2 K / (\hbar d_c)}{\hbar \omega \frac{m^*(\omega)}{m} + i \frac{\hbar}{\tau(\omega)}}$$

Planckian dynamics!

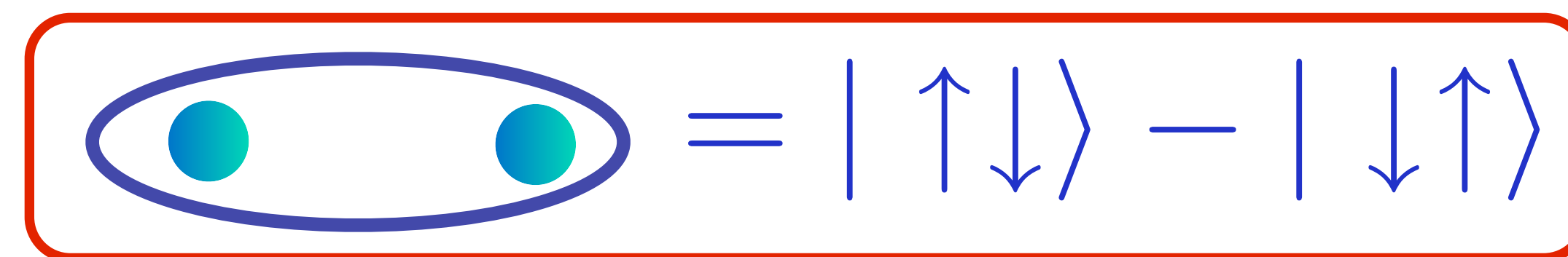
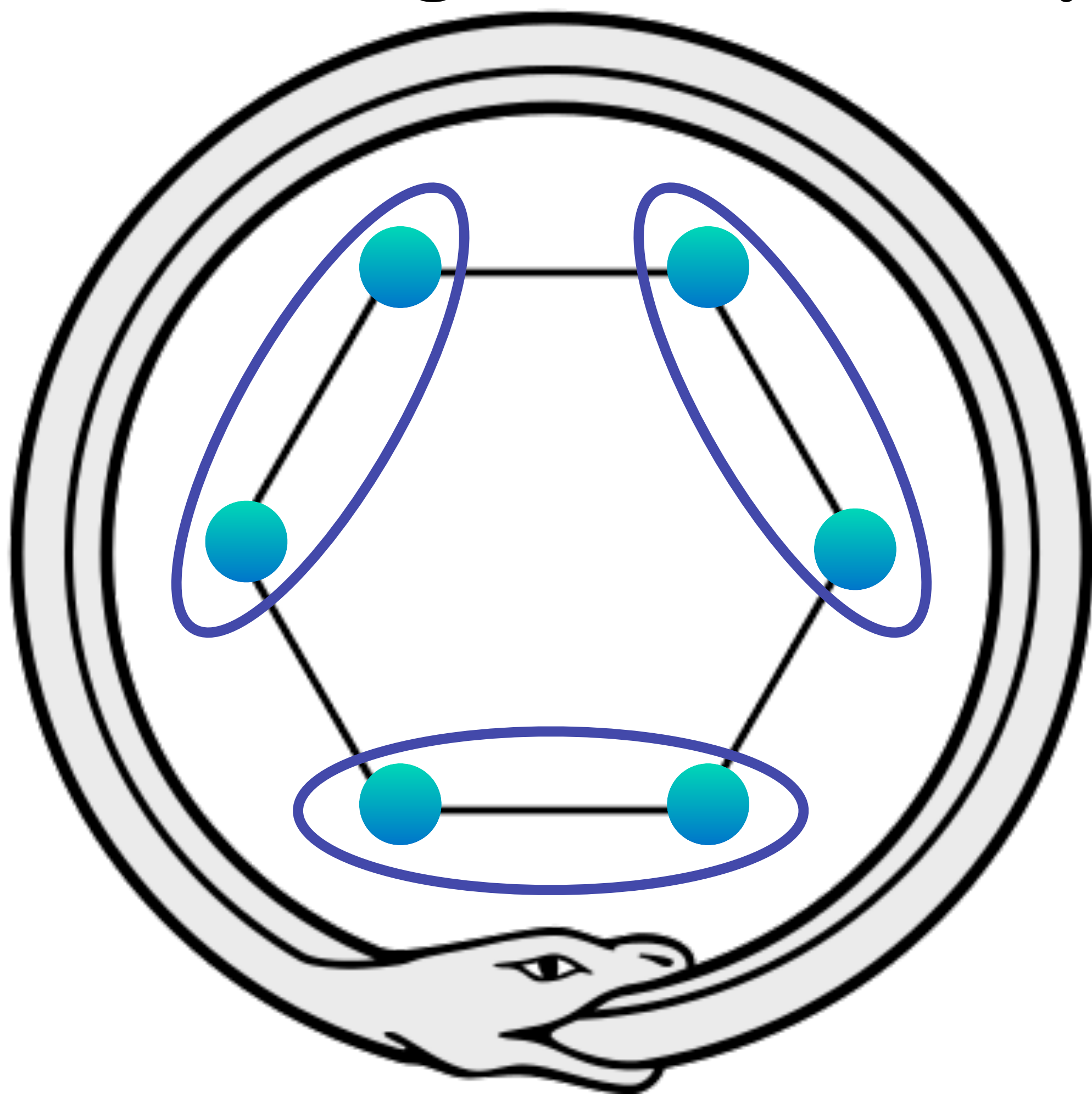
$$\tau(\omega) = \frac{\hbar}{k_B T} F\left(\frac{\hbar \omega}{k_B T}\right)$$



Quantum
entanglement
(1865)

Kekulé's spooky dream (1865)

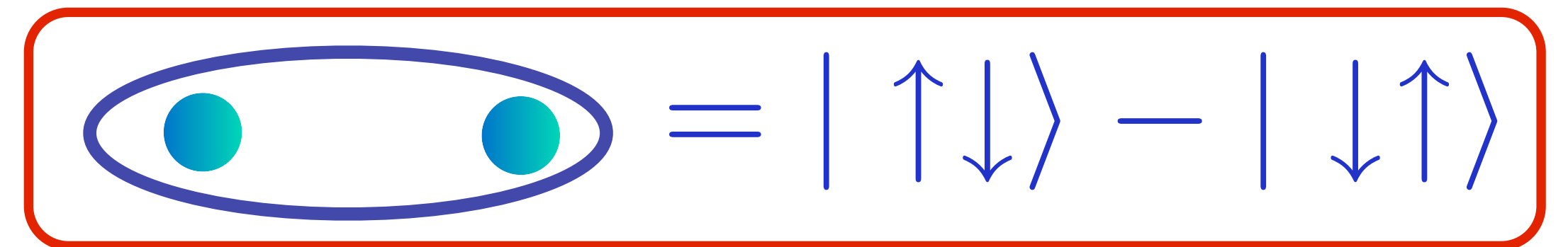
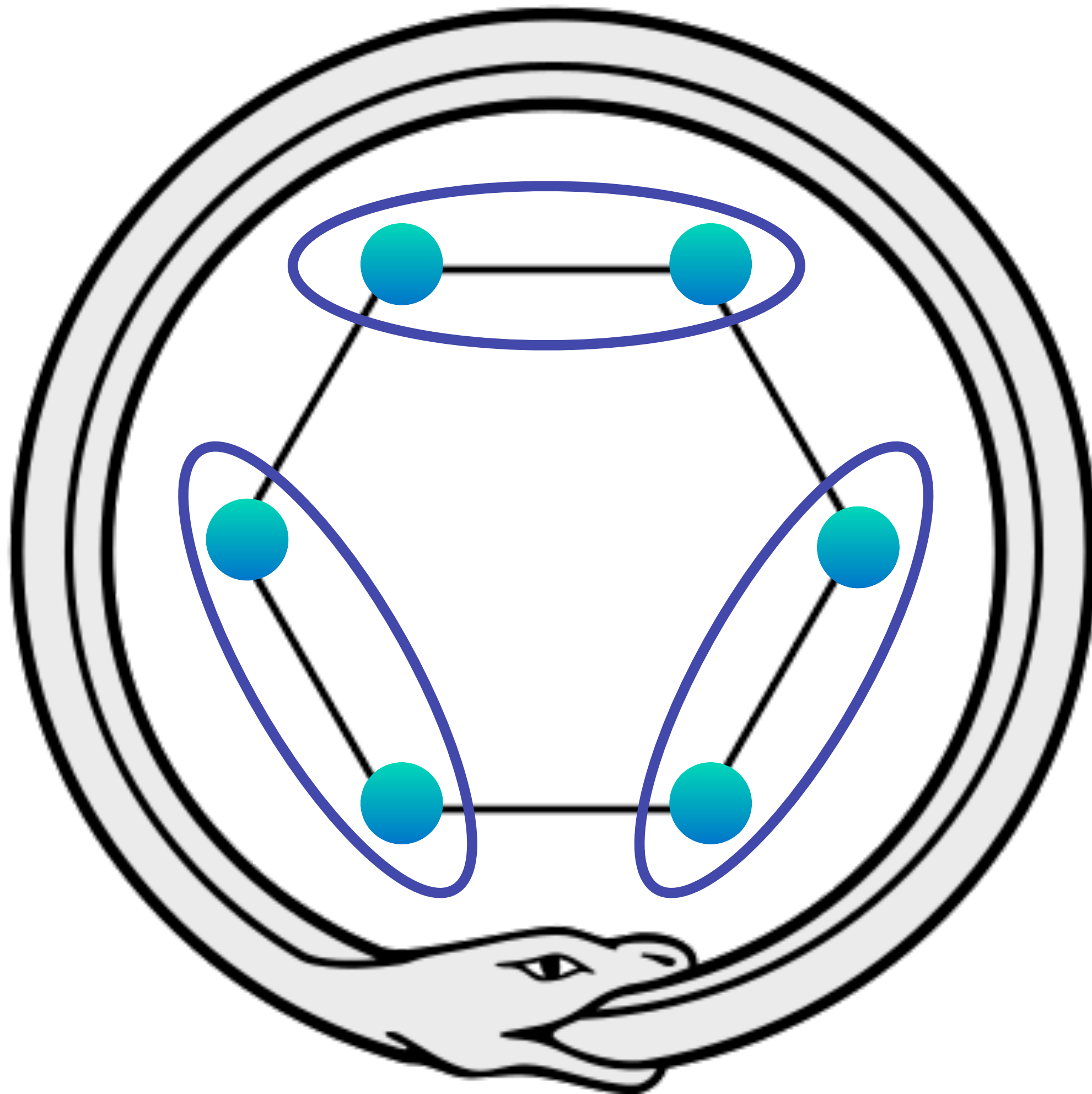
Kekulé spoke of the creation of the theory. He said that he had discovered the ring shape of the benzene molecule after having a reverie or day-dream of a snake seizing its own tail^{*}



Benzene

Kekulé's spooky dream (1865)

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Benzene

MAY 15, 1935

PHYSICAL REVIEW

VOLUME 47

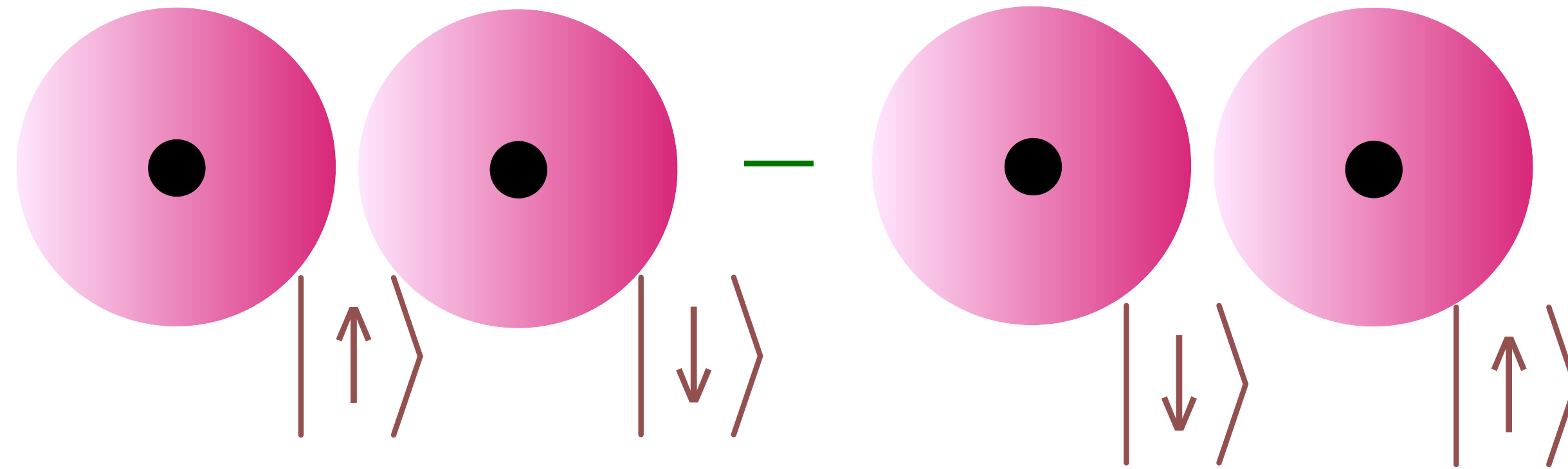
Can Quantum-Mechanical Description of Physical Reality Be Considered Complete?

A. EINSTEIN, B. PODOLSKY AND N. ROSEN, *Institute for Advanced Study, Princeton, New Jersey*

(Received March 25, 1935)

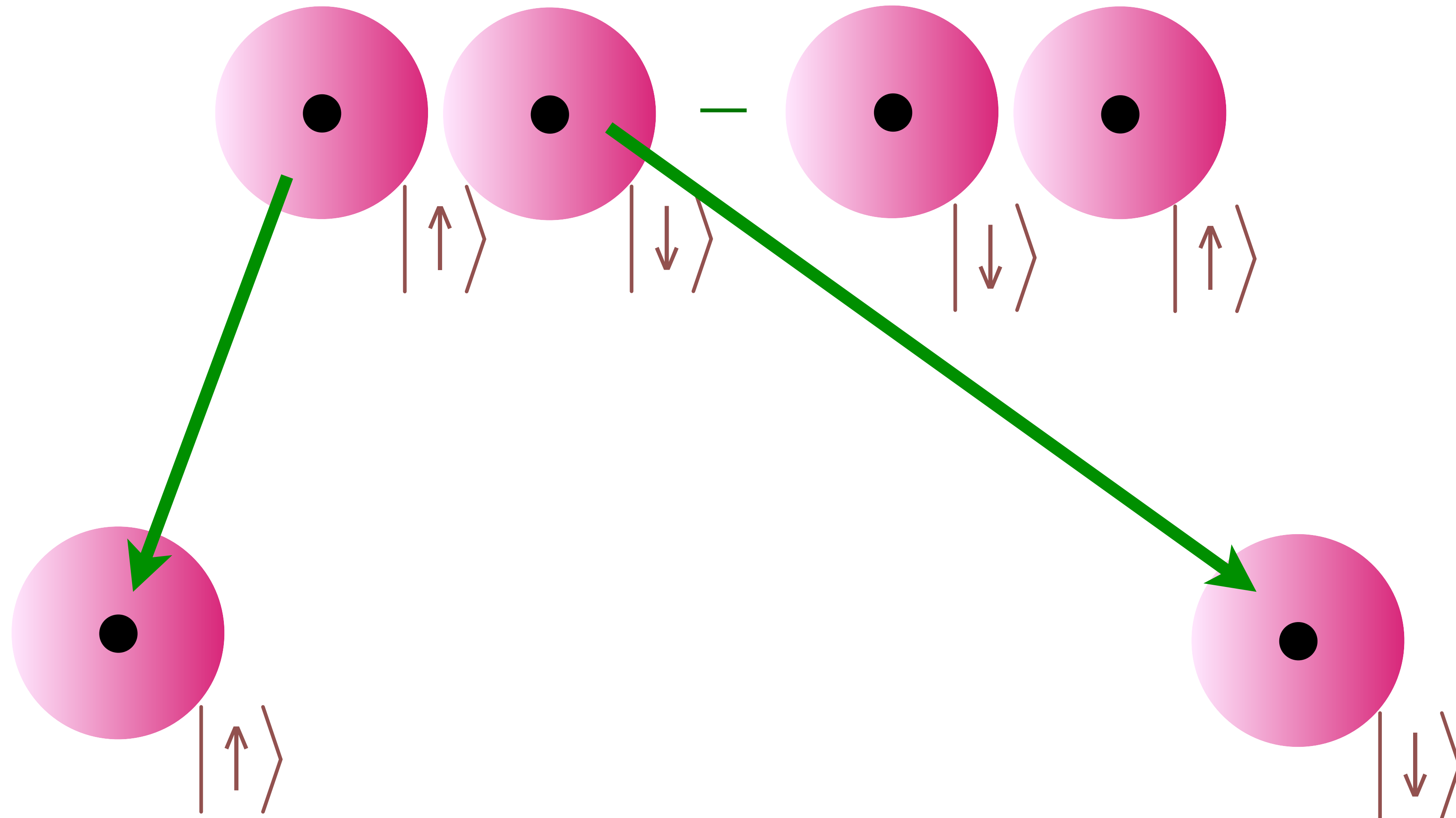
Quantum Entanglement

Einstein, Podolsky, Rosen (1935)



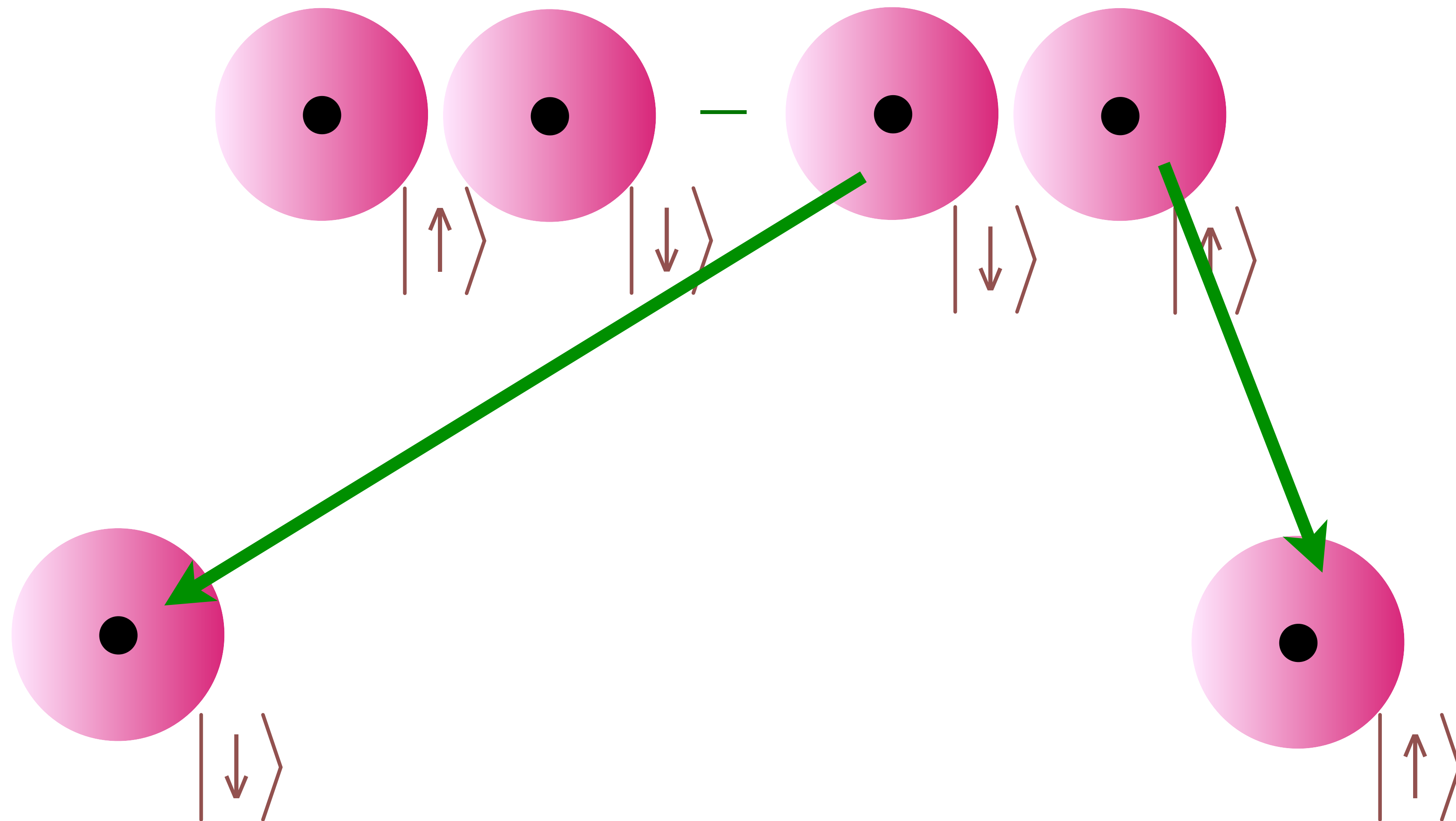
Quantum Entanglement

Einstein, Podolsky, Rosen (1935)



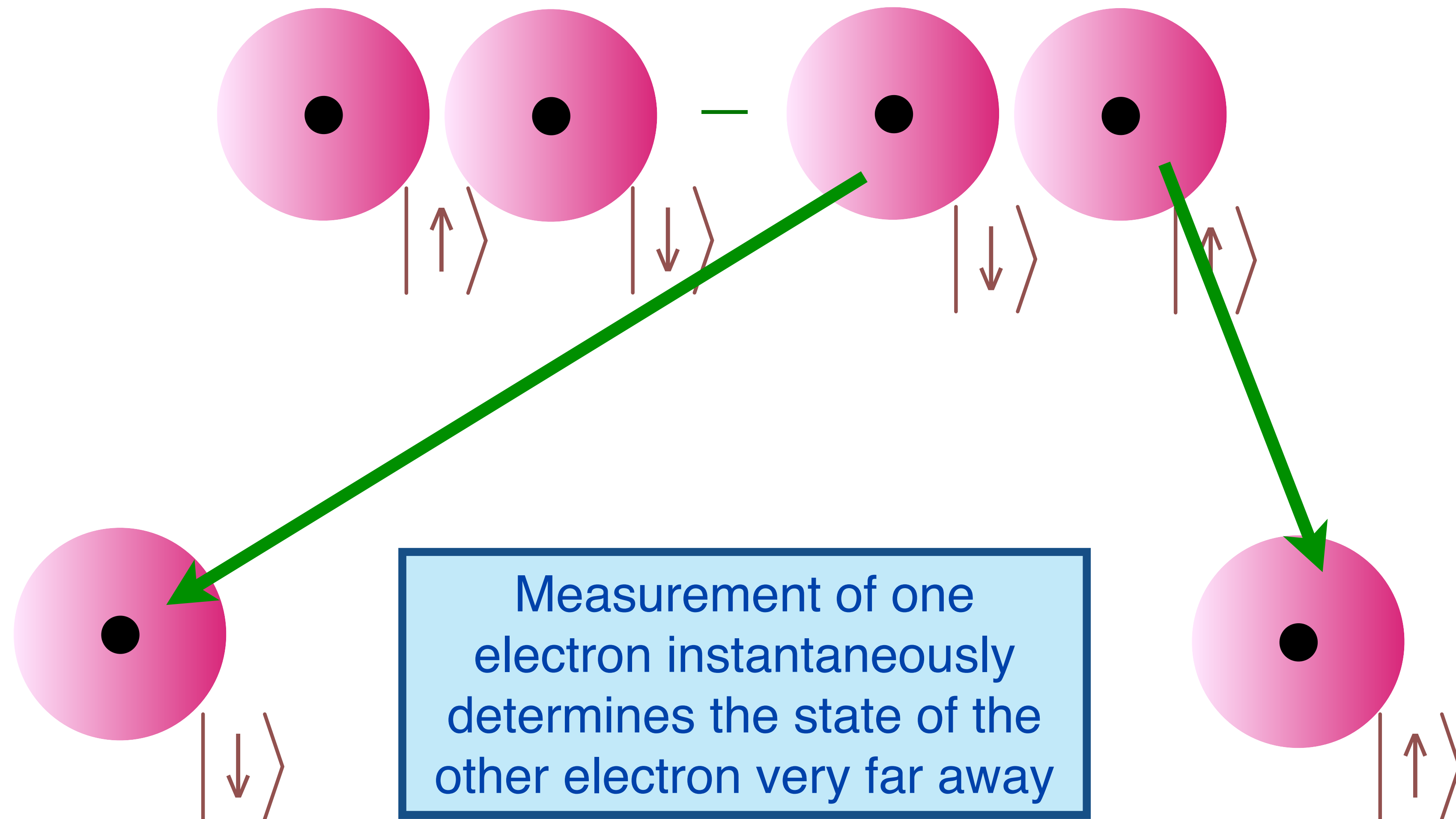
Quantum Entanglement

Einstein, Podolsky, Rosen (1935)



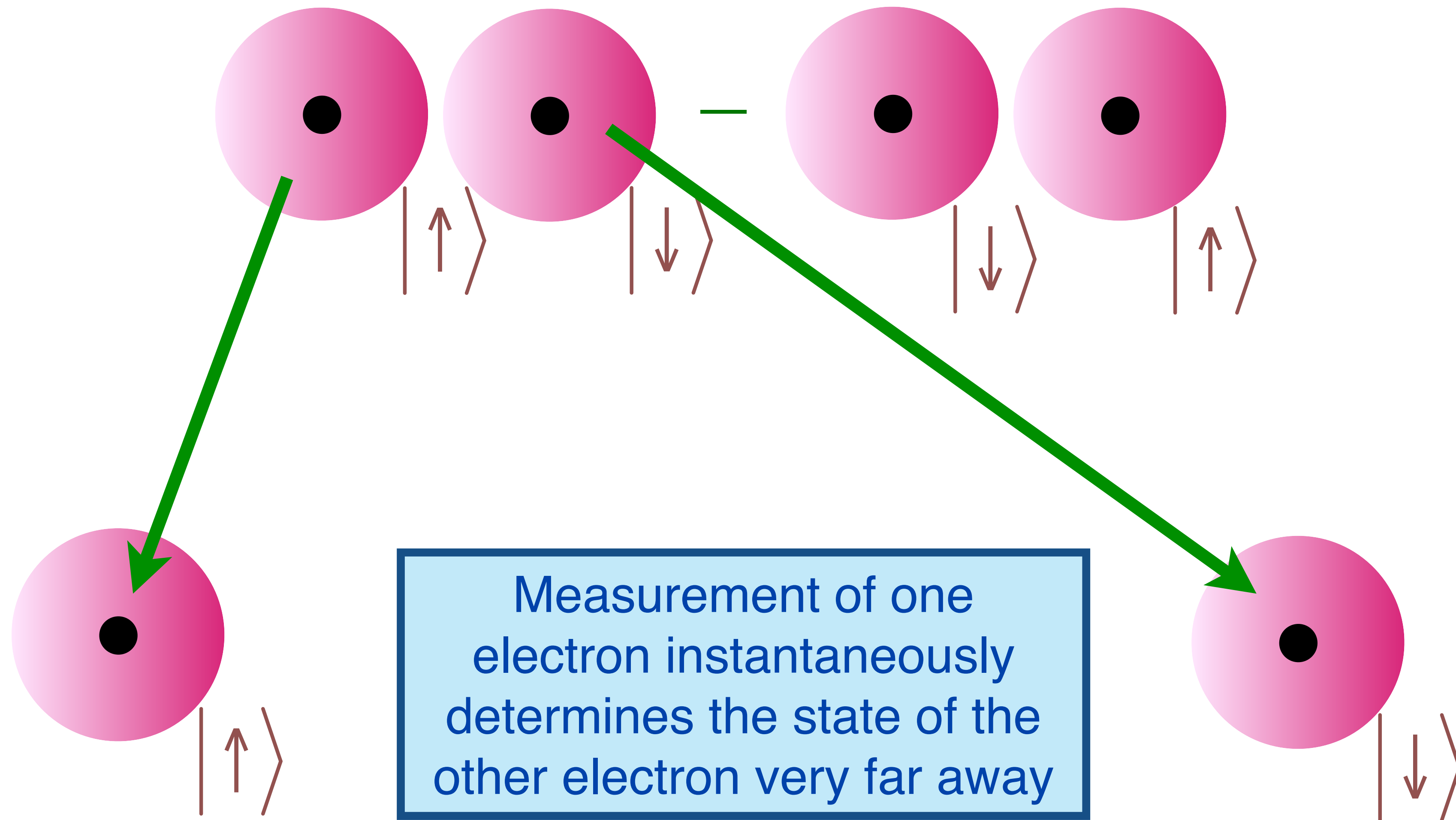
Quantum Entanglement

Einstein, Podolsky, Rosen (1935)



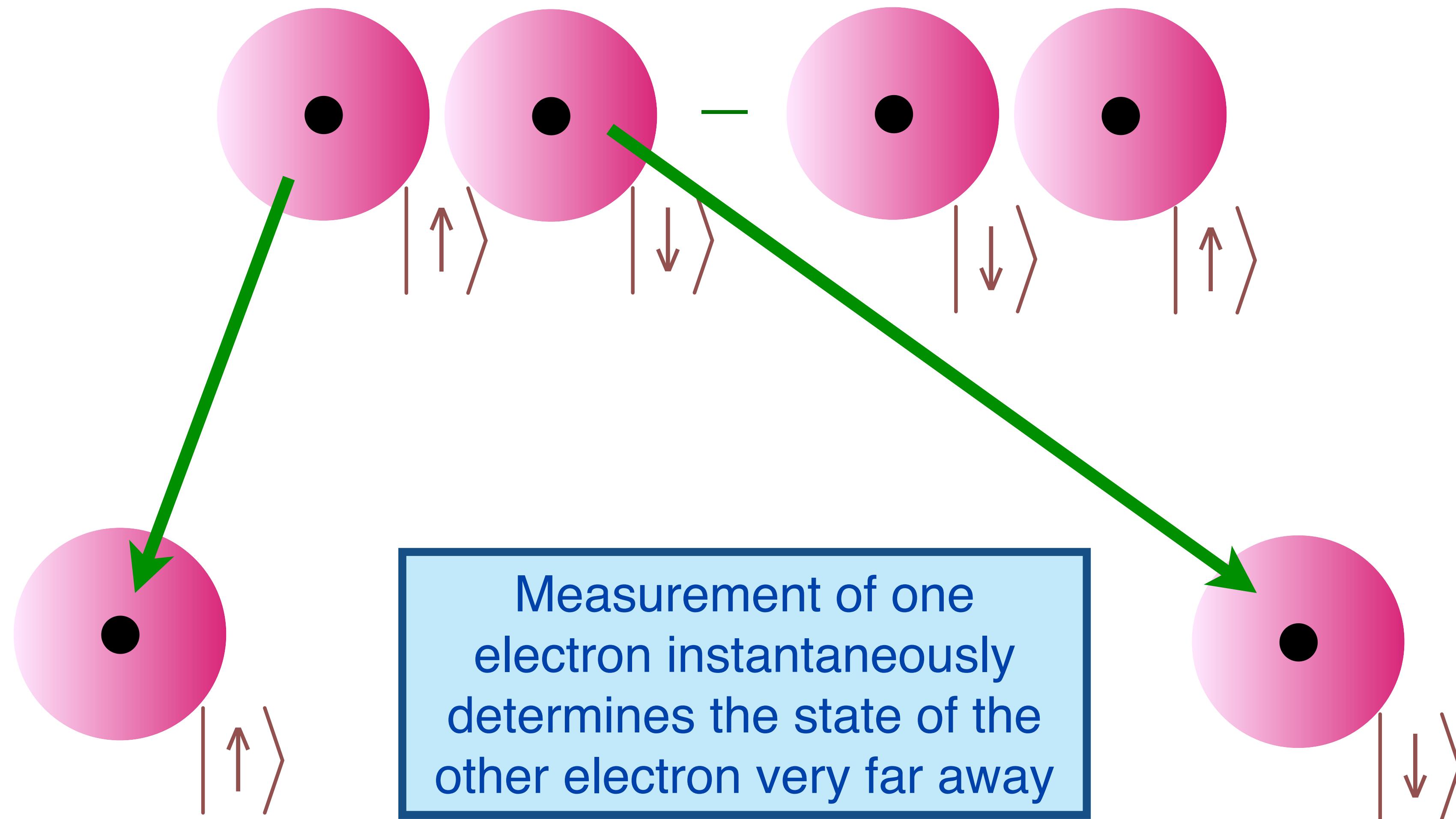
Quantum Entanglement

Einstein, Podolsky, Rosen (1935)



Quantum Entanglement

Einstein, Podolsky, Rosen (1935)



Spooky action at a distance !

natürlicher
deren Notwendigkeit im Raum
mus ja zuerst von Dir klar erkannt wurde, einen Bedeutung
Wahrheitsgehalt hat. Ich kann aber deshalb nicht ernsthaft dar-
an glauben, weil die Theorie mit dem Grundsatz unvereinbar
ist, daß die Physik eine Wirklichkeit in Zeit und Raum darstel-
len soll, ohne spukhafte Fernwirkungen. Allerdings bin ich
überzeugt, daß es wirklich mit der Theorie

I cannot seriously believe in it because the
theory cannot be reconciled with the idea that
physics should represent a reality in time and
space, free from spooky actions at distance

Albert Einstein to Max Born, 3 March 1947

Needed,
to solve open problems in the theory of
superconductivity and black holes:

A solvable model of quantum entanglement
of $3, 4, 5, \dots \infty$ particles

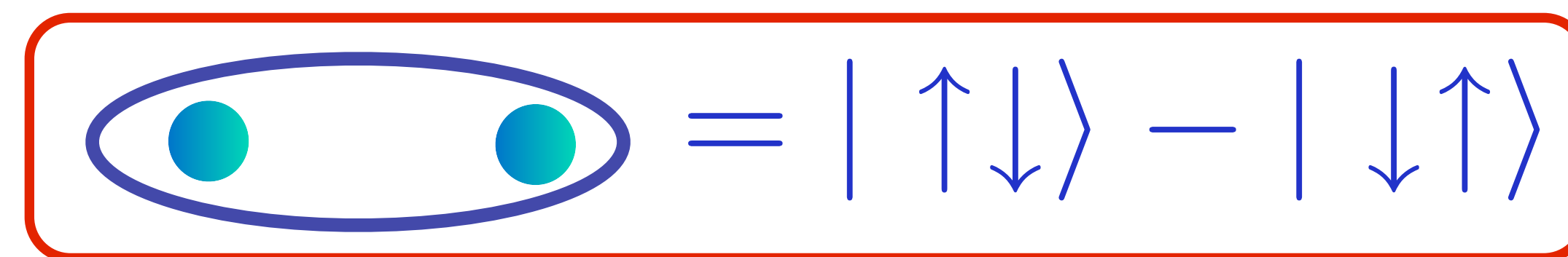
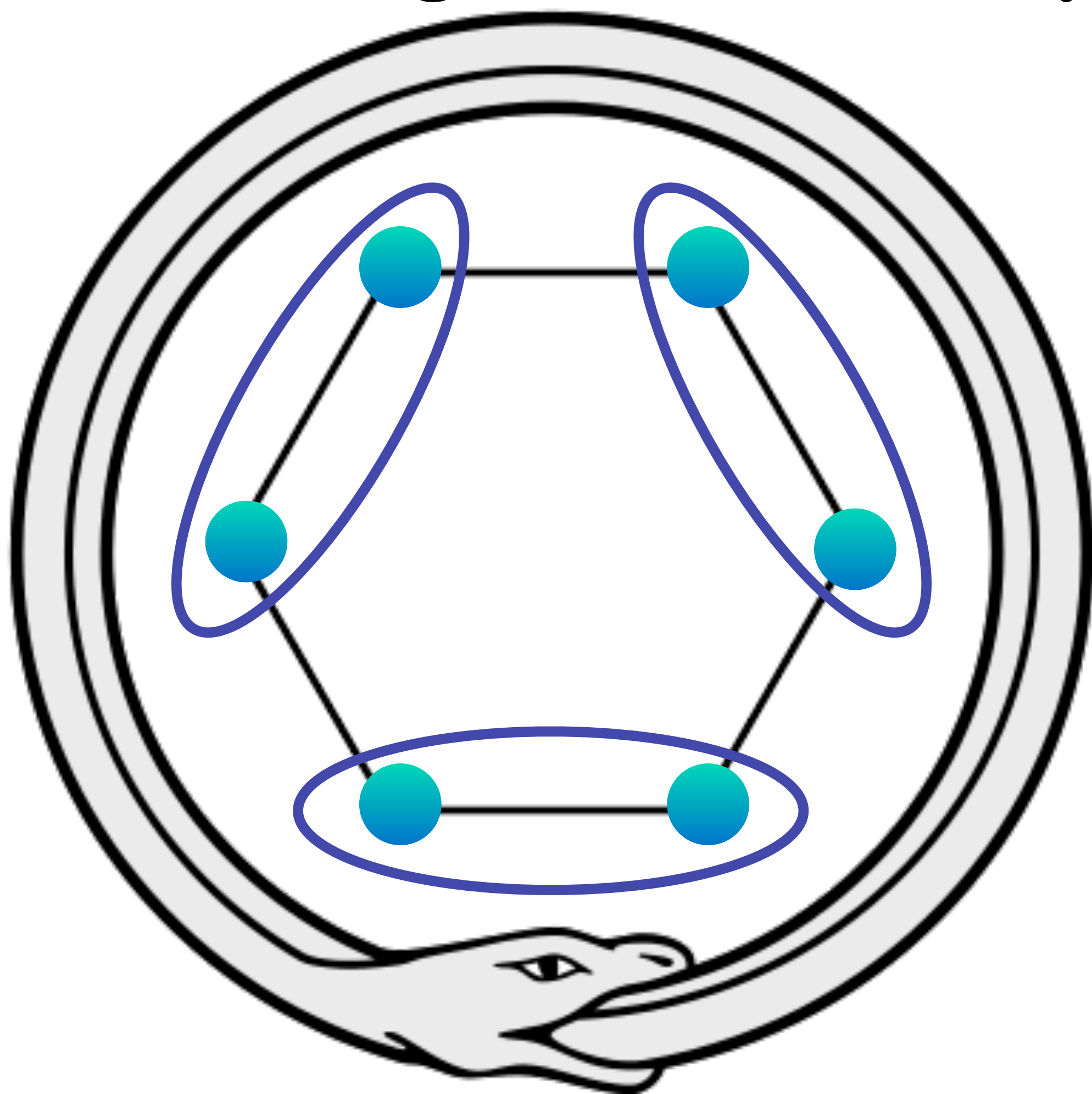
Needed,
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A solvable model of quantum entanglement
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**The Sachdev-Ye-Kitaev model
of many-particle entanglement**

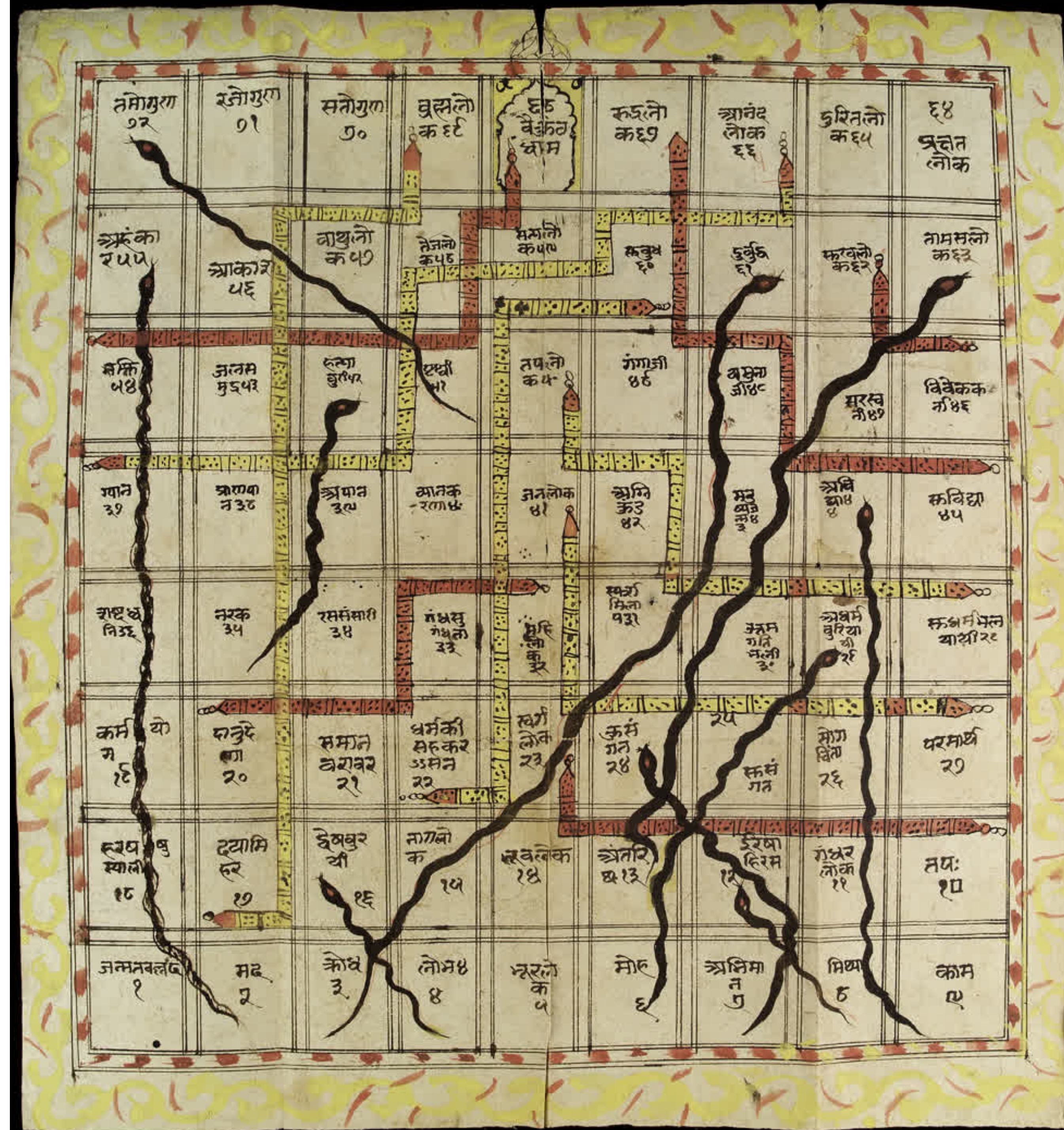
Kekulé's spooky dream (1865)

Kekulé spoke of the creation of the theory. He said that he had discovered the ring shape of the benzene molecule after having a reverie or day-dream of a snake seizing its own tail^{*}



Benzene

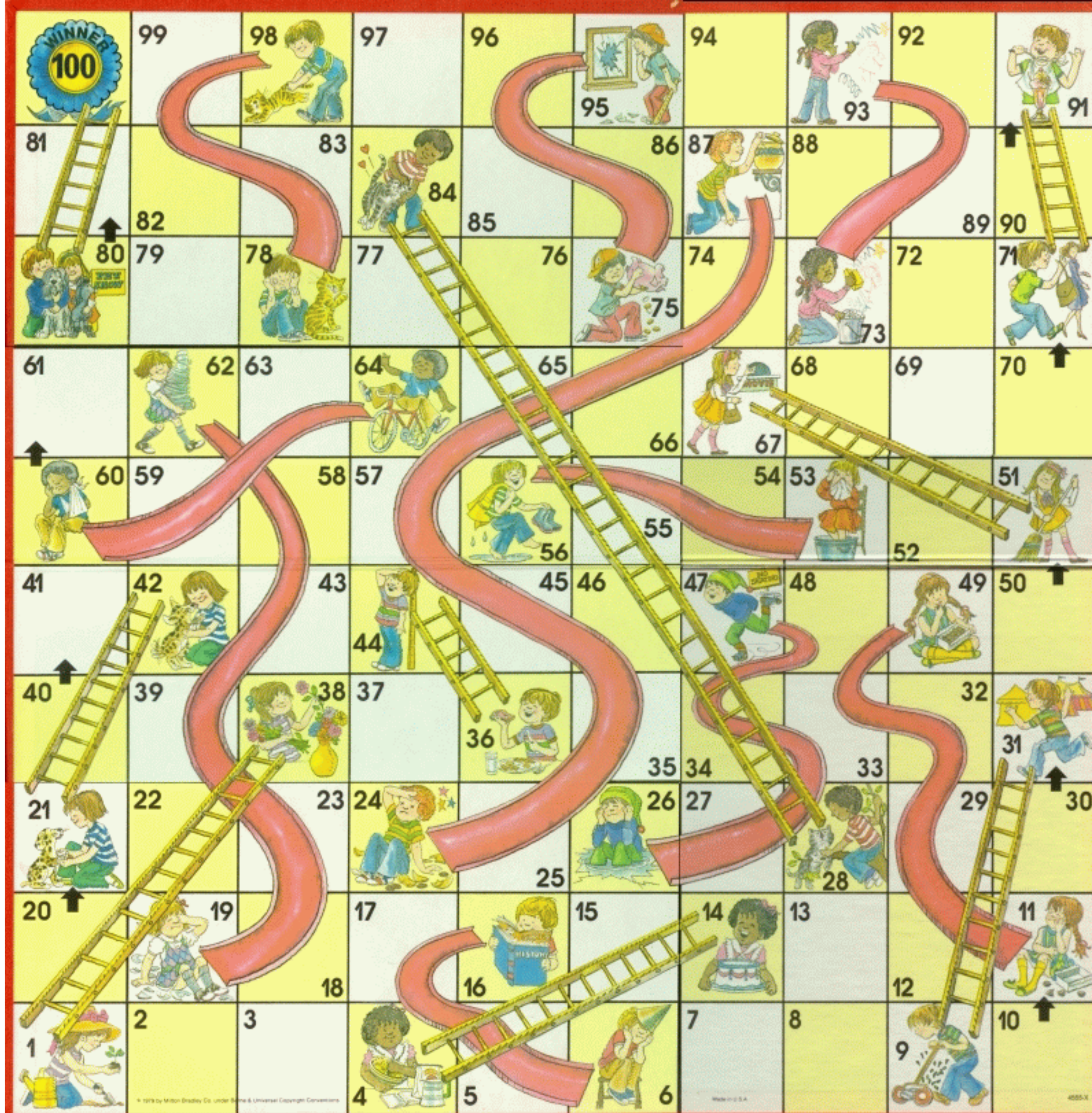
*Not true



My
spooky
dream
(1992)*

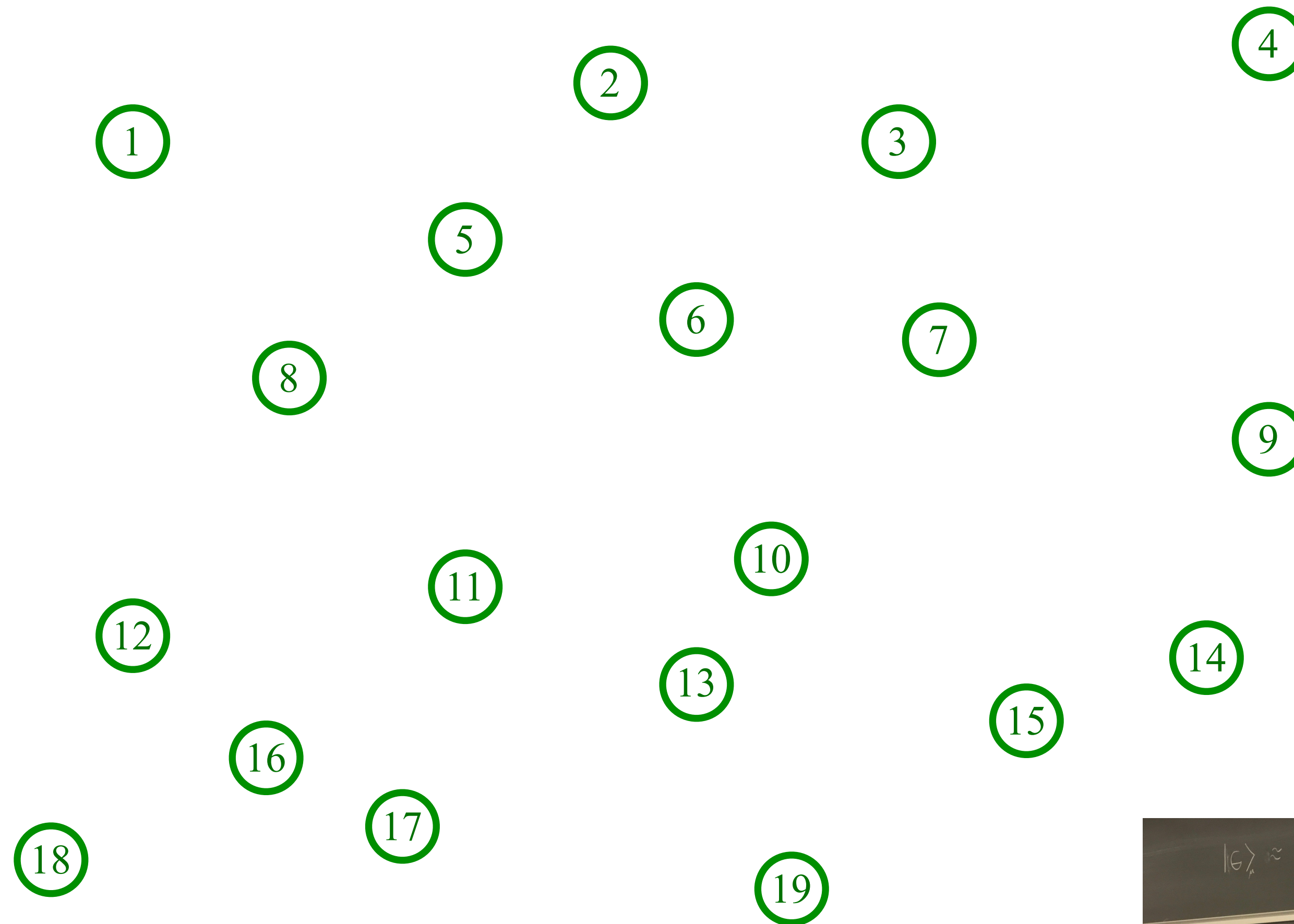
Hasbro
game of
Chutes
and
Ladders

*Not true

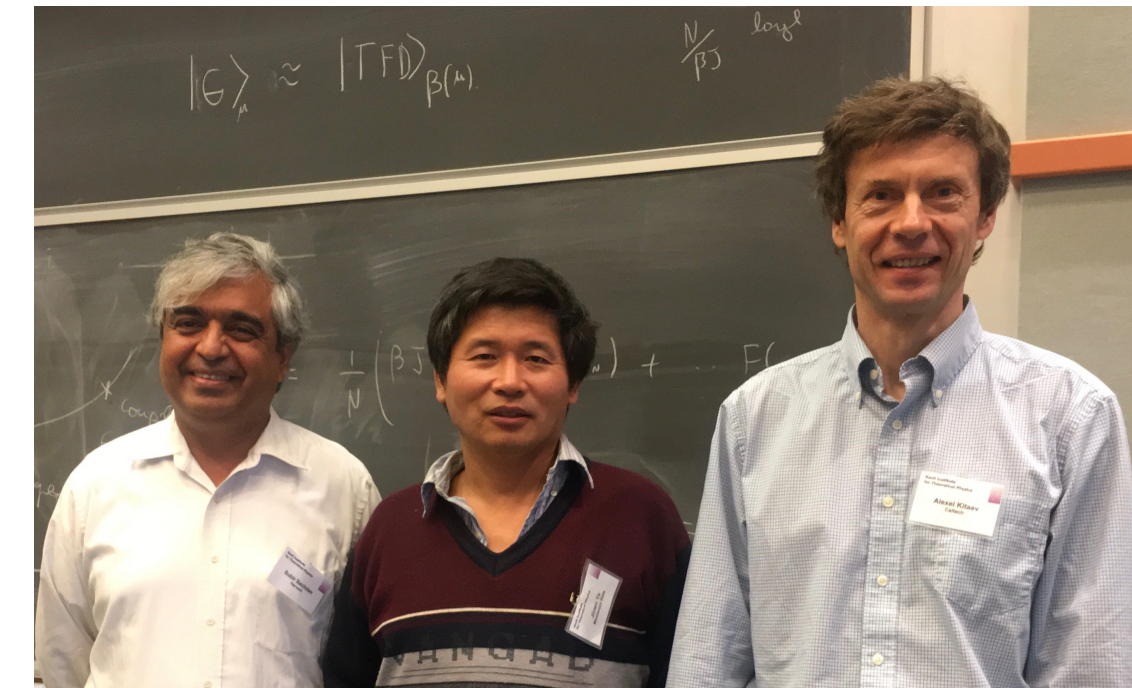


The Sachdev-Ye-Kitaev (SYK) model

Sachdev, Ye (1993); Kitaev (2015)

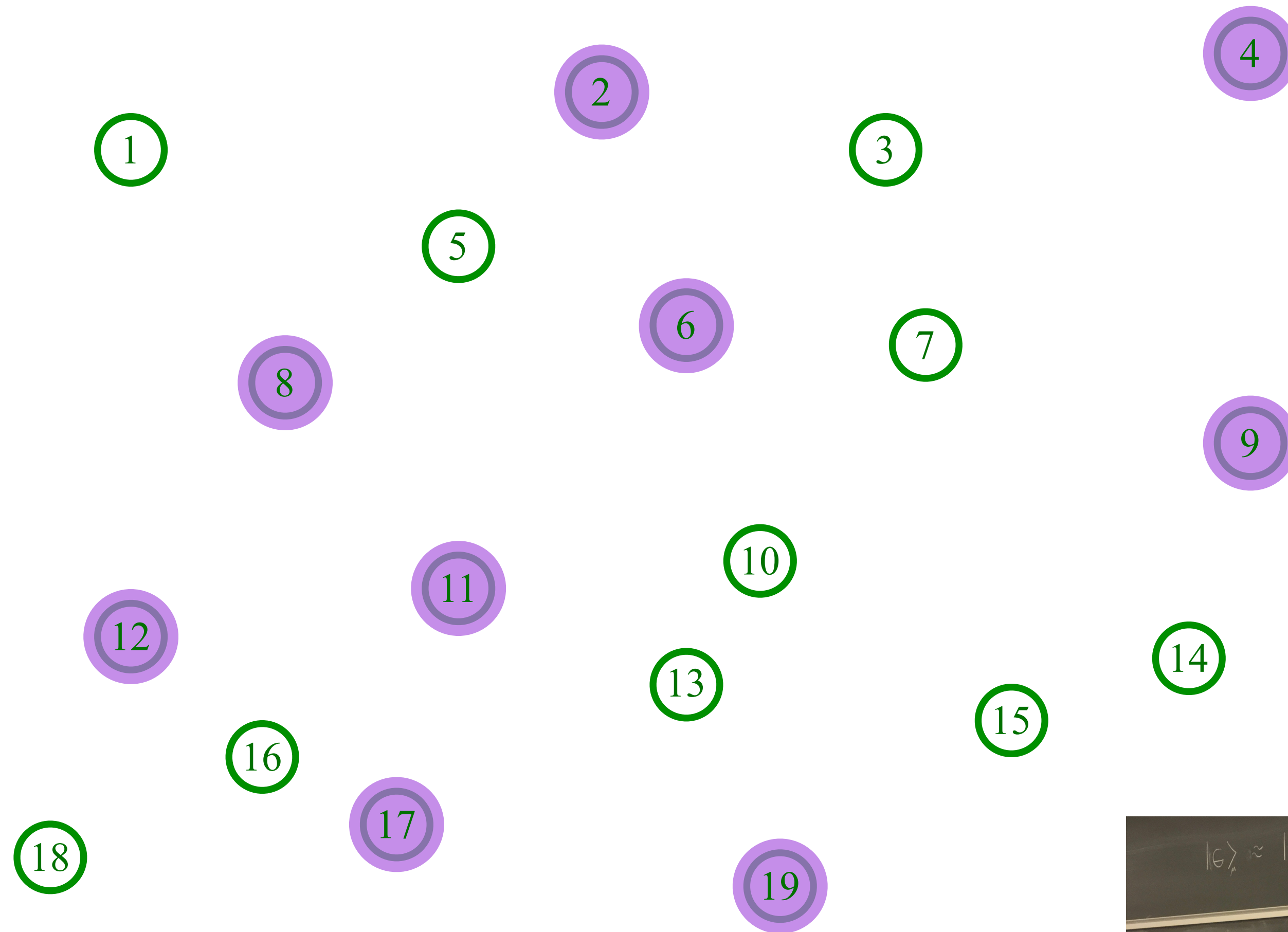


Pick a set of random positions

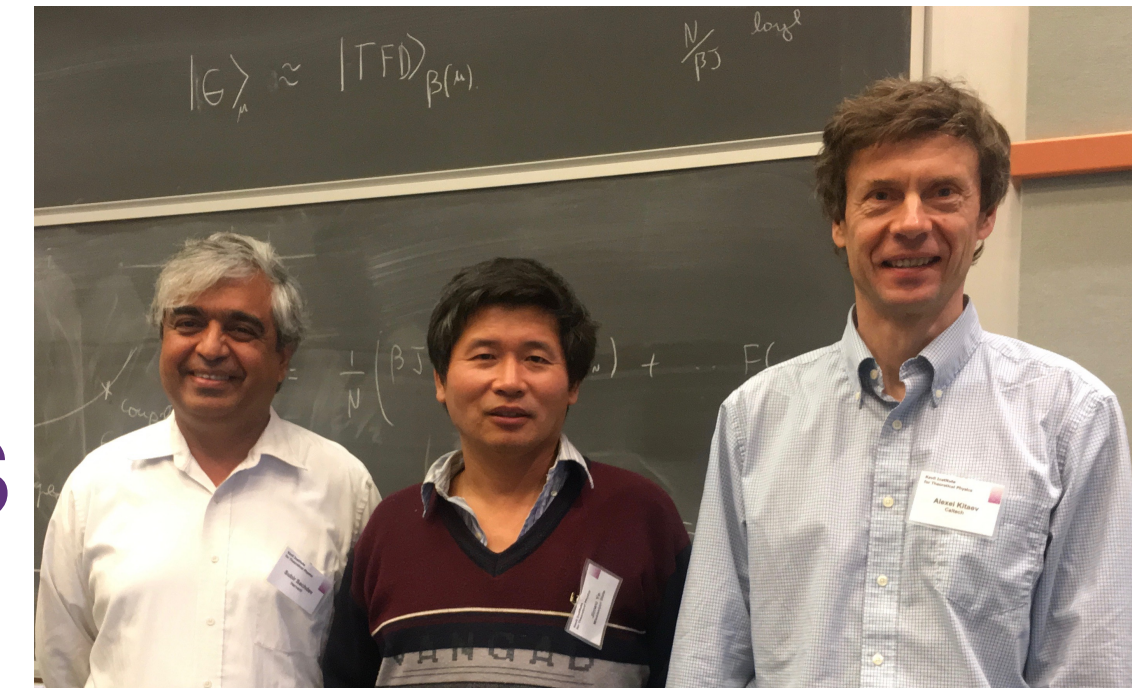


The Sachdev-Ye-Kitaev (SYK) model

Sachdev, Ye (1993); Kitaev (2015)

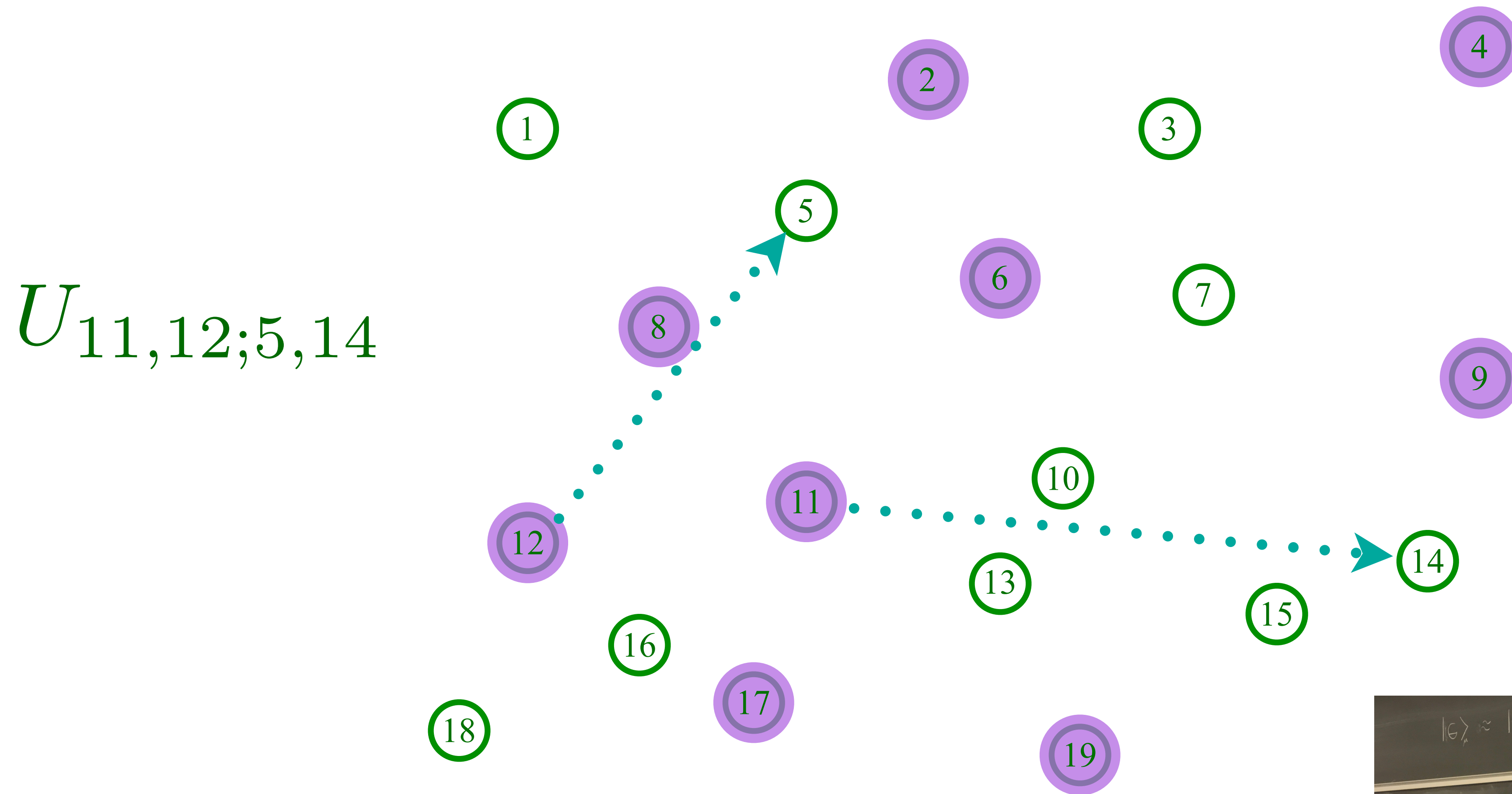


Place electrons randomly on some sites

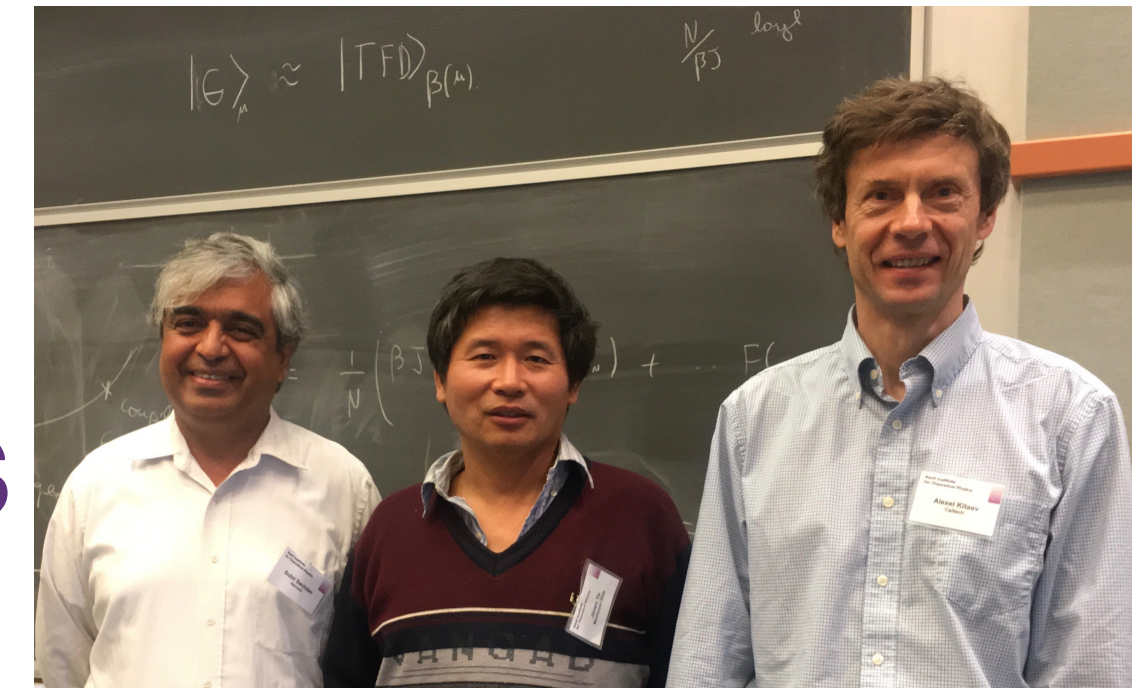


The Sachdev-Ye-Kitaev (SYK) model

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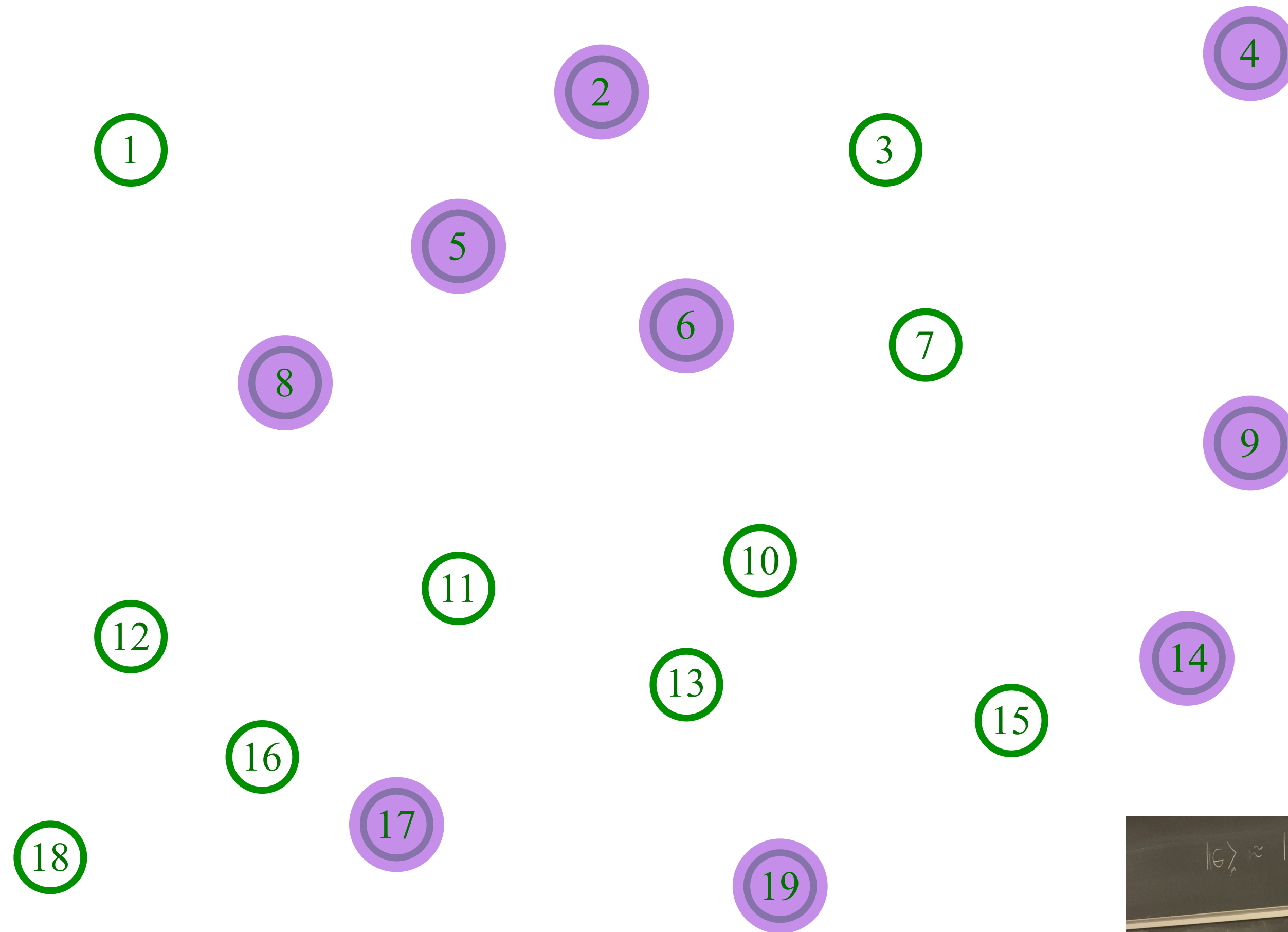
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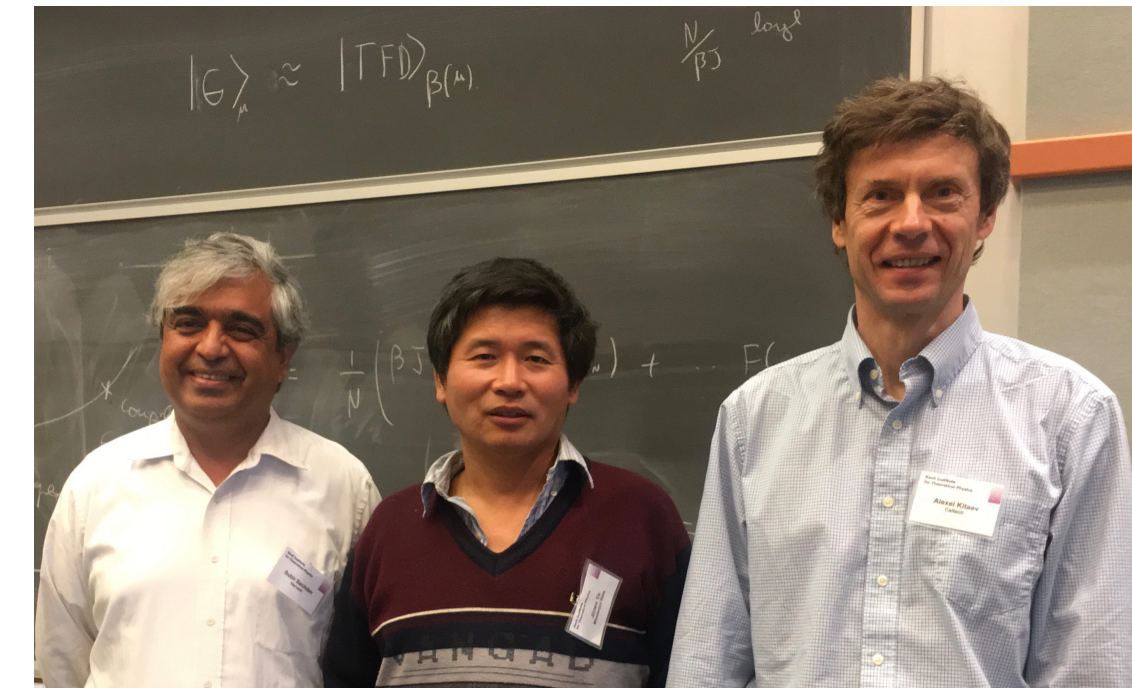
The Sachdev-Ye-Kitaev (SYK) model

Sachdev, Ye (1993); Kitaev (2015)

$$U_{11,12;5,14}$$

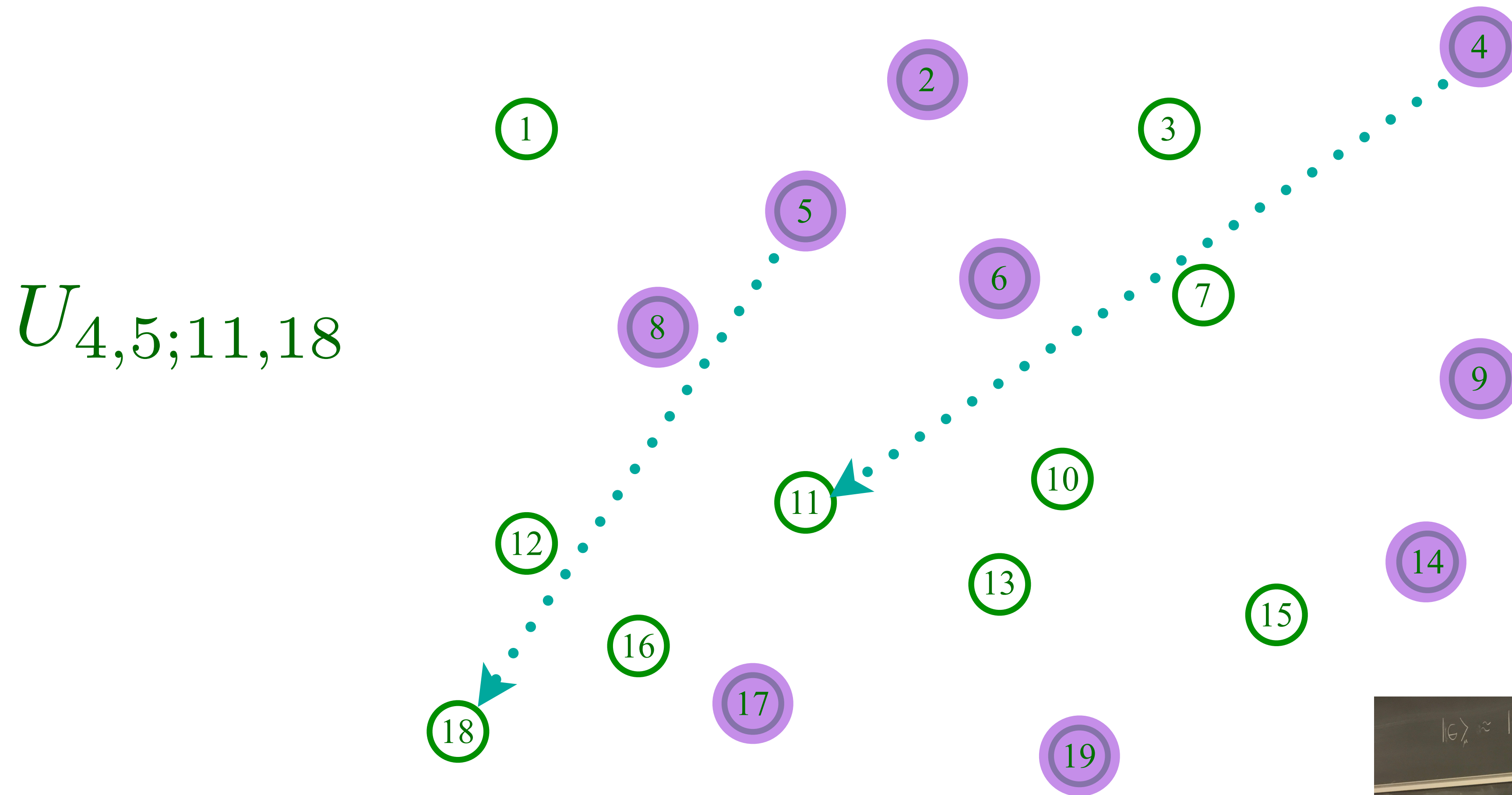


Entangle electrons pairwise randomly

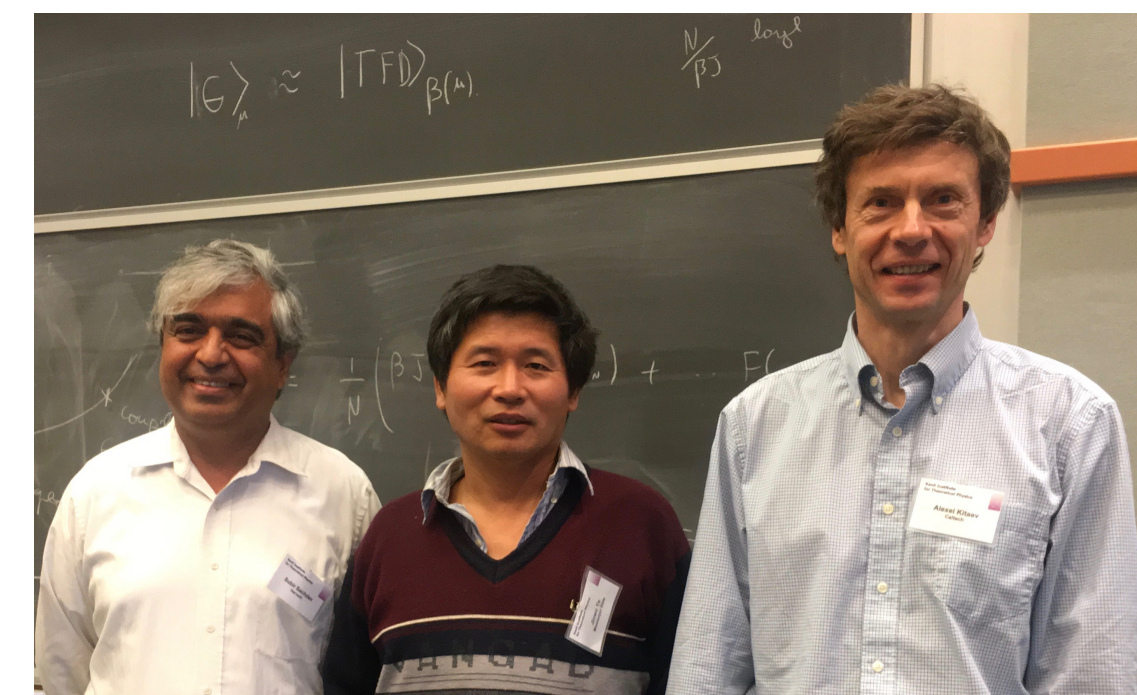


The Sachdev-Ye-Kitaev (SYK) model

Sachdev, Ye (1993); Kitaev (2015)



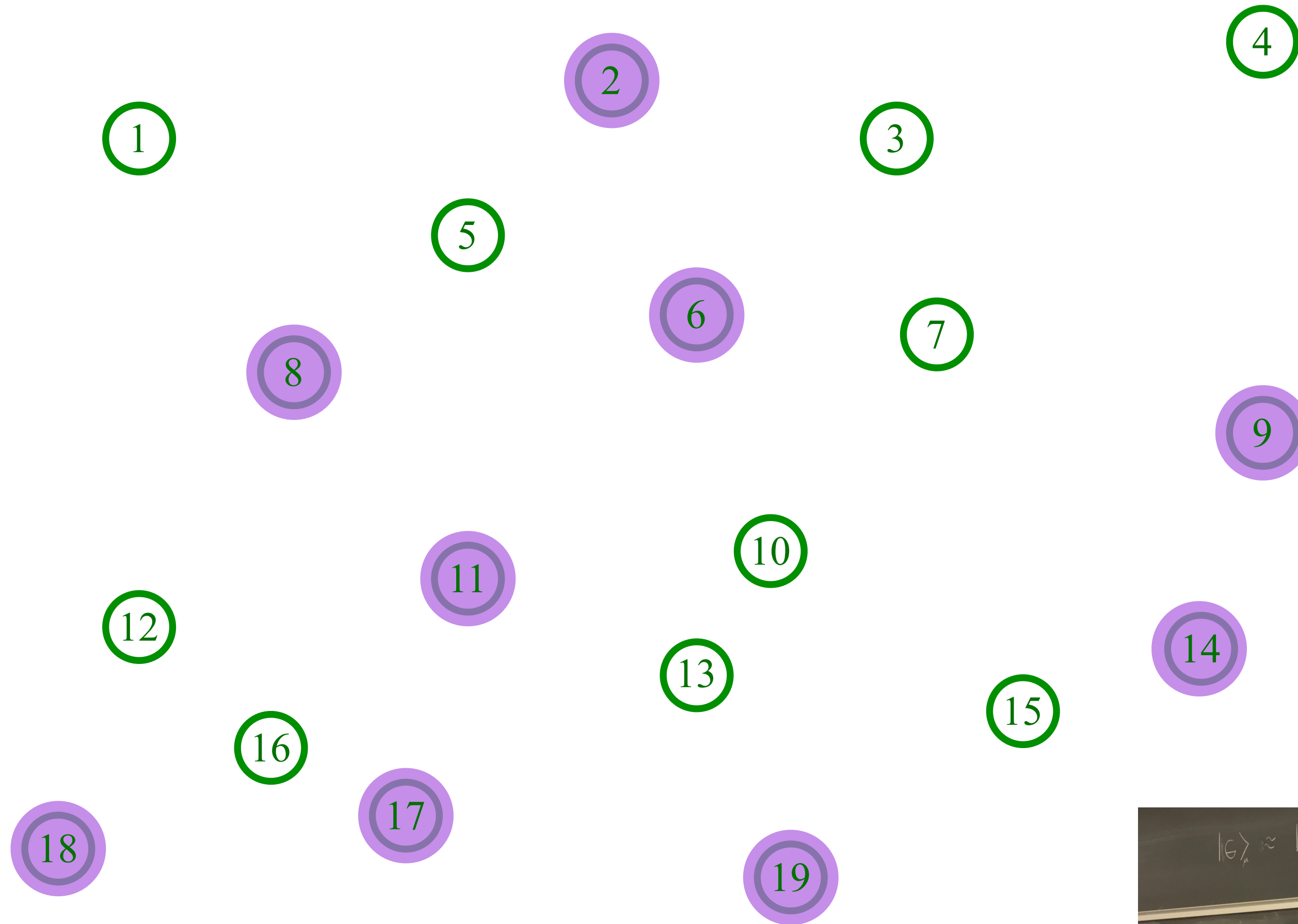
Entangle electrons pairwise randomly



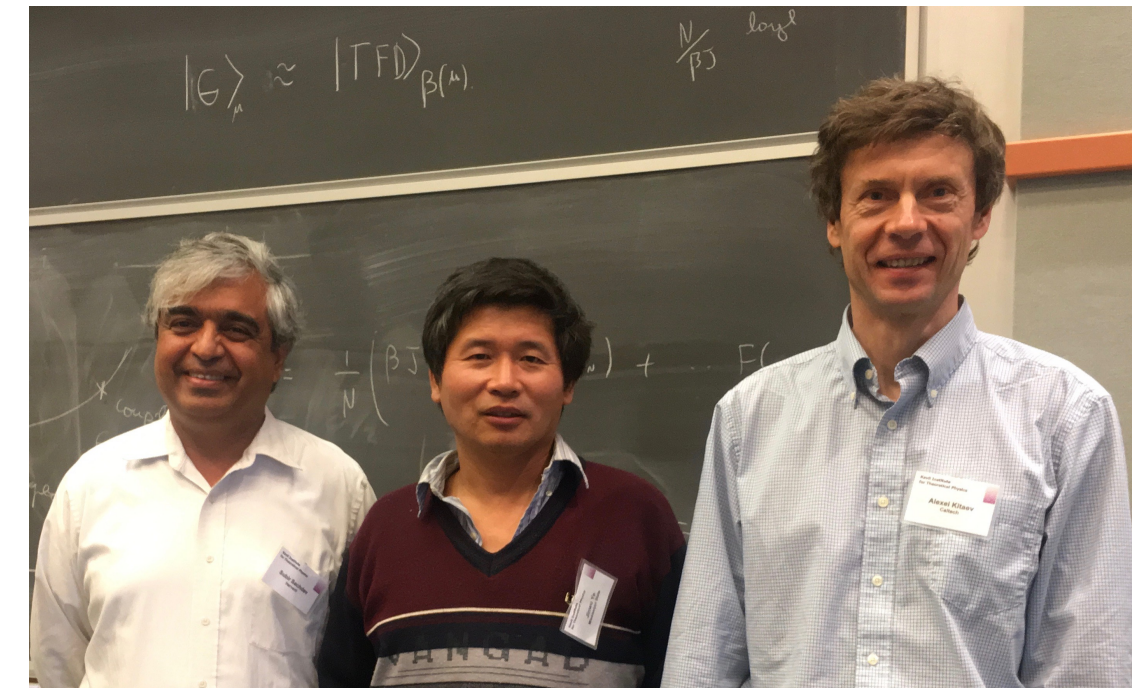
The Sachdev-Ye-Kitaev (SYK) model

Sachdev, Ye (1993); Kitaev (2015)

$$U_{4,5;11,18}$$

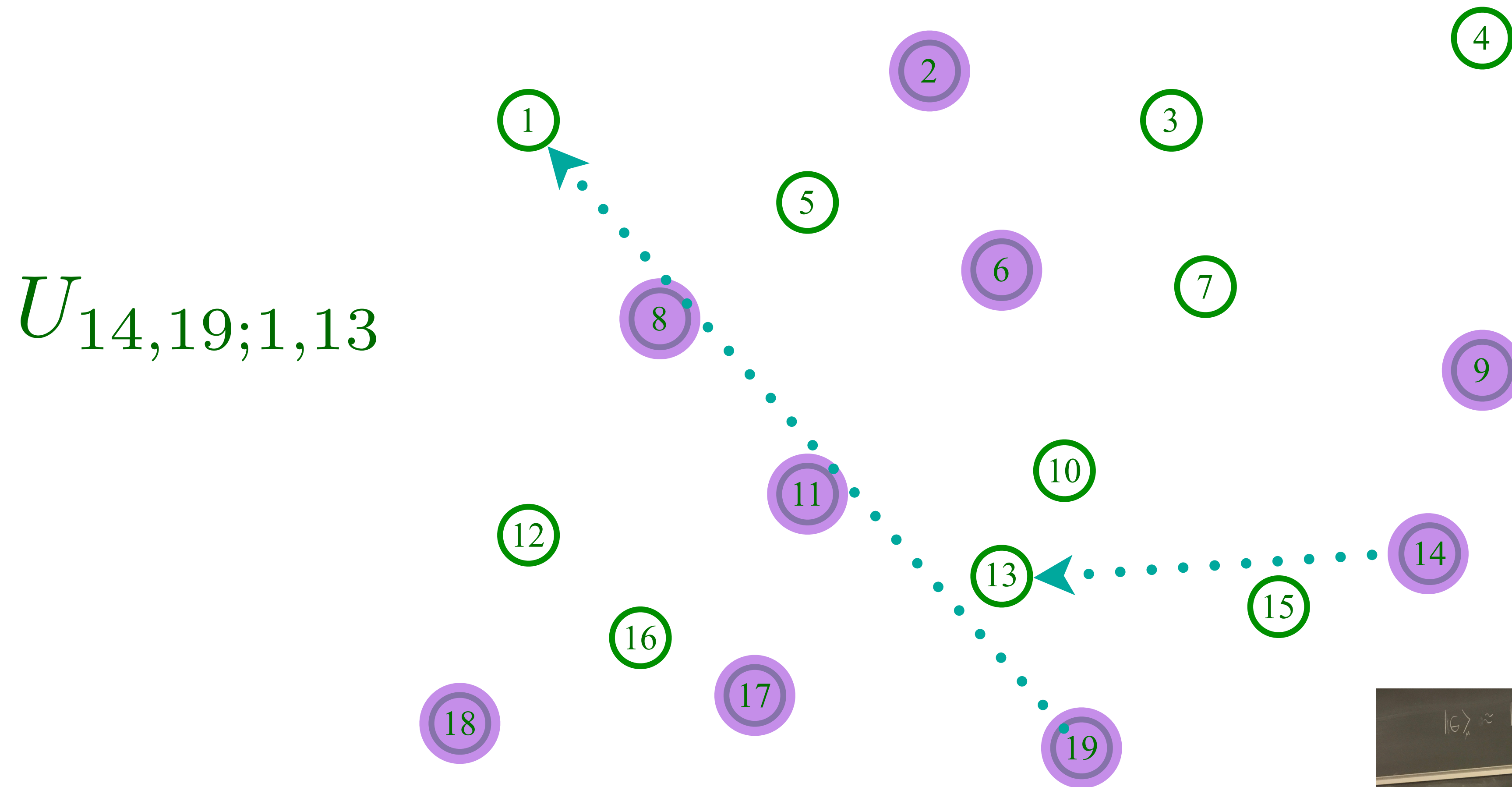


Entangle electrons pairwise randomly

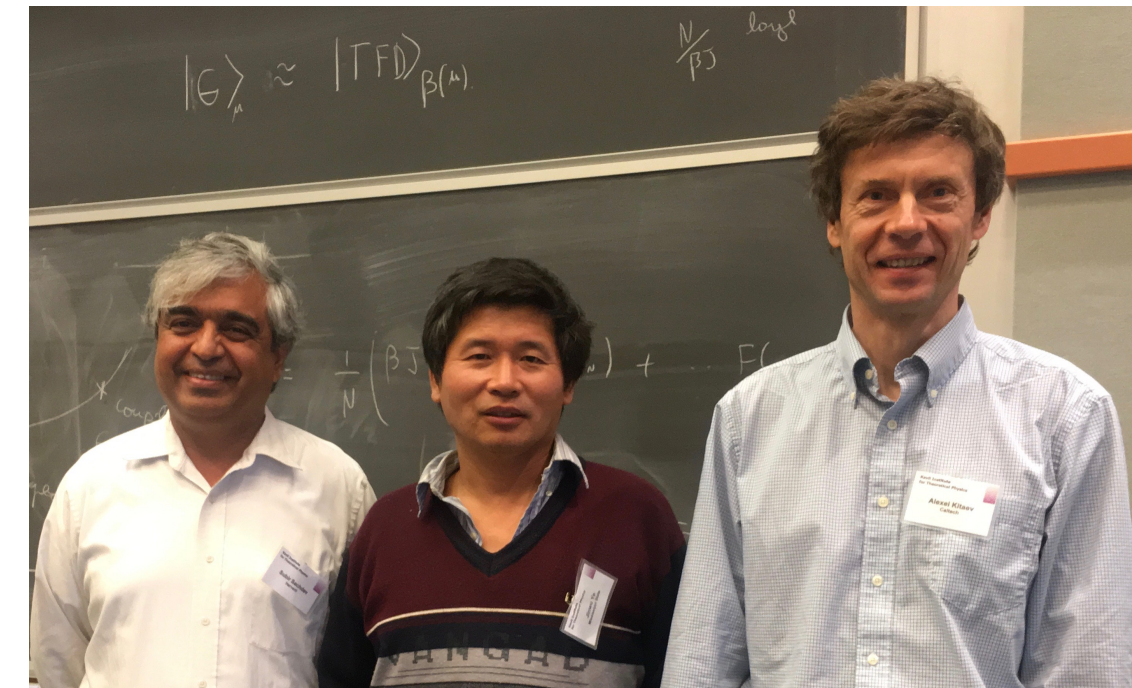


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Sachdev, Ye (1993); Kitaev (2015)



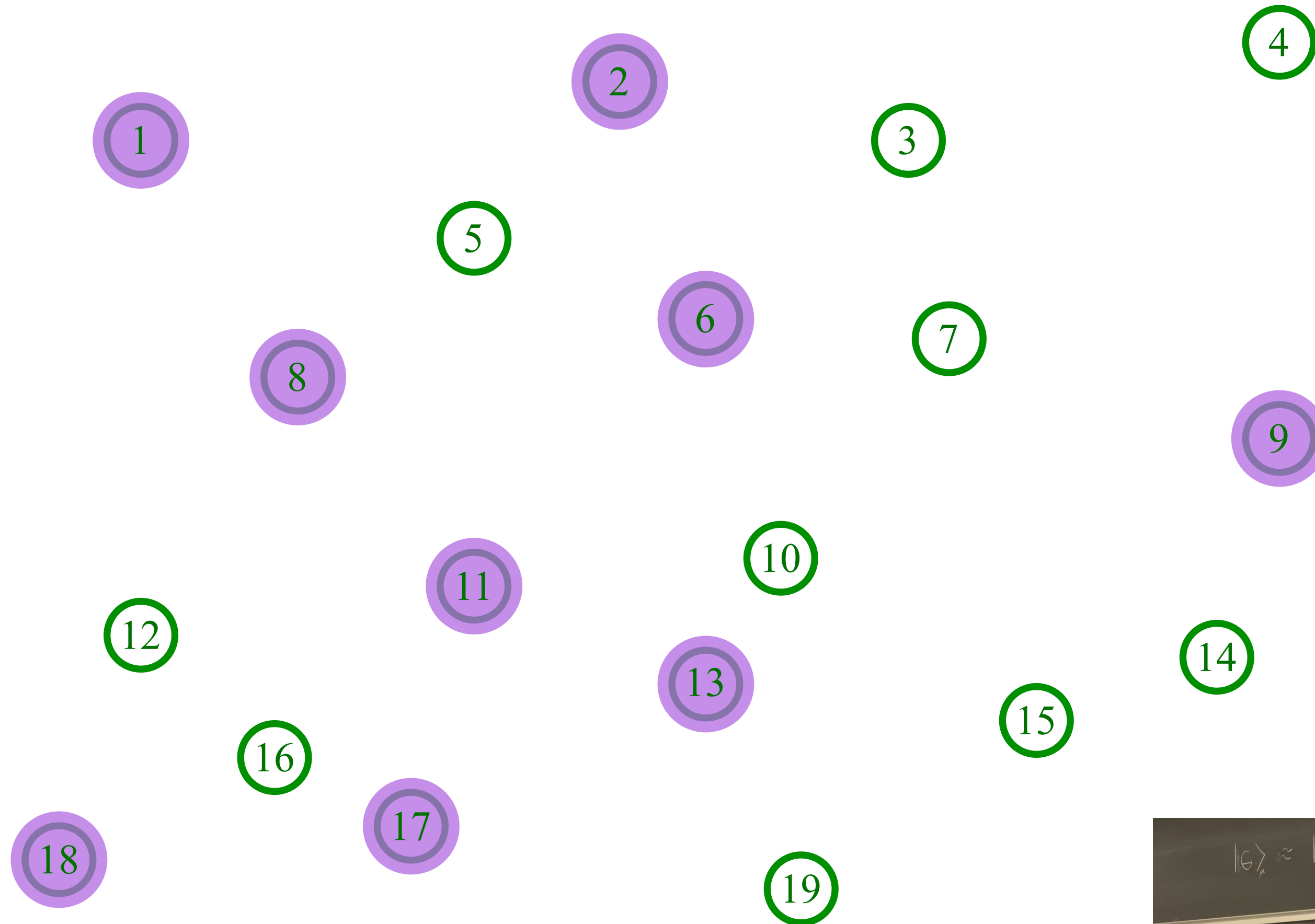
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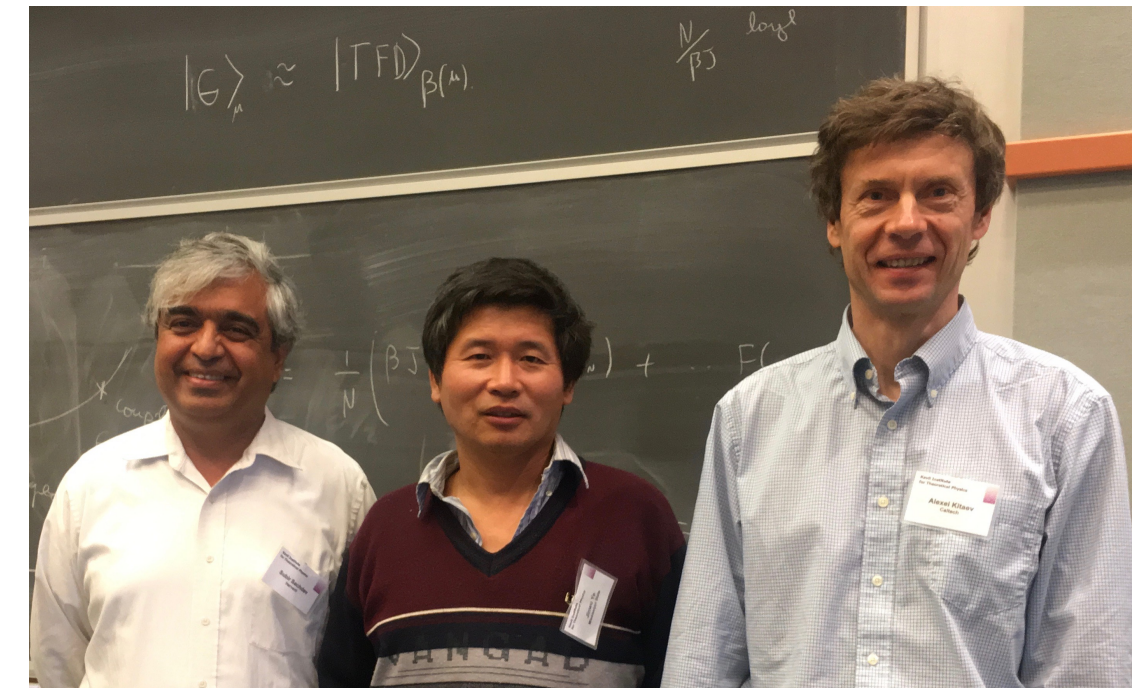
The Sachdev-Ye-Kitaev (SYK) model

Sachdev, Ye (1993); Kitaev (2015)

$$U_{14,19;1,13}$$



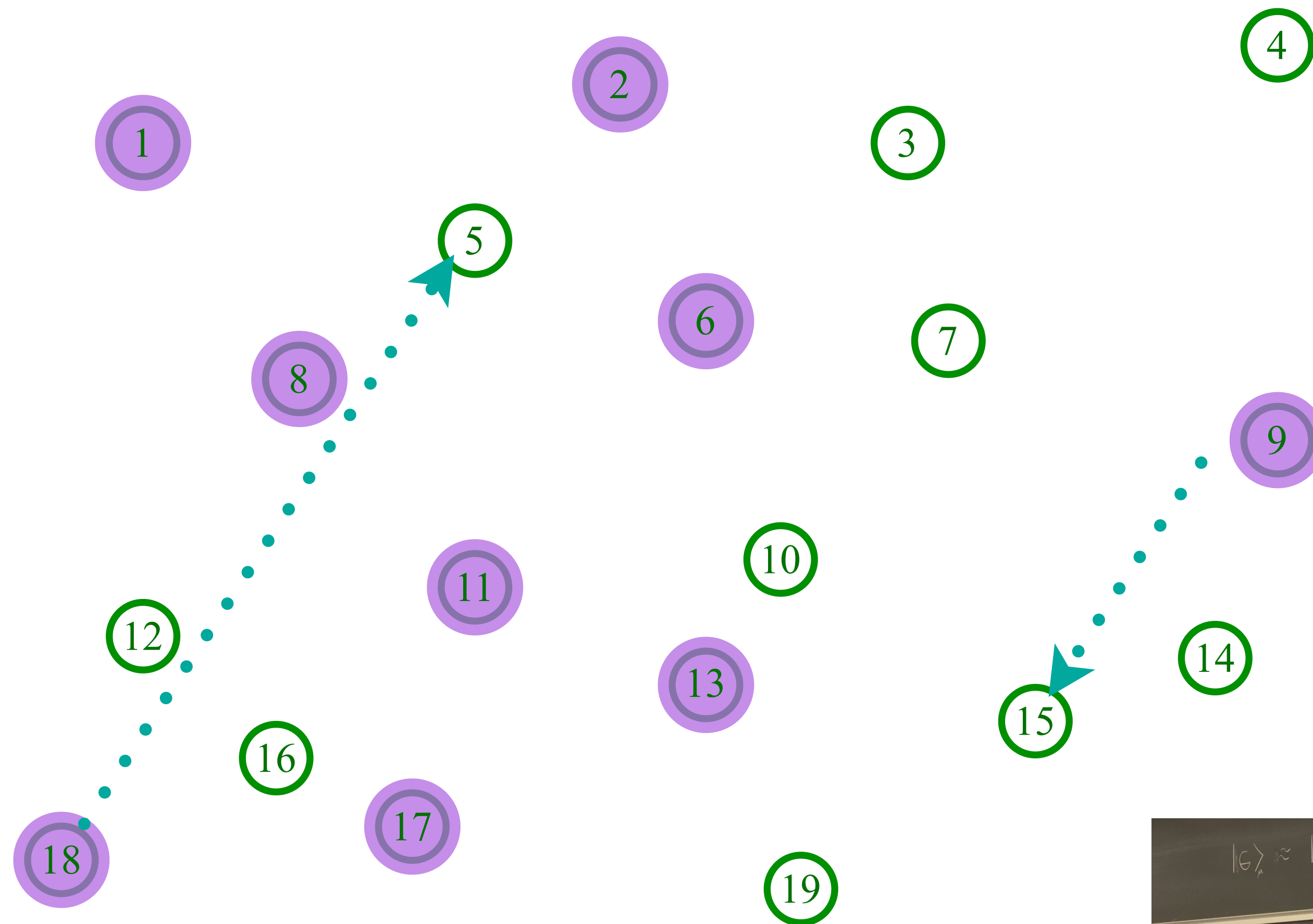
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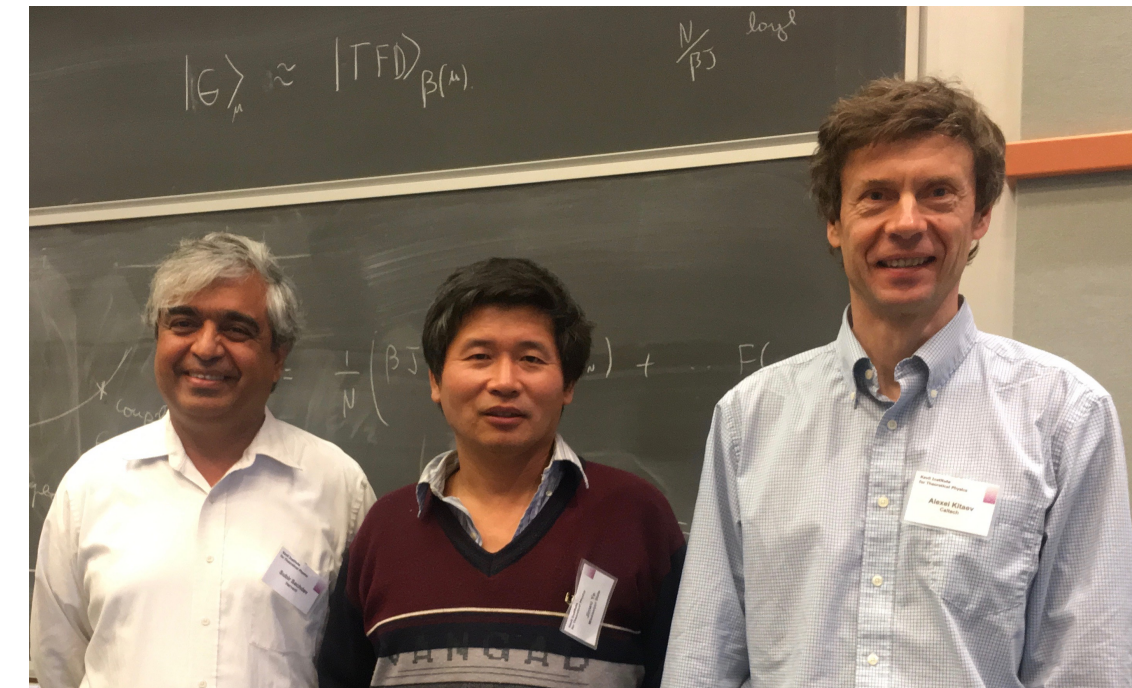
The Sachdev-Ye-Kitaev (SYK) model

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$$U_{9,18;5,15}$$



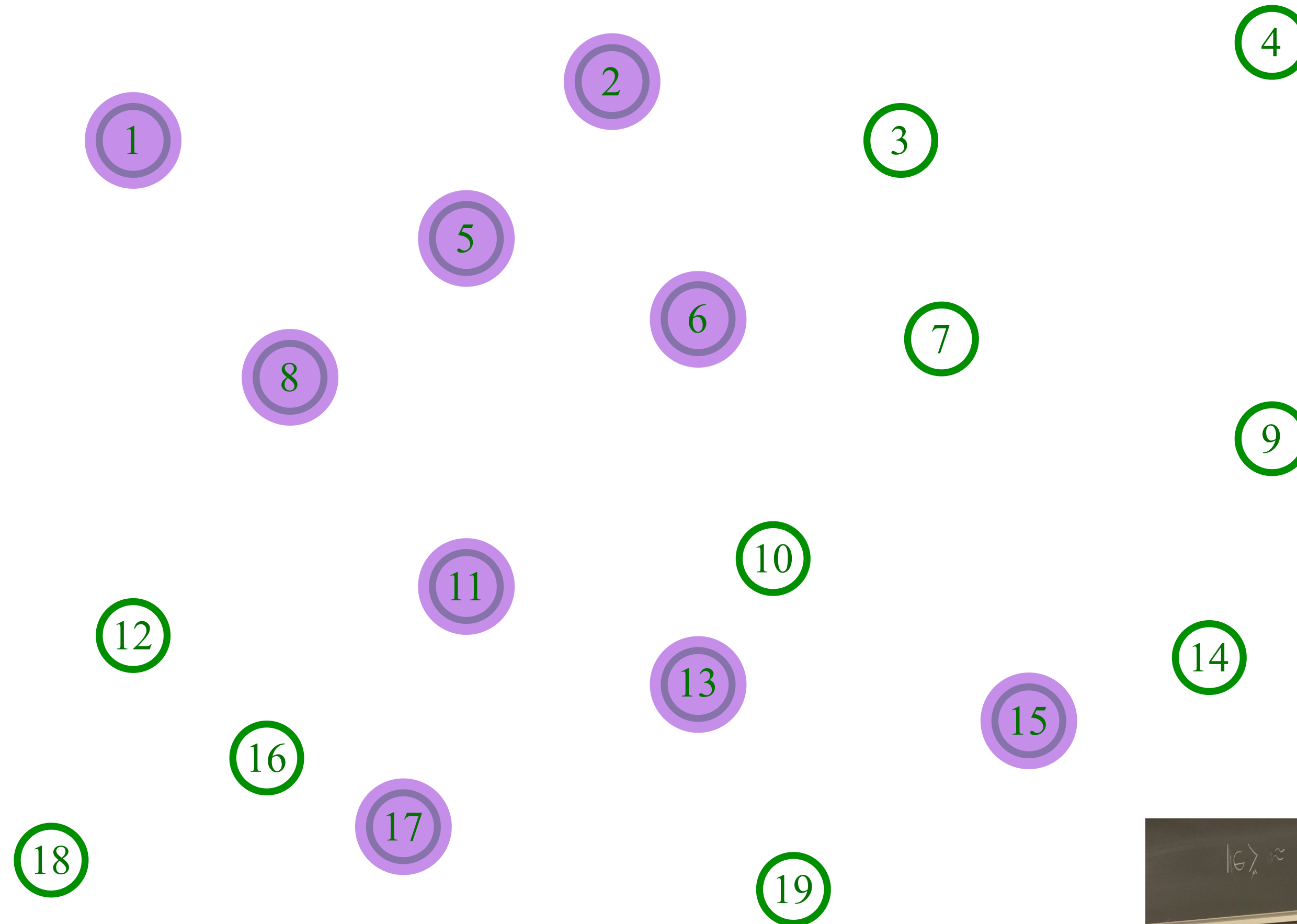
Entangle electrons pairwise randomly



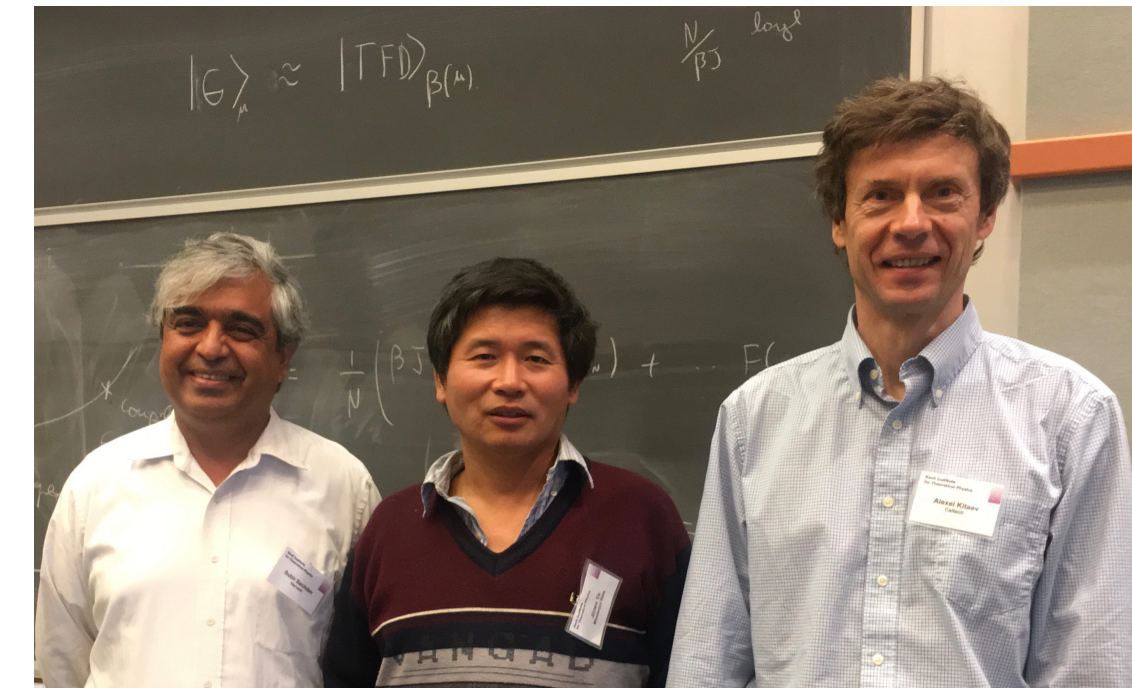
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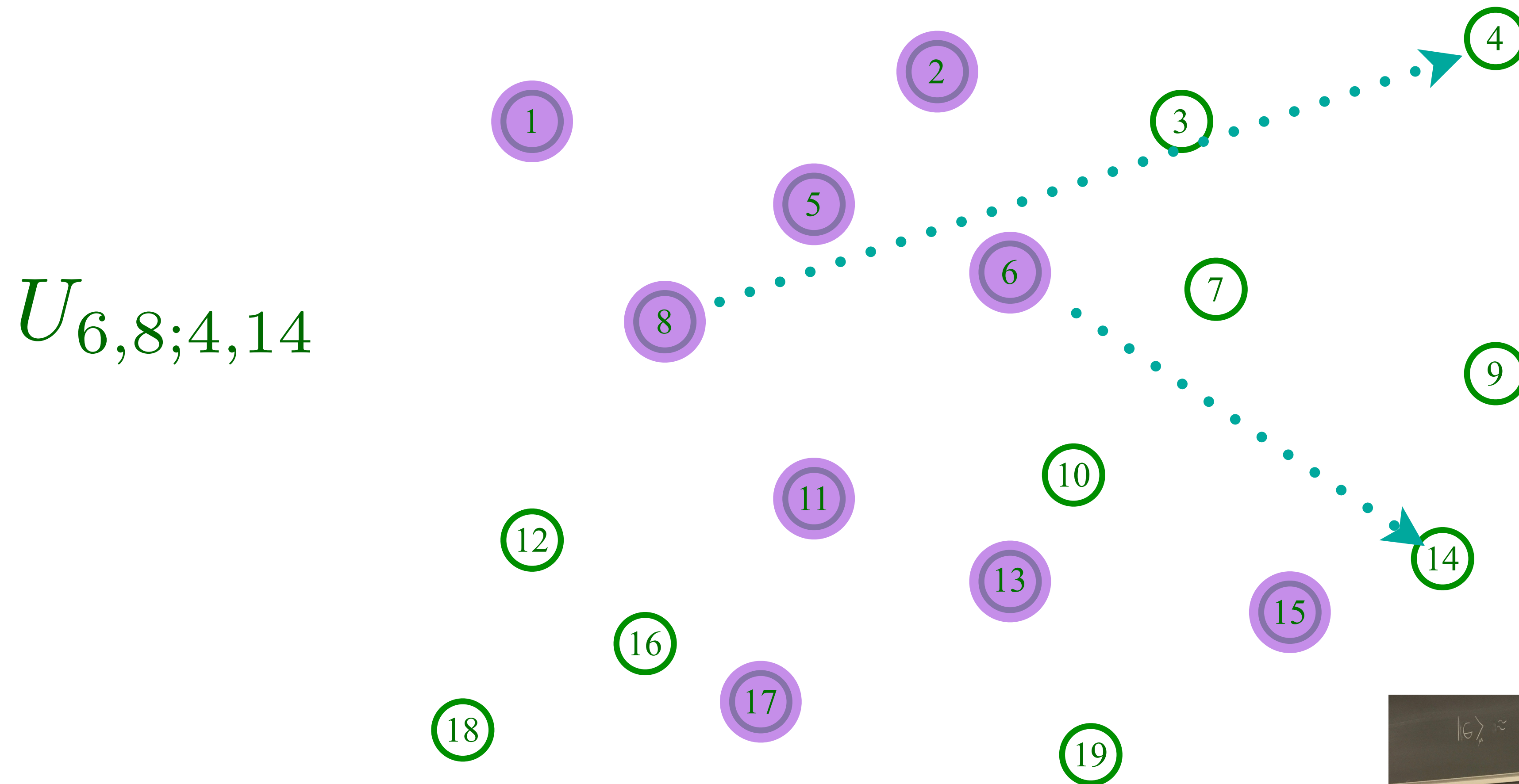


Entangle electrons pairwise randomly

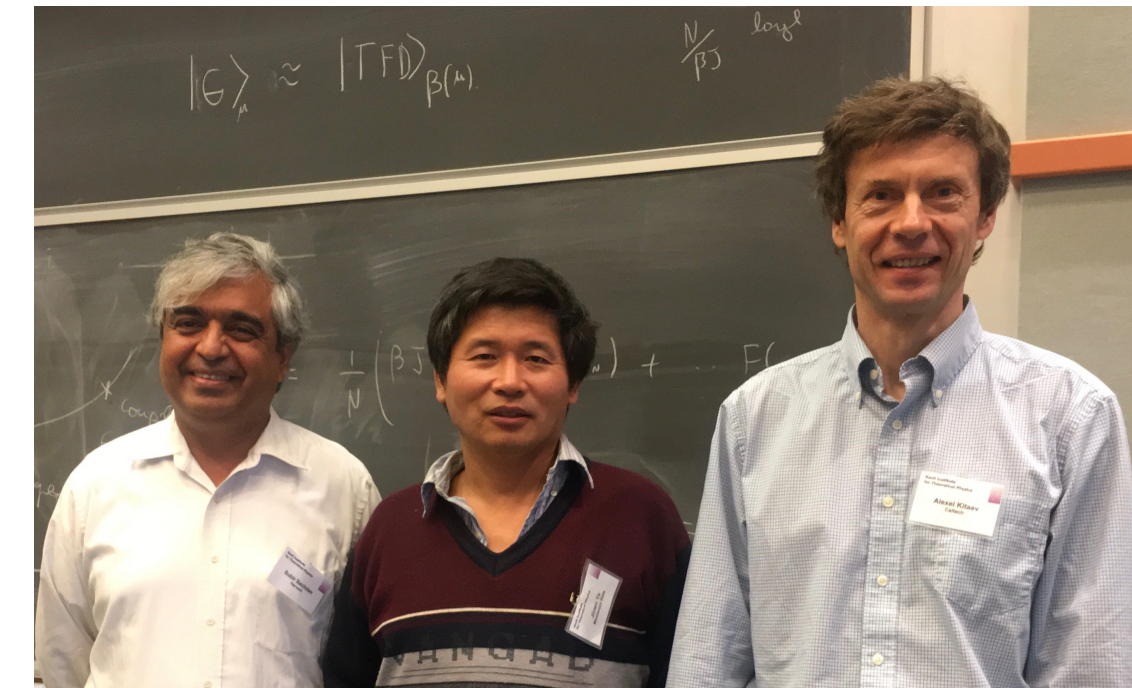


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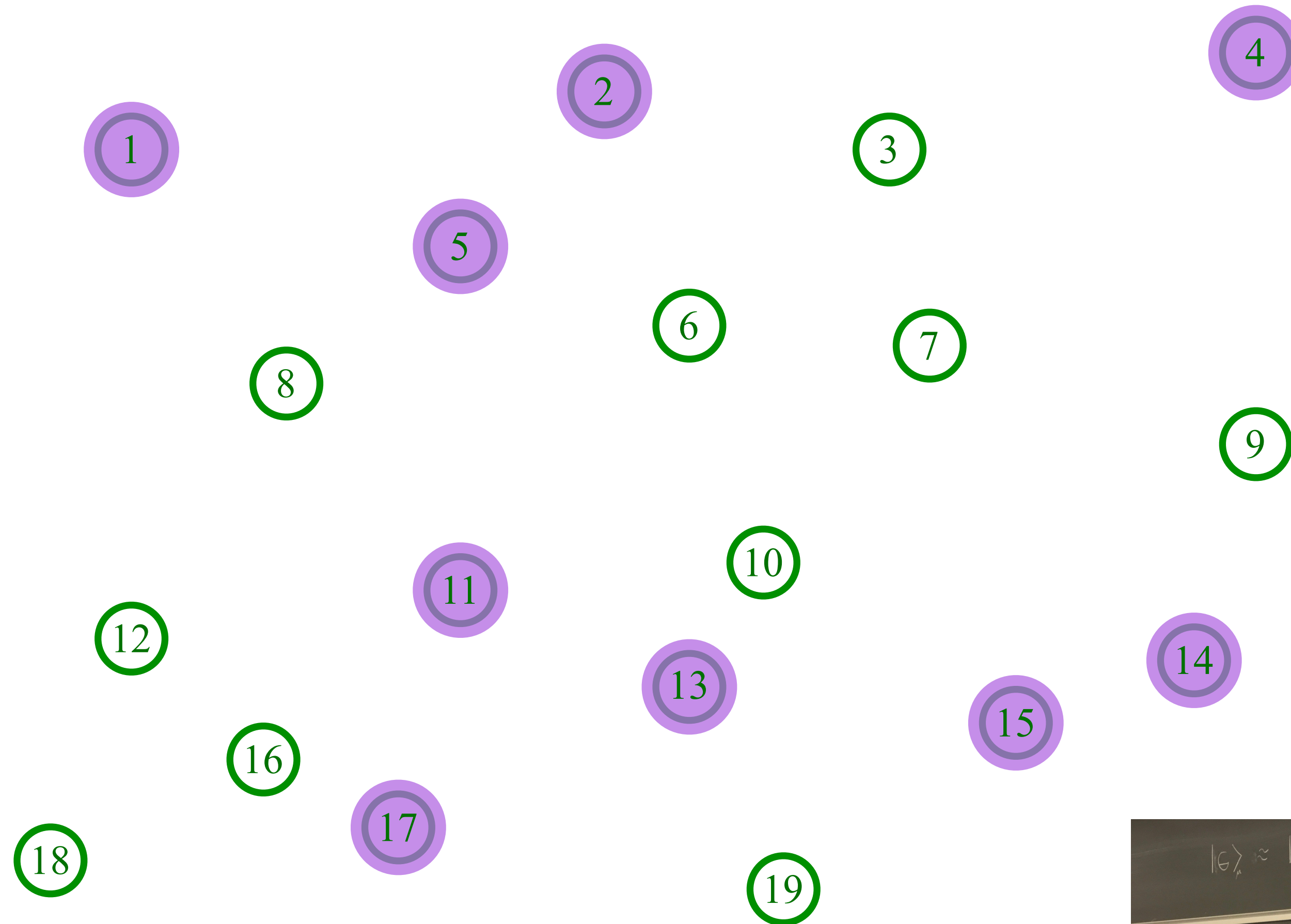
Entangle electrons pairwise randomly



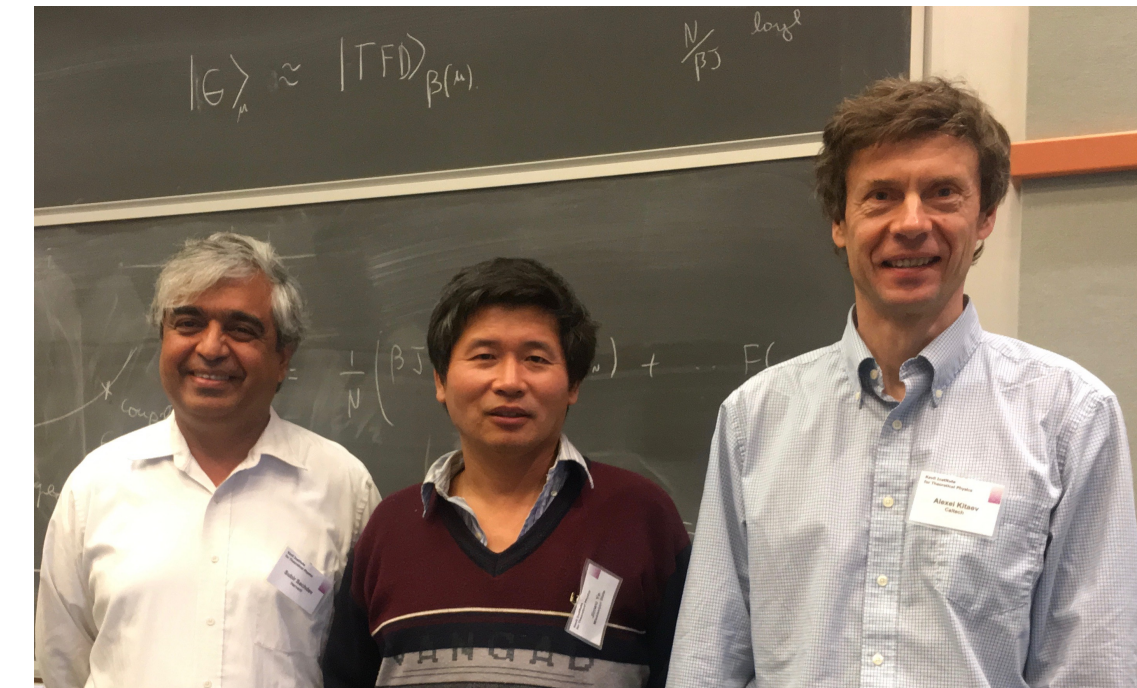
The Sachdev-Ye-Kitaev (SYK) model

Sachdev, Ye (1993); Kitaev (2015)

$$U_{6,8;4,14}$$



Entangle electrons pairwise randomly



The Sachdev-Ye-Kitaev (SYK) model

(See also: the “2-Body Random Ensemble” in nuclear physics; did not obtain the large N limit;
T.A. Brody, J. Flores, J.B. French, P.A. Mello, A. Pandey, and S.S.M. Wong, Rev. Mod. Phys. **53**, 385 (1981))

$$\mathcal{H} = \frac{1}{(2N)^{3/2}} \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta;\gamma\delta} c_{\alpha}^{\dagger} c_{\beta}^{\dagger} c_{\gamma} c_{\delta} - \mu \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

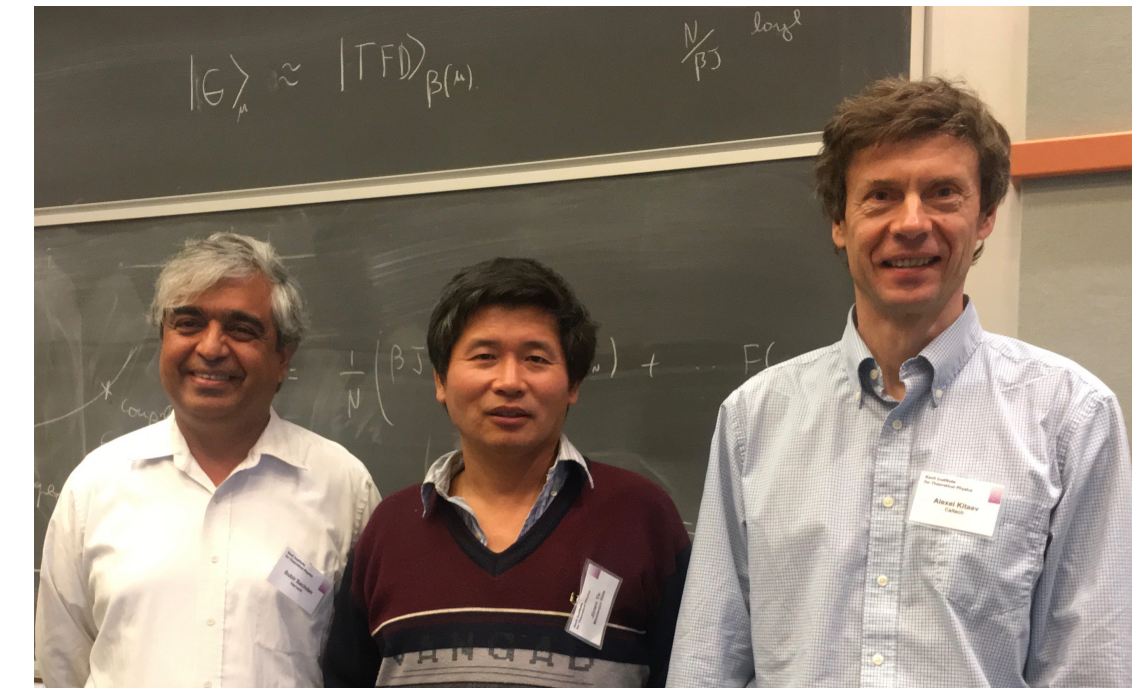
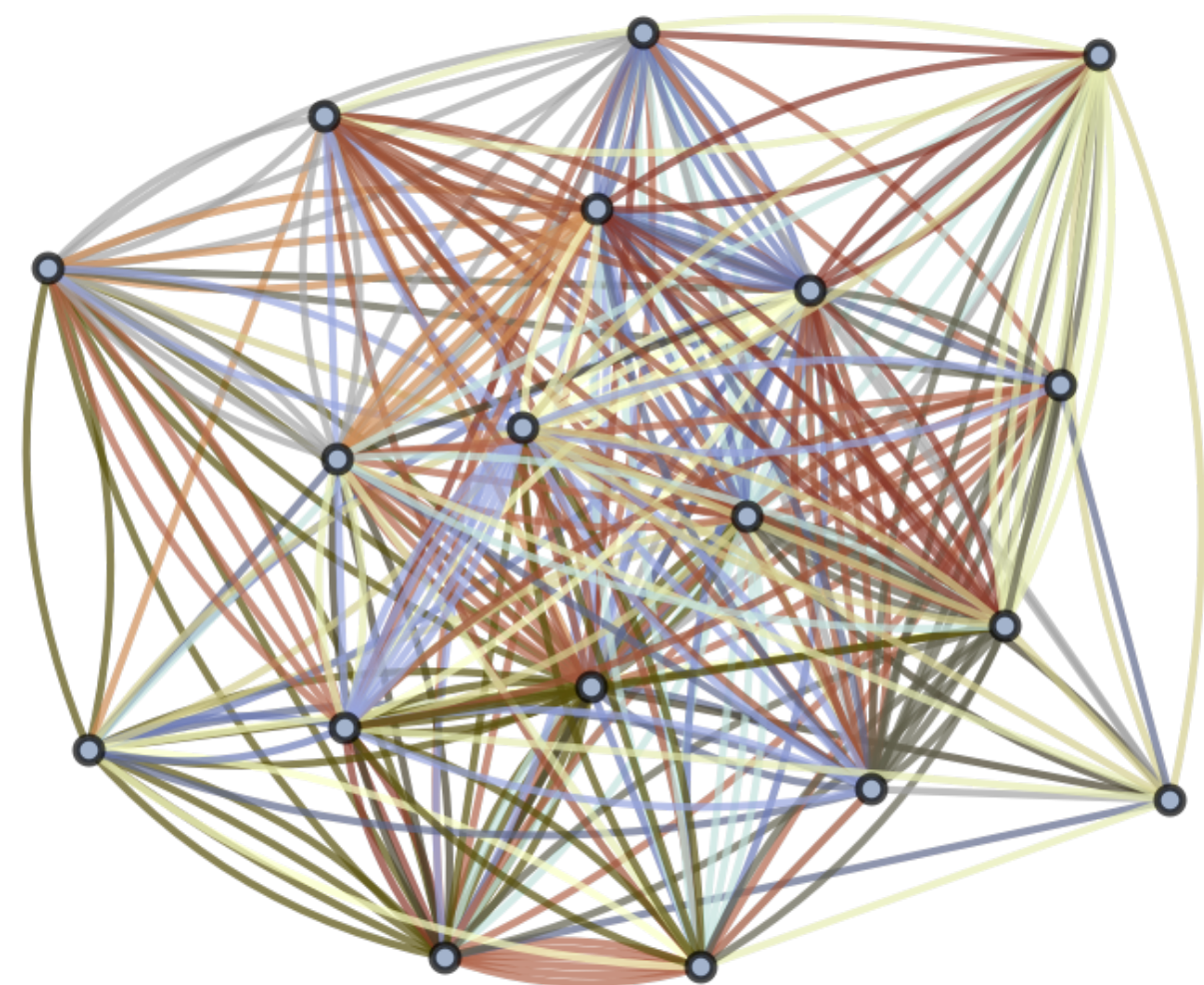
$$c_{\alpha} c_{\beta} + c_{\beta} c_{\alpha} = 0 \quad , \quad c_{\alpha} c_{\beta}^{\dagger} + c_{\beta}^{\dagger} c_{\alpha} = \delta_{\alpha\beta}$$

$$\mathcal{Q} = \frac{1}{N} \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha} ; \quad [\mathcal{H}, \mathcal{Q}] = 0 ; \quad 0 \leq \mathcal{Q} \leq 1$$

$U_{\alpha\beta;\gamma\delta}$ are independent random variables with $\overline{U_{\alpha\beta;\gamma\delta}} = 0$ and $\overline{|U_{\alpha\beta;\gamma\delta}|^2} = U^2$
 $N \rightarrow \infty$ yields critical strange metal.

S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

A. Kitaev, unpublished; S. Sachdev, PRX **5**, 041025 (2015)

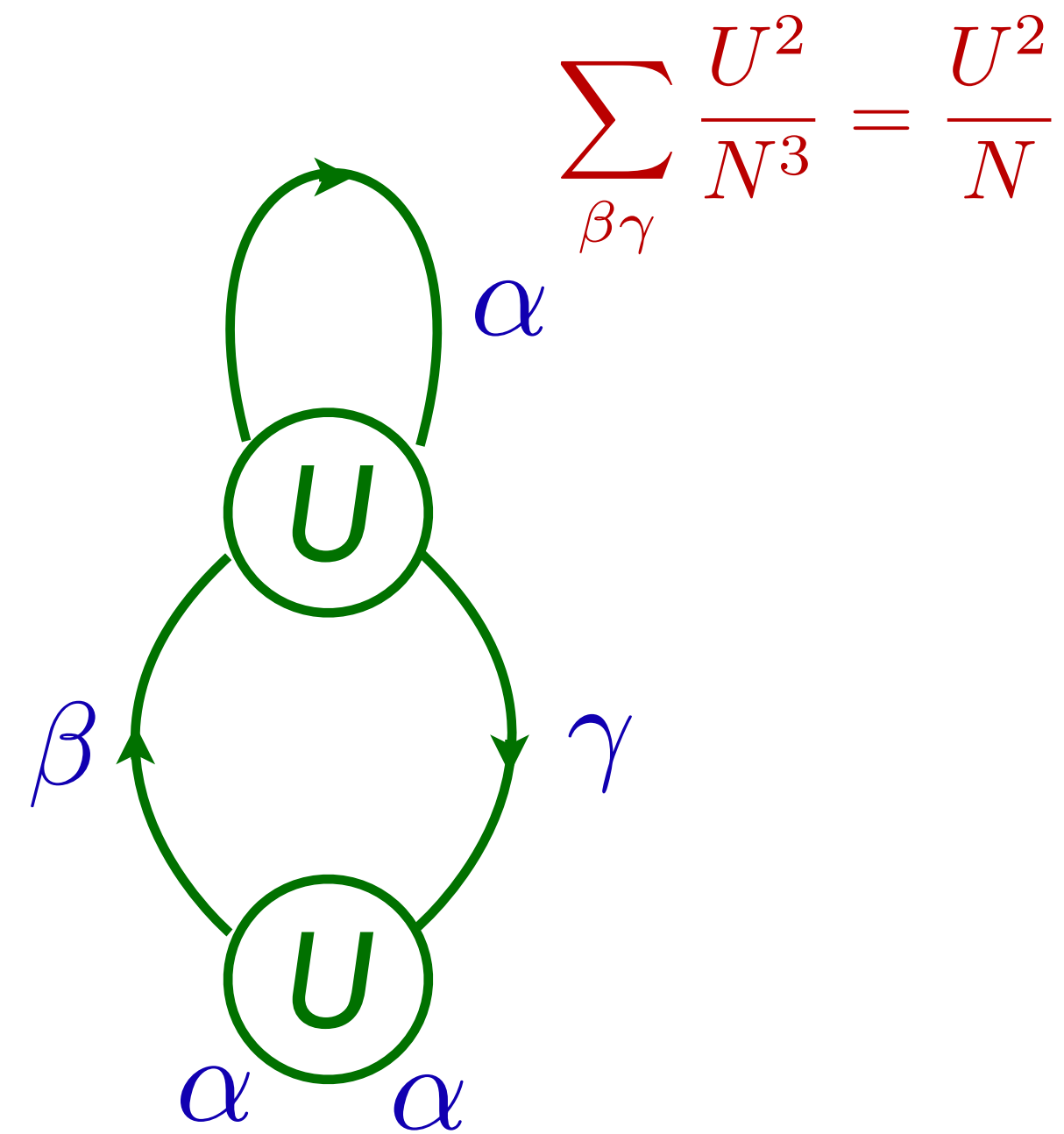
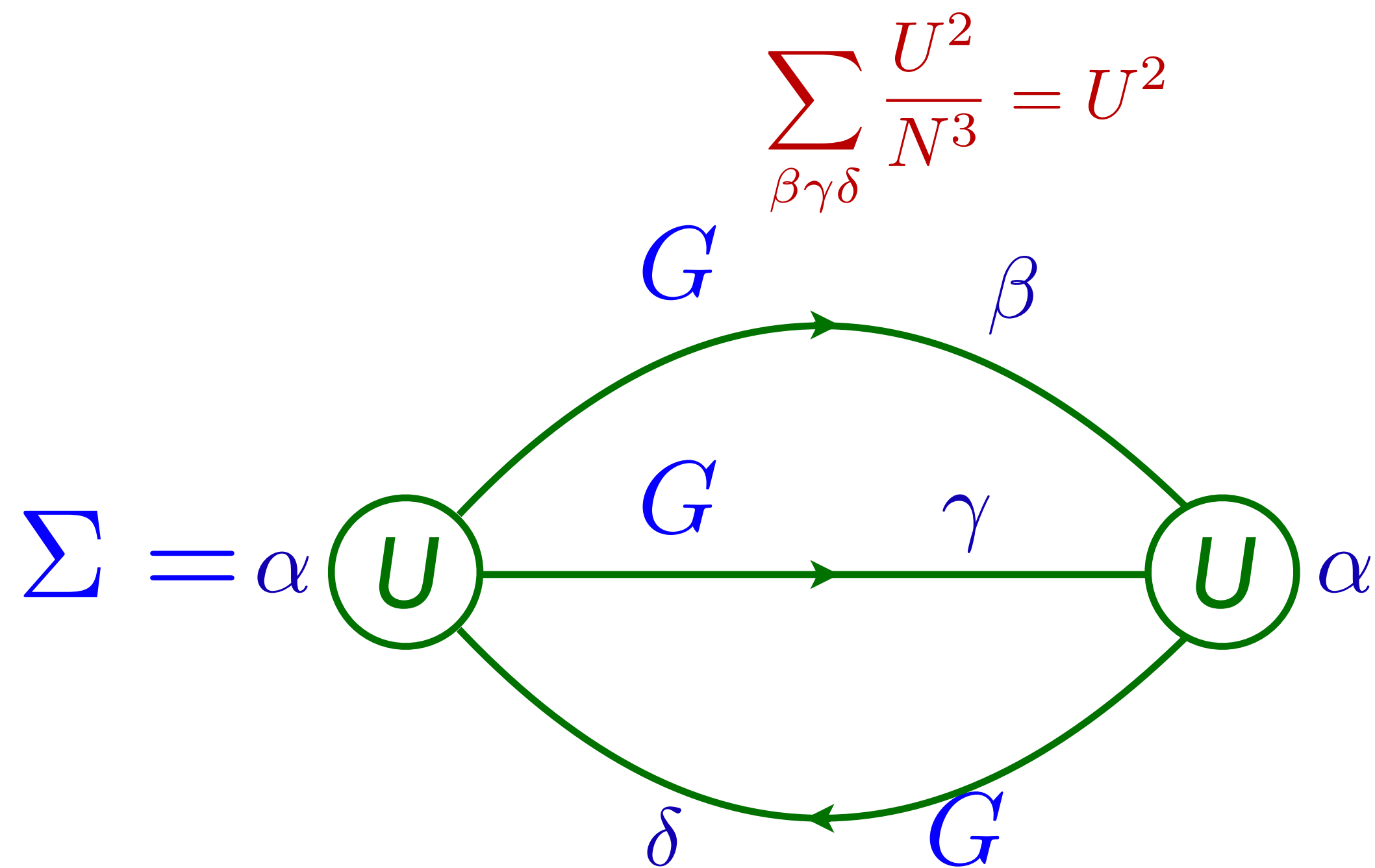


The Sachdev-Ye-Kitaev (SYK) model

Feynman graph expansion in $U_{\alpha\beta;\gamma\delta}$, and graph-by-graph average, yields exact equations in the large N limit:

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$

$$G(\tau = 0^-) = \mathcal{Q}.$$



S. Sachdev and J. Ye,
PRL **70**, 3339 (1993)



The Sachdev-Ye-Kitaev (SYK) model

Sachdev, Ye (1993); Kitaev (2015)

A solvable model of multi-particle
quantum entanglement.

No quasiparticles: yields a metal in which
current is carried
not by individual electrons,
but by an entangled “quantum soup”

The SYK model

Consequences of emergent time-reparameterization and conformal symmetries
in low-energy theory in 0+1 spacetime dimensions:

1. Planckian dynamics!

$$\tau(\omega) = \frac{\hbar}{k_B T} F\left(\frac{\hbar\omega}{k_B T}\right)$$

The SYK model

Consequences of emergent time-reparameterization and conformal symmetries
in low-energy theory in 0+1 spacetime dimensions:

1. Planckian dynamics!

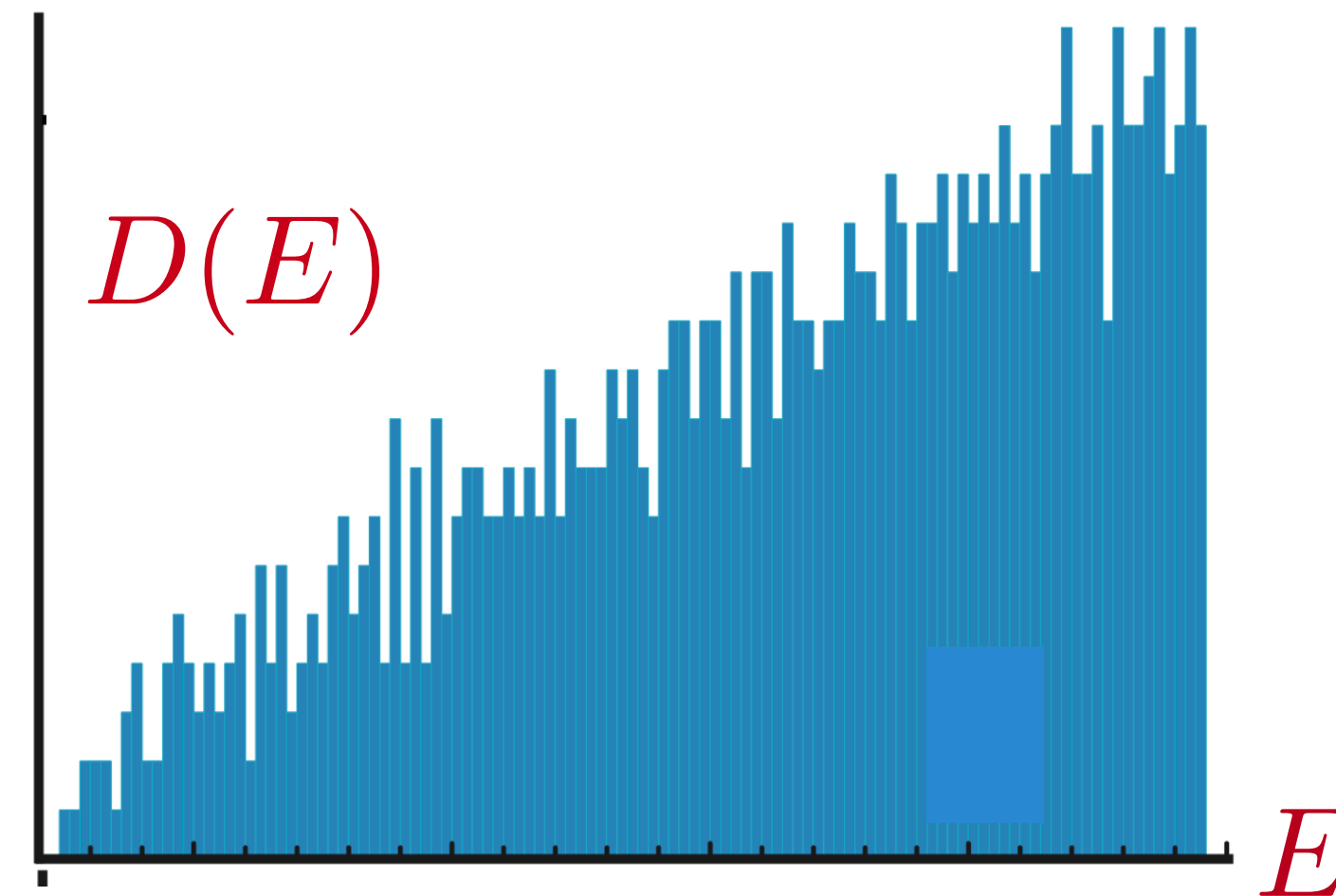
$$\tau(\omega) = \frac{\hbar}{k_B T} F\left(\frac{\hbar\omega}{k_B T}\right)$$



2. Zero temperature entropy without exponential ground state degeneracy!

$$\lim_{T \rightarrow 0} \lim_{N \rightarrow \infty} \frac{1}{N} S(T) = s_0 \quad , \quad D(E \rightarrow 0) = e^{N s_0} f_{\text{smooth}}(E)$$

$$s_0 = 0.46484769917080510749\dots \text{ for } Q = 1/2.$$



From the SYK model to the universal 2d-YSYK theory of strange metals

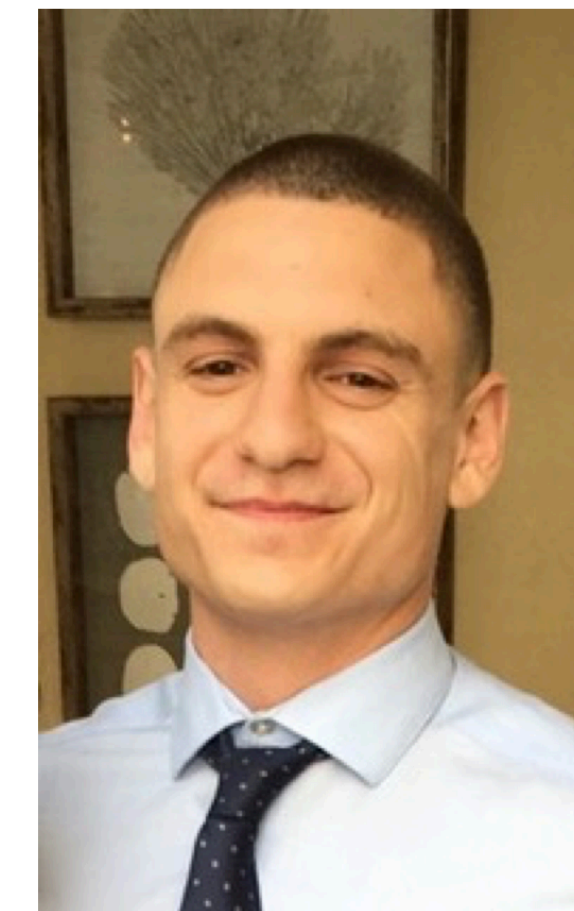
Aavishkar A. Patel, Haoyu Guo, Ilya Esterlis, S. Sachdev, *Science* **381**, 790 (2023)



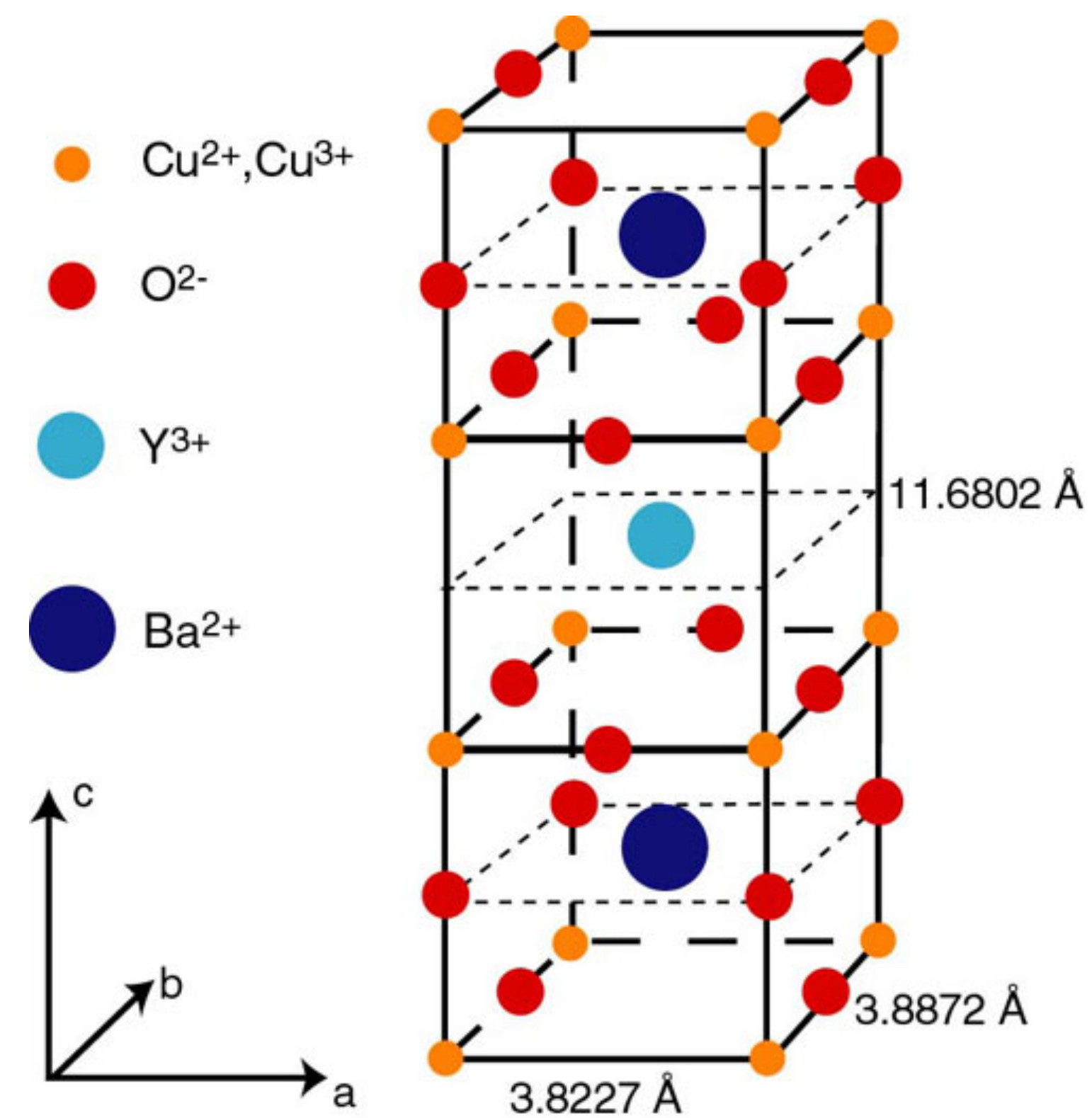
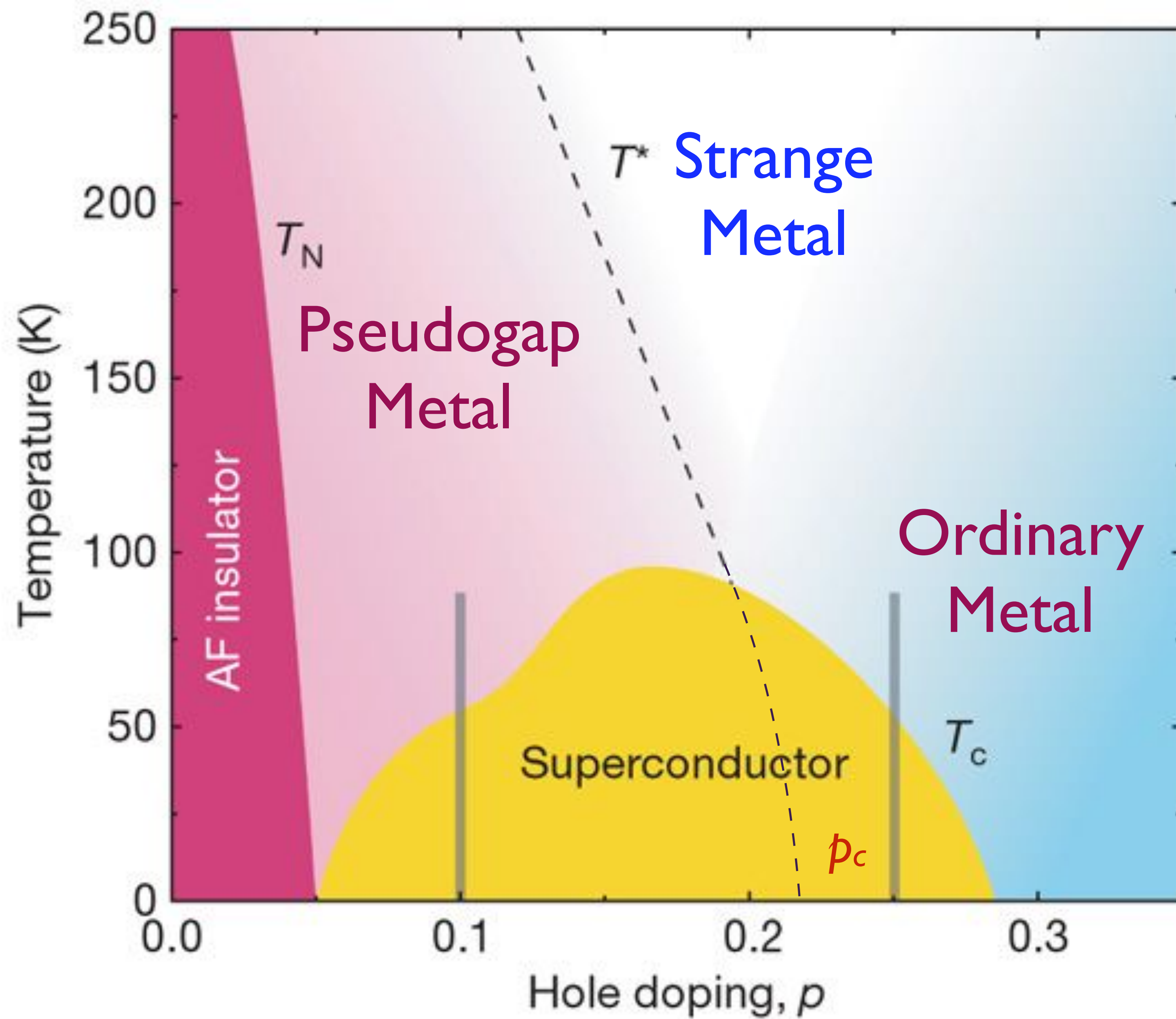
Aavishkar Patel
Flatiron

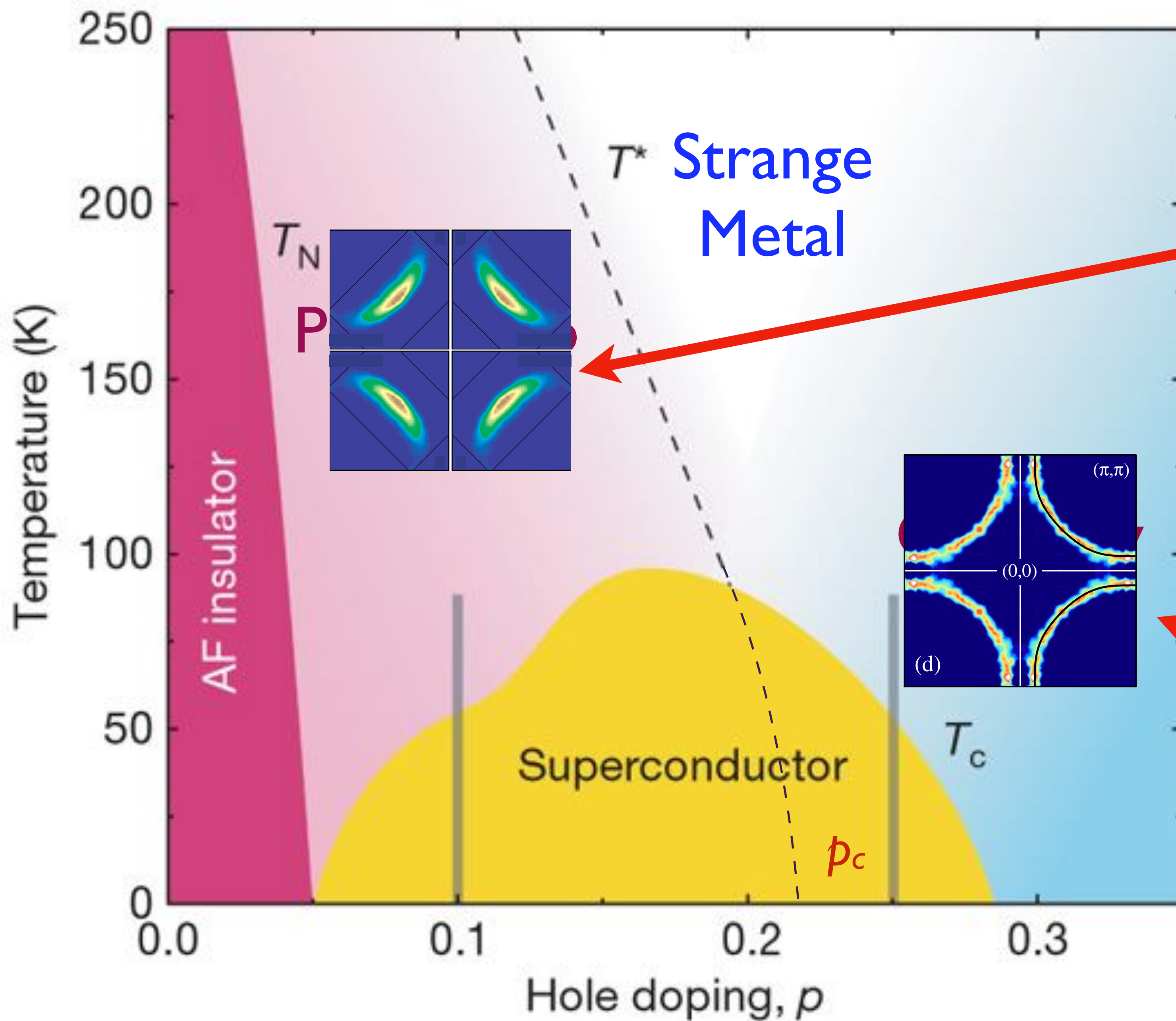


Haoyu Guo
Cornell



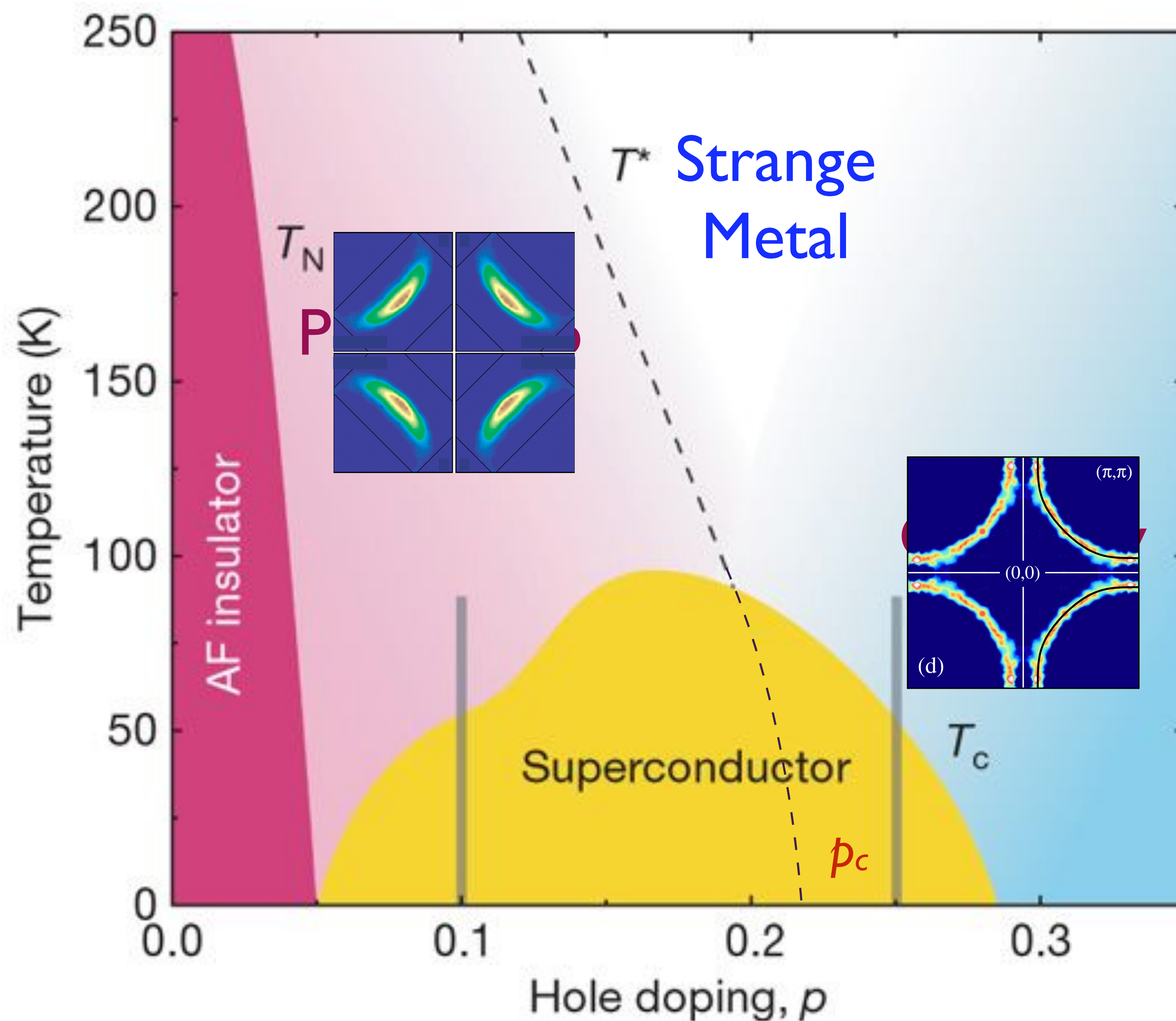
Ilya Esterlis
Wisconsin





“Pseudogap metal”
Fermi surface
modified by
electron-electron
interactions

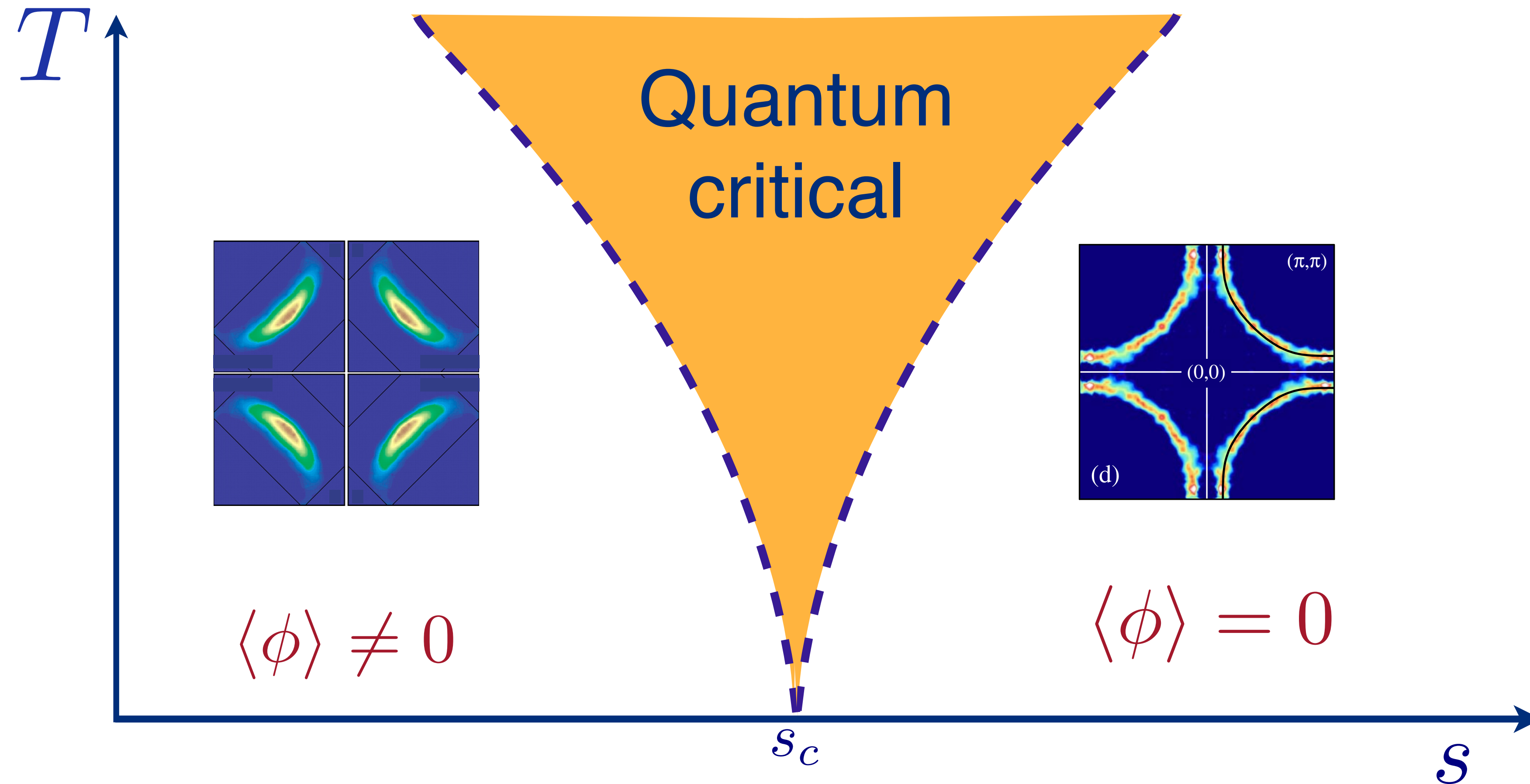
Fermi surface
as expected
in a model
of free electrons



View the strange metal as a property of a $T = 0$ quantum phase transition involving change in the Fermi surface.

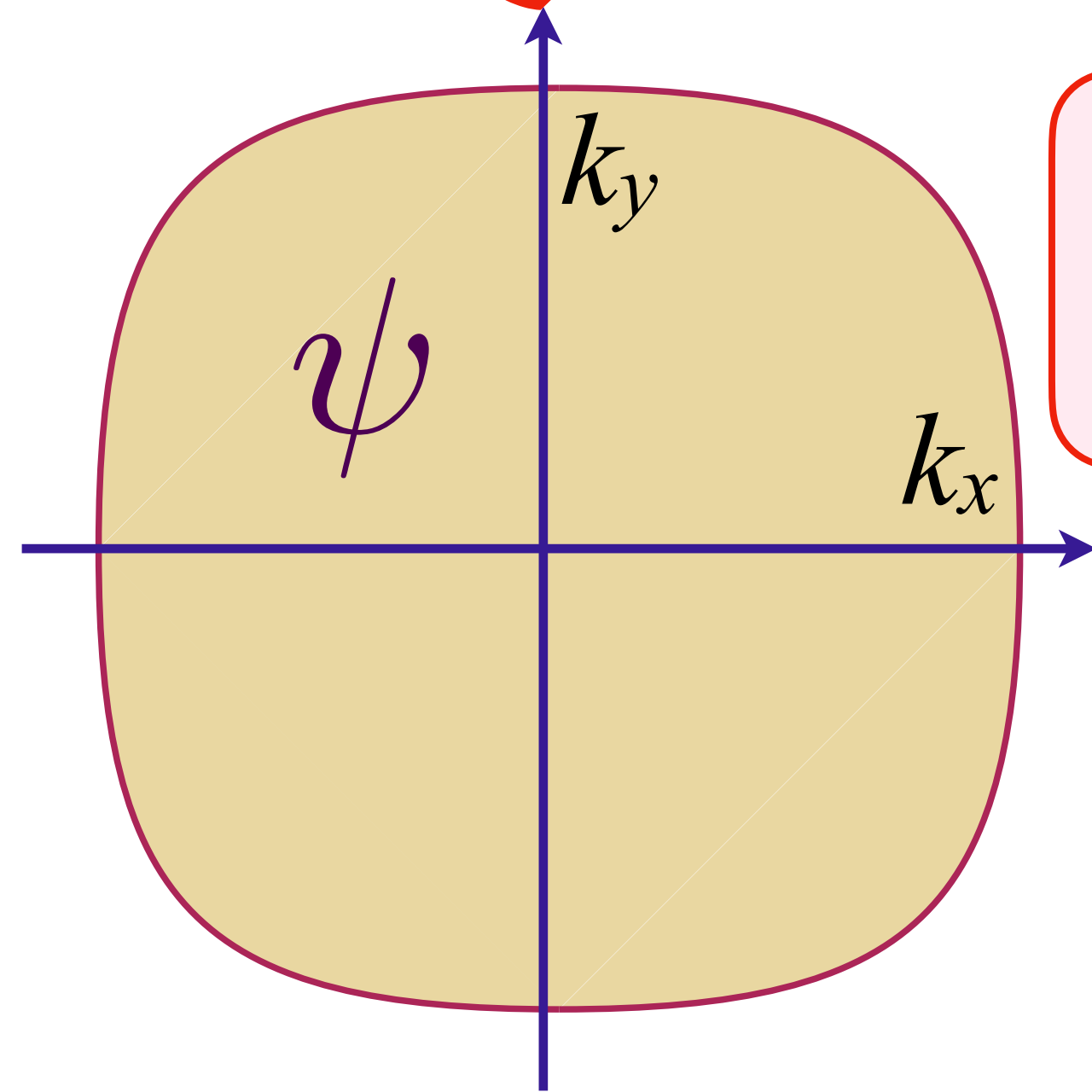
The onset of superconductivity may “hide” this quantum transition.

Quantum phase transitions of Fermi surface change



Fermi surface
+
a boson ϕ
with a 'mass' s
and
a boson-fermion
Yukawa coupling g .

Quantum phase transitions of Fermi surface change



Universal theory: the two-dimensional Yukawa-Sachdev-Ye-Kitaev model

$$+ \mathcal{L}[\phi]$$

$$+ [g_{\alpha\beta\gamma} + g'_{\alpha\beta\gamma}(\mathbf{r})] \psi_{\alpha}^{\dagger}(\mathbf{r}) \psi_{\beta}(\mathbf{r}) \phi_{\gamma}(\mathbf{r})$$

Key ingredient:
spatial disorder in
quantum critical coupling
associated with $g'(\mathbf{r})$.

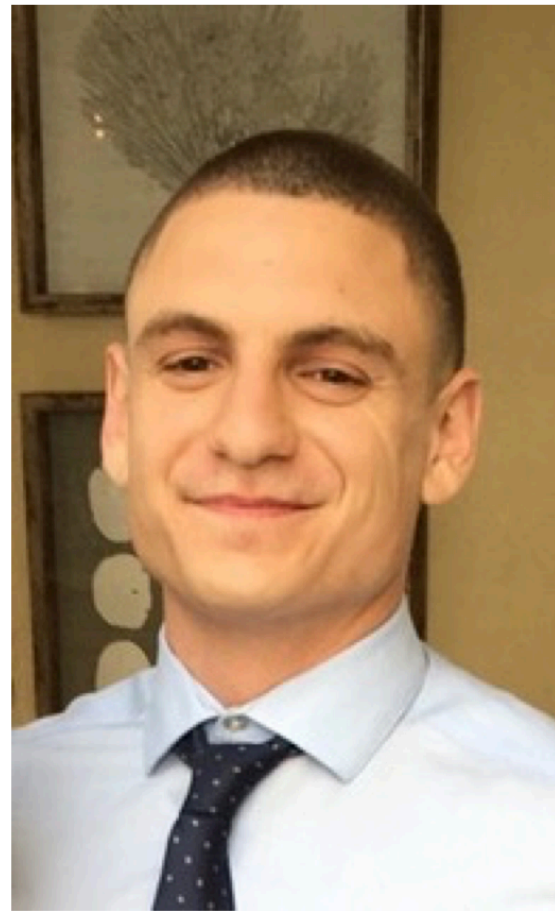
Spatially uniform Yukawa coupling g
with $\overline{g_{\alpha\beta\gamma}} = 0$, $\overline{g_{\alpha\beta\gamma} g_{abc}} = g^2 \delta_{\alpha a} \delta_{\beta b} \delta_{\gamma c}$

Spatially uniform theory
does *not* yield
a strange metal;
but a perfect metal.

Spatially random Yukawa coupling $g'(\mathbf{r})$
with $\overline{g'_{\alpha\beta\gamma}(\mathbf{r})} = 0$, $\overline{g'_{\alpha\beta\gamma}(\mathbf{r}) g'_{abc}(\mathbf{r}')} = g'^2 \delta_{\alpha a} \delta_{\beta b} \delta_{\gamma c} \delta(\mathbf{r} - \mathbf{r}')$

Strange metal and superconductor in the two-dimensional Yukawa-Sachdev-Ye-Kitaev model

Chenyuan Li, Aavishkar A. Patel, Haoyu Guo, Davide Valentinis, Jorg Schmalian, S.S., Ilya Esterlis, PRL in press; arXiv:2406.07608



Ilya Esterlis
Wisconsin



Haoyu Guo
Cornell



Aavishkar Patel
Flatiron



Chenyuan Li
Harvard → Rice



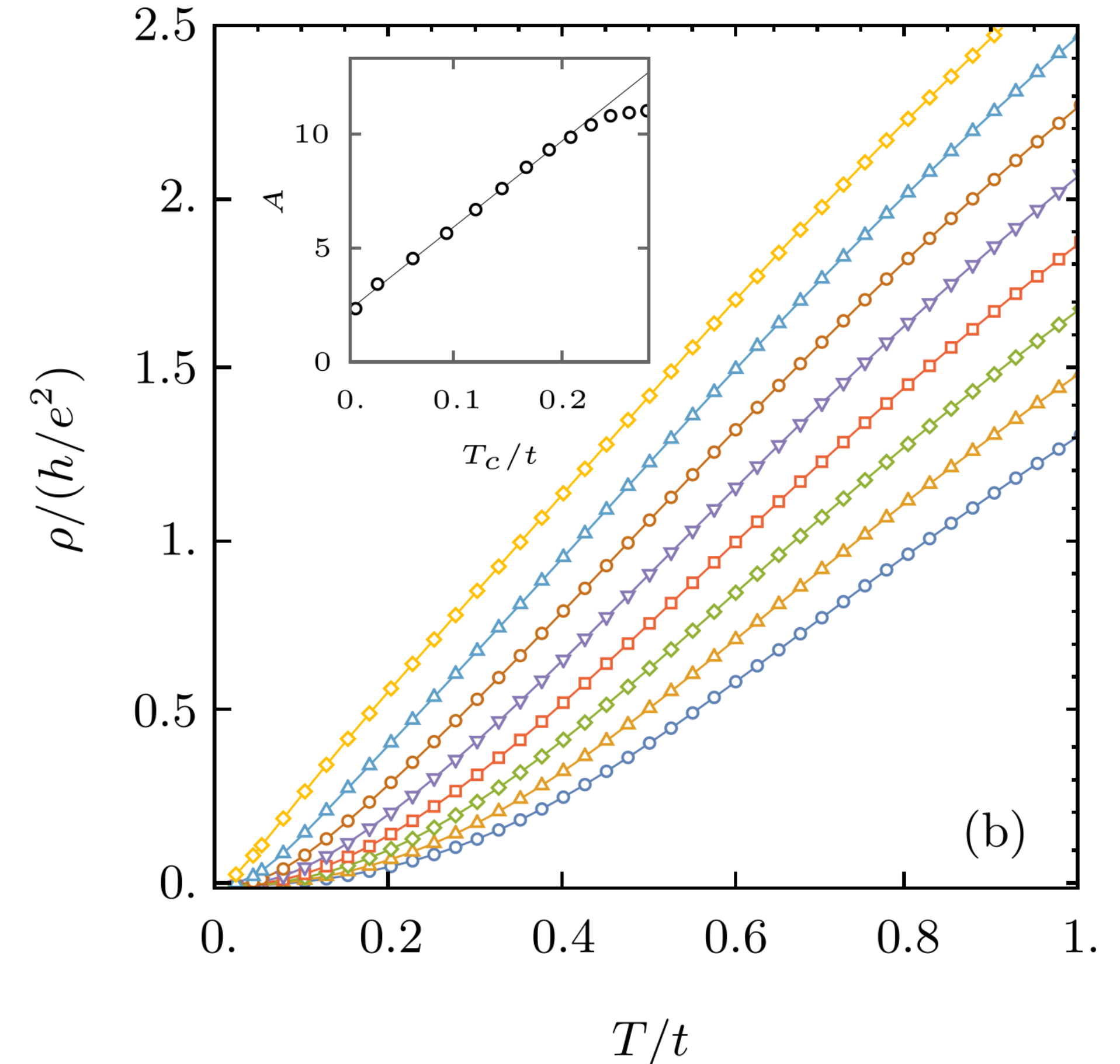
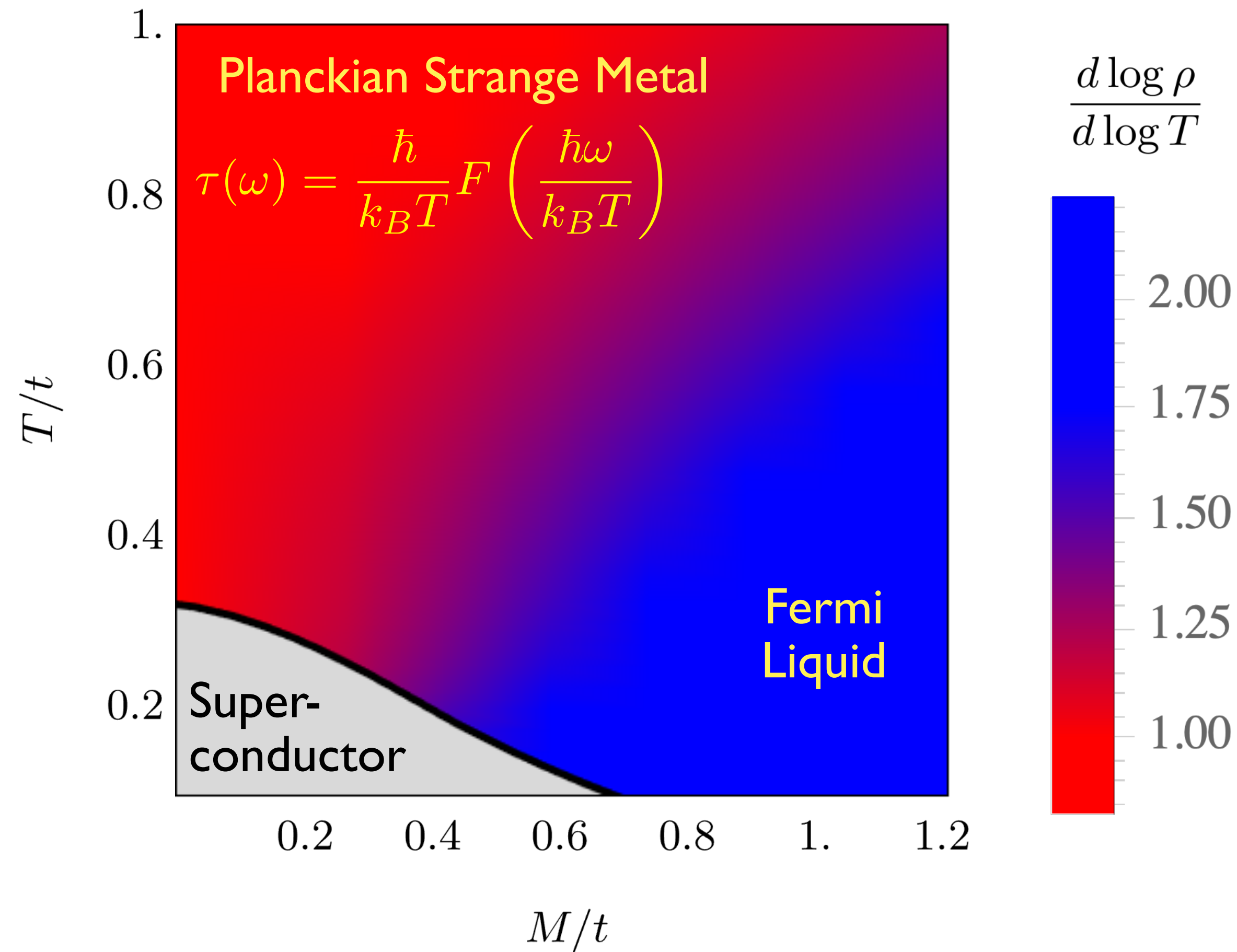
Davide Valentinis
KIT



Joerg Schmalian
KIT

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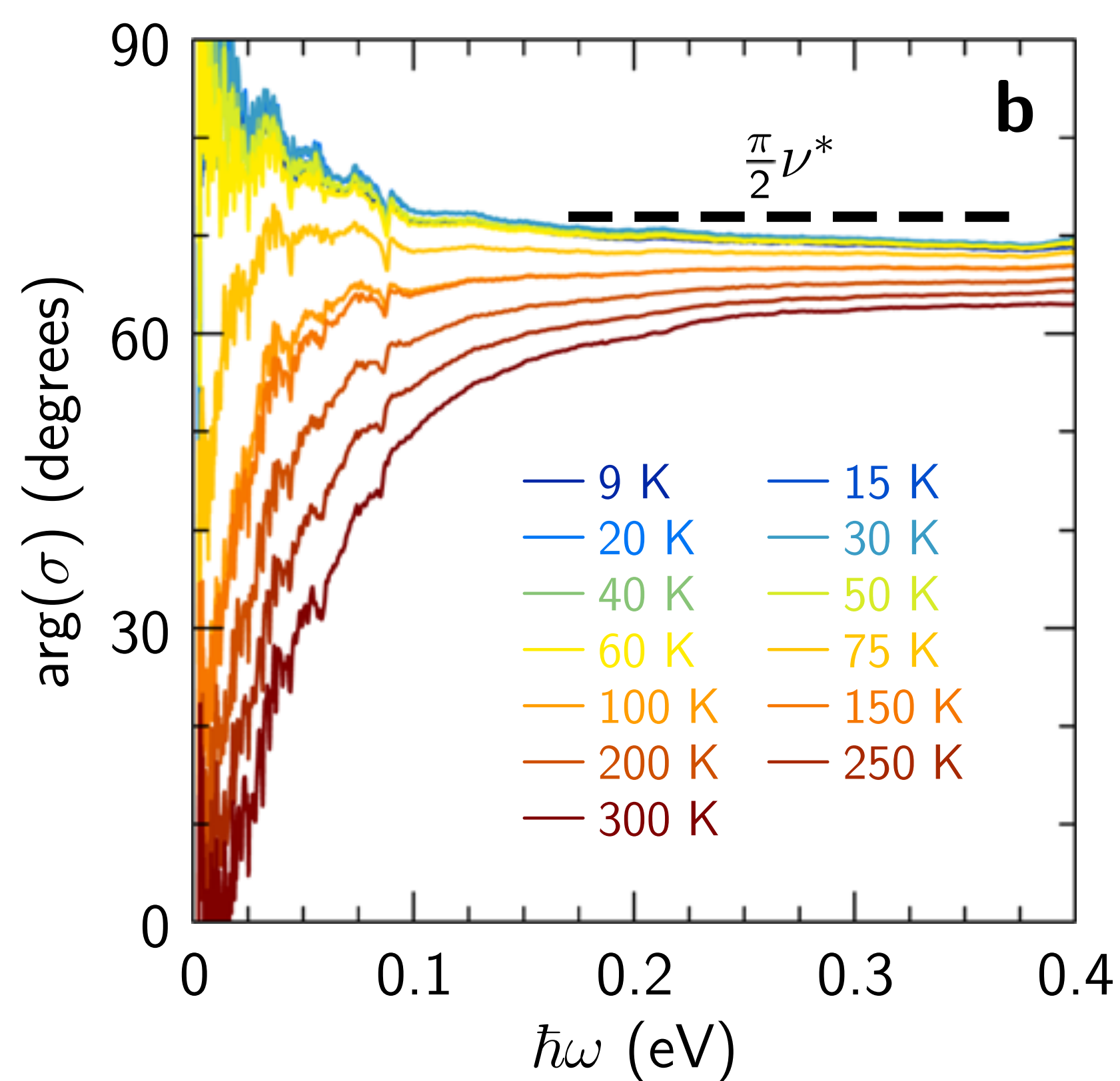
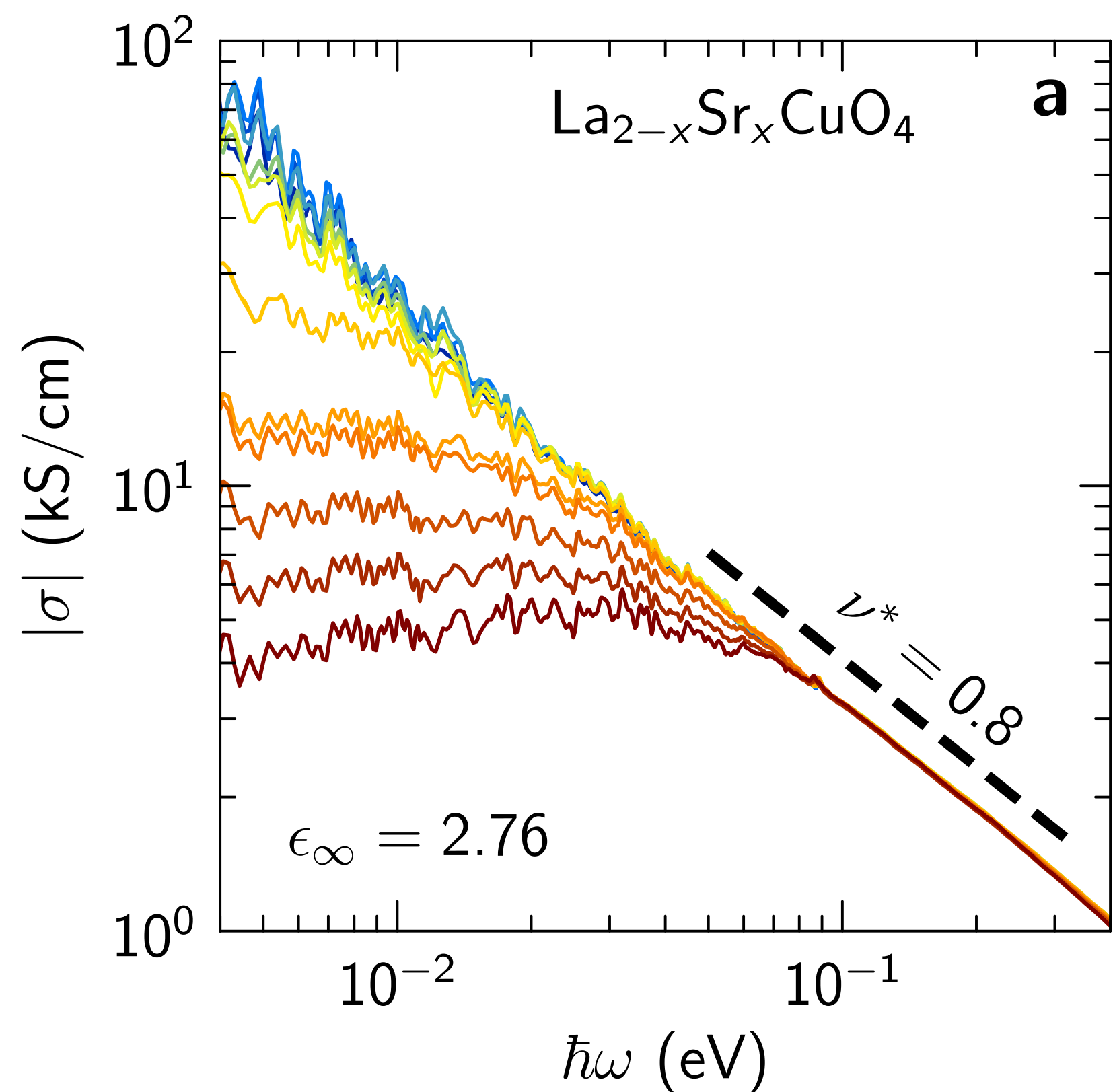


Reconciling scaling of the optical conductivity of cuprate superconductors with Planckian resistivity and specific heat

B. Michon, C. Berthod, C. W. Rischau, A. Ataei, L. Chen, S. Komiya, S. Ono, L. Taillefer, D. van der Marel, A. Georges

Nature Communications **14**, Article number: 3033 (2023)

$$\sigma(\omega) = i \frac{e^2 K / (\hbar d_c)}{\hbar \omega \frac{m^*(\omega)}{m} + i \frac{\hbar}{\tau(\omega)}}$$

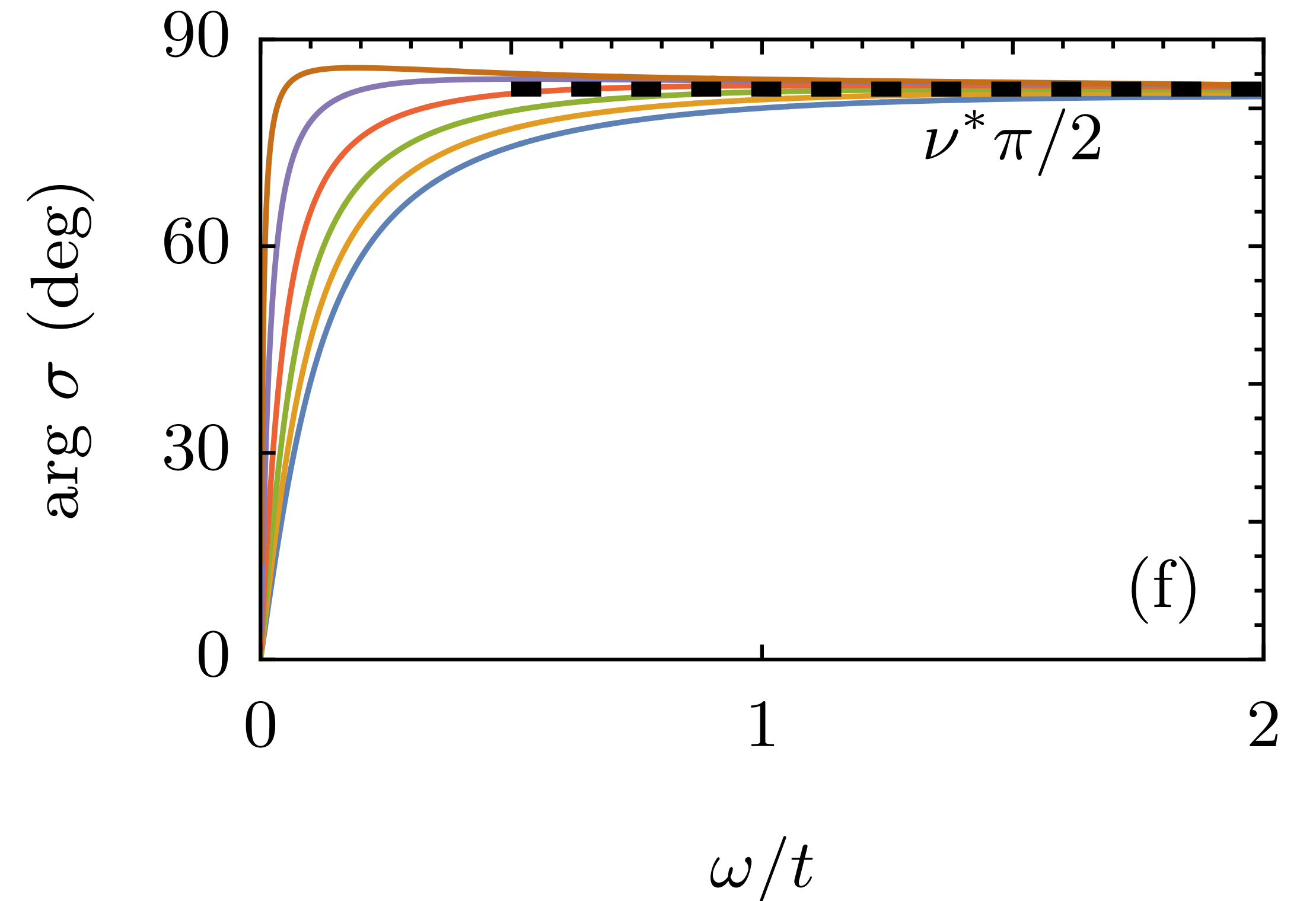
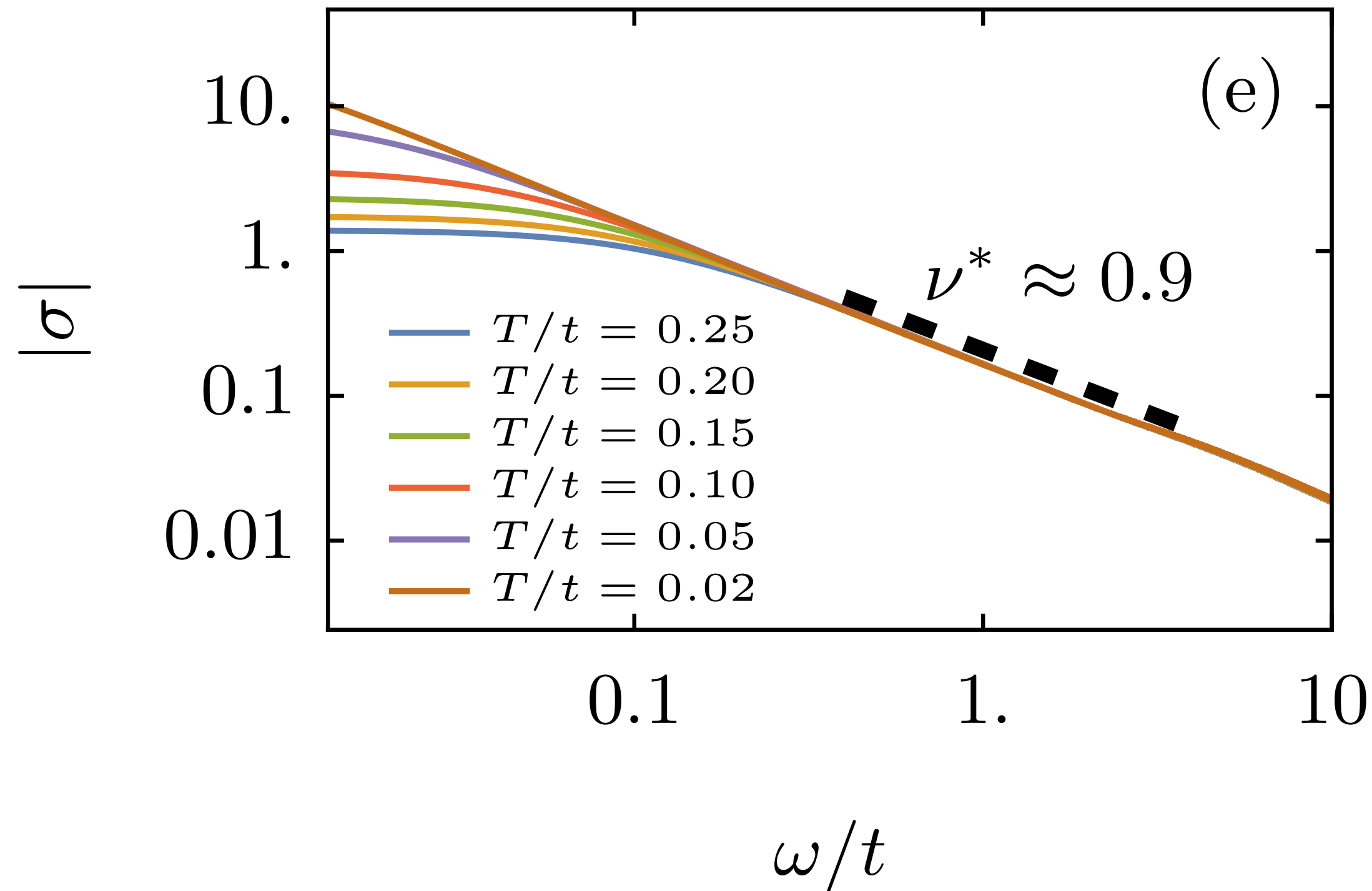


$\text{La}_{2-x}\text{Sr}_x\text{CuO}_4$
 $p = 0.24$
 $T_c = 19 \text{ K}$

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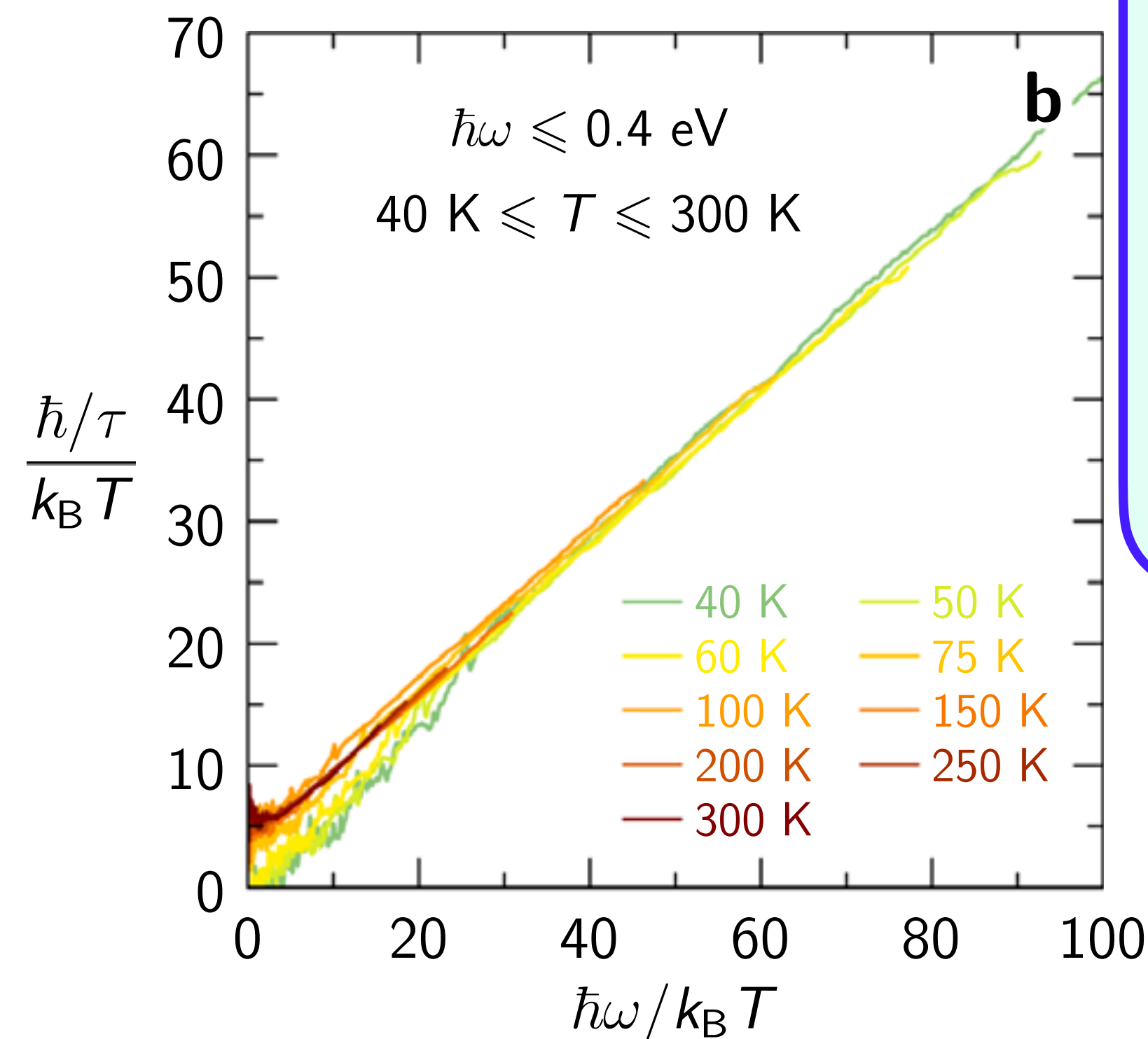
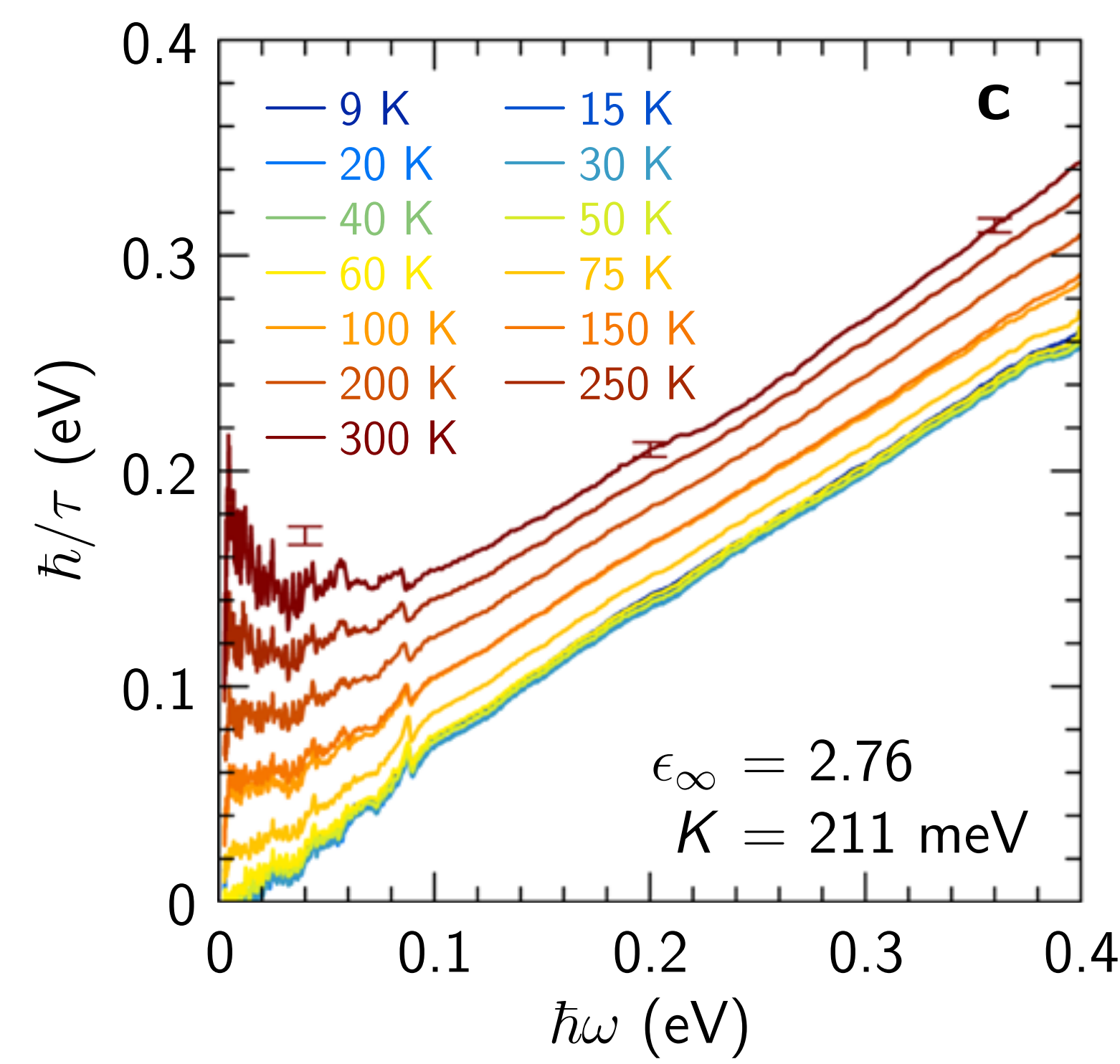


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Planckian dynamics!

$$\tau(\omega) = \frac{\hbar}{k_B T} F\left(\frac{\hbar\omega}{k_B T}\right)$$

and entropy

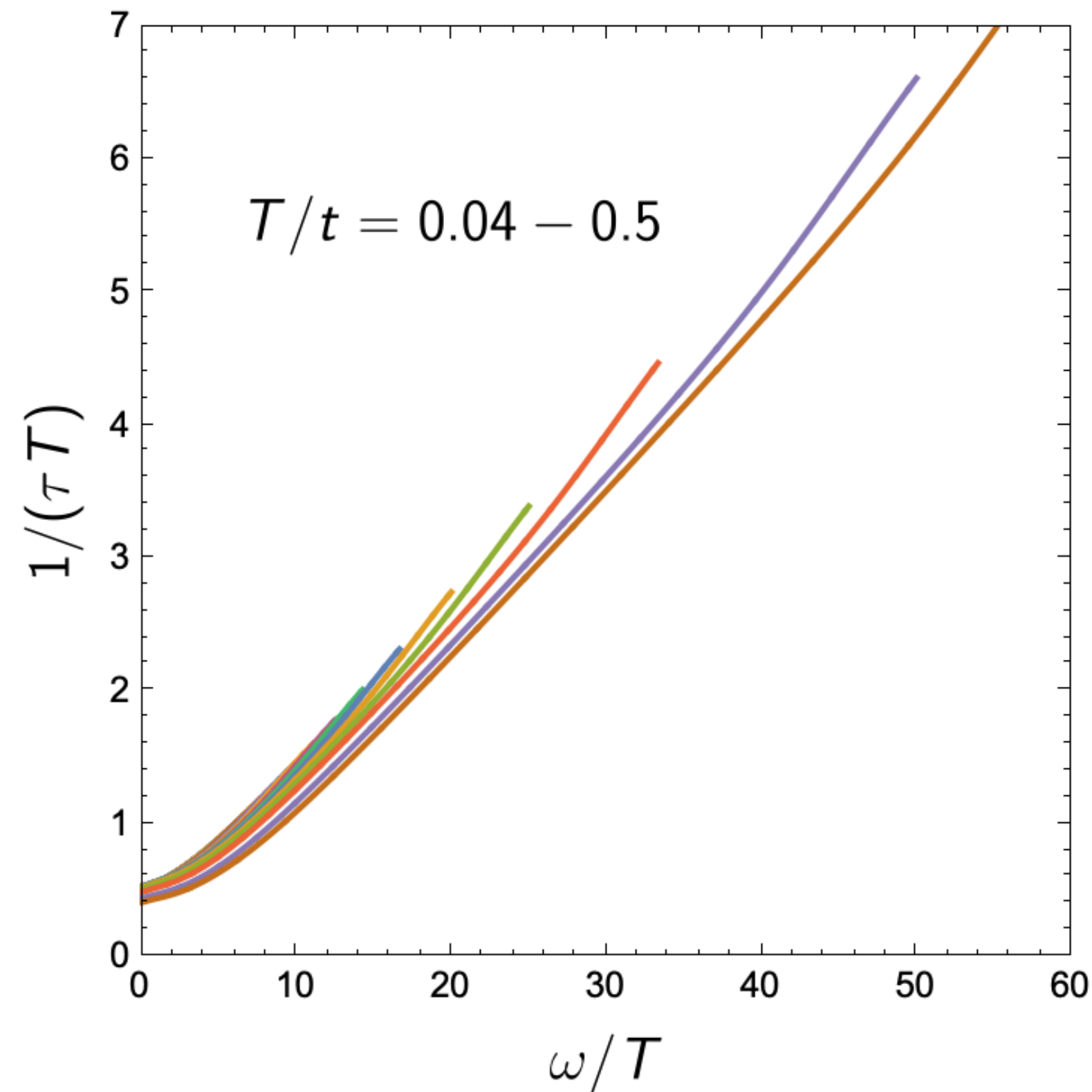
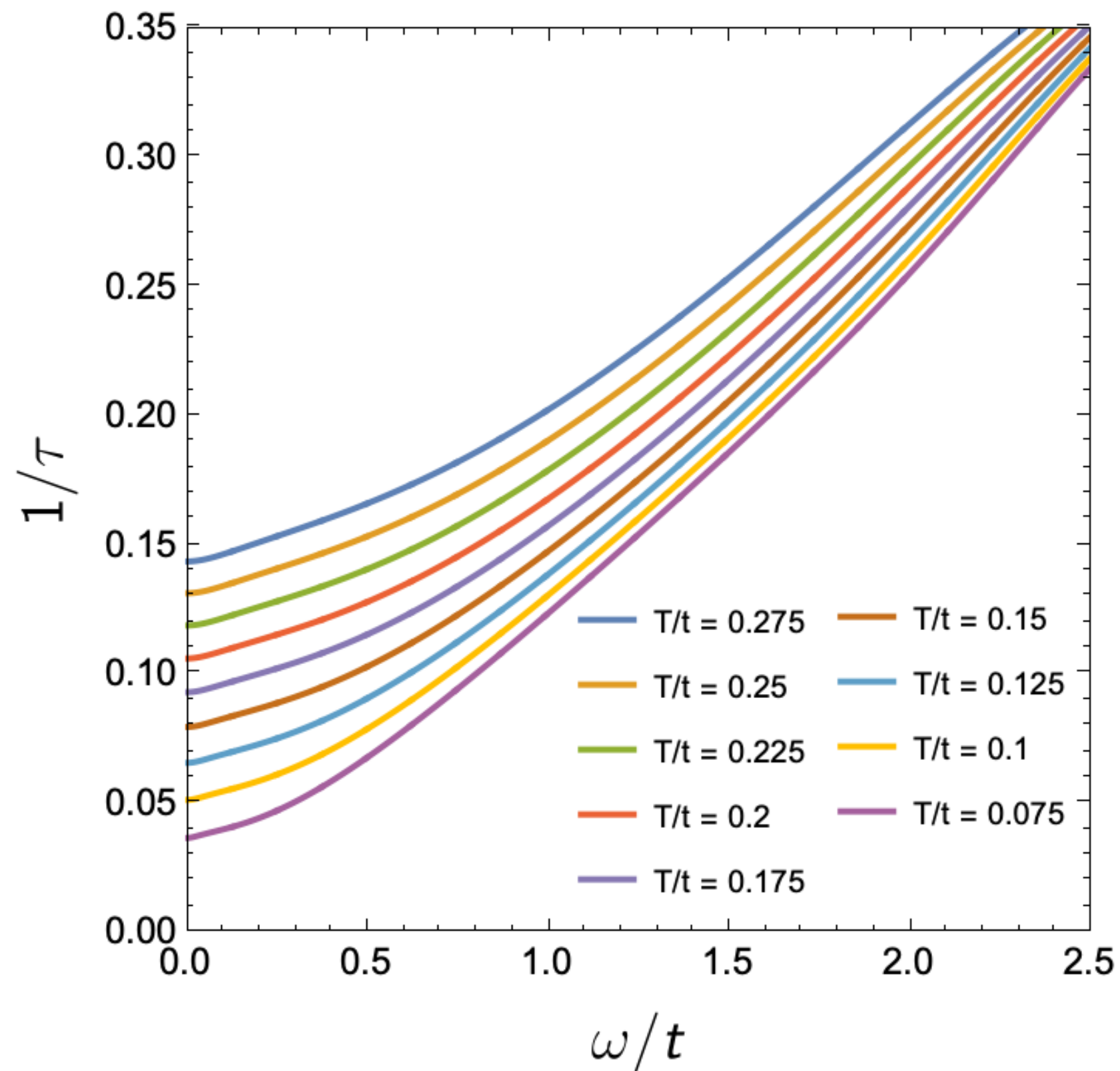
$$S(T \rightarrow 0) \sim T \ln(1/T).$$

$\text{La}_{2-x}\text{Sr}_x\text{CuO}_4$
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$$\tau(\omega) = \frac{\hbar}{k_B T} F \left(\frac{\hbar \omega}{k_B T} \right)$$

and entropy

$S(T \rightarrow 0) \sim T \ln(1/T)$
in 2d-YSYK model
(unlike zero temperature
entropy in SYK model).

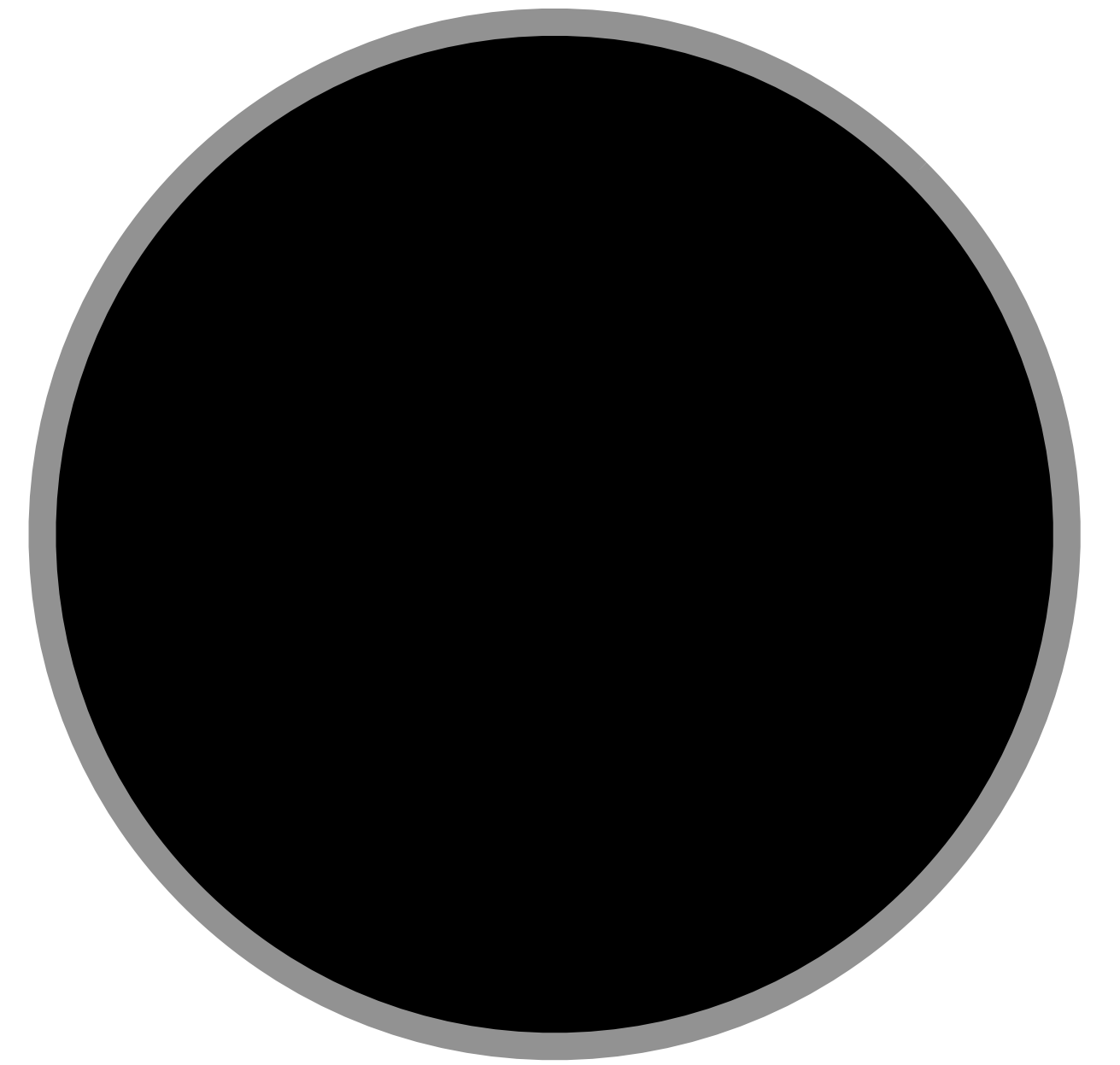
From the SYK model
to the
universal, low energy
(near-extremal)
density of quantum states
of charged and
rotating black holes

Black Holes

Objects so dense that light is gravitationally bound to them.



Horizon radius $R = \frac{2GM}{c^2}$

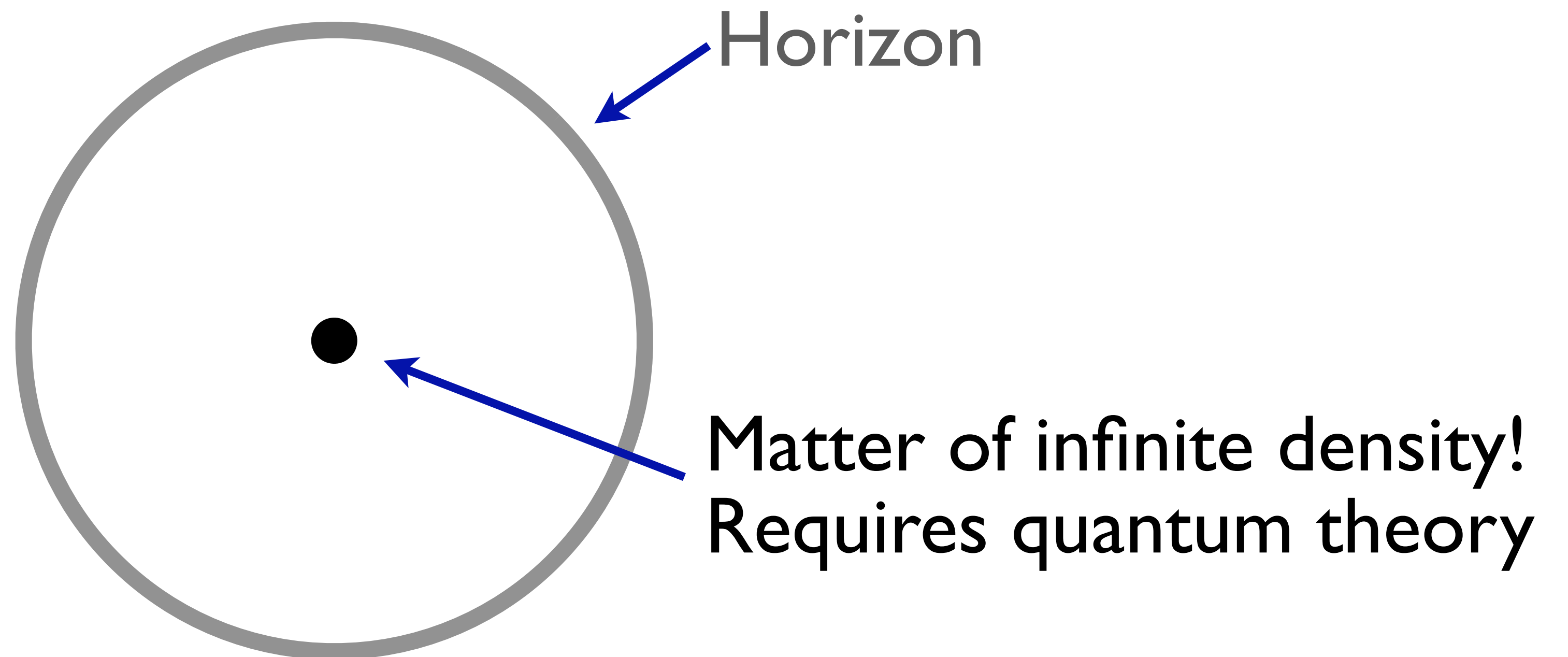


Karl Schwarzschild (1916)

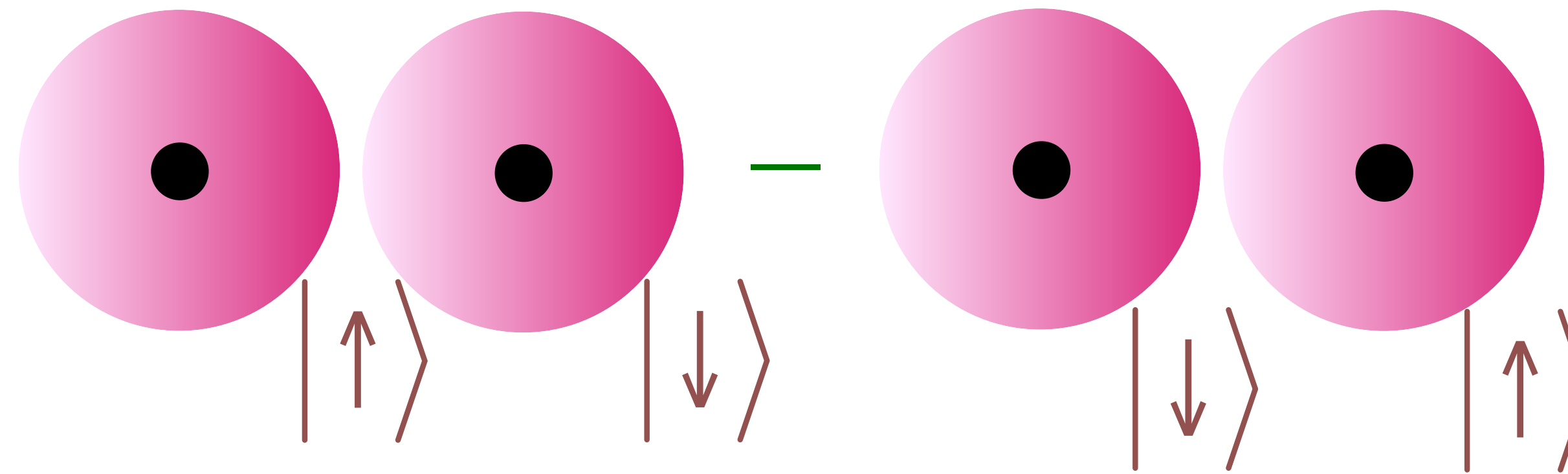
G Newton's constant, c velocity of light, M mass of black hole
For $M = \text{earth's mass}$, $R \approx 9 \text{ mm}$!

What is inside a black hole ???

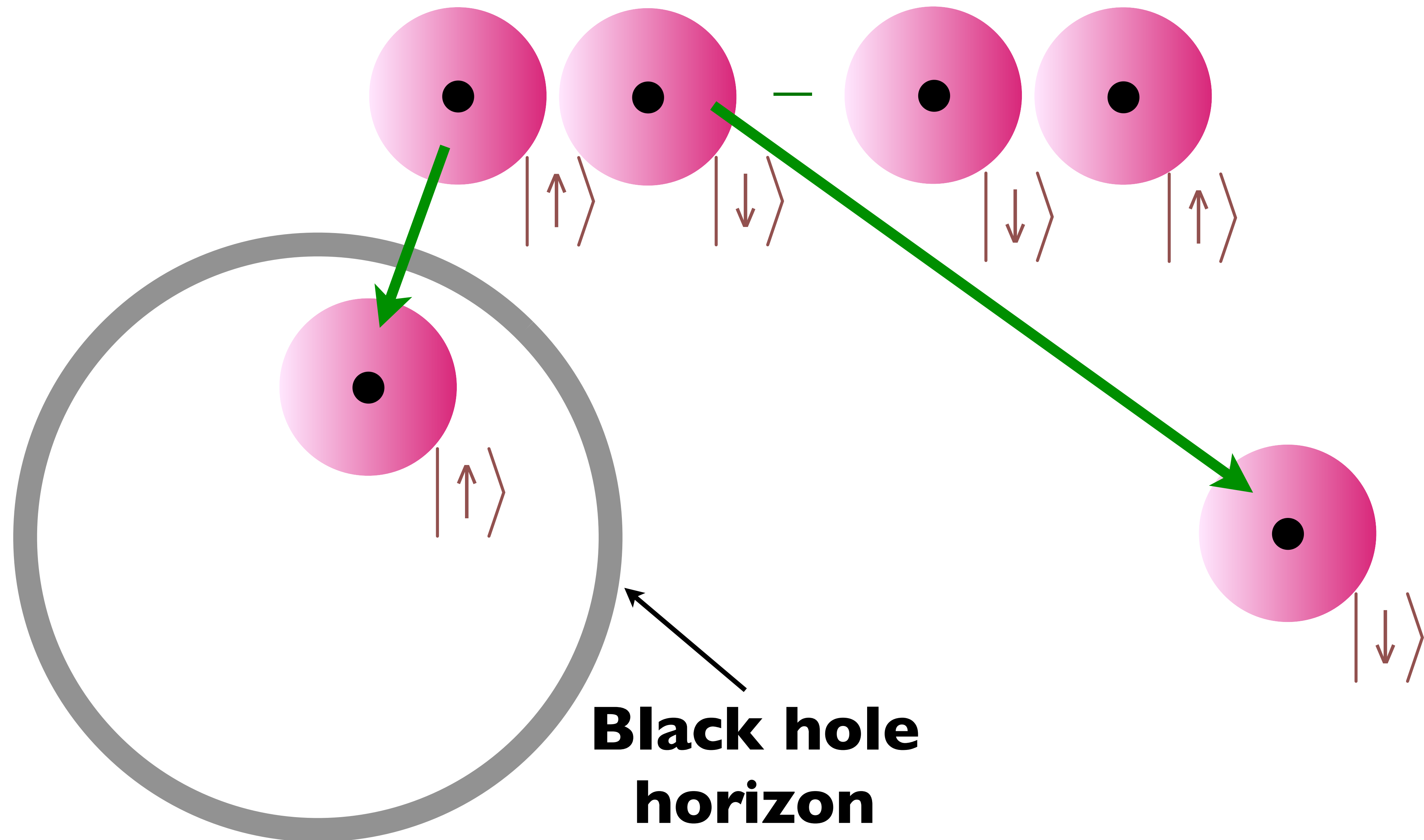
In Einstein's theory, all the matter in a black hole collapses to a singularity at the center of the black hole.



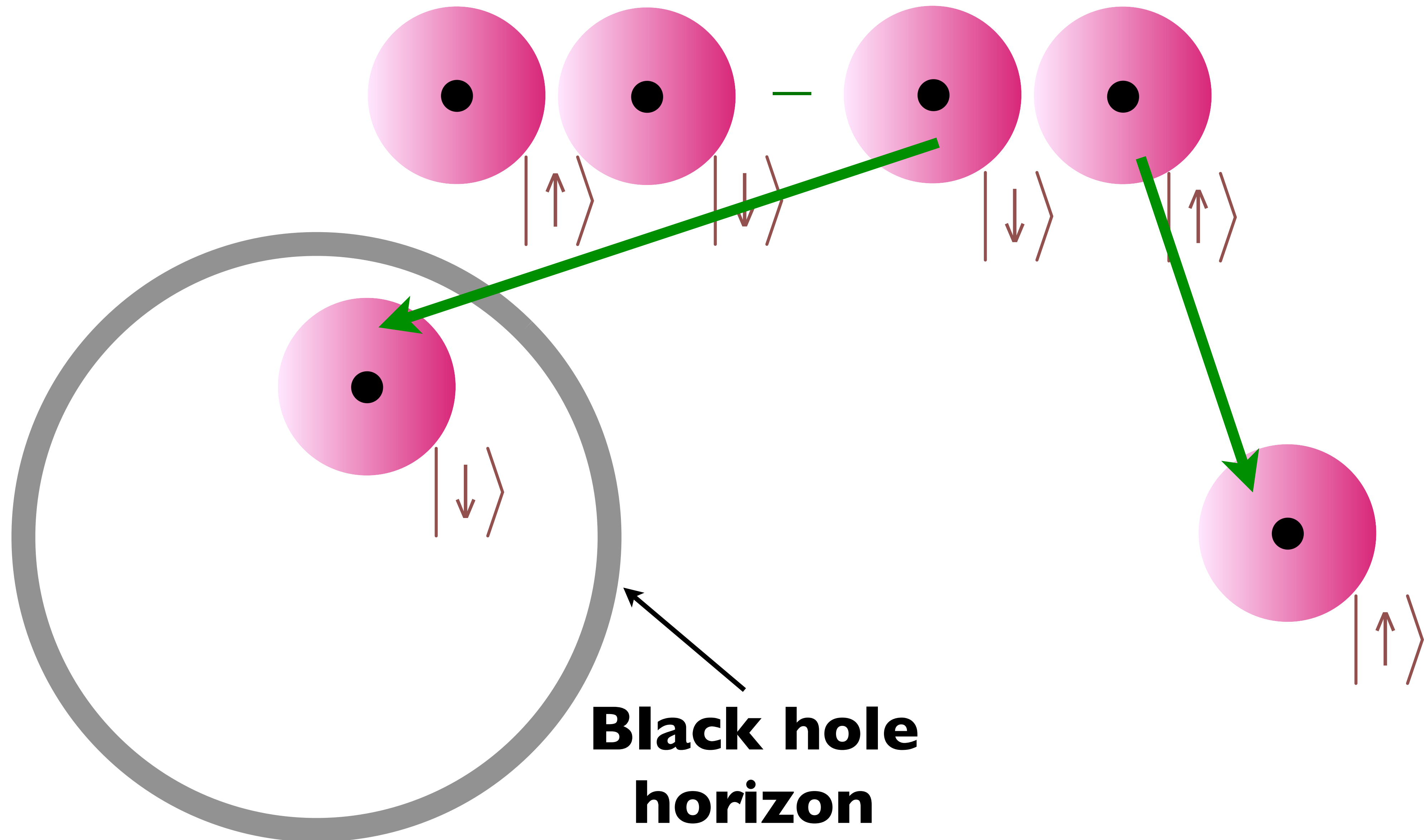
Quantum Entanglement across a black hole horizon



Quantum Entanglement across a black hole horizon



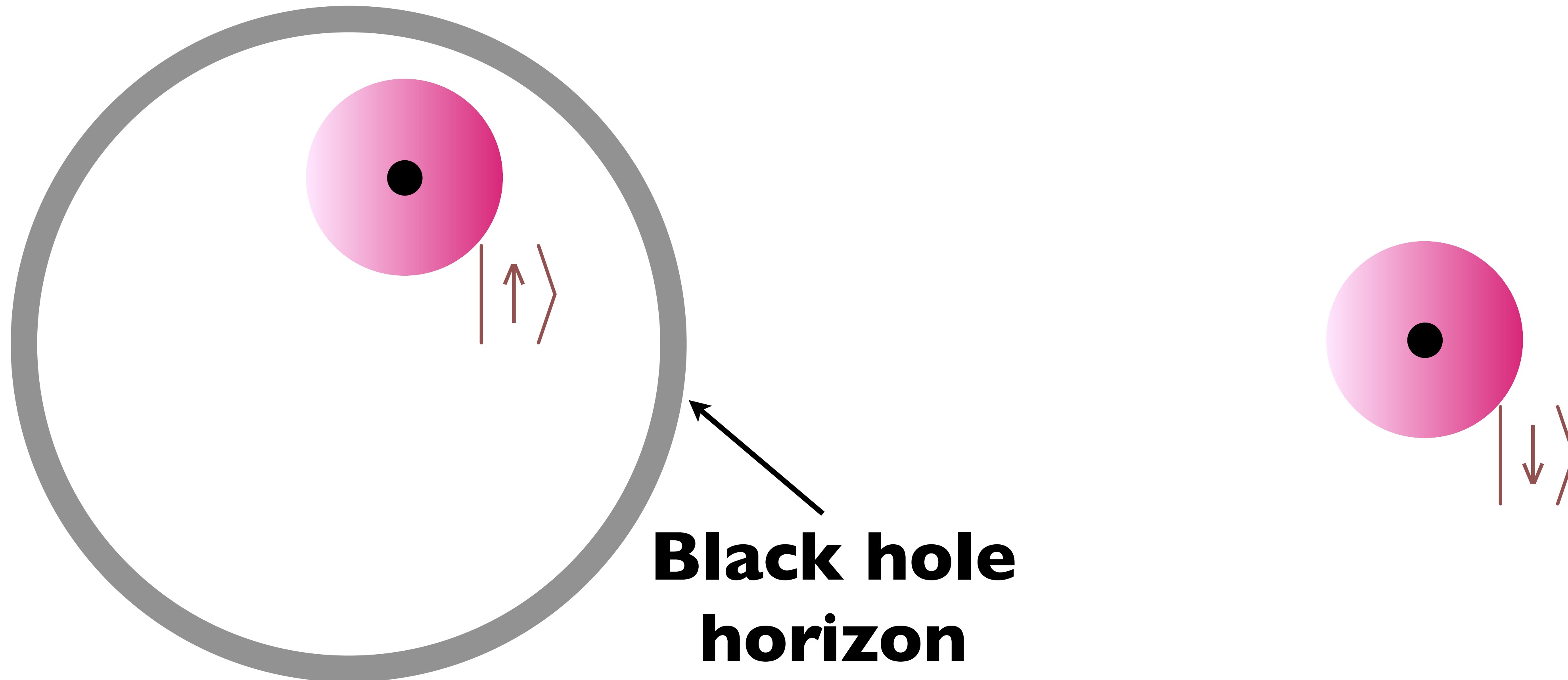
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Quantum Entanglement across a black hole horizon

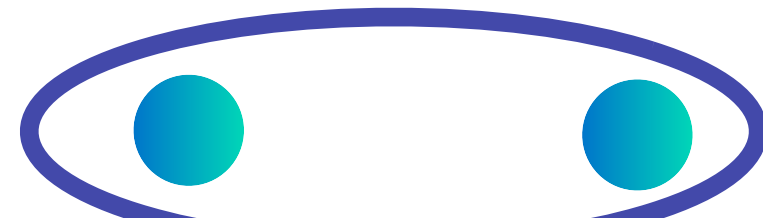
Bekenstein, Hawking: Black holes have a temperature and an entropy!

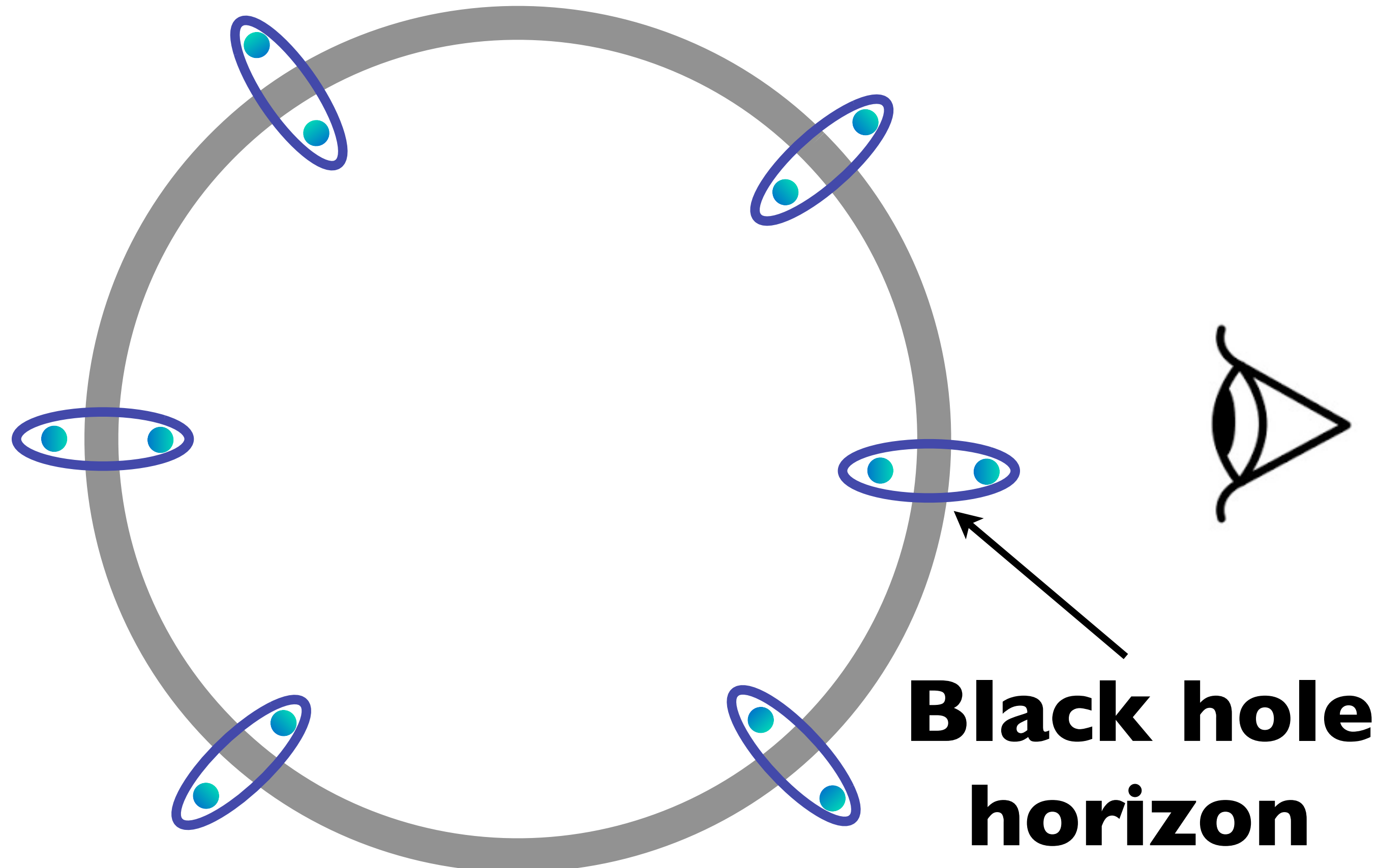
To an outside observer, the state of the electron inside the black hole cannot be known, and so the outside electron is in a random state.



Quantum Entanglement across a black hole horizon

Quantum entanglement
on the surface


$$= |\uparrow\downarrow\rangle - |\downarrow\uparrow\rangle$$



By computations *outside*
the black hole,
Bekenstein-Hawking obtained

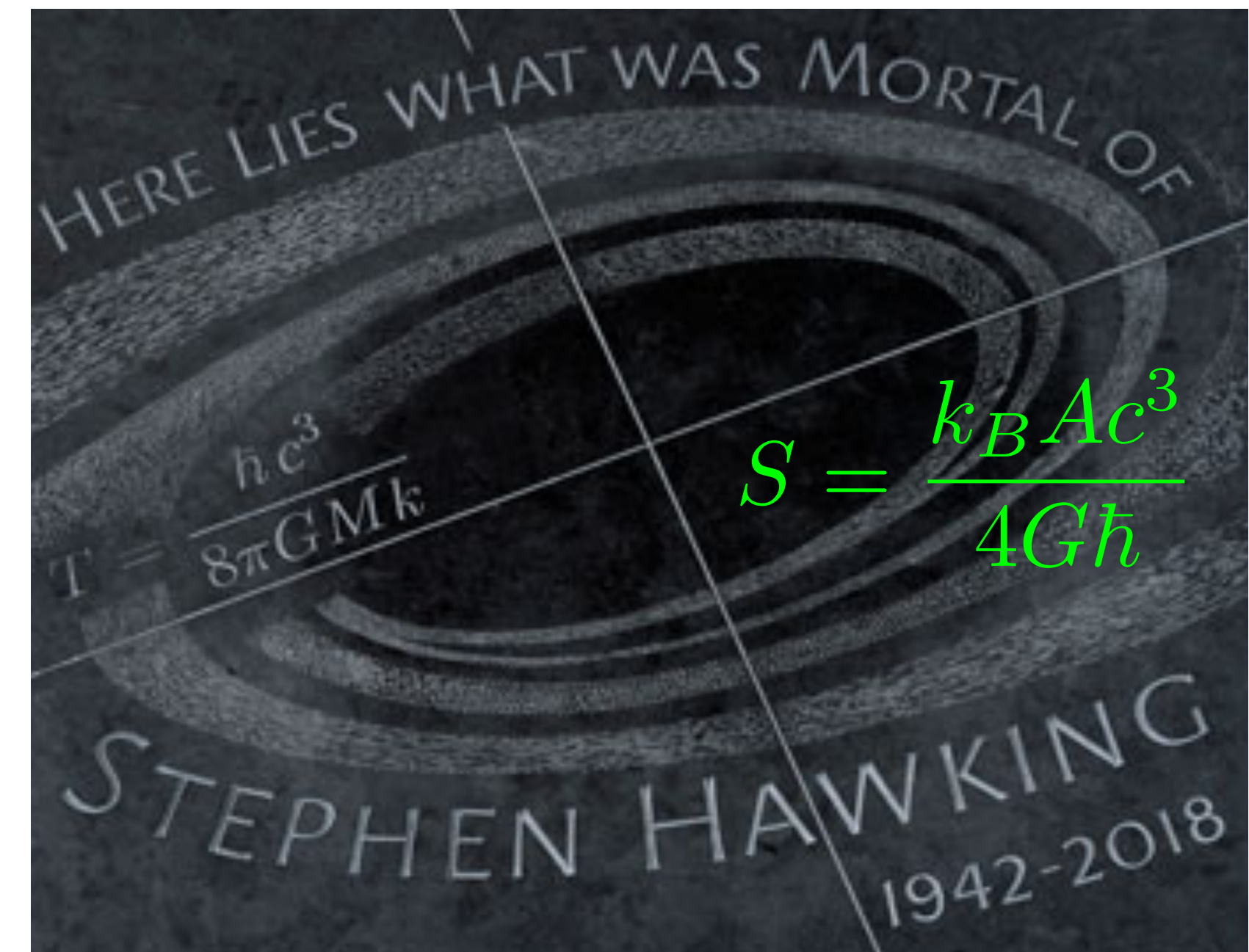
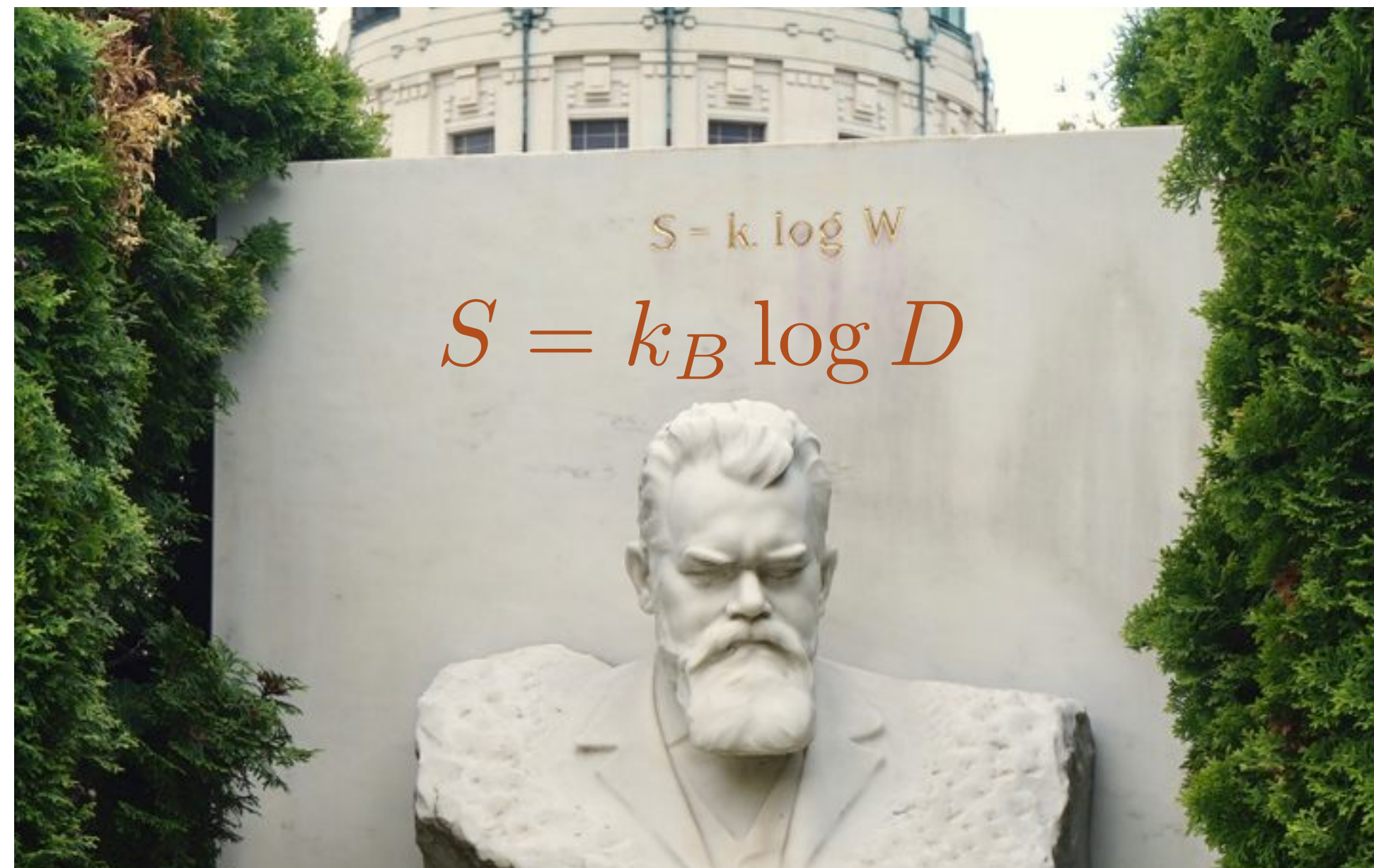
$$S = \frac{k_B A c^3}{4G\hbar}$$

where A is area of the black
hole horizon.

All other systems have en-
tropy proportional to their
volume.

Quantum Black Holes

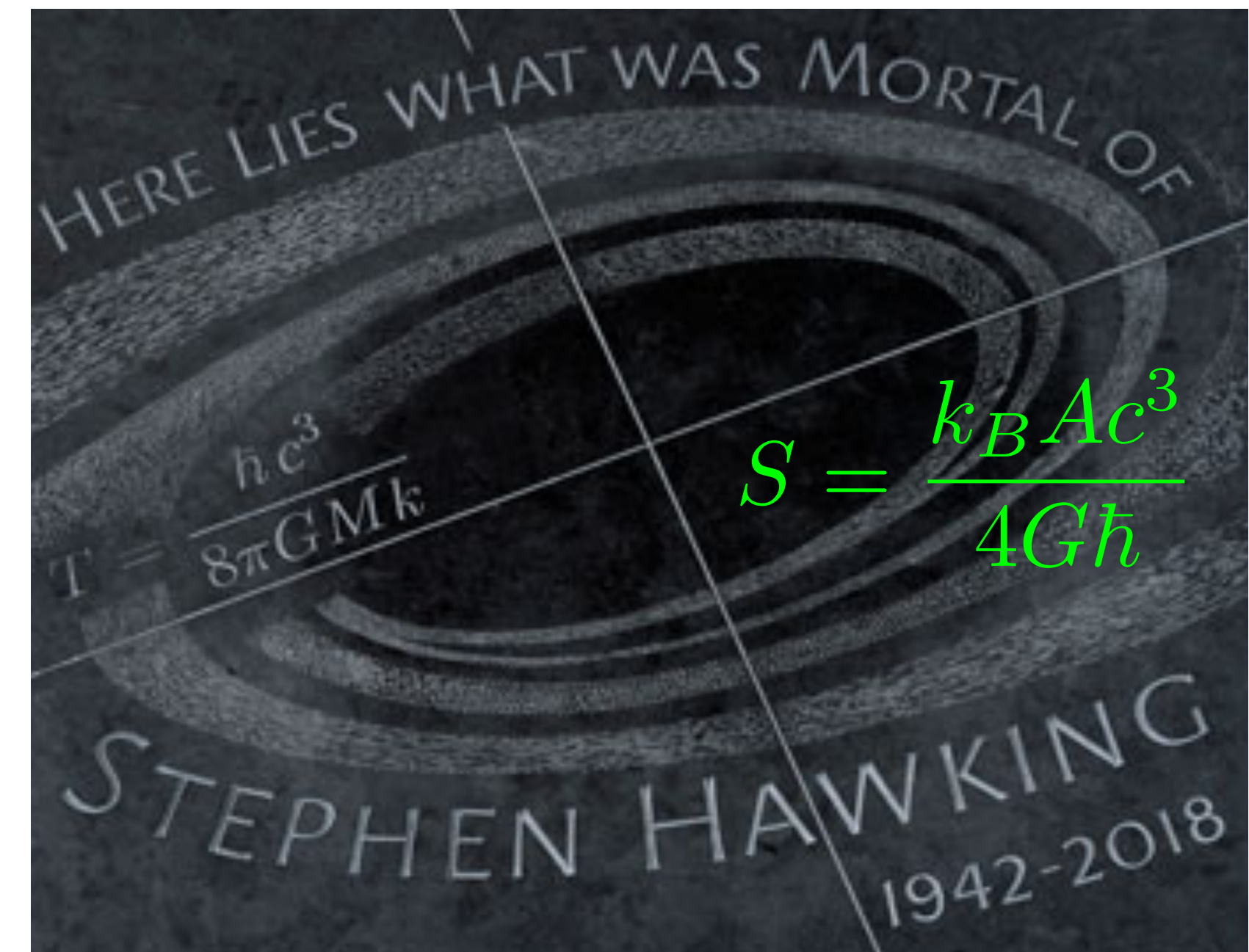
- Can we find a quantum theory for the collapsed matter at the center of the black hole, whose *density of quantum states* $D(E)$ [the quantum analog of Boltzmann's W] matches Bekenstein-Hawking entropy, in accordance with Boltzmann's principles of statistical mechanics, $S(E) = k_B \log D(E)$?



Quantum Black Holes

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For a black hole
with charge Q , the area
 $A_0 = 2GQ^2/c^4$ as $T \rightarrow 0$,
and so $S(T \rightarrow 0) > 0$.



Black Holes Obey Information-Emission Limits

April 22, 2021 • *Physics* 14, s47 –Christopher Crockett

G. Carullo, D. Laghi, J. Veitch, W. Del Pozzo, *Phys. Rev. Lett.* **126**, 161102 (2021)

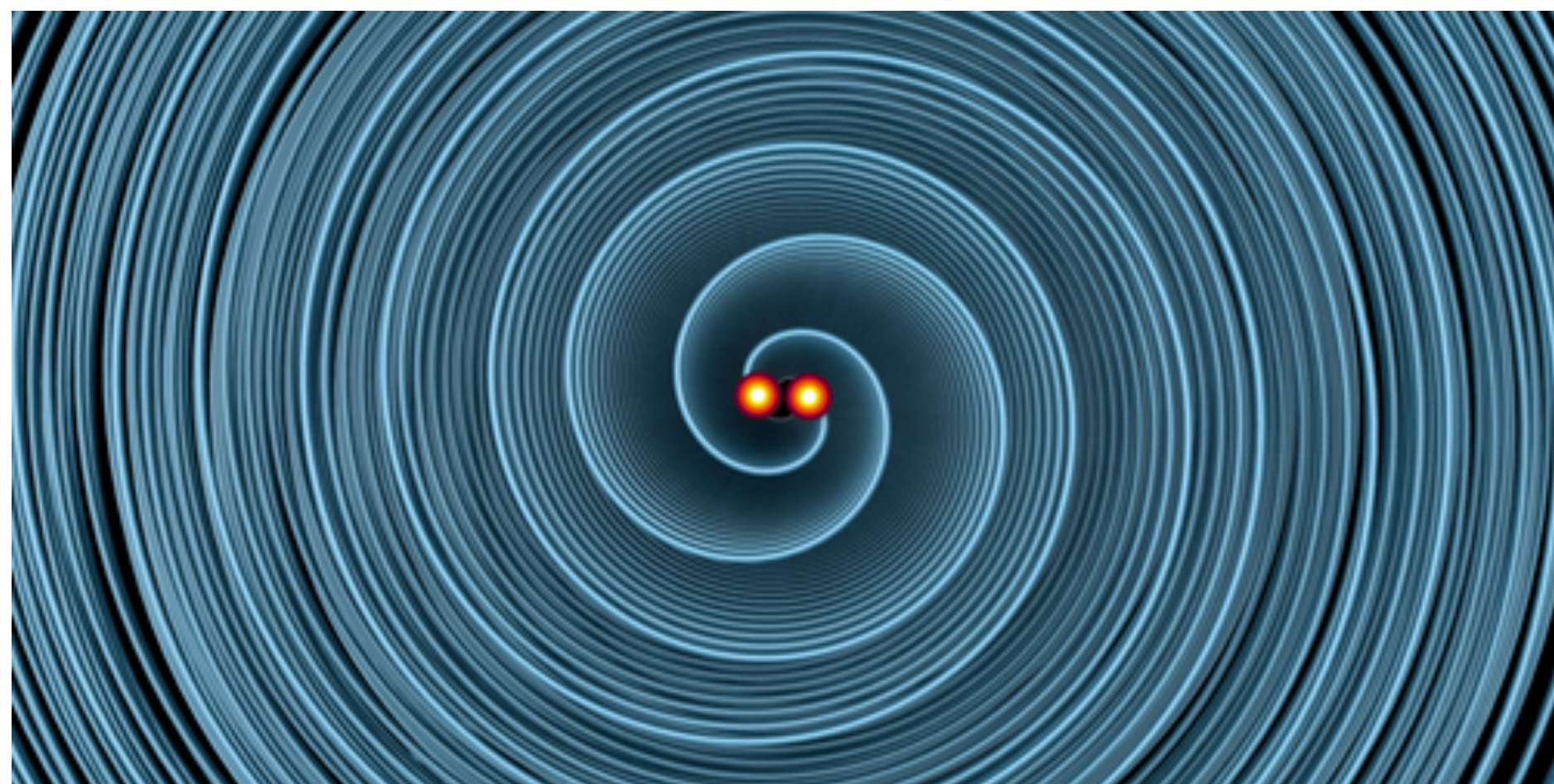
An analysis of the gravitational waves emitted from black hole mergers confirms that black holes are the fastest known information dissipaters.

Planckian dynamics!

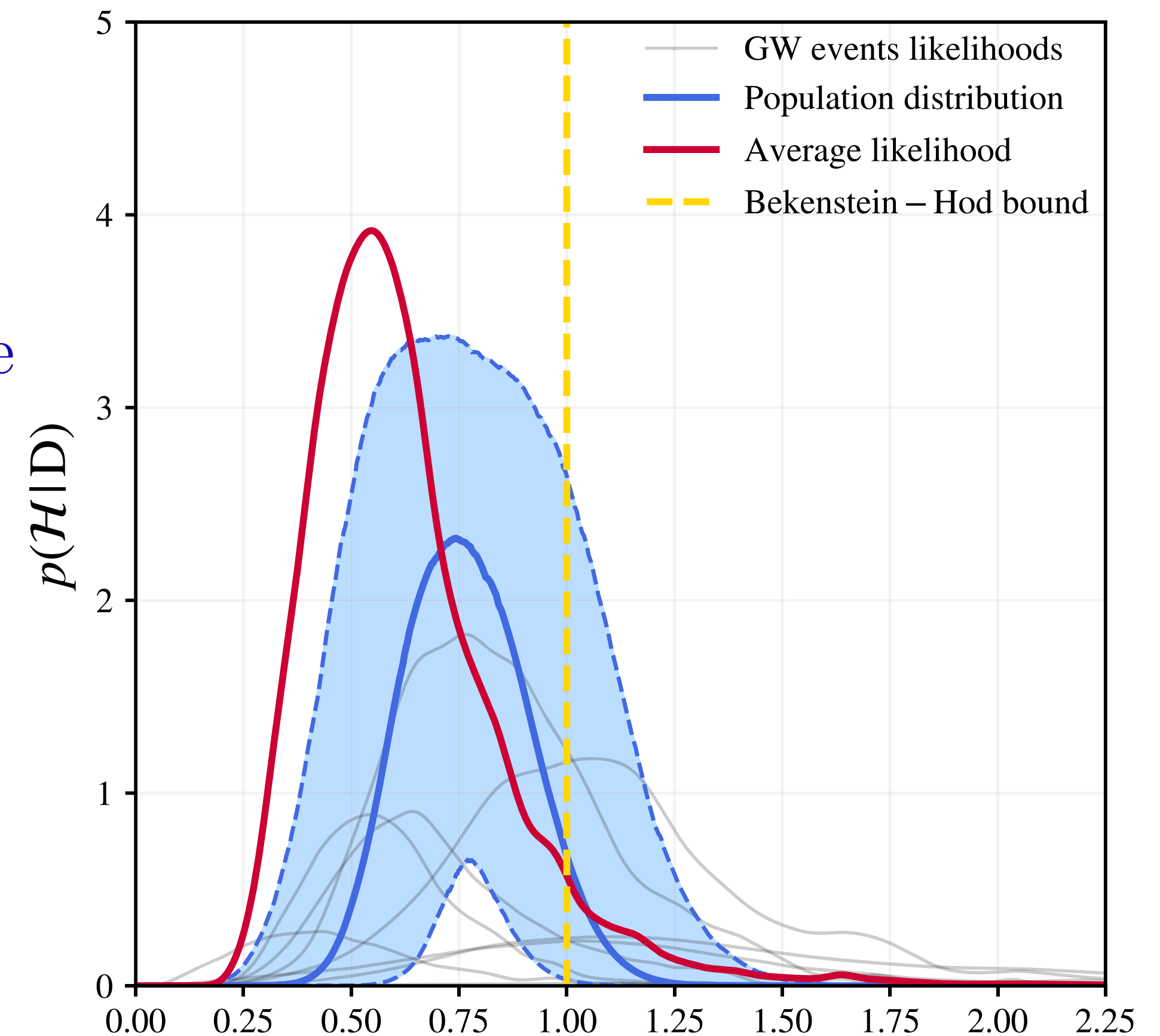
$$\tau(\omega) = \frac{\hbar}{k_B T} F\left(\frac{\hbar\omega}{k_B T}\right)$$

Gravity wave observations of 8 different black holes show a relaxation time

$$\tau \sim \frac{\hbar}{k_B T}$$



C.V. Vishveshwara, *Nature* **227**, 936 (1970)



$$\mathcal{H} = \frac{1}{\pi} \frac{\hbar/\tau}{k_B T}$$

Connections between the SYK model and black holes

- Planckian time $\sim \hbar/(k_B T)$ relaxation dynamics ('chaos').

Connections between the SYK model and black holes

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- Charged black holes have a non-zero Bekenstein-Hawking entropy in the limit $T \rightarrow 0$:

$S_{BH} = A_0 c^3 / (4\hbar G)$ where $A_0 = 2G Q^2 / c^4$ is the area of the charged black hole horizon at $T = 0$.

This matches the $T \rightarrow 0$ entropy $N s_0$, of the SYK model. Similar remarks apply to rotating neutral black holes.

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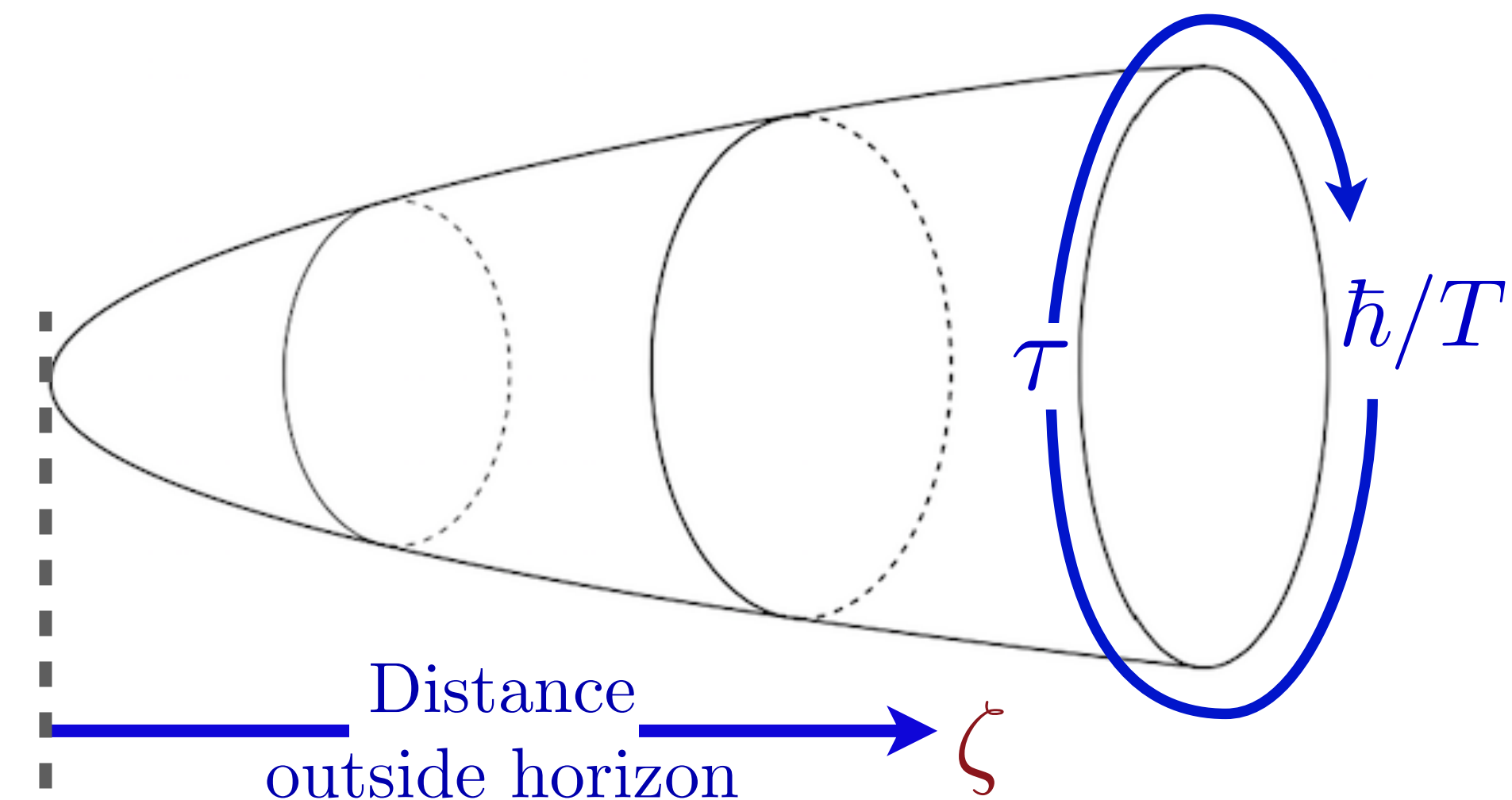
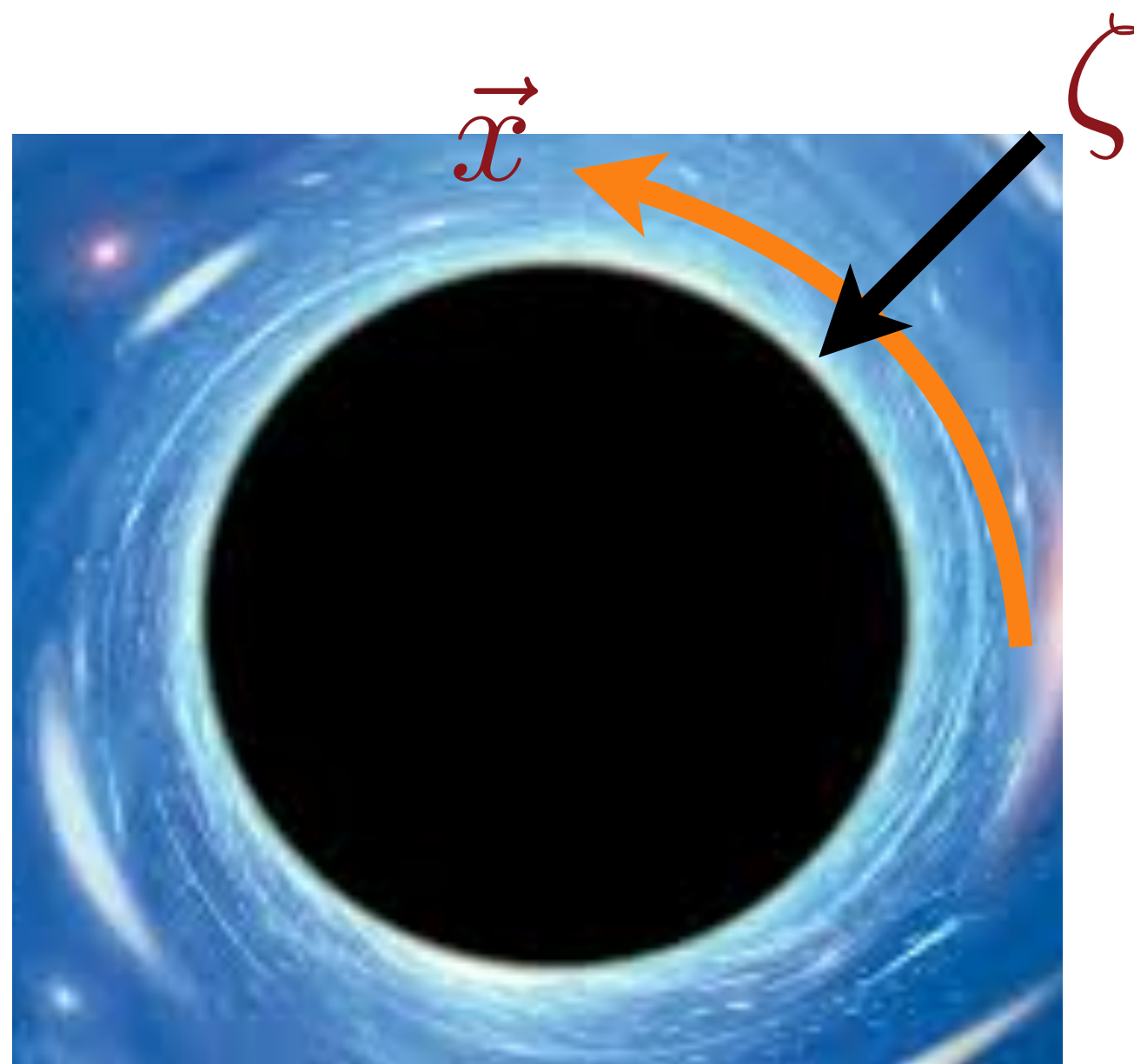
- This connection shows that S_{BH} is *not* realized by an exponentially large ground state degeneracy (as is the case in all earlier string-theoretic computations).

Thermodynamics of quantum black holes with charge Q :



$$\mathcal{Z}(Q, T) = \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_\mu \exp \left(-\frac{1}{\hbar} I_{\text{Einstein gravity} + \text{Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_\mu] \right)$$

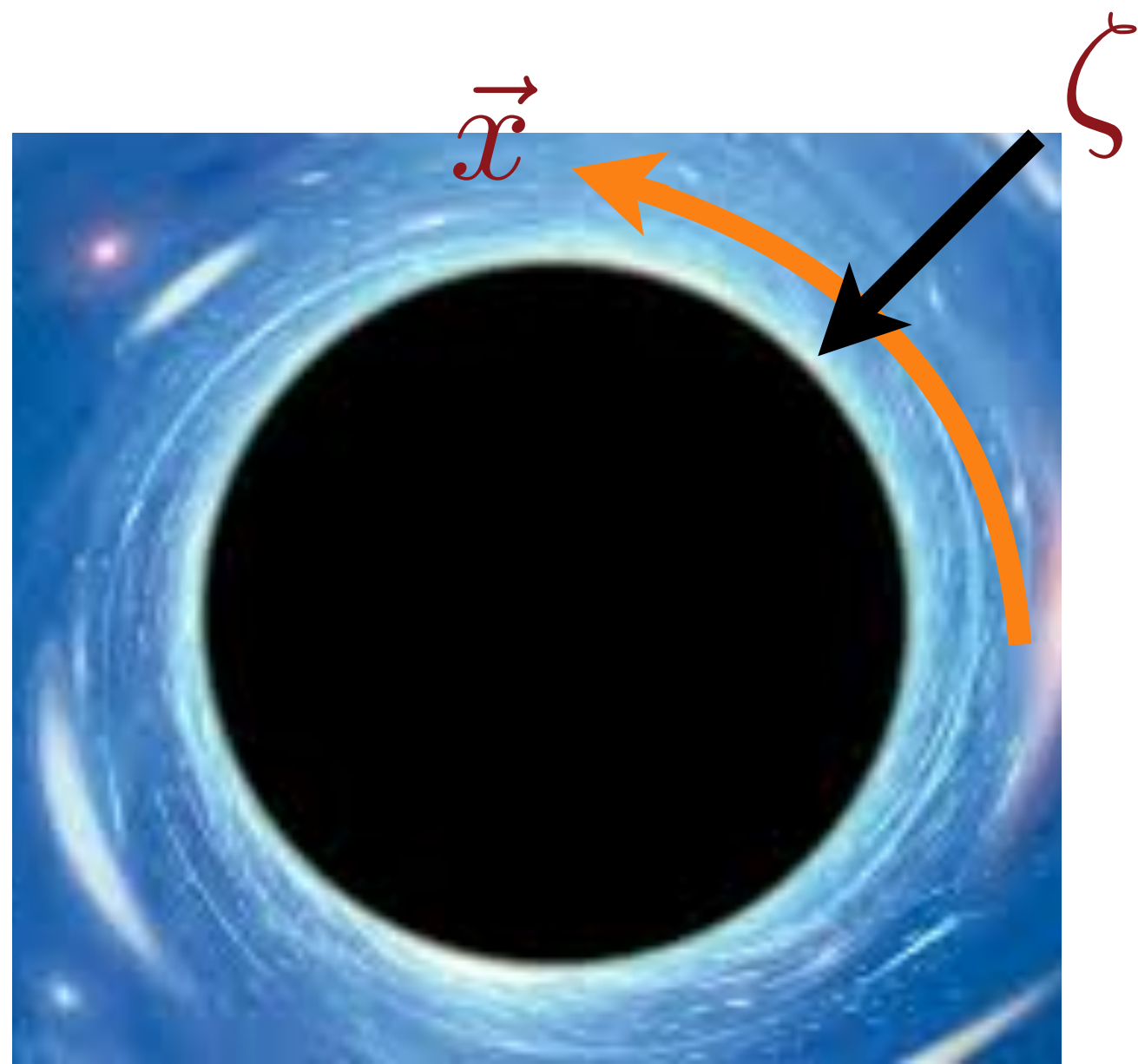
A. Chamblin, R. Emparan,
C.V. Johnson, and R.C. Myers,
PRD **60**, 064018 (1999)



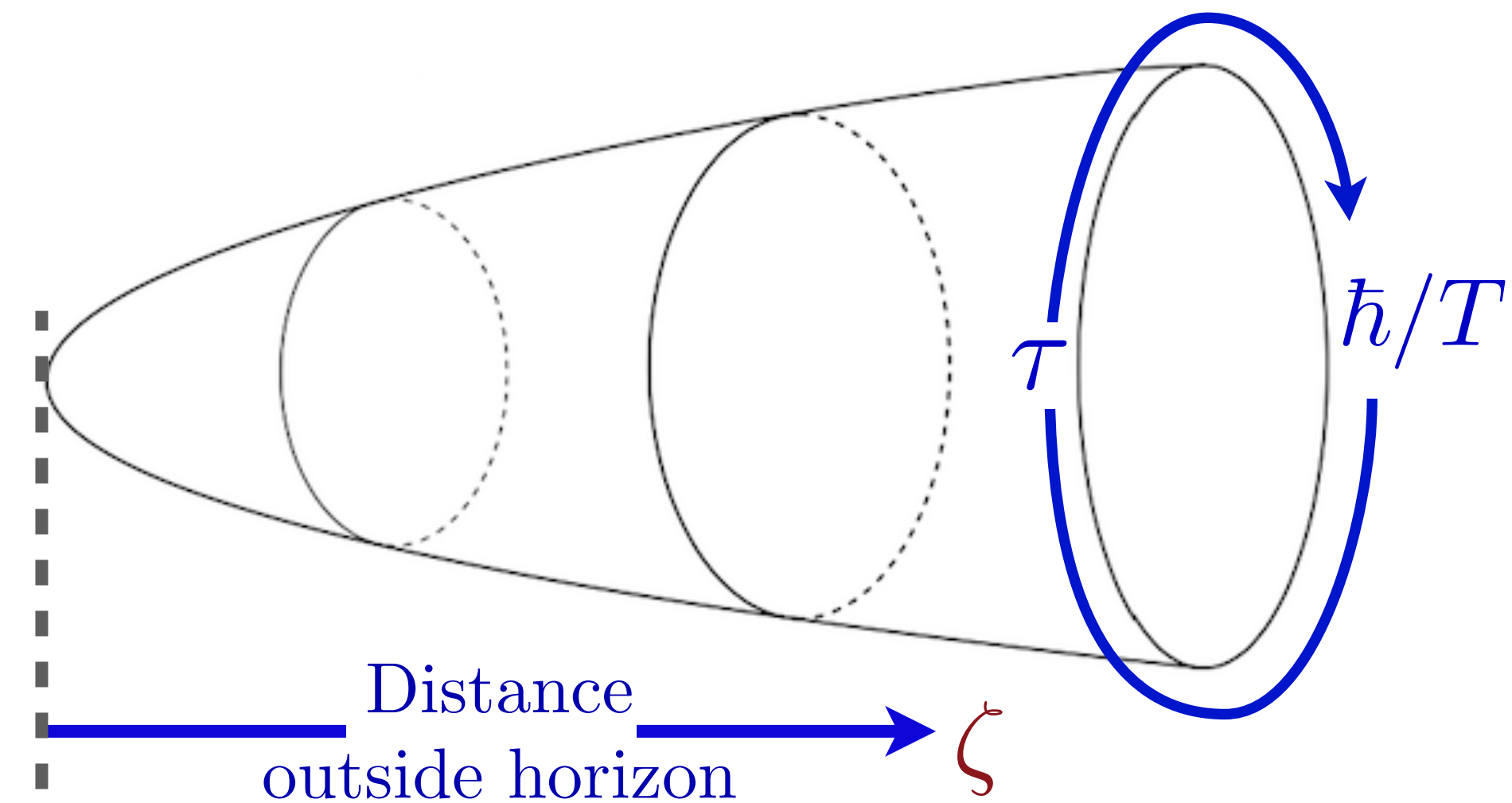
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$$\approx \exp \left(\frac{A_0 c^3}{4\hbar G} \right) \quad \text{as } T \rightarrow 0$$

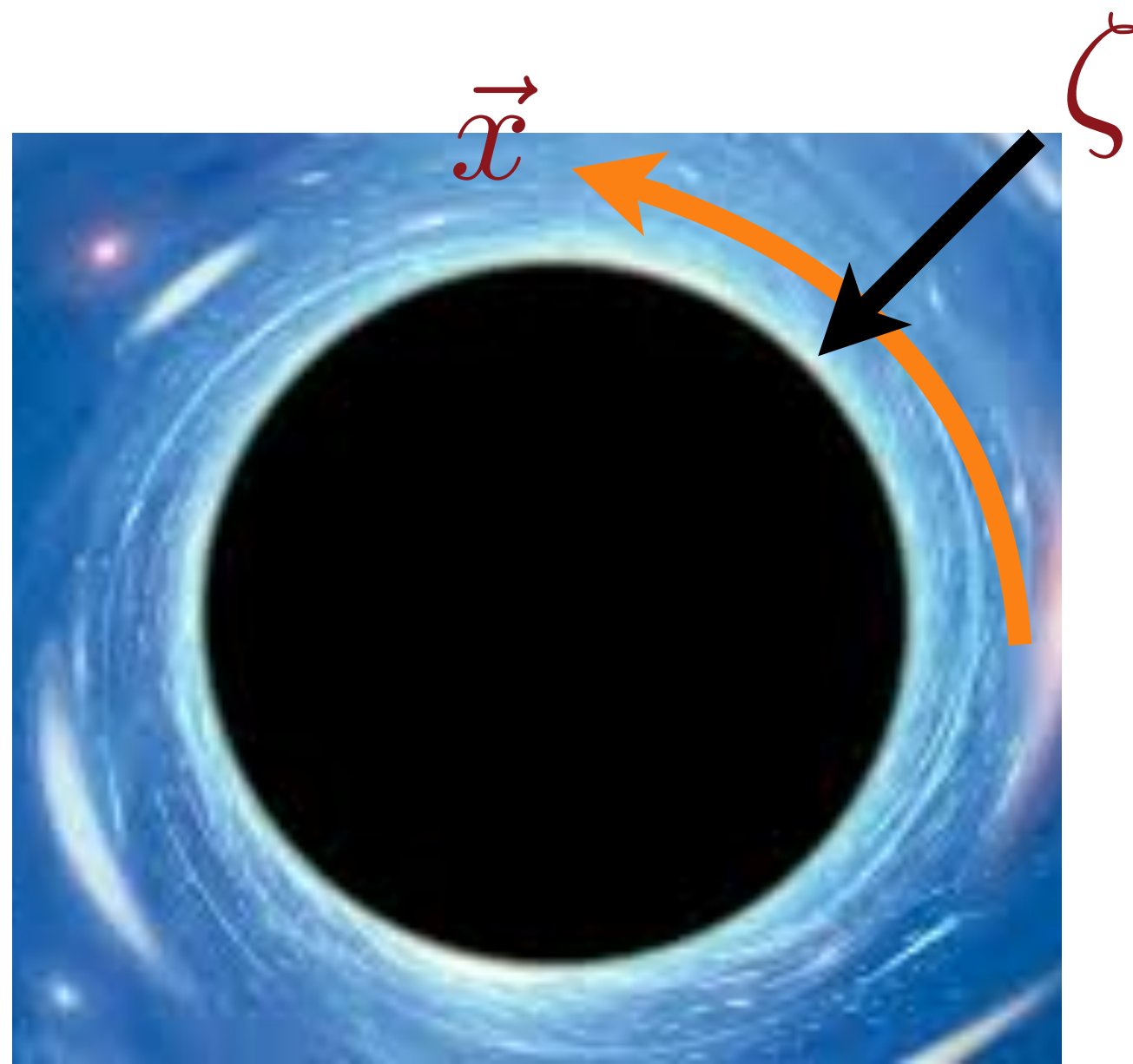


G.W. Gibbons and S.W. Hawking, PRD **15**, 2572 (1977)

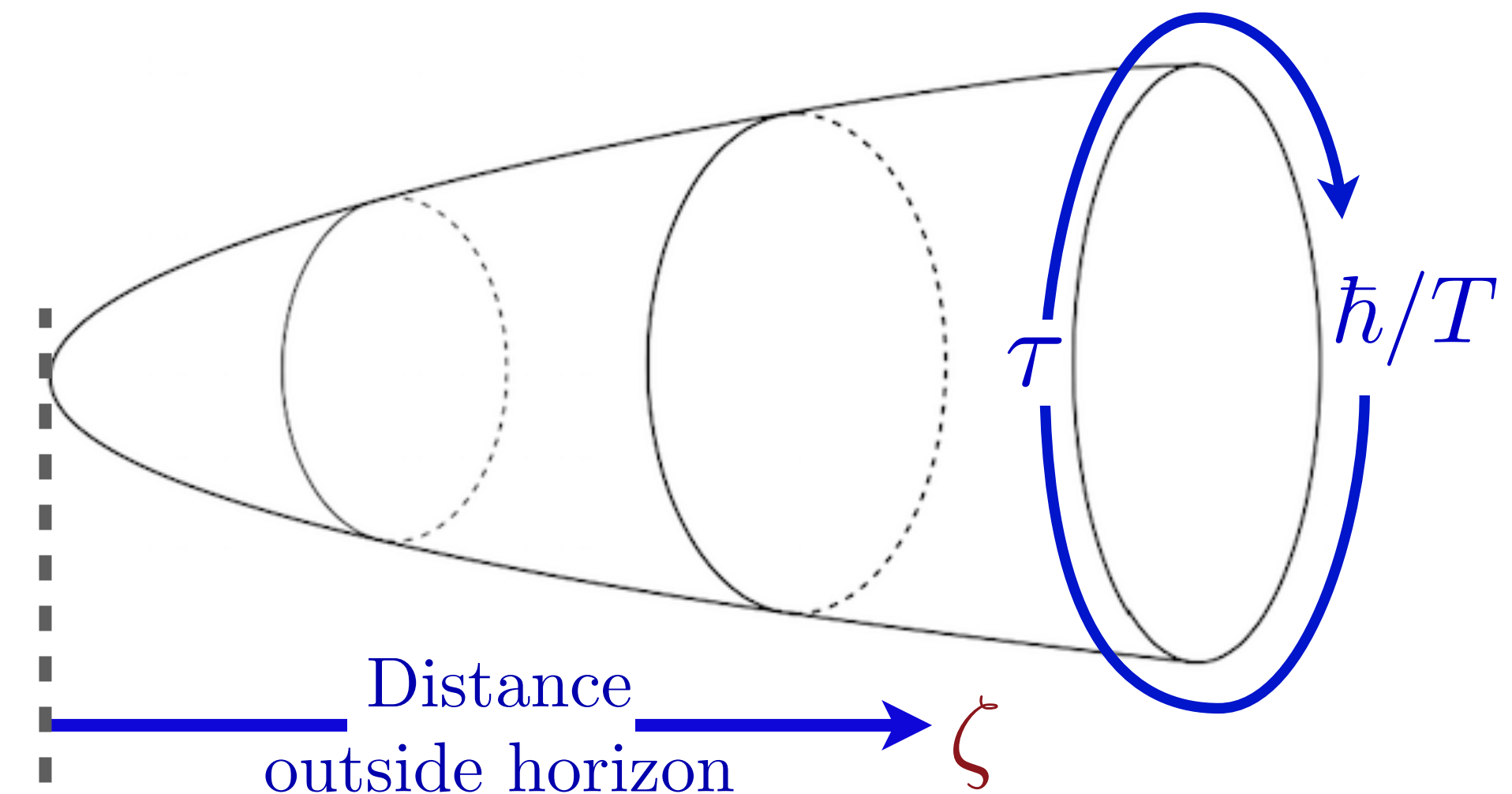


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Kitaev (2015); Maldacena, Stanford, Yang (2016)

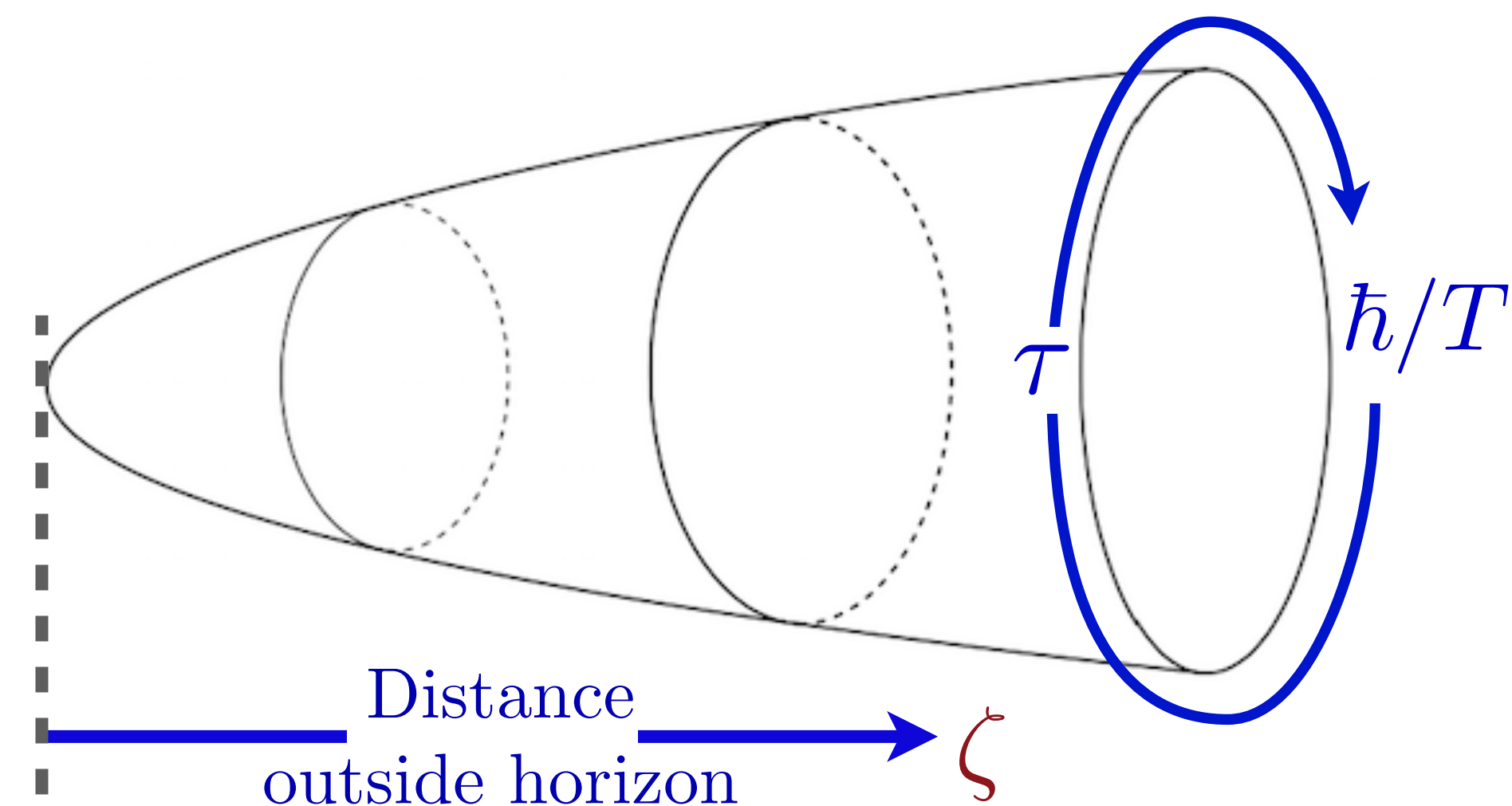
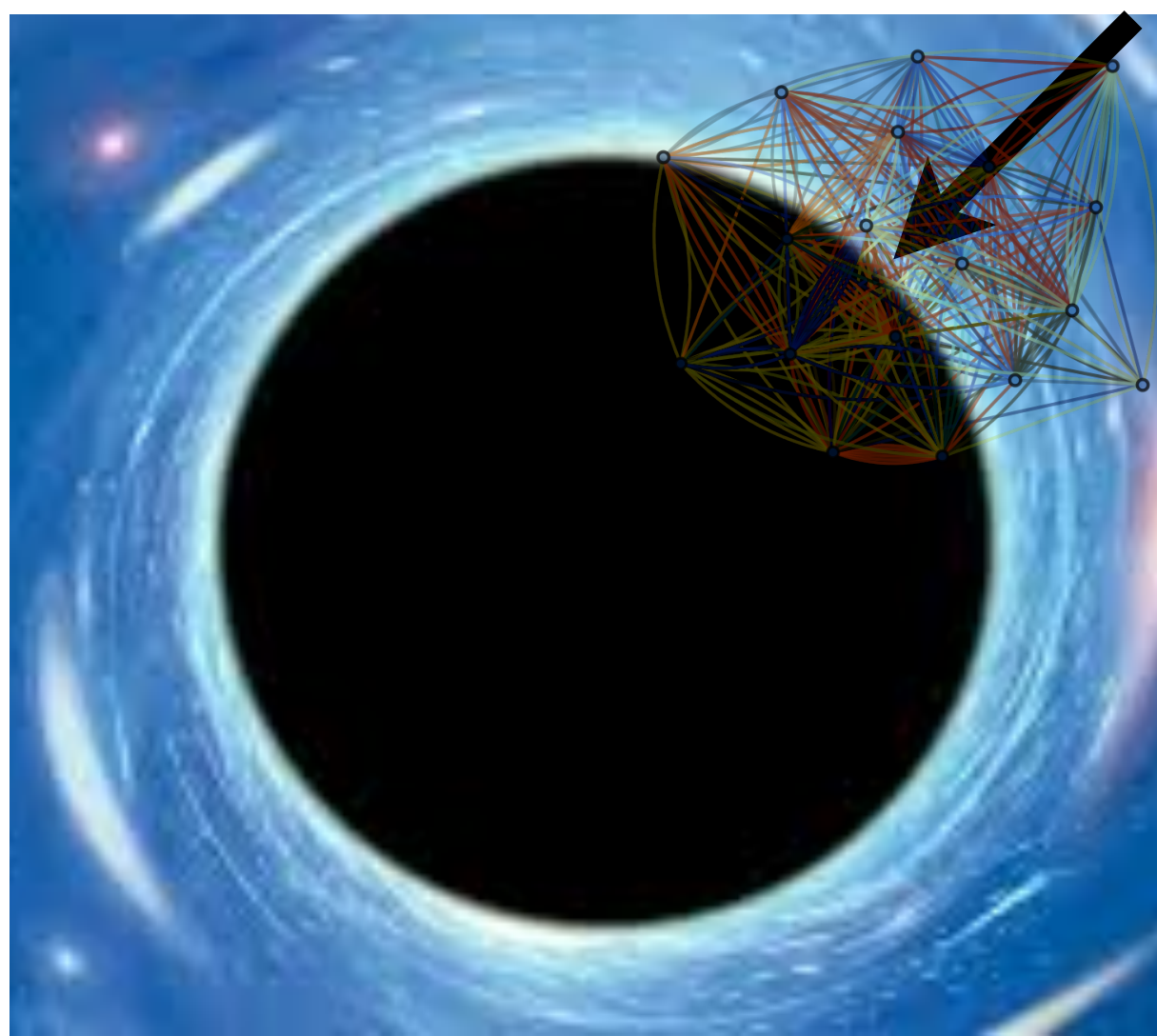


Thermodynamics of quantum black holes with charge \mathcal{Q} :

$$\mathcal{Z}(\mathcal{Q}, T) = \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_\mu \exp \left(-\frac{1}{\hbar} I_{\text{Einstein gravity} + \text{Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_\mu] \right) \\ \approx \exp \left(\frac{A_0 c^3}{4\hbar G} \right) \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_\mu \exp \left(-\frac{1}{\hbar} I_{\text{JT gravity of AdS}_2 + \text{boundary graviton}}^{(1+1)}[g_{\mu\nu}, A_\mu] \right)$$

Holography: quantum entanglement on the surface

$$= \int \mathcal{D}f(\tau) \mathcal{D}\phi(\tau) \exp \left(-\frac{1}{\hbar} I_{\text{SYK}}^{(0+1)}[\text{time reparameterizations } f(\tau), \text{phase rotations } \phi(\tau)] \right)$$



Kitaev (2015); Maldacena, Stanford, Yang (2016); Cotler et al. (2017)

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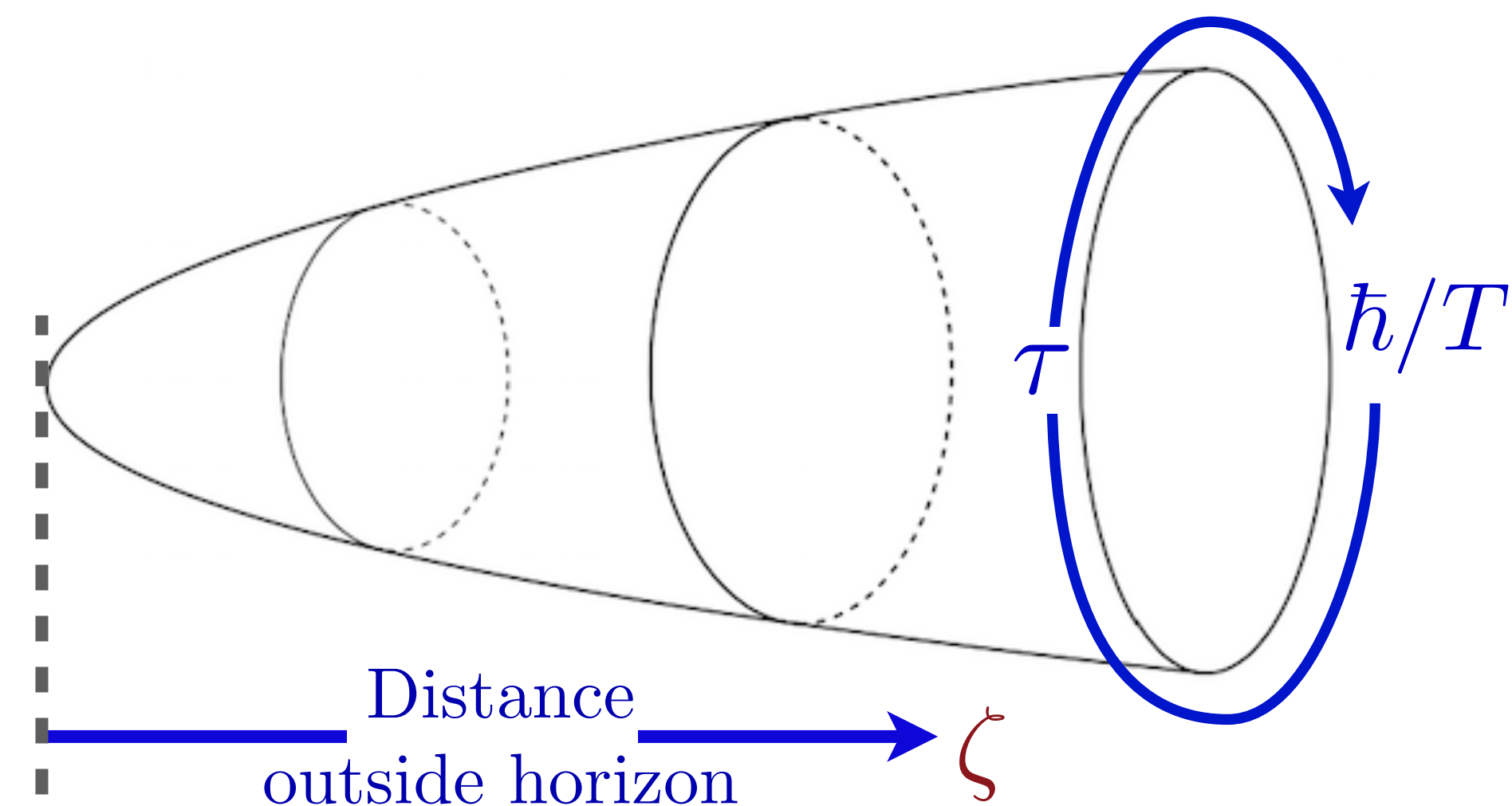
$$= \int \mathcal{D}f(\tau) \mathcal{D}\phi(\tau) \exp \left(-\frac{1}{\hbar} I_{\text{SYK}}^{(0+1)}[\text{time reparameterizations } f(\tau), \text{phase rotations } \phi(\tau)] \right)$$

The path integral over the action $I_{\text{SYK}}^{(0+1)}$ can be evaluated exactly,

and leads to a computation of $D(E)$

$$\mathcal{Z}(\mathcal{Q}, T) = \int dE D(E) \exp \left(-\frac{E}{k_B T} \right)$$

Kitaev (2015); Maldacena, Stanford, Yang (2016); Cotler et al. (2017)



D(E) of charged black holes from the SYK model

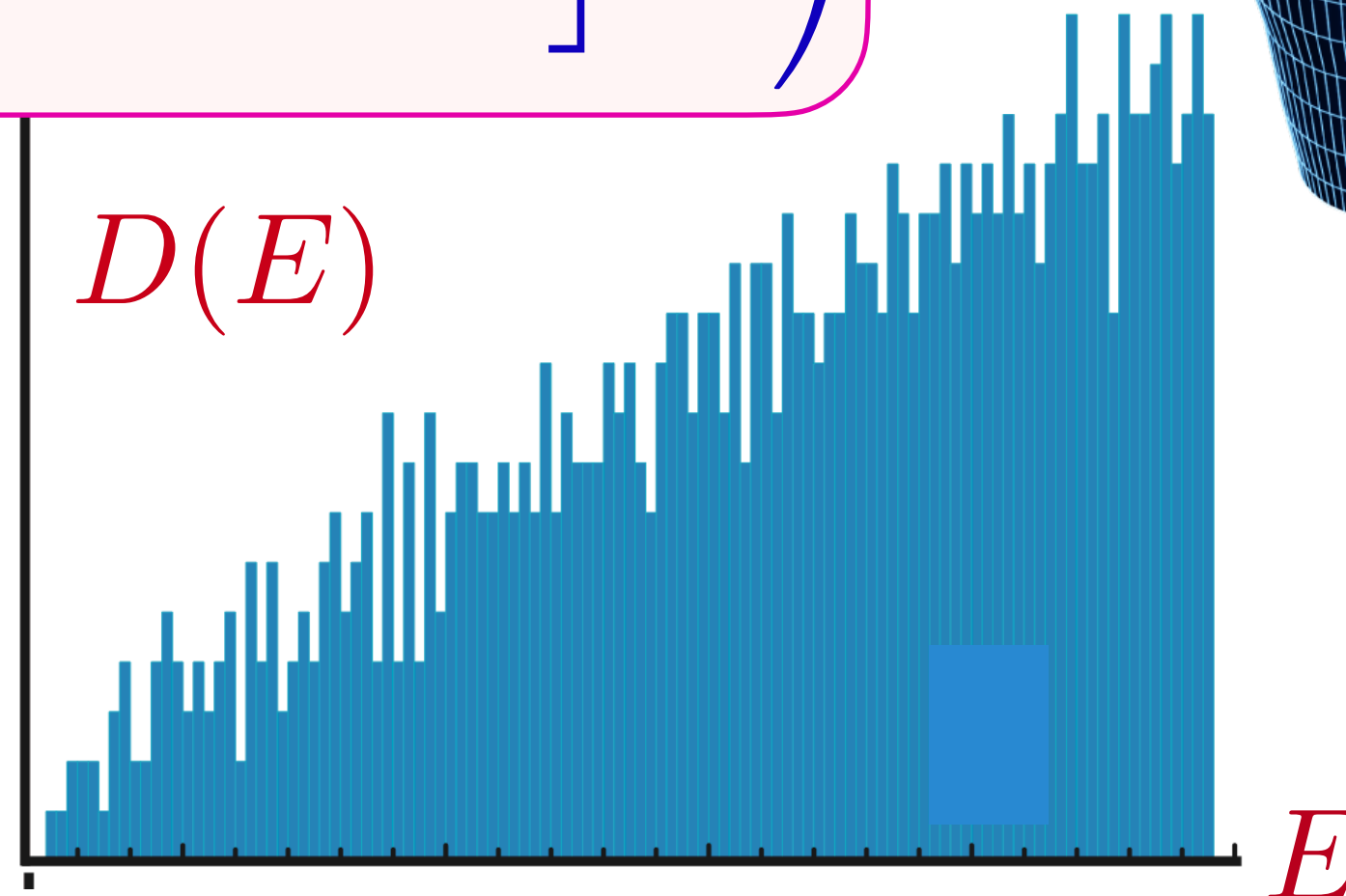
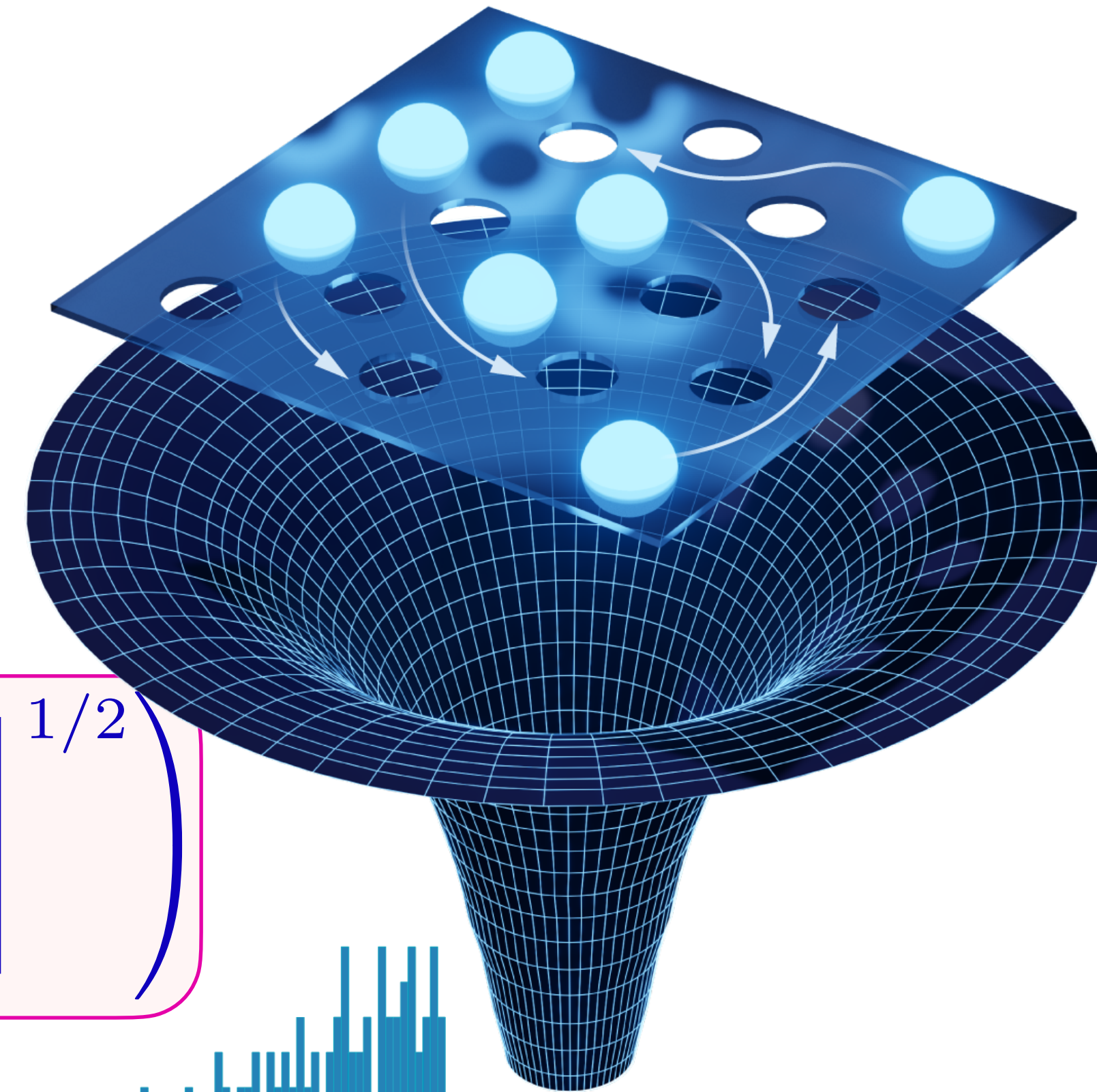
- For generic charged black holes in 3+1 dimensions with horizon area A_0 at $T = 0$ and fixed charge Q ($A_0 = 2GQ^2/c^4$), the density of quantum states at small energy E is

$$D(E) \sim \left(\frac{A_0 c^3}{\hbar G} \right)^{-347/90} \exp \left(\frac{A_0 c^3}{4\hbar G} \right) \sinh \left(\left[\frac{\sqrt{\pi} A_0^{3/2} c^2}{\hbar^2 G} E \right]^{1/2} \right)$$

Iliesiu, Murthy, Turiaci (2022)

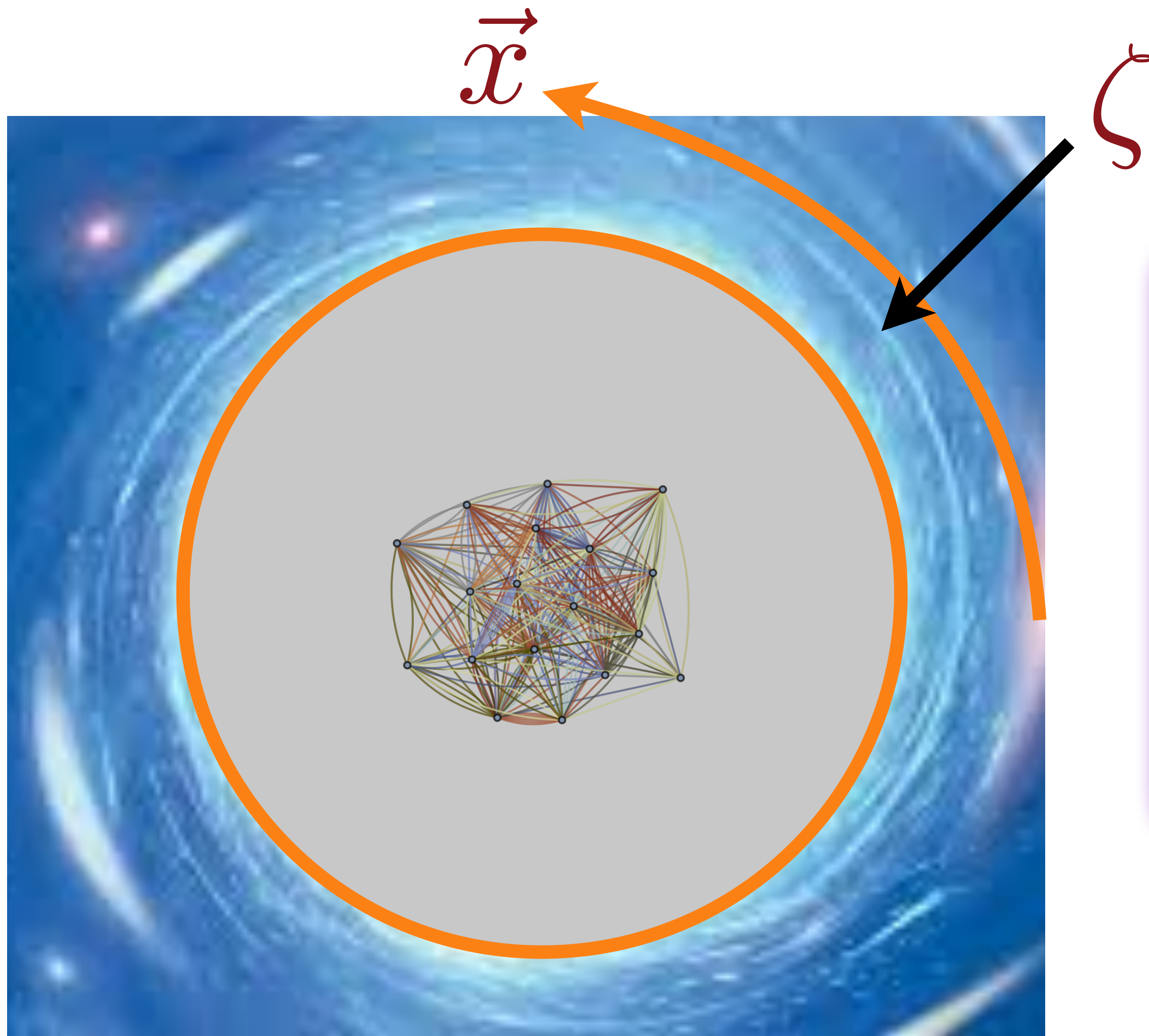
Developments from the SYK model

Bekenstein-Hawking



Similar remarks apply to rotating neutral black holes.

Quantum simulation of charged black holes by the SYK model

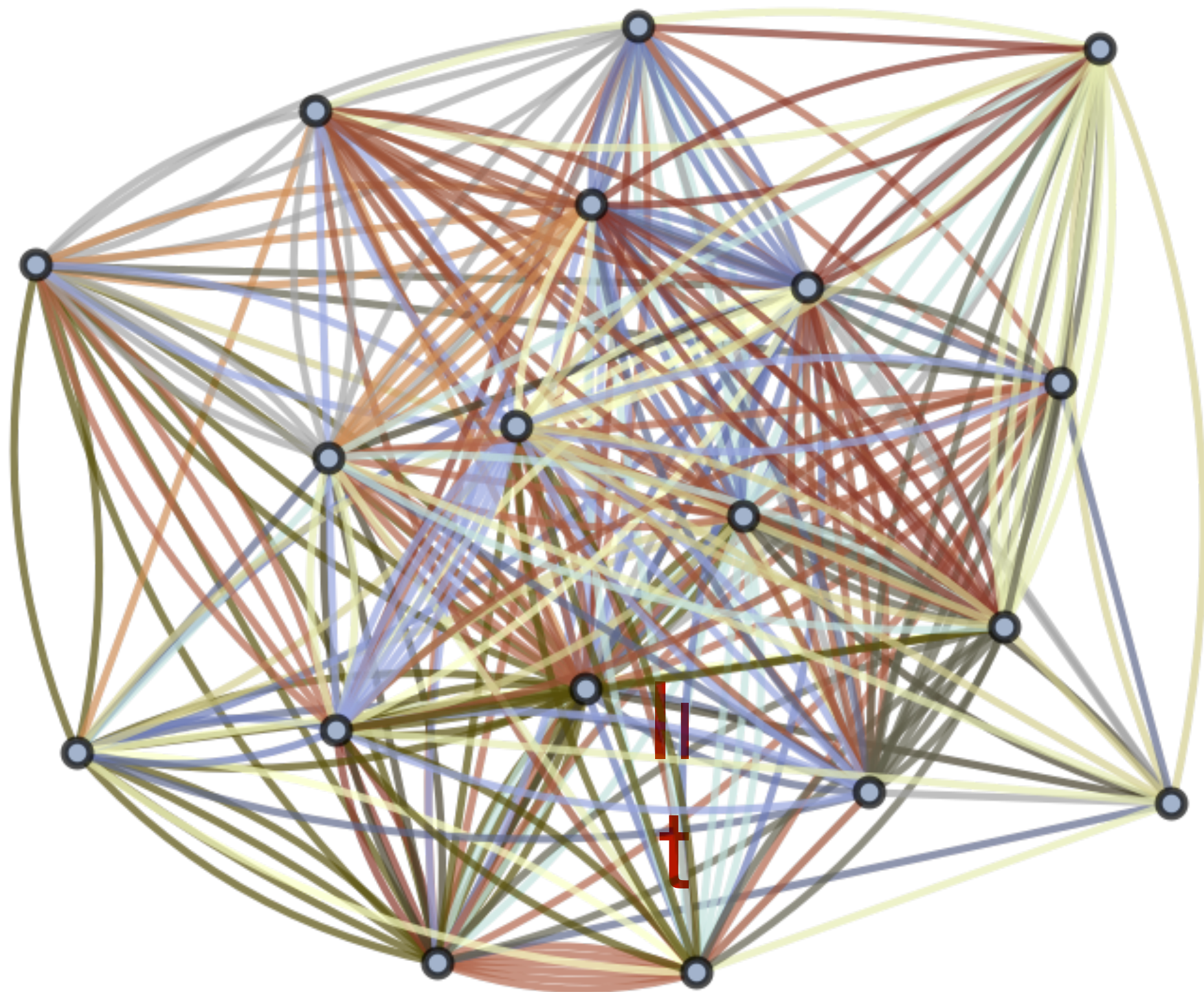


The SYK model simulates the low energy properties of the interior of the black hole for an outside observer in ζ - τ co-ordinates.

Recap

The Sachdev-Ye-Kitaev (SYK) model

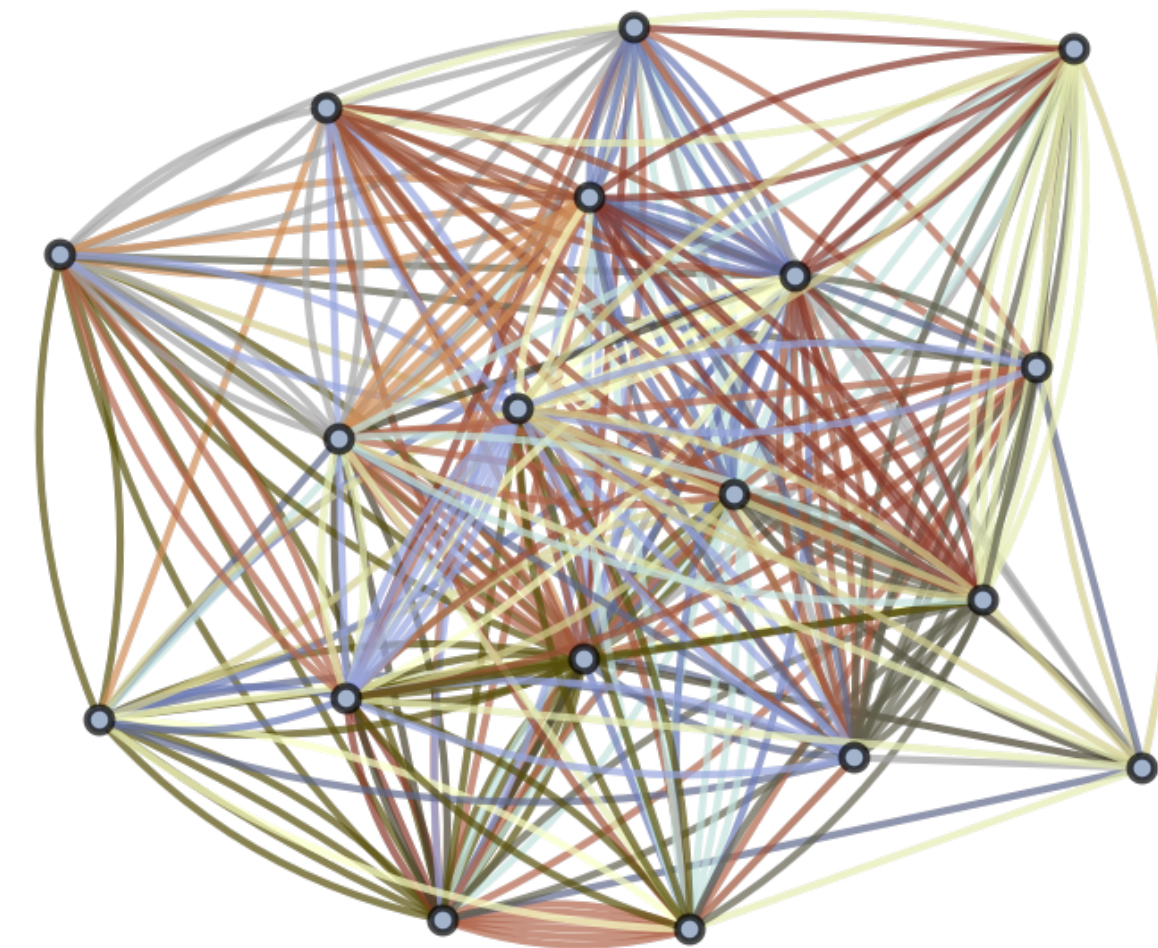
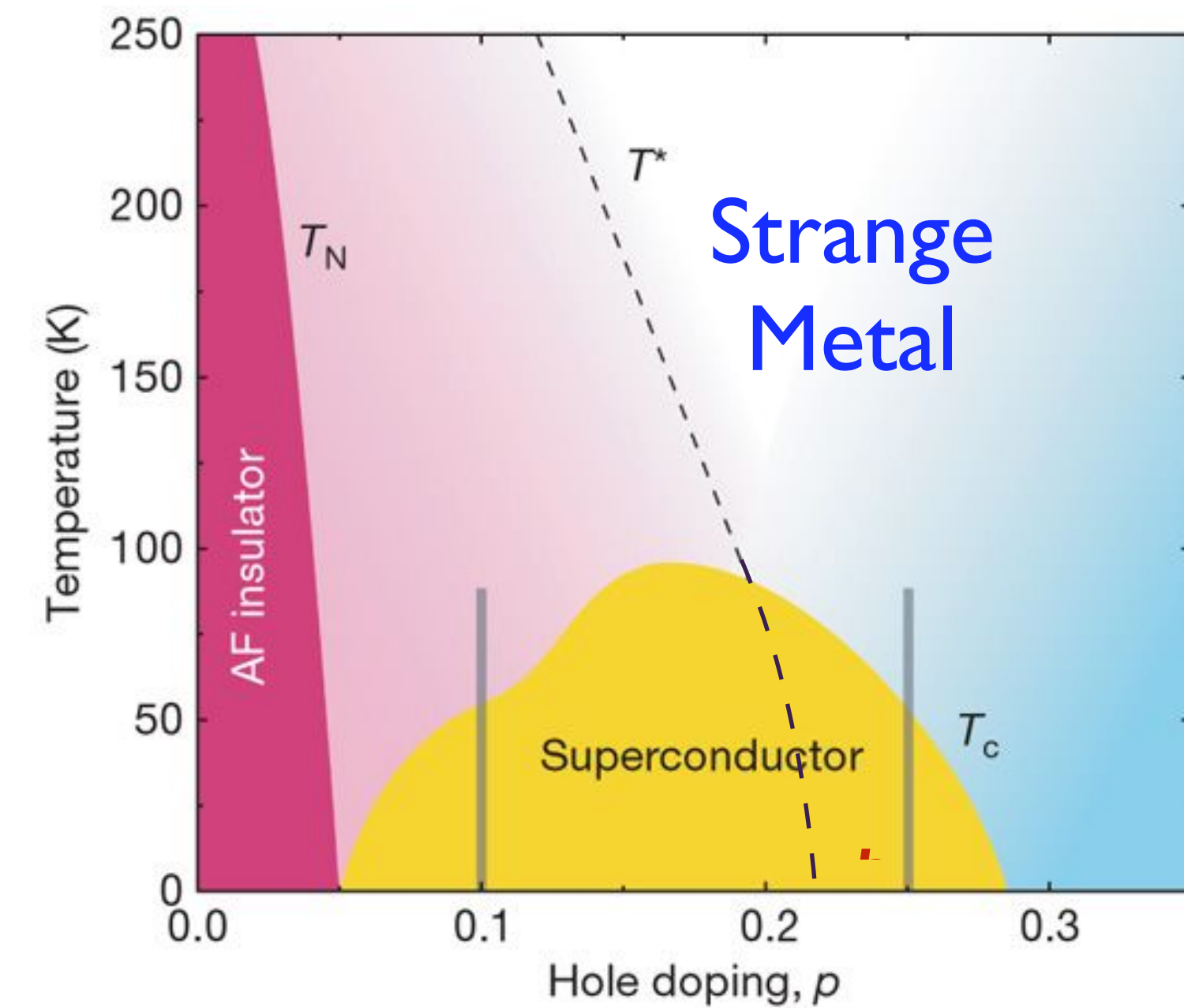
The SYK model describes multi-particle quantum entanglement resulting in the loss of identity of the particles



The Sachdev-Ye-Kitaev (SYK) model

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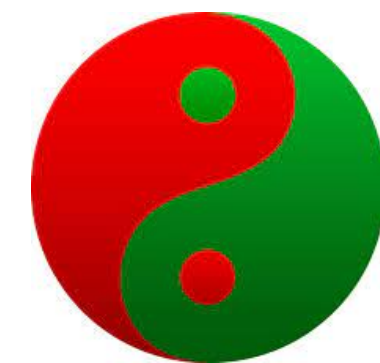
A 2d-YSYK theory describes the **strange metal** behavior of numerous quantum materials



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In a *dual* set of variables the SYK model has led to the computation of the low energy density of states of ***charged black holes***

