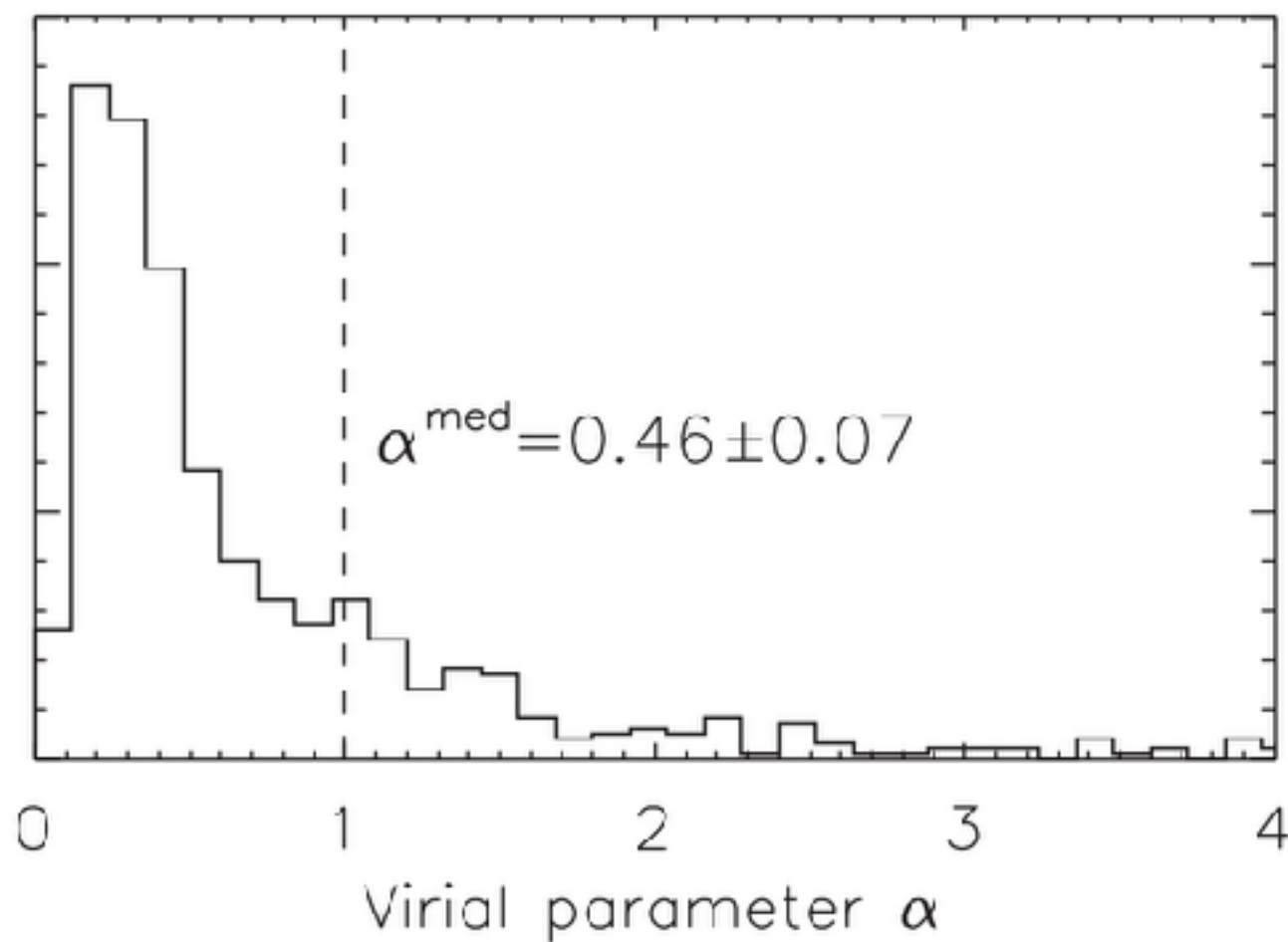


Can the virial parameter reliably trace the dynamical state of molecular clouds?

Camilo H Peñaloza

Paul Clark, Simon Glover and Ralf Klessen

Are clouds bound?



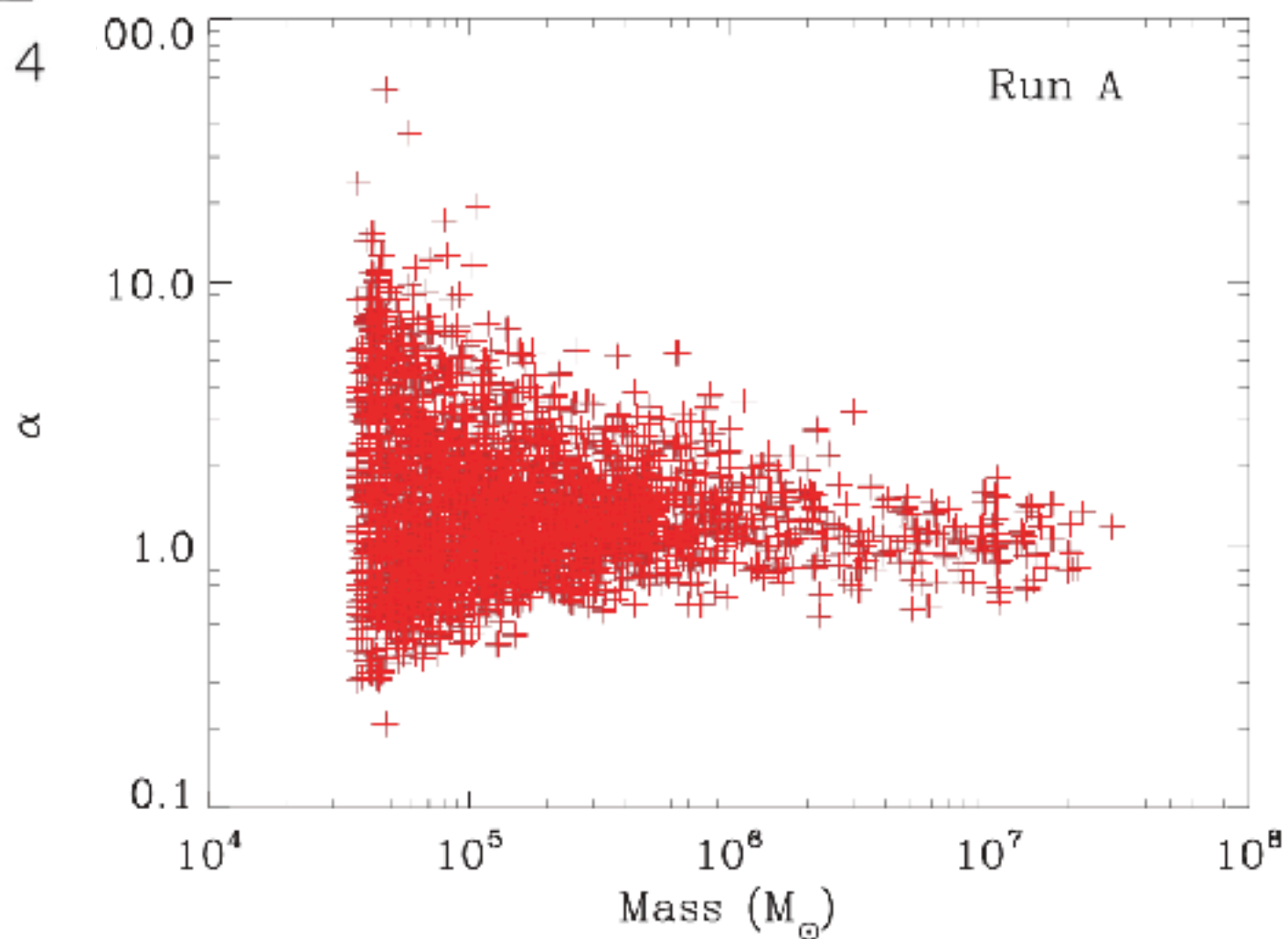
Roman-Duval et al. 2010

$$\alpha = \frac{2E_{\text{kin}}}{E_{\text{pot}}}$$

$$\alpha = \frac{5\sigma^2 R}{GM}$$

Bertoldi & McKee 1992

Dobbs et al. 2011



Are clouds bound?

- Are these expressions equivalent?

$$\alpha = \frac{5\sigma^2 R}{2GM}$$

Dobbs et al. 2011

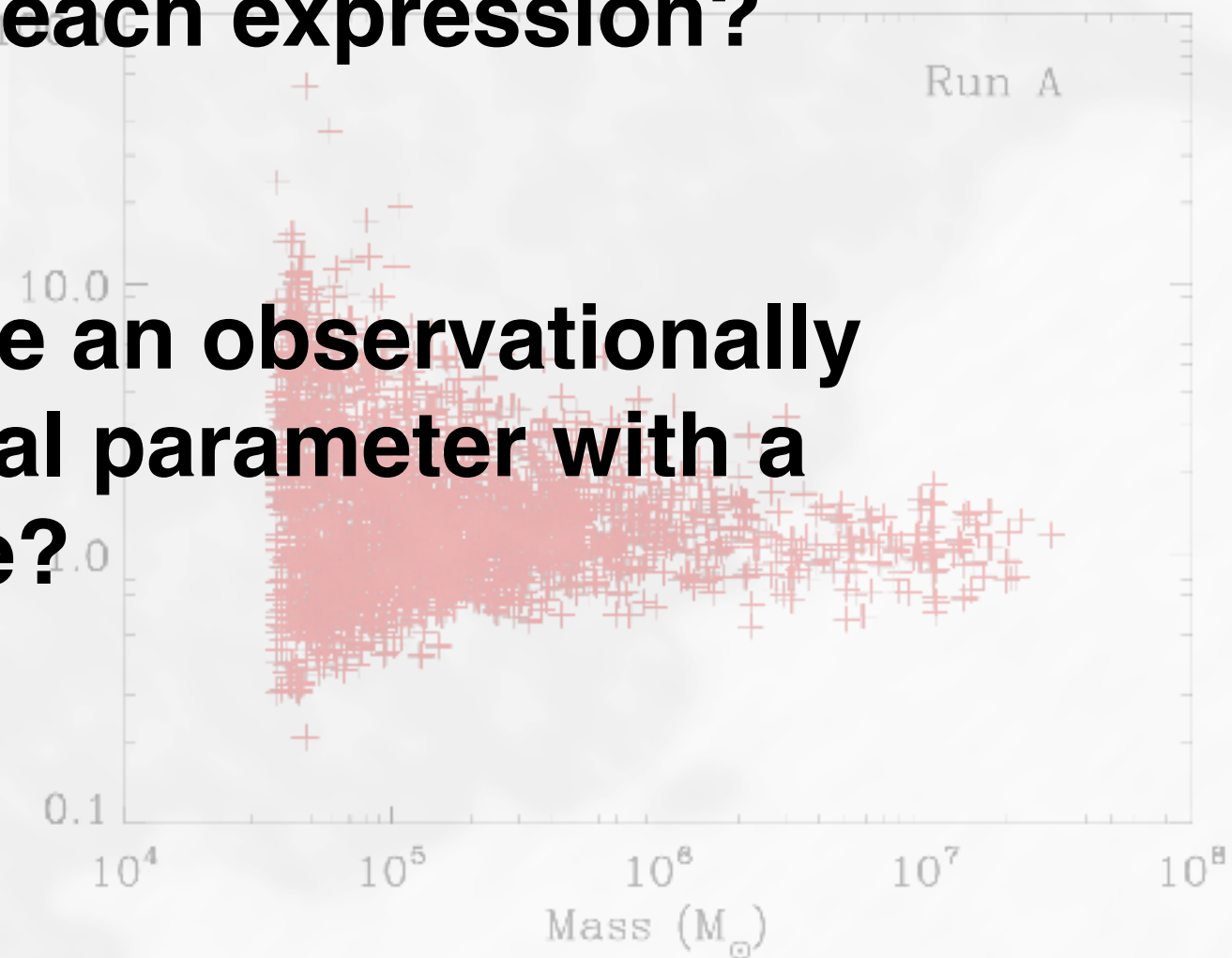
- What are the biases for each expression?

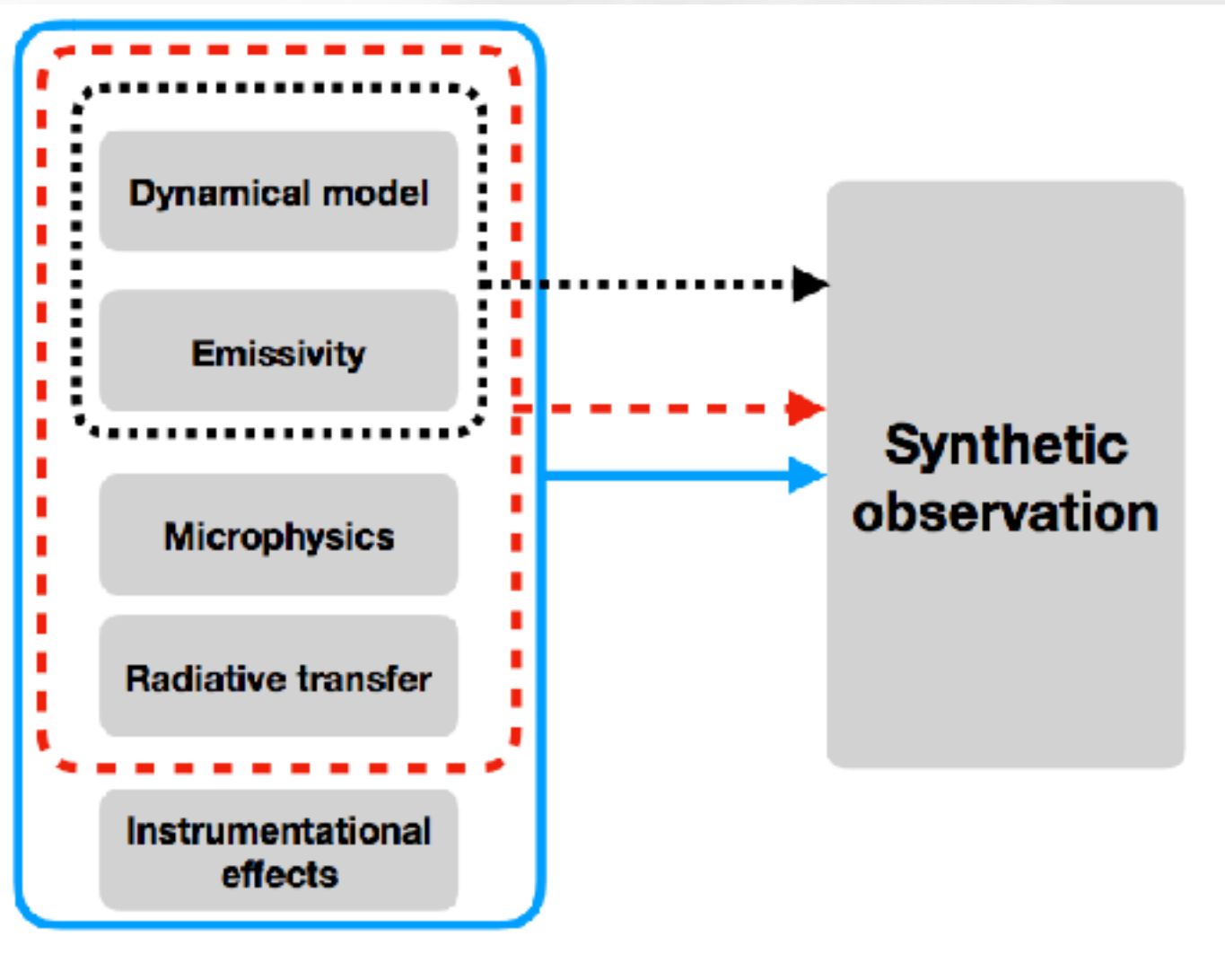
Virial parameter α

Roman-Duval et al. 2010

- Can we directly compare an observationally derived value of the virial parameter with a numerically derived one?

$$\alpha = \frac{2E_{\text{kin}}}{E_{\text{pot}}}$$

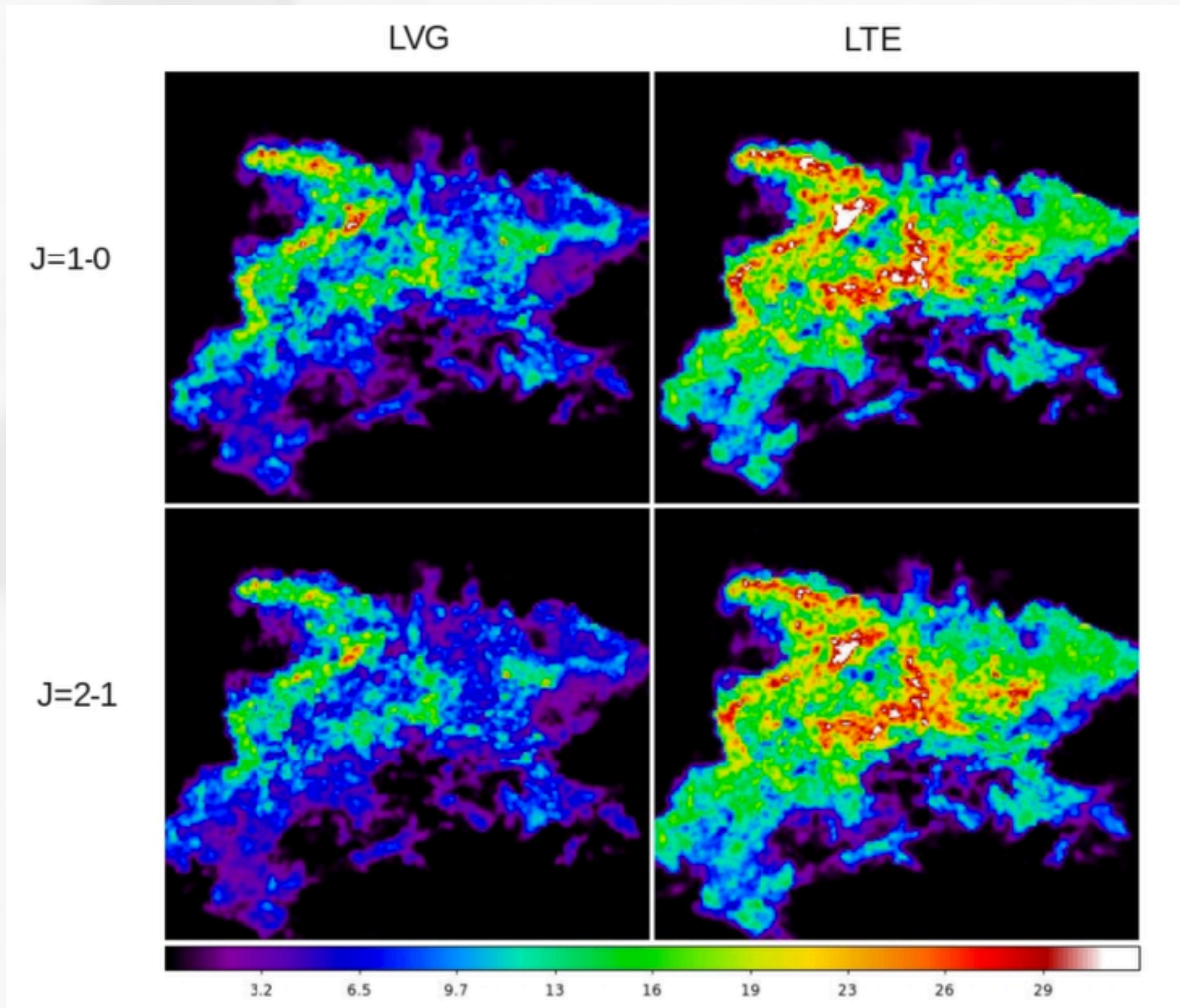




Haworth et al. 2017

- Hydrodynamics:
 - Modified version of GADGET-2, (Springel 2005)
 - Time dependant chemistry
 - Attenuation of the ISRF
 - Stop before SF
 - No magnetic fields
 - No feedback
- Radiative Transfer:
 - RADMC-3D (Dullemond 2012)
 - AMR-grid (Accounts for all SPH particles)
 - Sobolev-Gnedin Approximation(LVG+)
 - CO's first two rotational transition lines

LTE vs Non-LTE



- Hydrodynamics:

- Modified version of GADGET-2, (Springel 2005)
- Time dependant chemistry
- Attenuation of the ISRF
- Stop before SF
- No magnetic fields
- No feedback

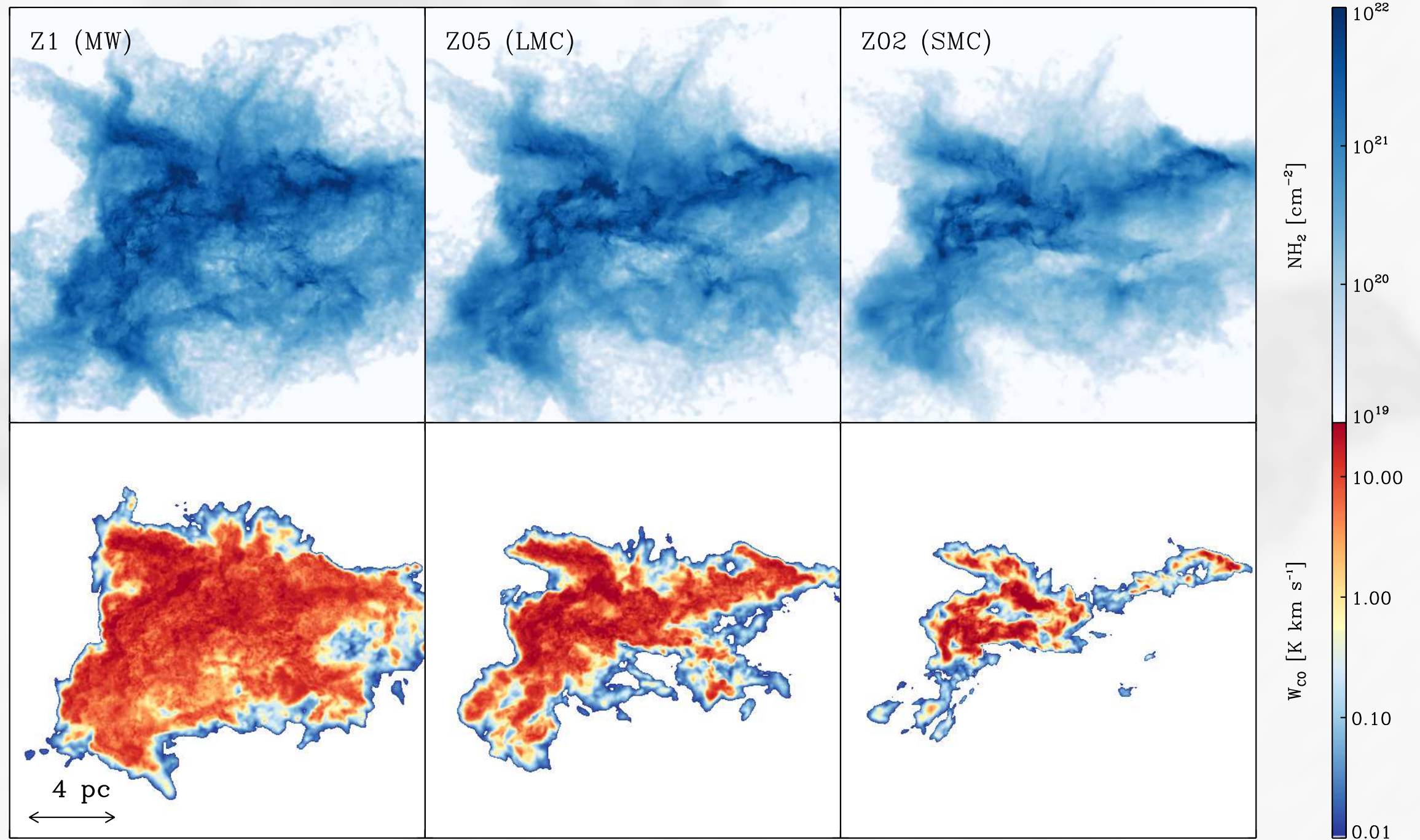
- Radiative Transfer:

- RADMC-3D (Dullemond 2012)
- AMR-grid (Accounts for all SPH particles)
- Sobolev-Gnedin Approximation(LVG+)
- CO's first two rotational transition lines

- Initial Conditions:

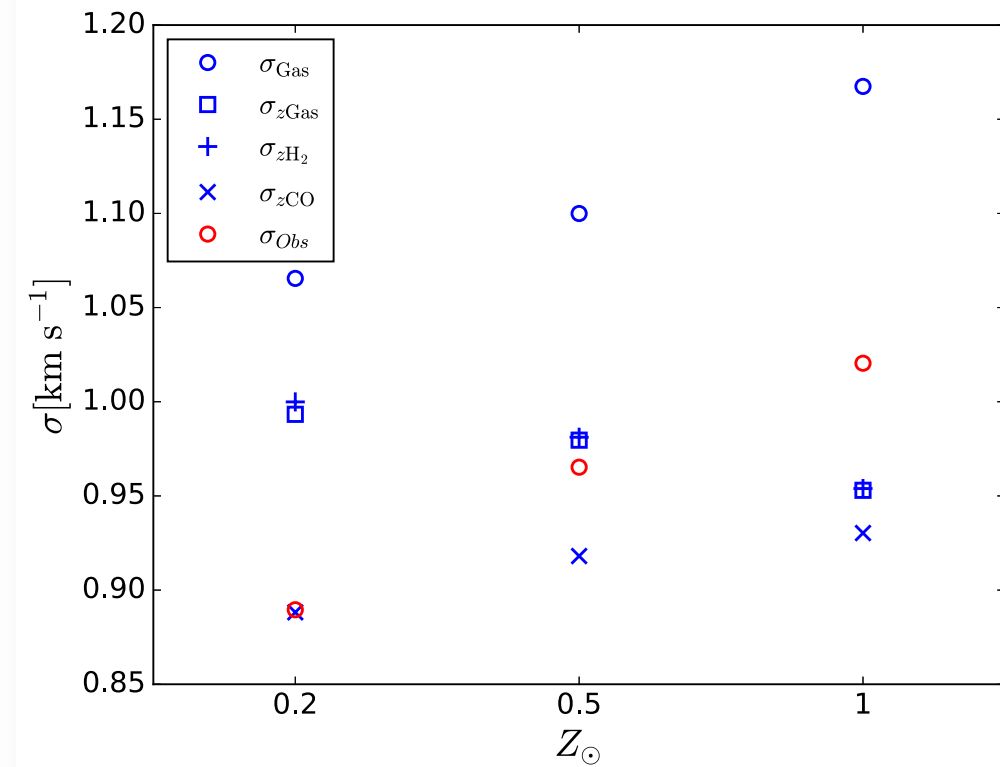
- $10^4 M_{\odot}$ uniform sphere
- $n = 276 \text{ cm}^{-3}$
- $\nabla \cdot \mathbf{u} = 0$
- $G_0 = 1.7$ in Habing(1968)
- $Z = Z_{\odot} / 0.5 \text{ } Z_{\odot} / 0.2 \text{ } Z_{\odot}$

Observations vs Simulations

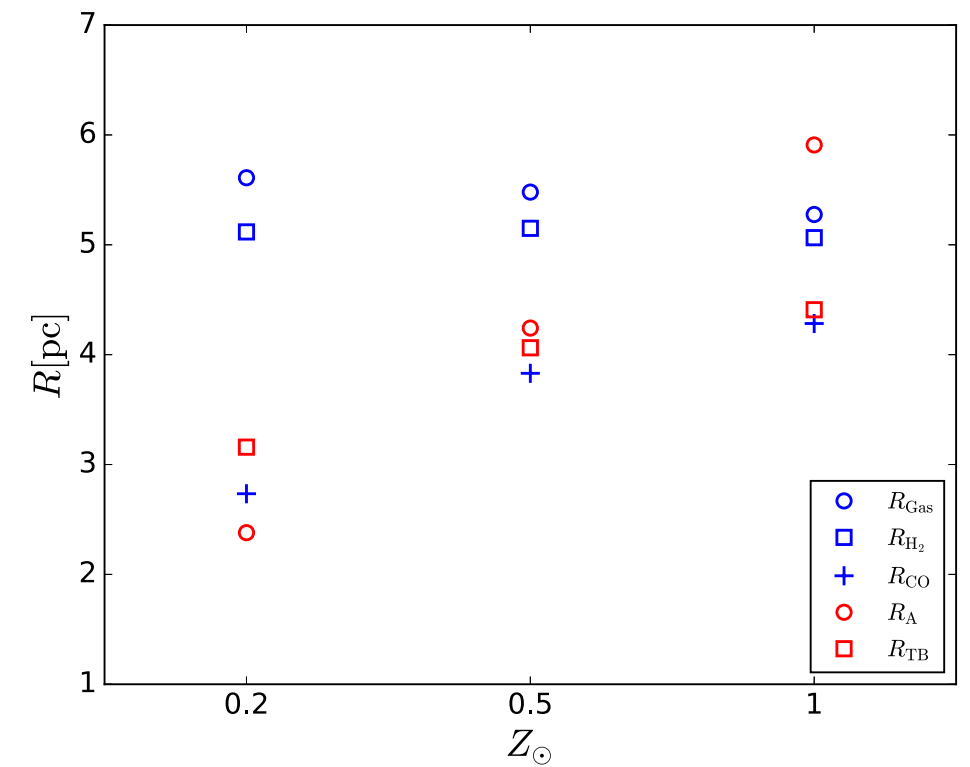


The components of the virial parameter

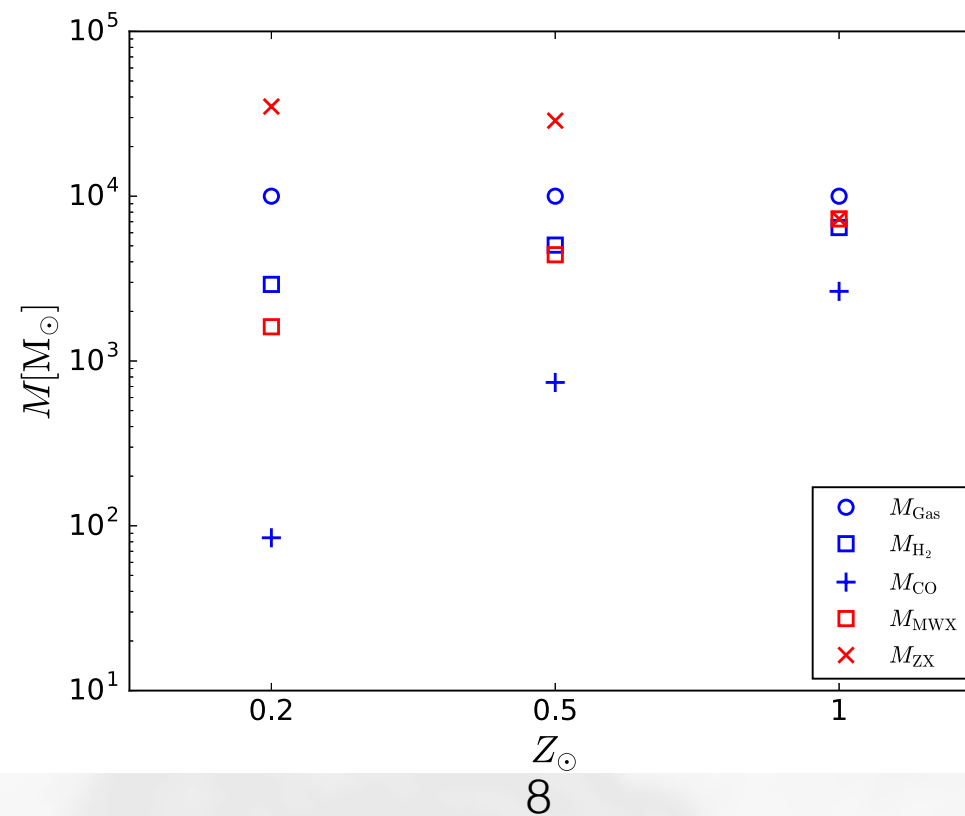
Velocity Dispersion



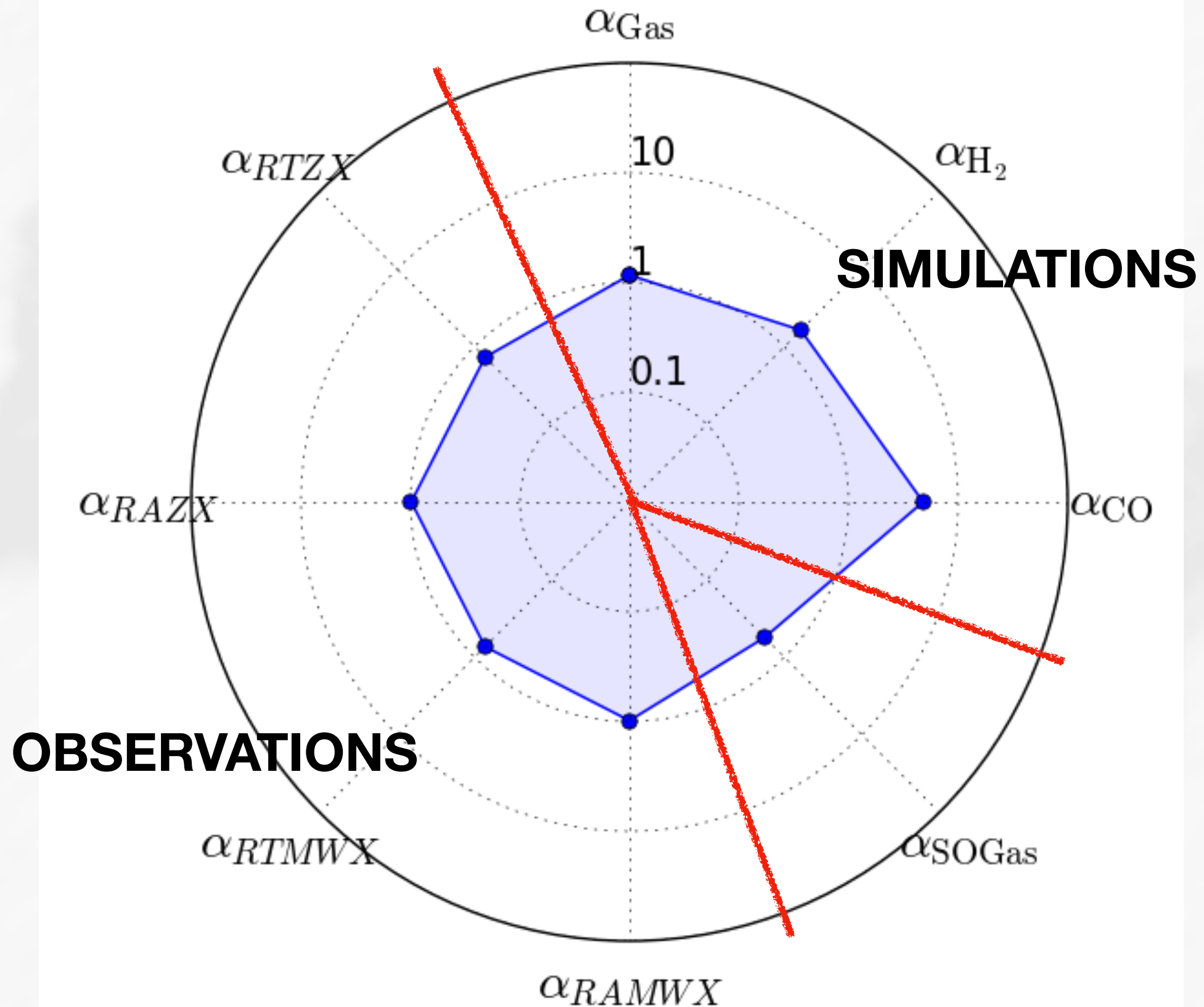
Radius



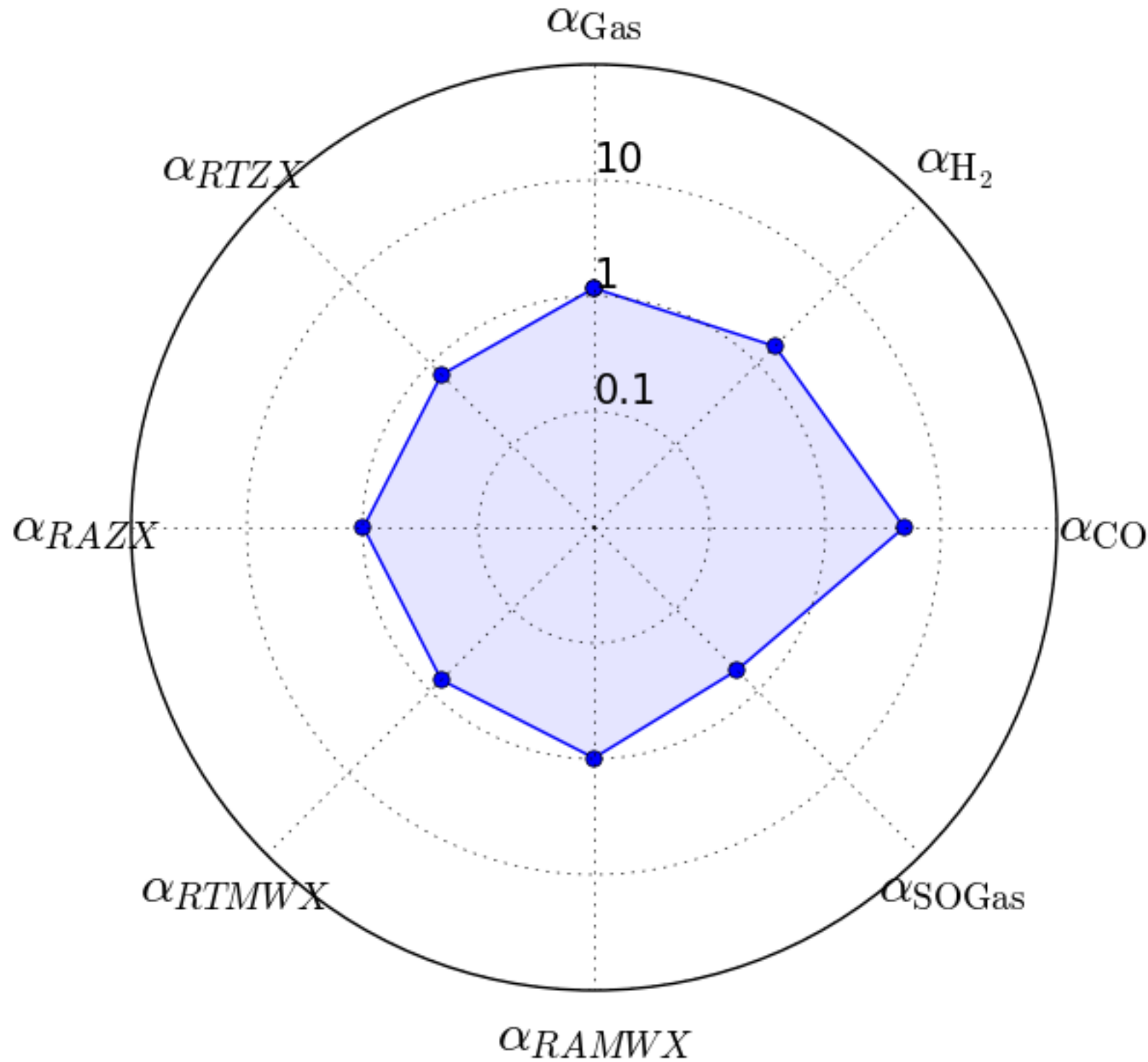
Mass



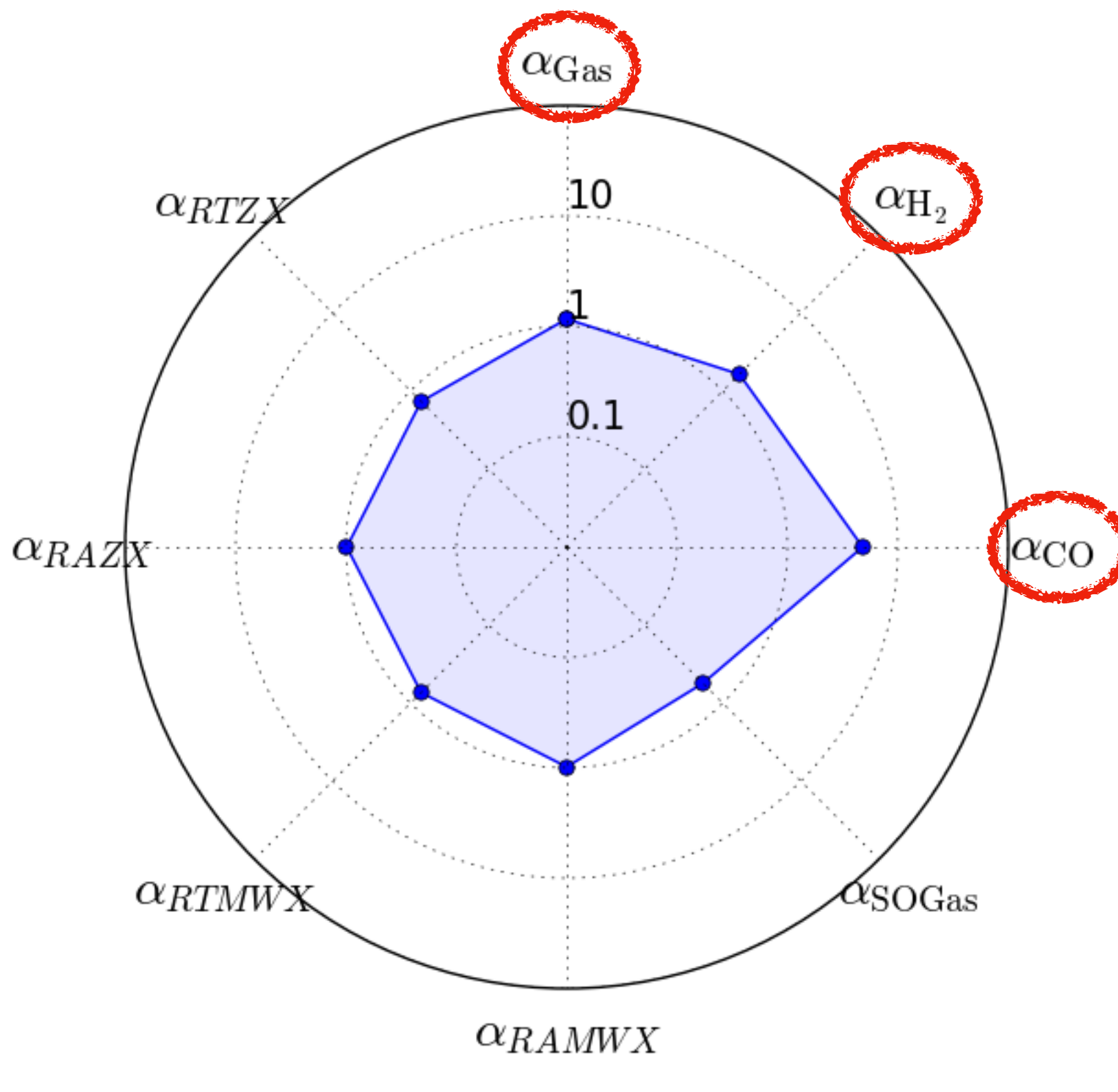
The virial parameter (Z_1)



The virial parameter (Z_1)



The virial parameter (Simulations)

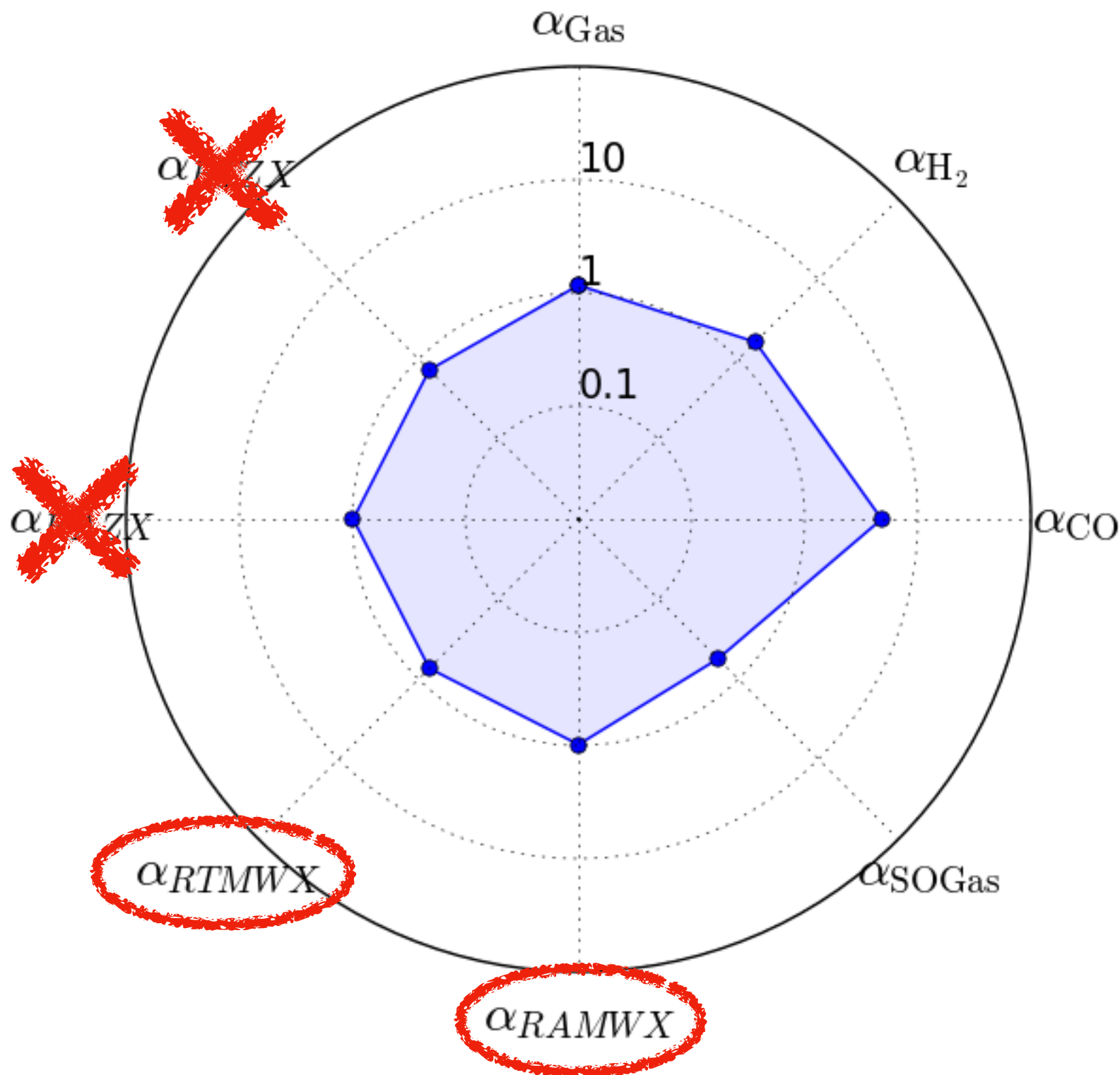


$$\alpha_{\text{gas}} = \frac{2E_g}{E_k}$$

$$\alpha_{\text{gas}} = \frac{\sum 5GM_n/3R_n}{\sum m_i v_i^2 / 2t}$$

$$n = \text{gas}, \text{H}_2, \text{CO}$$

The virial parameter (Observations)

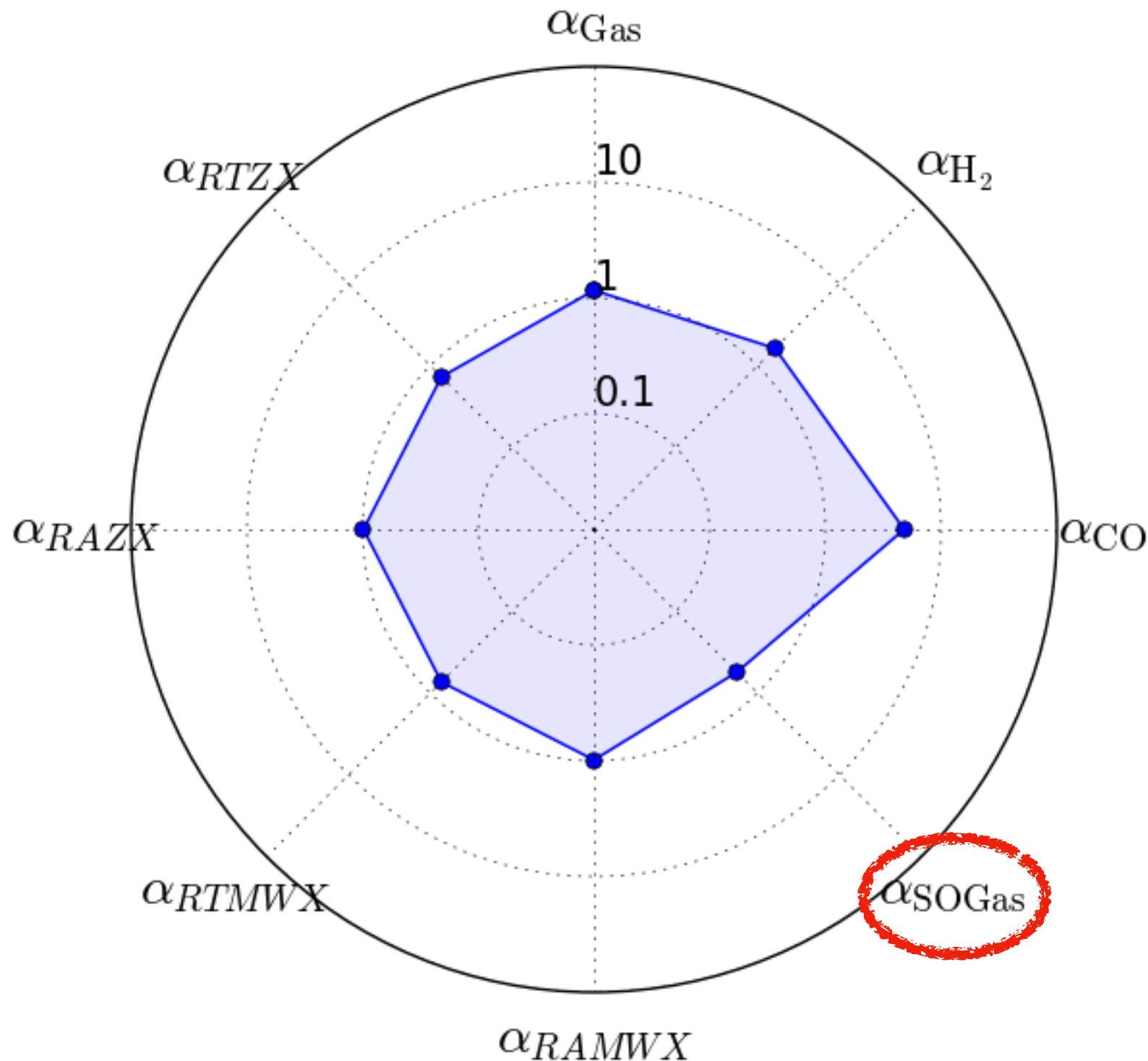


$$\alpha = \frac{5\sigma^2 R}{GM}$$

$$R_A = \sqrt{\frac{A}{\pi}}$$

$$R_{\text{T}_B} = \sqrt{\frac{\sum_i^N T_{\text{B},i} (R_{ij} - \langle R_j \rangle)^2}{\sum_i^N T_{\text{B},i}}}$$

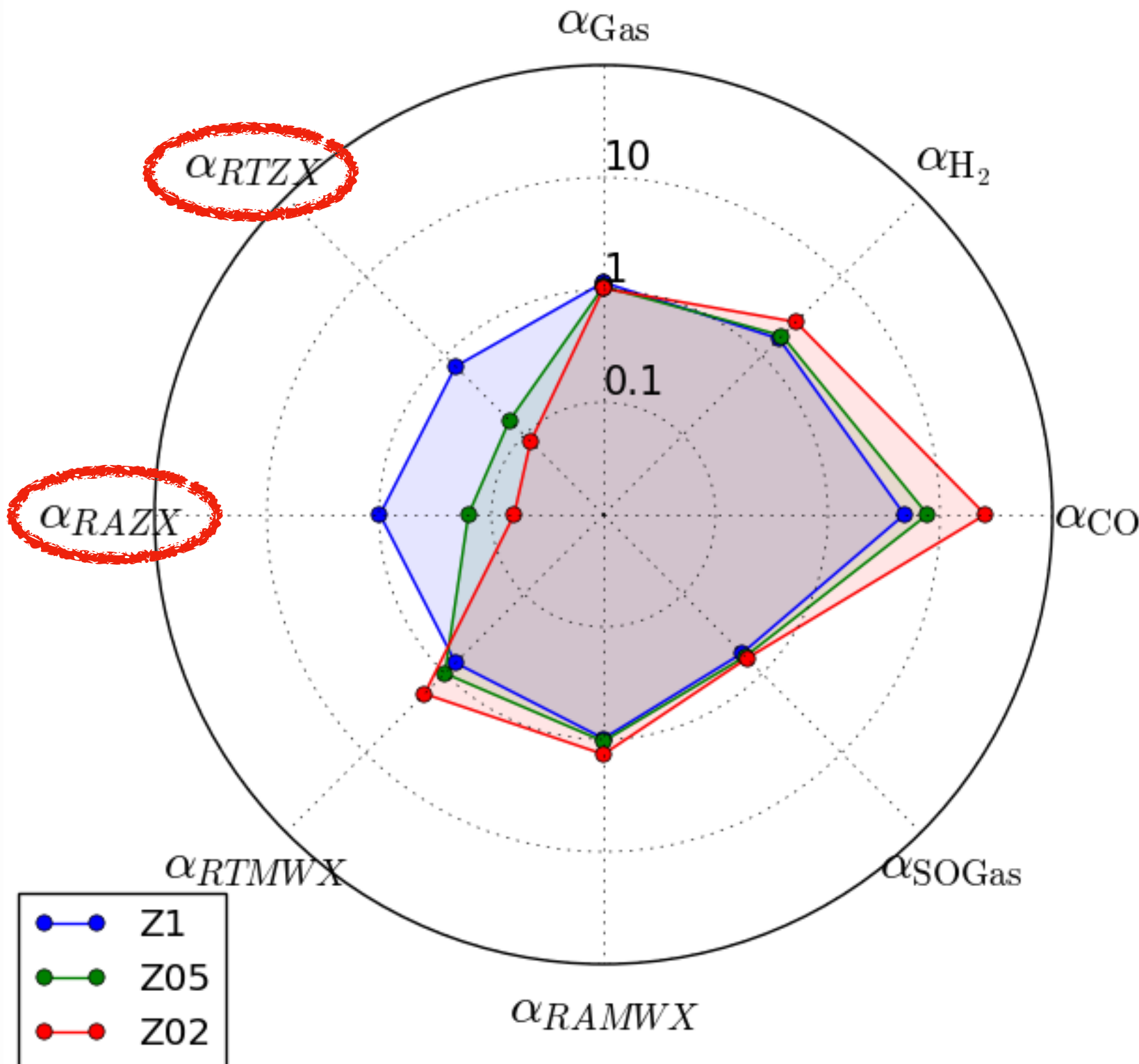
The virial parameter



$$\alpha = \frac{5\sigma^2 R}{GM}$$

R , M and σ^2
are all calculated
from simulations

The virial parameter



$$\alpha = \frac{5\sigma^2 R}{GM}$$

$$M = \alpha_{\text{CO,MW}} L_{\text{CO}}$$

$$M = \alpha_{\text{CO,z}} L_{\text{CO}}$$

- The observationally derived virial parameter is highly dependent on how the velocity dispersion, size and mass are calculated.
- Our results show that the same cloud can appear unbound in simulations and bound through synthetic observations.
- When comparing observations to simulations its important to consider what techniques are used and how different quantities are calculated. Synthetic observations can really help!

THANKS :)